Warped Penguins

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Based on arXiv:1004.2037 In collaboration with Csaba Csáki, Yuval Grossman, Philip Tanedo



LEPP Particle Theory Seminar

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Why are **Penguins** important in **RS**?



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Randall-Sundrum in one slide



Randall, Sundrum (99);

Yuhsin Tsai, Cornell University/LEPP

Randall-Sundrum in one slide



Randall, Sundrum (99); Davoudiasl, Hewett, Rizzo (99);

Yuhsin Tsai, Cornell University/LEPP Wai

Randall-Sundrum in one slide



Randall, Sundrum (99); Davoudiasl, Hewett, Rizzo (99); Grossman, Neubert (00); Gherghetta, Pomarol (00); **Bulk Higgs:** Agashe, Contino, Pomarol (04); Davoudiasl, Lille, Rizzo (05)

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Flavor changing & the wavefunction overlap



4D Yukawa Coupling: $Y_*^{ij} \bar{L}_i HE_j \times f_{Li}(R') f_{Ej}(R')$ 4D Gauge Coupling: $g_{ij} \bar{L}_i ZL_i \times \int_R^{R'} dz \left(\frac{R}{z}\right)^4 f_{Li}(z) f_Z(z) f_{Li}(z)$

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Before we proceed,

Two assumptions for an interesting model:

- M_{kk} is not too heavy. Give me a number! $M_{kk} \sim 3 \text{TeV}$. \Rightarrow Can be seen in LHC!
- Y_*^{ij} are ancharic, $\mathcal{O}(1)$ numbers. \Rightarrow No tuning is required!

FCNC from the loop & tree

 $\mu \rightarrow \boldsymbol{e}\gamma$: μ e μ е $\propto Y^3_{\star}$ $\mu
ightarrow$ 3e & $\mu
ightarrow$ e: $\mu \rightarrow$ $\propto Y_*^{-1}$

The opposite Y_* bounds



Constraint Y_* from both sides!

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The opposite Y_* bounds



Want to get the precise bounds!

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Tree level constraints from



K. Agashe, A. E. Blechman, and F. Petriello, hep-ph/0606021.

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The origin of the flavor changing

Taking the zero mode Z as an example, the nonuniversal part of the wave function $h^{(0)}(z)$ near IR gives the flavor changing,

Wave function:

$$h^{(0)}(z) = \frac{1}{\sqrt{R \ln R'/R}} \left[1 + \frac{M_Z^2}{4} z^2 \left(1 - 2 \ln \frac{z}{R} \right) \right]$$

Gauge coupling:

$$g^{zf_i f_j} = g^z_{SM} \left[1 + rac{(M_Z R')^2 \ln R'/R}{2(3-2c)} f_i f_j
ight]$$

The γ' and Z' have the similar form.

The Yukawa bound

The tree level result,

$$Br(\mu \to 3e) \simeq 10^{-13} \left(\frac{3\text{TeV}}{M_{KK}}\right)^4 \left(\frac{2}{Y_*}\right)^2$$
$$Br(\mu \to e)_{\text{Ti}} \simeq 2 \cdot 10^{-12} \left(\frac{3\text{TeV}}{M_{KK}}\right)^4 \left(\frac{2}{Y_*}\right)^2$$

Comparing to the experimental bounds,

•
$$Br(\mu
ightarrow 3e) < 10^{-12}$$

•
$$Br(\mu
ightarrow e)_{
m Ti} < 6.1 \cdot 10^{-13}$$

The $\mu \rightarrow e$ gives the most stringent bound (for $M_{KK} = 3$ TeV).

*Y*_{*} > 3.7

$\mu \rightarrow e \gamma$ in a warped XD



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Warped penguins

Warped Penguin looks weird...

- They live in 5D.
- Previous analyses suggested the brane Higgs case is log-divergent.



People thought they are divergent because of the KK sum.

Want to do a full 5D calculation to check the finiteness!

Warped penguins

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- They live in 5D.
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Can show $\mu \rightarrow \boldsymbol{e}\gamma$ is finite!

$\mu ightarrow e \gamma$ in 5D

The diagrams of $\mu \rightarrow e \gamma$:

- Get the leading order result by doing mass insertions.
- Rotate all the flavor mixing to the Yukawa.
- It is an L-R operator: $a \bar{L} \sigma^{\mu\nu} F_{\mu\nu} E$
- From the gauge symmetry:

 $\epsilon_\mu \mathcal{M}^\mu \sim \epsilon_\mu ar{u} \left(\, {m p}^\mu + {m p}^\mu_e - (m_\mu + m_e) \gamma^\mu \,
ight) u$

• To get **a**, we only need to calculate the coefficient of p^{μ} .



The diagrams:



The Result

The result of the amplitude can be written as follows:

$$\mathbf{a_{k\ell}} \times R'^2 \frac{e}{16\pi^2} \frac{v}{\sqrt{2}} \left(f_{L_i} Y_{ik} Y_{k\ell}^{\dagger} Y_{\ell j} f_{-E_j} \right) \bar{L}_i^{(0)} \sigma^{\mu\nu} E_j^{(0)} F_{\mu\nu}^{(0)}$$

The loop integral gives $a_{k\ell} \simeq 0.5$.



$$\begin{aligned} & \mathsf{Br}(\mu \to e \, \gamma)_{\rm thy} > 8.2 \cdot 10^{-7} a^2 \left(\frac{3 \mathrm{TeV}}{M_{KK}}\right)^4 \left(\frac{Y_*}{2}\right)^4 \\ & \mathsf{Br}(\mu \to e \, \gamma)_{\rm exp} < 1.2 \cdot 10^{-11} \\ & \text{for } M_{kk} = 3 \, \mathrm{TeV}, \, \text{we have} \\ & \mathbf{Y}_* < \mathbf{0.2} \end{aligned}$$

The Yukawa bounds

There is a tension between the two bounds!

 $Y_*(\mu o e) > 3.7$ $Y_*(\mu o e \gamma) < 0.2$

To release the tension:

- Make M_{KK} heavier. $M_{KK} \sim 10 \, \text{TeV}$
- Give some structure to Y_*^{ij} .
- Introduce some symmetry to suppress the FCNC.

K.Agashe, R.Contino, L. Da Rold, and A. Pomarol, hep-ph/0605341. M.E.Albrecht, M.Blanke, A.J.Buras, B.Duling, and K.Gemmler, 0903.2415.

Introducing a custodial symmetry for leptons

 $SU(2)_L \times SU(2)_R \times U(1)_X \times P_{L,R}$

This model gives:

• γ , Z, γ' , Z' & Z_H.

 The coupling of Z, Z' to the LH fermions becomes flavor universal.

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This model gives:

• γ , Z, γ' , Z' & Z_H.

• The coupling of *Z*, *Z'* to the LH fermions becomes flavor universal.

The LH FCNC is suppressed.



Strategy

- Make the RH leptons towards UV.
- Make the LH leptons towards IR.



This reduces the Y_{*} bound to

 $Y_{*} > 1$

 $Y_* > 3.7$ for the non-custodial case.

Strategy

- Make the RH leptons towards UV.
- Make the LH leptons towards IR.

Tension between the two bounds

It's not enough...

 $Y_*(\mu \rightarrow e) > 1$ $Y_*(\mu \rightarrow e \gamma) < 0.2$

A mild tension still exists.

Very interesting detail for the 5D loops!



The fermion propagator in 5D

$$(-p + i\gamma^5\partial_5 + m)\Delta(p, z, z') = i\delta(z - z')$$

Trick:

$$\left(p^2-\partial_5^2+m^2\right)F(p,z,z')=i\delta(z-z')$$

$$\Delta(\boldsymbol{\rho}, \boldsymbol{z}, \boldsymbol{z}') = (\boldsymbol{p} - i\gamma^5\partial_5 + \boldsymbol{m})F(\boldsymbol{\rho}, \boldsymbol{z}, \boldsymbol{z}')$$

In flat XD case:

$$F(\boldsymbol{p}, \boldsymbol{z}, \boldsymbol{z}') = A(\boldsymbol{p}, \boldsymbol{z}') \cos(\boldsymbol{p} \, \boldsymbol{z}) + B(\boldsymbol{p}, \boldsymbol{z}') \sin(\boldsymbol{p} \, \boldsymbol{z})$$

The chiral BC

The chiral BC constraints the form of the amplitude!



How to calculate the 5D loop?

Feynman's trick with Bessel functions?? We don't know that...

However, we can

- Tayler expand the propagator into powers of the external momentum (p^{μ} , q^{μ}). $\frac{1}{(k^2 + 2k \cdot q)} = \frac{1}{k^2} \left(1 \frac{2k \cdot q}{k^2} + ... \right)$.
- Isolate the p^{μ} terms. Get the coefficient. $p\gamma^{\mu} = 2p^{\mu} \gamma^{\mu}p$.
- Solve the numerical integral, get the coefficient of $\mu \to e \gamma$. $\mathbf{a} \times R'^2 \frac{e}{16\pi^2} \frac{v}{\sqrt{2}} \left(f_L Y Y^{\dagger} Y f_E \right) \bar{u} \notin u; \quad \mathbf{a} = \int dx \int dy \text{ (scalar function)}.$

Why is $\mu \rightarrow e \gamma$ is finite?

Lorentz invariance + chiral BC gives the finiteness:

The propagators are composed of two parts

- The Dirac operator part (gives the γ^{μ} structure).
- The Bessel function part (contains the 5D profile).

In the UV limit, the photon vertex is pulled back to the IR brane.

Can just look at the Dirac operator part.

Each loop contains two sectors:

- o photon emission
- mass insertion



Why is $\mu \rightarrow e \gamma$ is finite?

- γ emission: Lorentz inv + chiral BC gives $(\not k \gamma^{\mu} \not k k^2 \gamma^{\mu})$
- mass insertion: The chiral BC gives ₭.
- Combining the two m-insertion amp, the leading order becomes $\mathcal{M}_{(a)} + \mathcal{M}_{(b)} \sim \not k \left(\not k \gamma^{\mu} \not k - k^{2} \gamma^{\mu} \right) + \left(\not k \gamma^{\mu} \not k - k^{2} \gamma^{\mu} \right) \not k = \mathbf{0}$
- From NDA, only the leading order term can be divergent.

 $\mu \rightarrow \boldsymbol{e} \gamma$ is finite!

5D V.S. KK



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The mass matrix:

$$M = egin{pmatrix} m^{11} & 0 & m^{13} \ m^{21} & M_{{
m KK},1} & m^{23} \ 0 & 0 & M_{{
m KK},2} \end{pmatrix}$$

In the mass basis, the Yukawa takes the form ($\epsilon = m/M_{KK}$)

$$\hat{y} \sim \begin{pmatrix} 1 & 1+\epsilon & -1+\epsilon \\ 1+\epsilon & \cdots & \cdots \\ 1-\epsilon & \cdots & \cdots \end{pmatrix}$$

 $\mathcal{M} \propto (1+\epsilon)(1+\epsilon) + (-1+\epsilon)(1-\epsilon) \sim \epsilon$

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The KK result should be the same as 5D For the finiteness:

In the mass basis, naive NDA gives $\mathcal{M} \propto M_{KK}^{-1}$. The Yukawa gives additional $\epsilon = m/M_{KK}$ and ensures finiteness.

$$\Rightarrow \mathcal{M} \propto \sum_{KK} rac{1}{\left(M_{KK}
ight)^2}$$

• In the M_{KK} basis, treat *m* as the mass insertion. The chiral BC effect becomes obvious.



For the *a* value:

Do the momentum integral to infinity and sum the KKs.

$$\mathcal{M} = \sum_{n=1}^{N} \int_{0}^{\infty} d^{4}k \ \hat{\mathcal{M}}^{(n)}(k)$$

The leading order result is

$${\cal M} \propto {1 \over \left(M_{{\cal K}{\cal K}}
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 !?

which is different from the 5D result $\mathcal{M} \propto R^{\prime 2}$!

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 $5D \neq KK !?$

The reason is,

$\label{eq:Wrong UV limit!}$ To match the 5D & KK, the momentum cutoff \lesssim KK cutoff.

The correct way of doing the sum and the integral is

$$\lim_{N\to\infty}\sum_{n=1}^N\int_0^{N\,M_{KK}}d^4k\,\,\hat{\mathcal{M}}^{(n)}(k)$$

This gives the same leading order contribution $\propto M_{KK}^{-2}$ as in 5D.

The 5D & KK results do match!



Outlook & Conclusion

One loop $\mu \rightarrow e\gamma$ for brane Higgs case is finite.

- The bulk higgs is also finite.
- How about the finiteness for higher loops?

A tension between the tree-level & loop-induced Yukawa bounds.

- Only having the custodial symmetry is not enough.
- Need some structure on Y_{*}.
- How about $b \rightarrow s\gamma$?

The nontrivial UV limit for KK is important for mathcing the 5D.

Does this happen in other loop-induced processes?

A long journey for Warped Penguins...



Yes, we worked very hard!



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