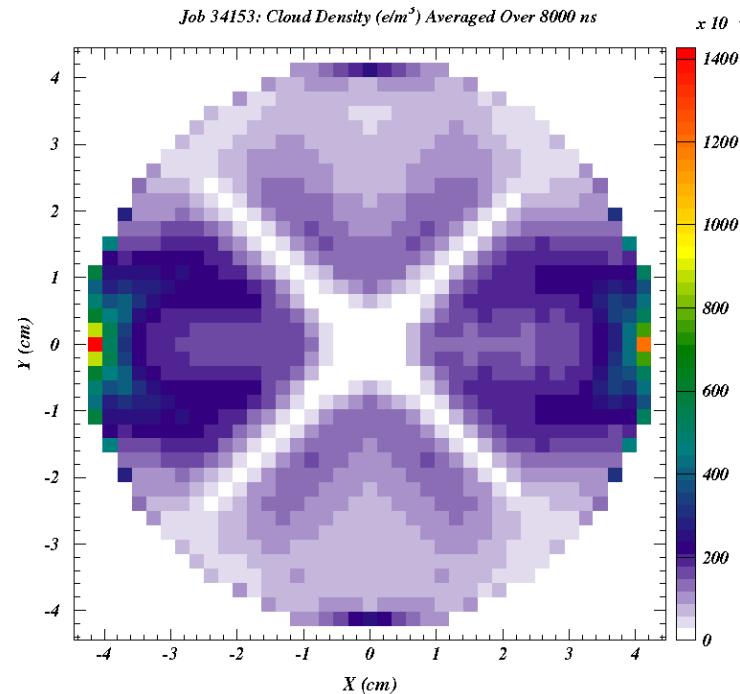




# ECLOUD Simulation of Electron Cloud Buildup in the Q48W Quadrupole Magnet for 30-Bunch Trains of 2.1 GeV Positrons With and Without a Precursor Bunch



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Plus bonus slide on  
optimal hole orientation  
for proposed  
time-resolved RFA in  
Q48W



Cornell Laboratory for Accelerator-Based Sciences and Education

Electron Cloud Meeting

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# Kiran's table of cloud dynamics measurements

	Train bunch current (mA)	Lead Bunch current (mA)	Energy (GeV)	Bunch spacing (ns)	Nr of bunches	1 <sup>st</sup> bunch blow up	2 <sup>nd</sup> bunch blowup	Date	Precursor bunch test
1	0.75	0.75	2	4	30	yes	yes	Apr 12	no
2	0.75	0.75	2	8	30	yes	yes	Apr 12 June 11	yes
3	0.75	0.75	2	12	30	yes	no	Apr 12	yes
4	0.75	0.75	2	14	30	no	no	Apr 12 June 11	N/A
5	0.75	0.75	2	16	30	no	no	Apr 12	N/A
6	0.75	0.75	2	20	30	no	no	Apr 12	N/A
7	0.75	0.75	2	24	30	no (?)	no	Apr 12	N/A
8	0.75	0.75	2	28	30	no(?)	yes(?)	Apr 12	no
9	0.75	0.75	4	4	30	no	no	June 11	N/A
10	0.75	0.75	4	4	45	yes	no	June 11	yes
11	0.50	0.50	2	8	45	no	no	June 11	N/A
12	0.50	0.75	2	8	45	yes	no	June 11	no
13	0.75	0.50	2	8	30	yes	yes (bigger than 2)	June 11	no

Can the model account for bunch blowup for 8-ns spacing, the effect of the precursor, and for no blowup with 14-ns spacing?



# Model Parameters for 2.1 GeV Q48W

**30-bunch trains, 0.75 mA/bunch**

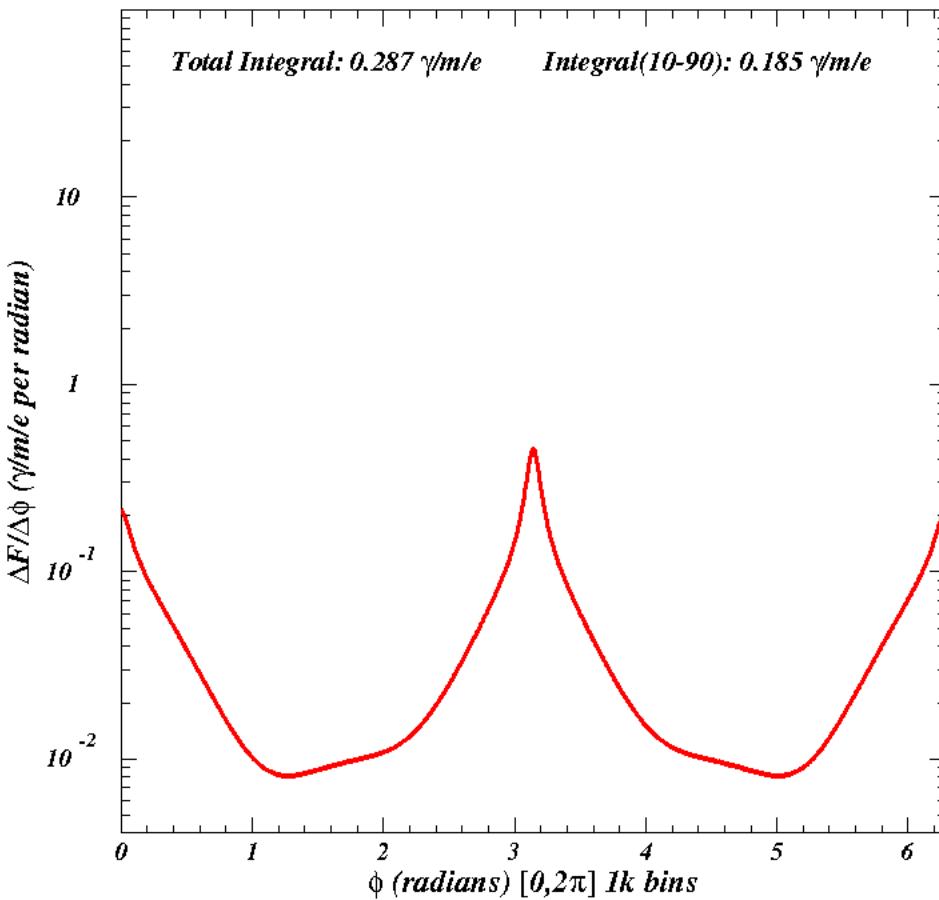
$\sigma_x = 350 \mu$ ,  $\sigma_y = 18 \mu$ ,  $\sigma_z = 9.2 \text{ mm}$

**Round aluminum v.c., diameter 89 mm**

**Field gradient 3.7 T/m**

**QE=14% (direct  $\gamma$ 's), 20% (reflected  $\gamma$ 's)**

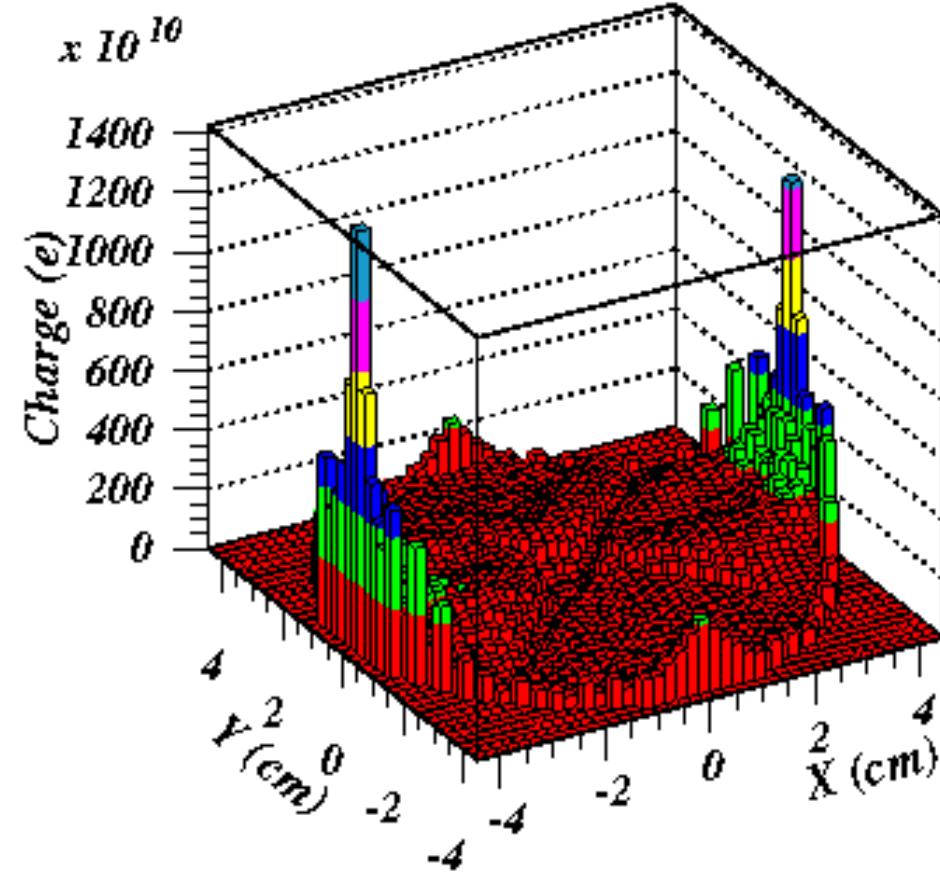
*Job 34153: Photoelectron Production Distributions*



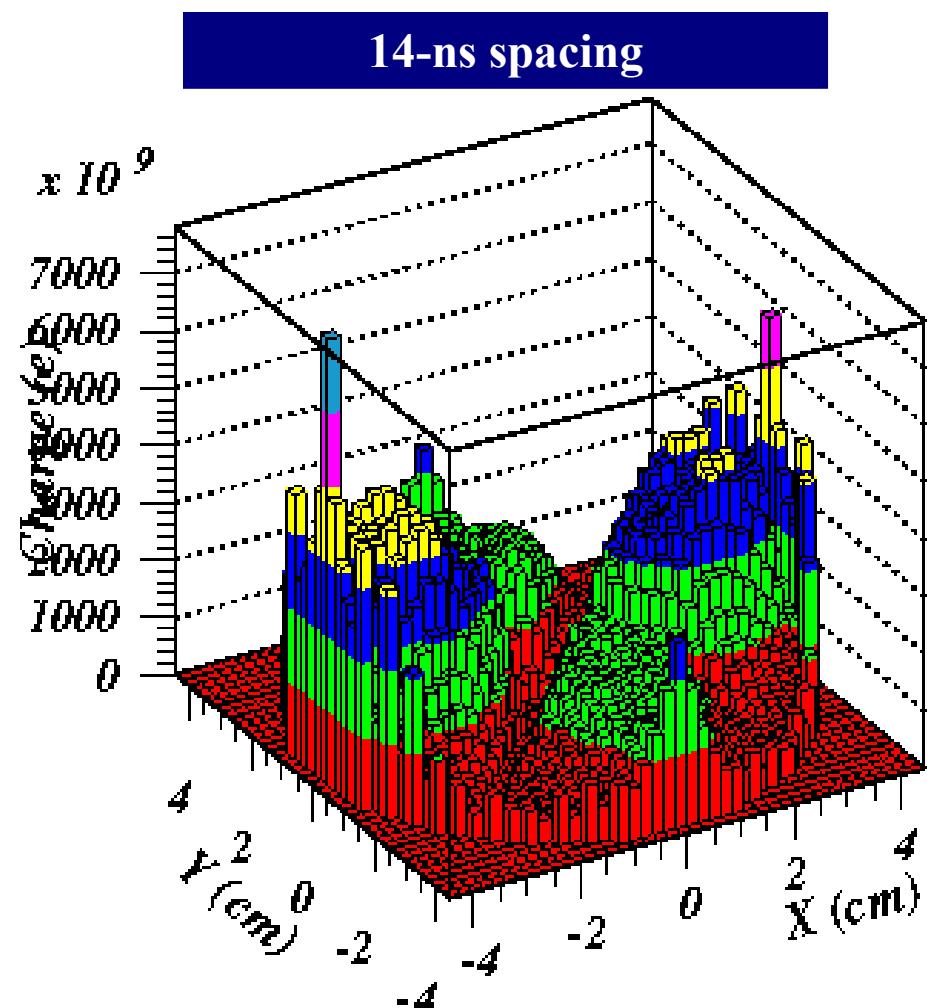
**NB: the quantum efficiency is poorly known.  
The critical energy is only 140 eV (!), so photoelectron energies of a few eV were assumed.**



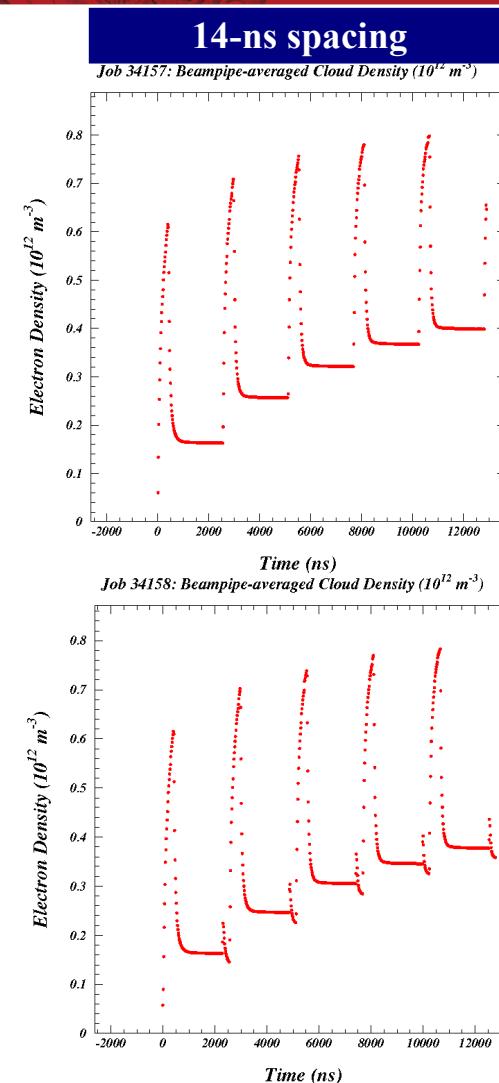
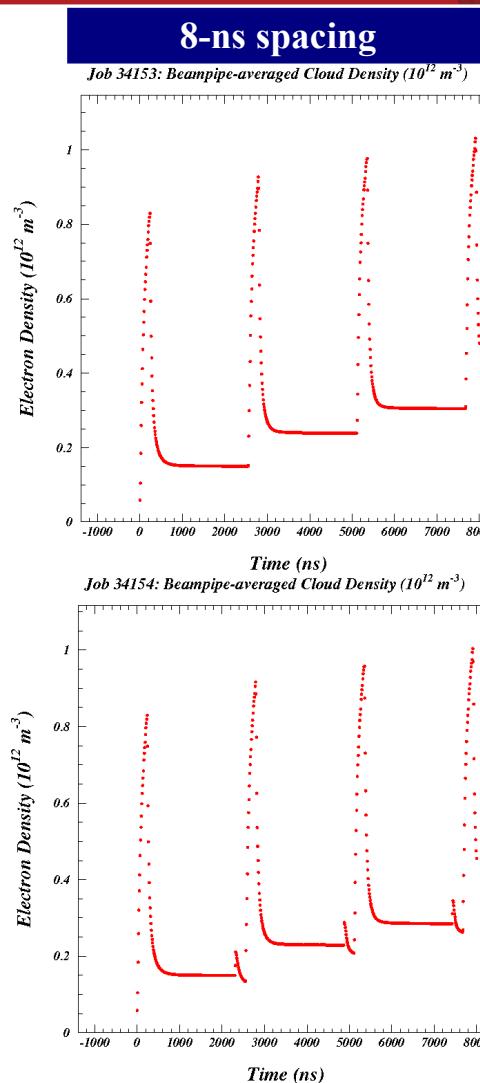
8-ns spacing



14-ns spacing



Clouds quite different. Integrated profiles do not depend on presence of precursor bunch.



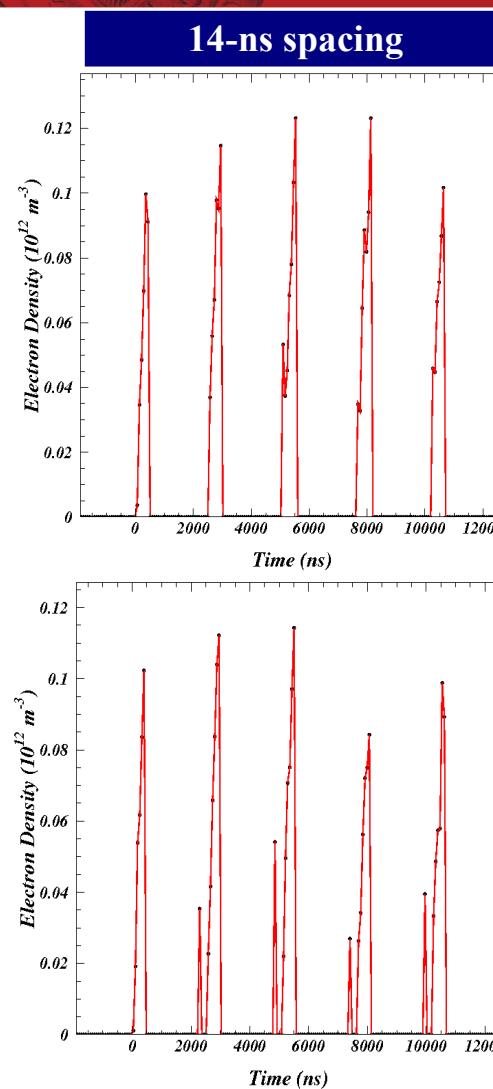
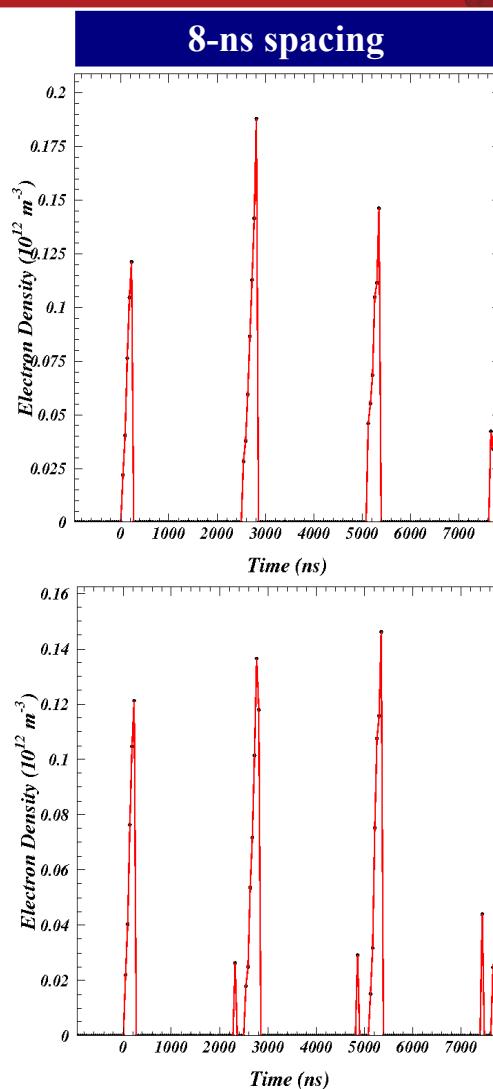
Without precursor bunch

With precursor bunch

Saturation not reached after four days of computing.



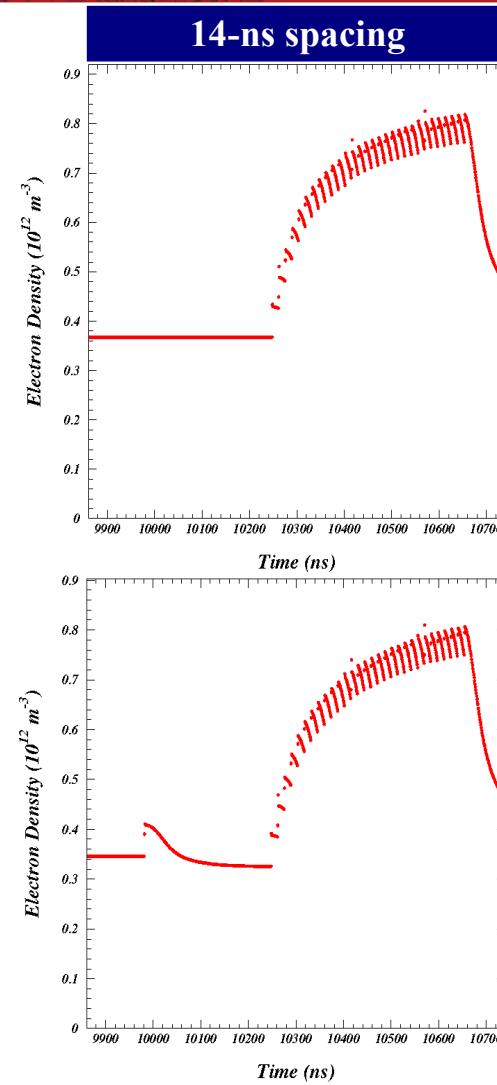
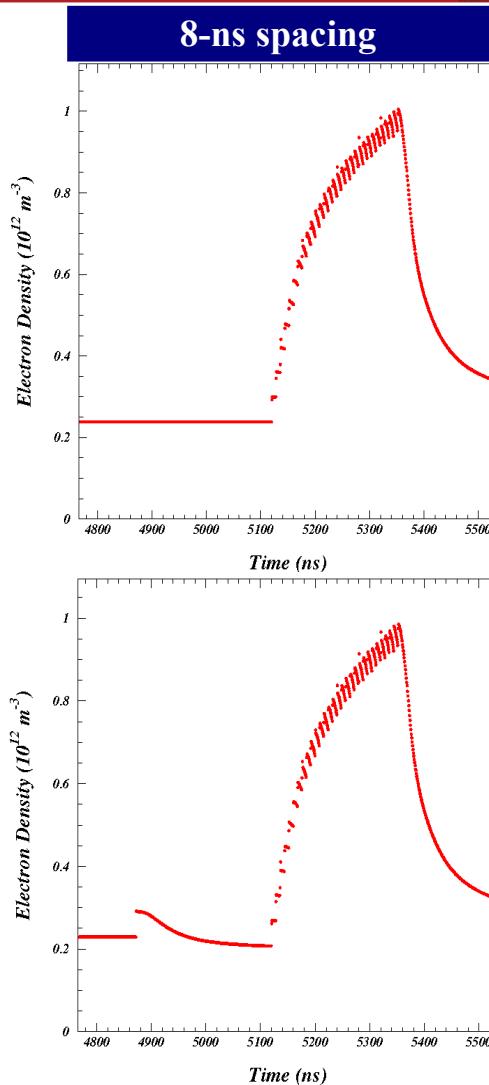
# 20-sigma cloud density prior to bunch arrival



**Without precursor bunch**

**With precursor bunch**

20-sigma density appears to saturate earlier than bp-avg. Now zoom in on the third and fifth trains.



**Without precursor bunch**

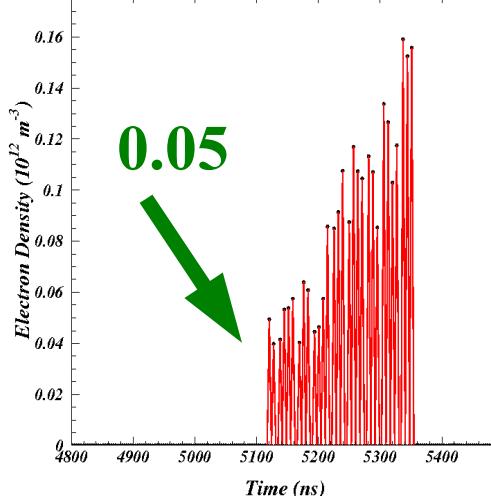
**With precursor bunch**

**Additional cloud from precursor bunch mostly gone after 250 ns.**

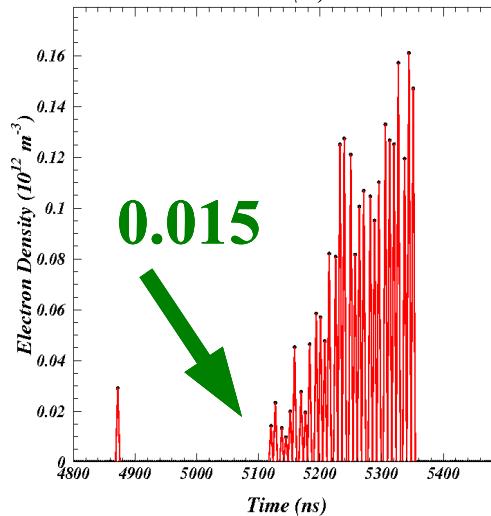


# 20-sigma cloud density prior to bunch arrival

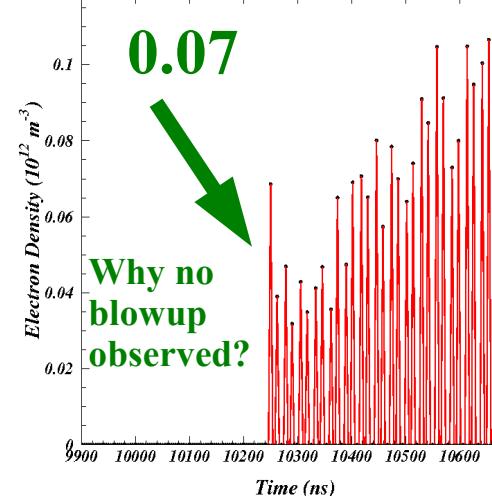
8-ns spacing third train



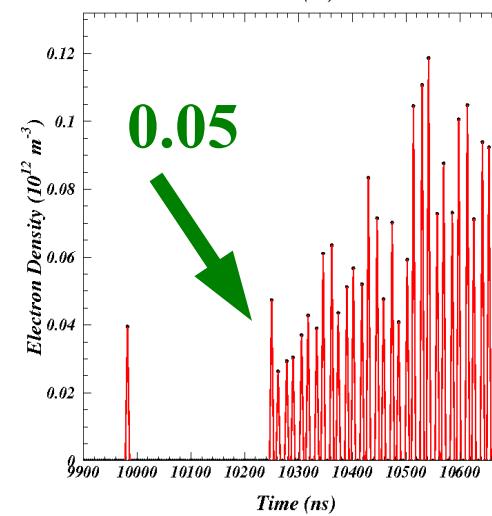
First modeling result showing reduction of cloud density by a precursor bunch!



14-ns spacing, fifth train



Without precursor bunch



With precursor bunch

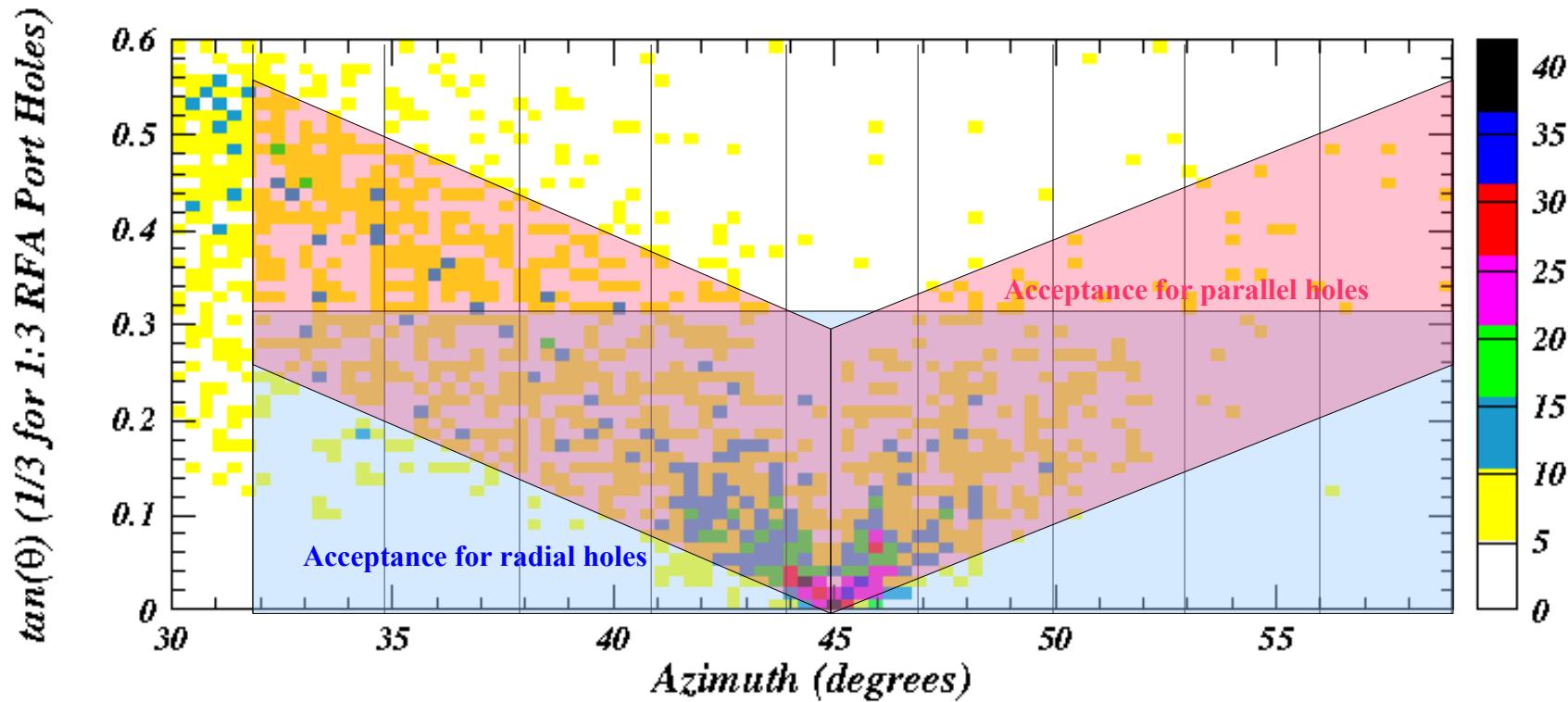
Some indication that the precursor bunch reduces cloud density more for 8-ns spacing than for 14-ns spacing.



# Time -Resolved RFA Design for Q48W

## How to orient the holes in the beampipe?

Cloud electron incident angle on wall relative to perpendicular near the pole tip at 45 degrees.  
 $\theta = 0$  corresponds to perpendicular incidence.



Collectors span about 3 degrees.  
Arrival angle along field lines has poor acceptance away from the central collector.  
Parallel holes are much better, nearly as good as parallel to the field lines.