



Cornell University
Laboratory for Elementary-Particle Physics



Two Papers for IPAC15

SYNCHROTRON RADIATION ANALYSIS OF THE SUPERKEKB POSITRON STORAGE RING

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INITIAL MODELING OF ELECTRON CLOUD BUILDUP IN THE FINAL-FOCUS QUADRUPOLE MAGNETS IN THE SUPERKEKB POSITRON RING

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-- Updated 27 April 1015 --

Jim Crittenden

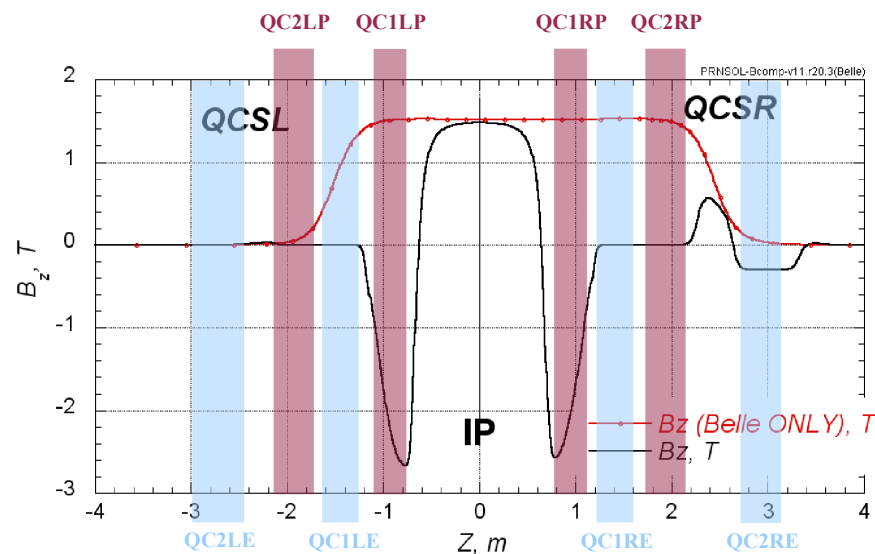
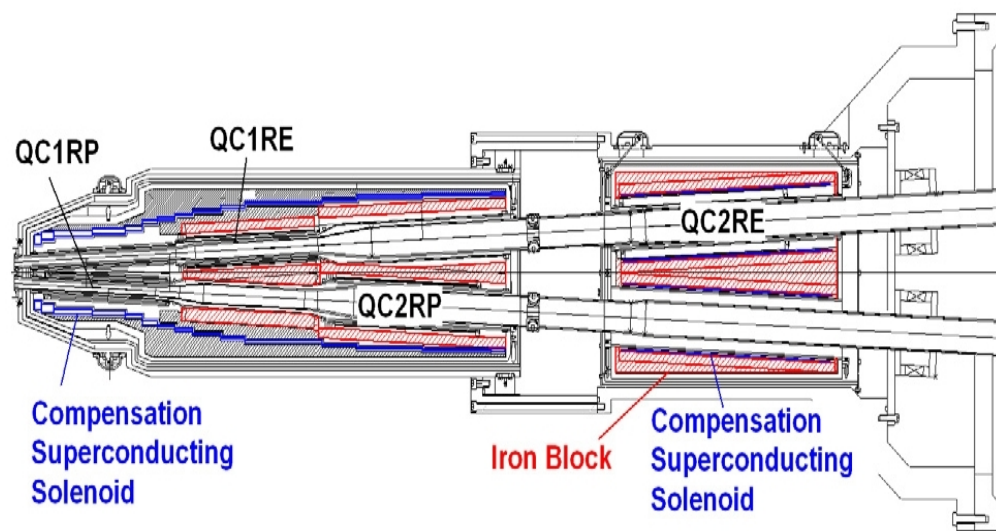
Electron Cloud Meeting

22 April 2015



SuperKEKB Interaction Region

Solenoid and quadrupole magnetic fields



Final-focus Quadrupole Magnets for the 4 GeV Positron Beam

QC2RP: 0.410 m 28 T/m

QC1RP: 0.334 m 69 T/m

QC2LP: 0.410 m 28 T/m

QC1LP: 0.334 m 69 T/m

The combined BELLE-II detector solenoid and compensation solenoid fields produce a field which varies along the length of QC1RP, ranging from about 0.6 T to about 2.4 T. The direction of the field is rotated $83/2$ mrad relative to the LER beam axis.



Electron cloud simulation package ECLCLOUD

- * Originated at CERN in the late 1990's
- * Widespread application for LHC, KEK, RHIC, ILC ...
- * Under active development at Cornell since 2008
- * Successful modeling of CESRTA tune shift measurements
- * Validated with CESRTA measurements of electron trapping in a quadrupole magnet (PRSTAB 2015)

I. Generation of photoelectrons

- A) Production energy, angle
- B) Azimuthal distribution (v.c. reflectivity)

II. Time-sliced cloud dynamics

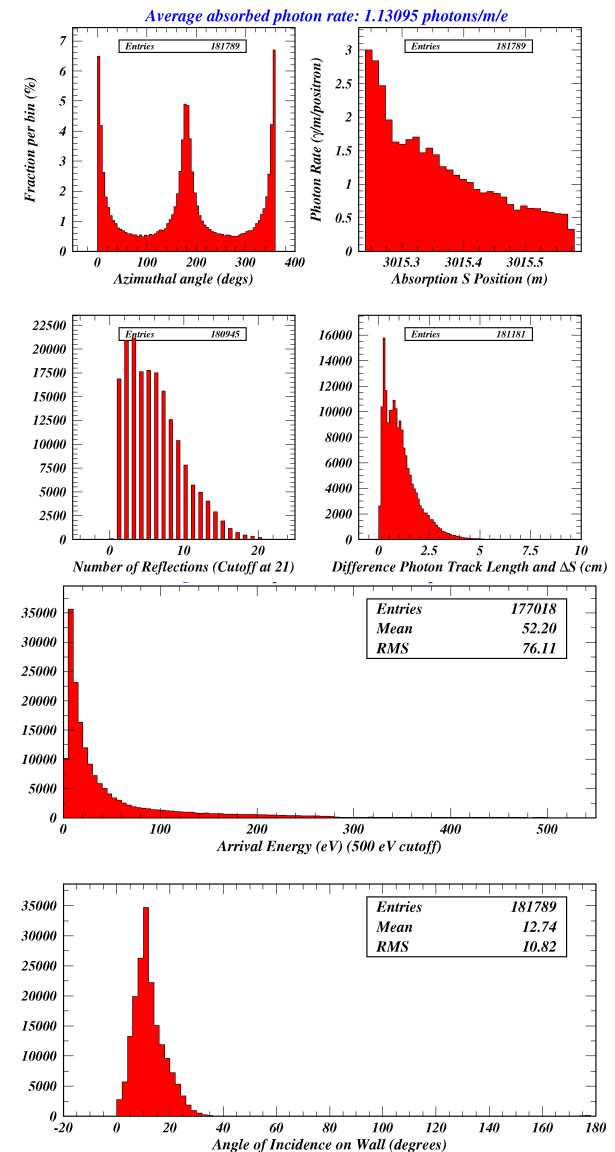
- A) Cloud space charge force
- B) Beam kick
- C) Magnetic fields

III. Secondary yield model

- A) True secondaries (yields > 1!)
- B) Rediffused secondaries (high energy)
- C) Elastic reflection (dominates at low energy)

IV. Model for a stripline detector in a quadrupole field

- A) Acceptance vs incident angle, energy, B-field
- B) Charge entering holes removed from cloud
- C) Charge hitting wall creates secondaries





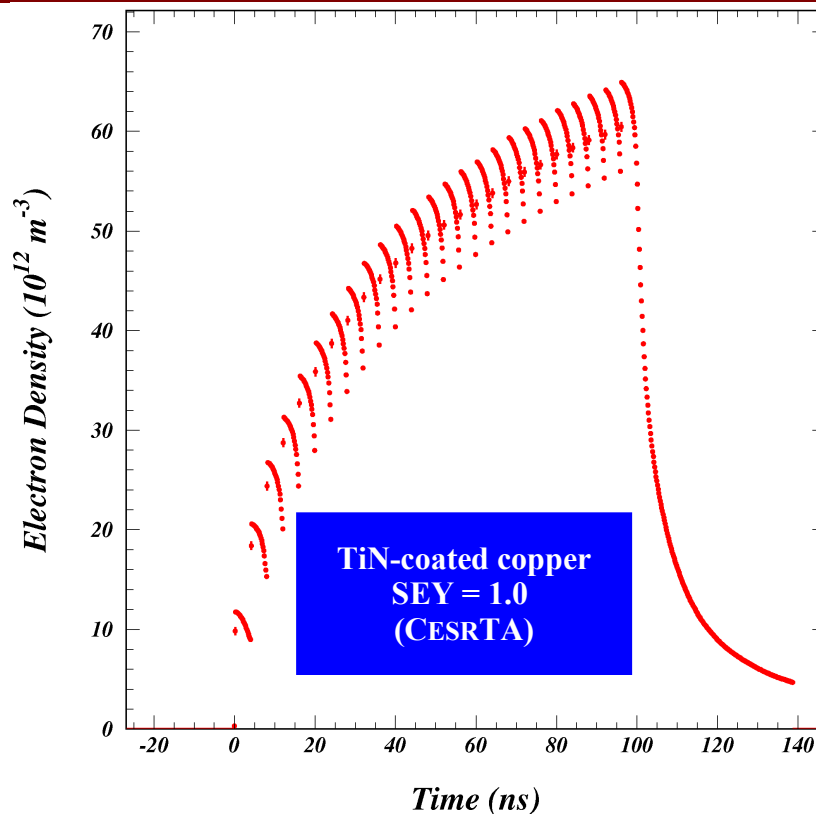
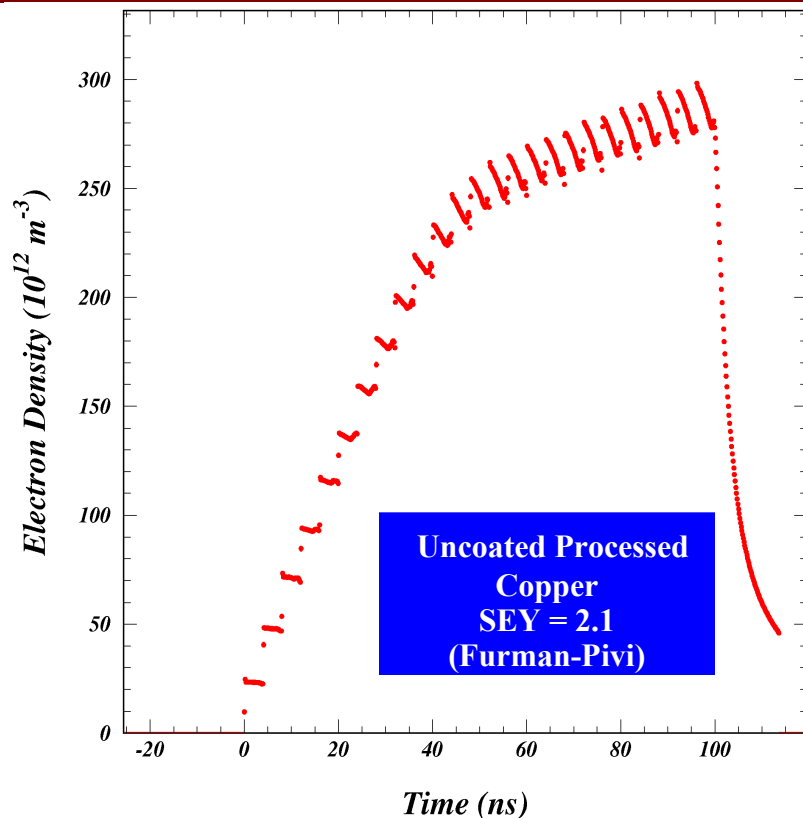
SuperKEKB Positron Ring Operating Parameters

4 GeV

2500 bunches @ 9.4×10^{10} e⁺

$\sigma_z = 6$ mm

4 ns spacing



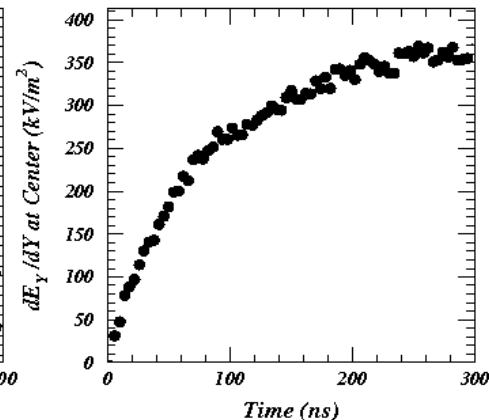
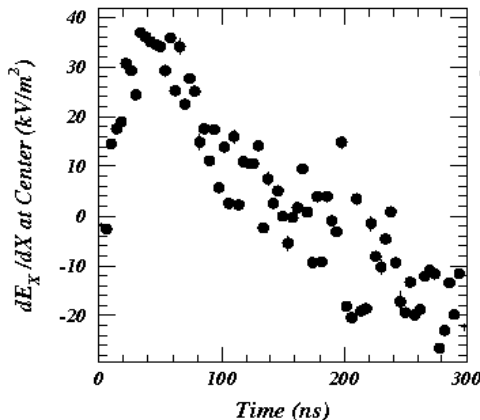
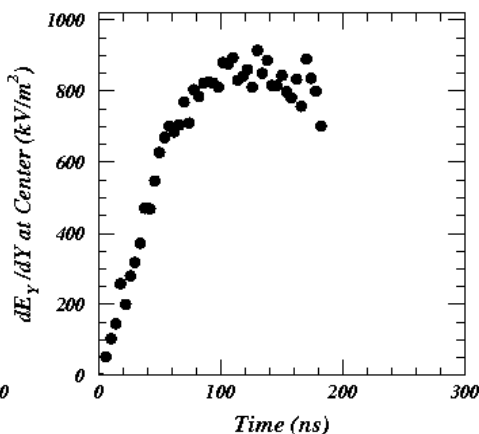
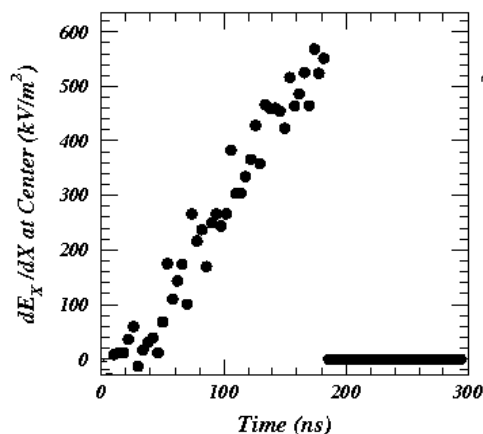
Such electron cloud densities are 3-4 orders of magnitude higher than the ring average estimated by the KEK vacuum group.



Huge Tune Shifts ! Factor >100 larger than Ohmi (IPAC14)

Uncoated Processed Copper
SEY = 2.1 (Furman-Pivi)

TiN-coating
SEY = 1.0
(CESRTA)



$$\text{Fractional tune shift} = \Delta L * dE_Y/dY * \beta_Y / (4\pi E/eV) = 2.0e-7 dE_Y/dY (V/m^2) = 0.16$$

$$(\beta_Y = 3000 \text{ m} !)$$

Ohmi calculated maximum values of 0.0009 for $\rho_e = 6e11 \text{ m}^{-3}$ about 30 m from the IP (IPAC14).

No electron cloud production in the final-focus quadrupoles was taken into account,
since no photon scattering simulation had been done.