



Rest-Frame-Radiation Power



$$\widetilde{\vec{B}}(\widetilde{t},\widetilde{\vec{x}}) = -\frac{1}{c^2} \widetilde{\Phi} \,\widetilde{\vec{n}} \times \dot{\vec{\beta}}_{r\perp} \qquad \widetilde{\vec{E}}(\widetilde{t},\widetilde{\vec{x}}) = -\frac{1}{c} \widetilde{\Phi} \,\widetilde{\vec{\beta}}_{r\perp} + \frac{1}{\widetilde{\Delta}} \widetilde{\Phi} \,\widetilde{\vec{n}}$$

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Energy flux instantanious rest frame

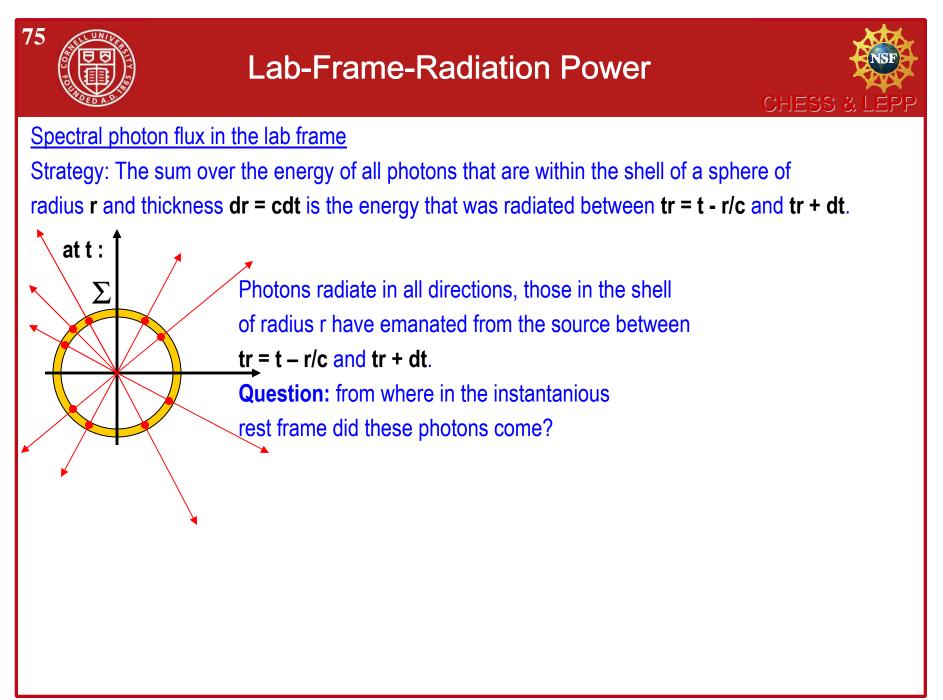
$$\begin{aligned} \widetilde{\vec{S}} &= \frac{1}{\mu_0} \widetilde{\vec{E}} \times \widetilde{\vec{B}} = \frac{1}{\mu_0} \frac{1}{c^2} \widetilde{\Phi}^2 [\frac{1}{c} \widetilde{\vec{\beta}}_{r\perp} \times (\widetilde{\vec{n}} \times \widetilde{\vec{\beta}}_{r\perp}) - \frac{1}{\Delta} \widetilde{\vec{n}} \times (\widetilde{\vec{n}} \times \widetilde{\vec{\beta}}_{r\perp})] \\ &= \frac{1}{\mu_0} \frac{1}{c^2} \widetilde{\Phi}^2 [\frac{1}{c} \widetilde{\vec{\beta}}_{r\perp}^2 \widetilde{\vec{n}} + \frac{1}{\Delta} \widetilde{\vec{\beta}}_{r\perp}^2] \end{aligned}$$

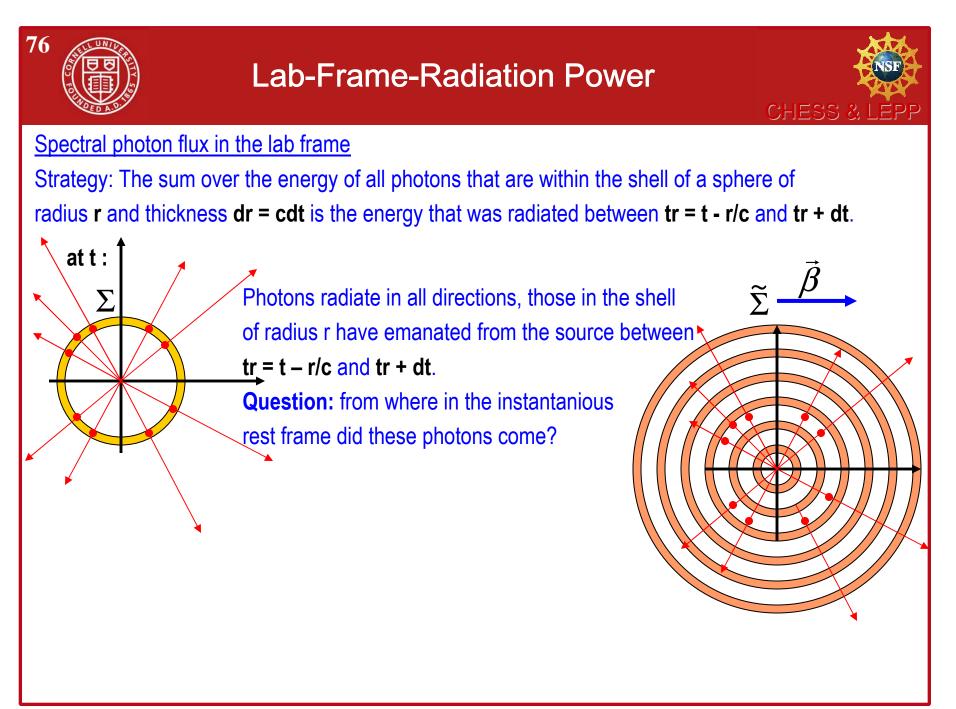
Radiation power in the instantanious rest frame

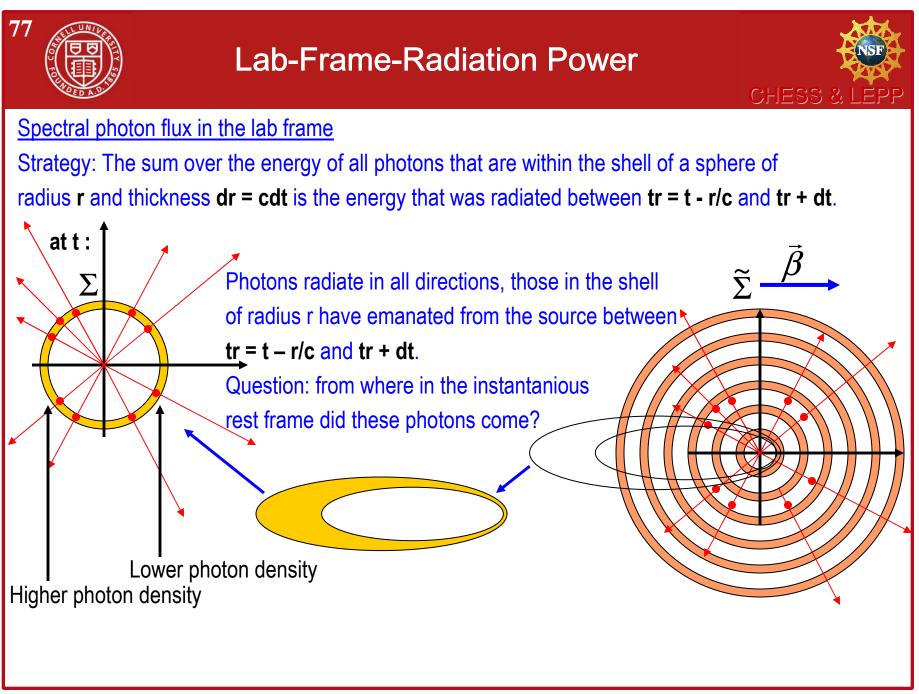
$$\widetilde{P} = \oint_{4\pi} \widetilde{\vec{S}} \cdot \widetilde{\vec{n}} \Delta^2 \sin \vartheta d\vartheta d\varphi = \frac{1}{\mu_0} \frac{\widetilde{\Delta}^2}{c^3} \widetilde{\Phi}^2 \vec{\vec{\beta}}_r^2 \oint_{4\pi} \sin^3 \vartheta d\vartheta d\varphi = \frac{2}{3c} \frac{q^2}{4\pi\varepsilon_0} \vec{\vec{\beta}}_r^2$$

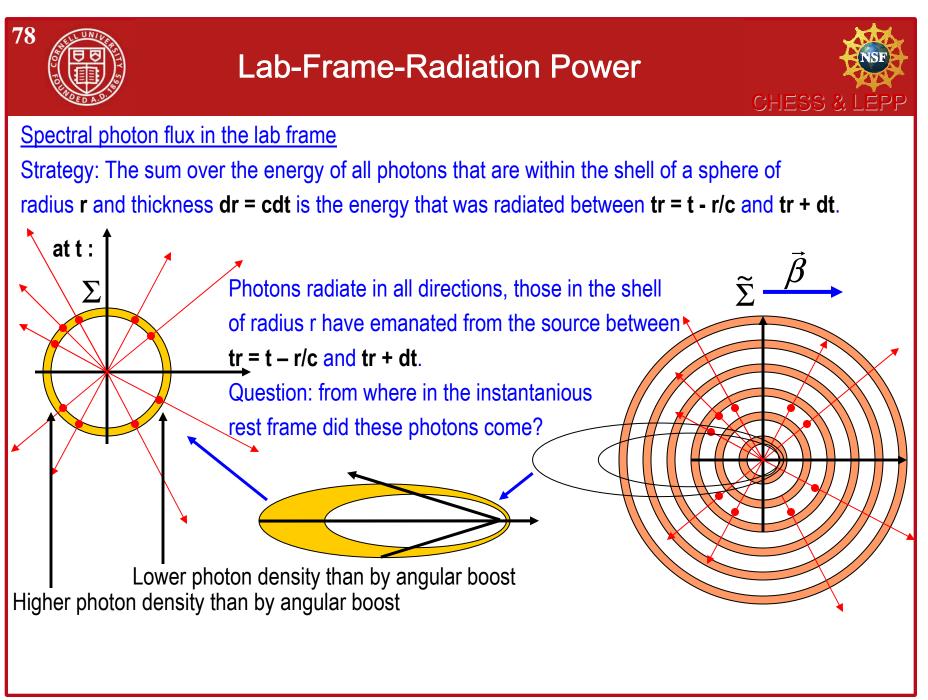
Now find the radiation power P in the lab frame Σ

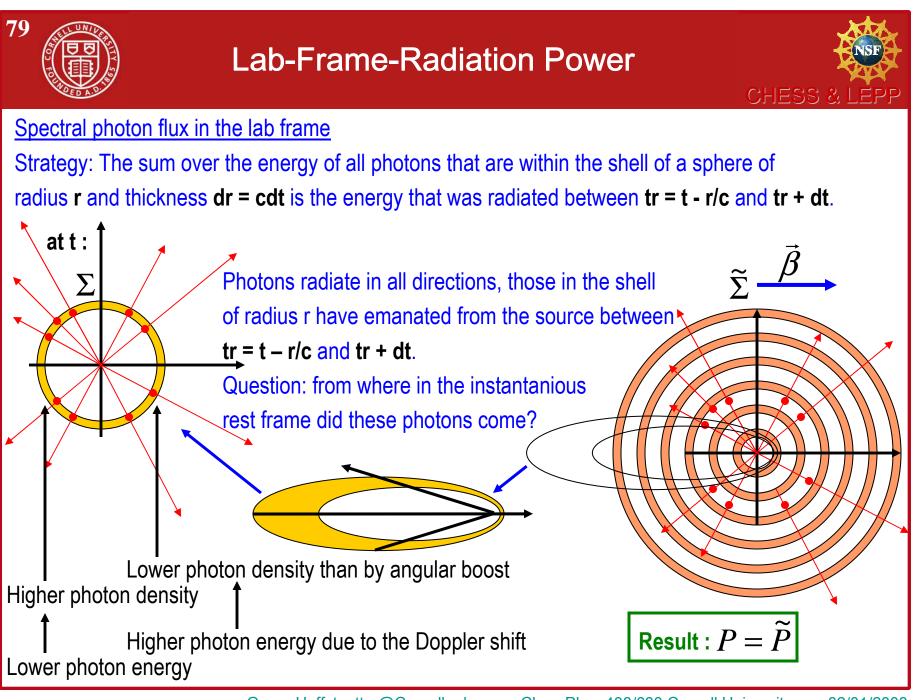
Option 1: Derive the much more complicated form of $\vec{E}, \vec{B}, \vec{S}$ and integrate over \vec{S} . Option 2: Express P in terms of 4-vectors and tensors and see how it transforms. Option 3: Determine how photons change their frequency by the Lawrence transformation (via Doppler shift) and sum over all photons.











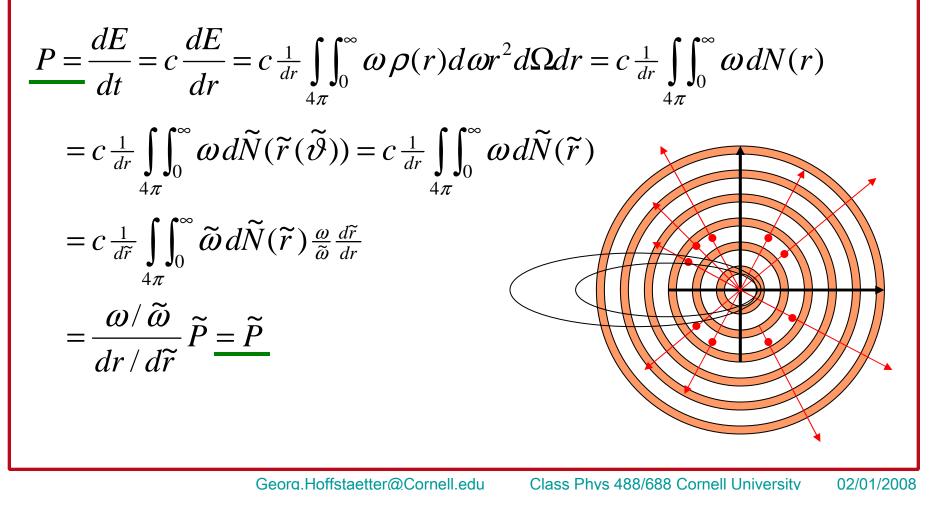


Lab-Frame-Radiation Power



Spectral photon flux in the lab frame

Strategy: The sum over the energy of all photons that are within the shell of a sphere of radius **r** and thickness $d\mathbf{r} = \mathbf{c}d\mathbf{t}$ is the energy that was radiated between $\mathbf{tr} = \mathbf{t} + \mathbf{r/c}$ and $\mathbf{tr} + d\mathbf{t}$.







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Strategy: The sum over the energy of all photons that are within the shell of a sphere of radius **r** and thickness $d\mathbf{r} = \mathbf{c}d\mathbf{t}$ is the energy that was radiated between $\mathbf{tr} = \mathbf{t} + \mathbf{r}/\mathbf{c}$ and $\mathbf{tr} + d\mathbf{t}$.

$$\underline{P} = \frac{dE}{dt} = c \frac{dE}{dr} = c \frac{1}{dr} \iint_{4\pi}^{\infty} \omega \rho(r) d\omega r^2 d\Omega dr = c \frac{1}{dr} \iint_{4\pi}^{\infty} \omega dN(r)$$

$$= c \frac{1}{dr} \iint_{4\pi}^{\infty} \omega d\tilde{N}(\tilde{r}(\tilde{\vartheta})) = c \frac{1}{dr} \iint_{4\pi}^{\infty} \omega d\tilde{N}(\tilde{r}) = c \frac{1}{d\tilde{r}} \iint_{4\pi}^{\infty} \tilde{\omega} d\tilde{N}(\tilde{r}) \frac{\omega}{\tilde{\omega}} \frac{d\tilde{r}}{dr}$$

$$= \frac{\omega/\tilde{\omega}}{dr/d\tilde{r}} \tilde{P} = \underline{\tilde{P}}$$
At the surface of the sphere with radius $\mathbf{r} = \mathbf{c}(\mathbf{t} - \mathbf{tr})$:
$$\begin{pmatrix} ct \\ \vec{r} \end{pmatrix} = \begin{pmatrix} r \\ r\vec{n} \end{pmatrix} \text{ and } \begin{pmatrix} \frac{1}{c}\omega \\ \vec{k} \end{pmatrix} = \frac{1}{c} \begin{pmatrix} \omega \\ \omega\vec{n} \end{pmatrix}$$
The Lawrence transformation thus leads to the same linear transformation for \mathbf{r} and ω , and therefore $\frac{\tilde{\omega}}{\omega} = \frac{d\tilde{r}}{dr}$



Synchrotron Radiation Power



Power in the instantanious rest frame:
$$\tilde{P} = \frac{2r_e}{3c} mc^2 \dot{\beta}_r^2$$

Because it equals the power in the lab frame, there must be a Lorentz-invariant quantity that has the value $\tilde{\beta}_r^2$ in the rest frame.

The 4-vector of acceleration:
$$b^{\mu} = \frac{d}{d\tau} P^{\mu} = (\frac{1}{c} \frac{d}{d\tau} E, \frac{d}{d\tau} \vec{p})$$

$$b^{\mu} = \frac{d}{d\tau} P^{\mu} = \gamma \frac{d}{dt} P^{\mu} = \gamma (\frac{1}{c} \dot{E}, \dot{\vec{p}}) \quad \Rightarrow \quad \gamma^2 (\dot{\vec{p}}^2 - \frac{1}{c^2} \dot{E}^2) = \text{const.}$$

Evaluation in the instantanious rest frame:

$$\left[\frac{d}{dt}\gamma\right]_{\vec{\beta}=0} = 0 \quad \Rightarrow \quad \left[\gamma^2(\dot{\vec{p}}^2 - \frac{1}{c^2}\dot{E}^2)\right]_{\vec{\beta}=0} = (mc)^2\dot{\beta}_r^2$$

Radiation power in the laboratory frame

$$P = \frac{2r_e}{3mc} \gamma^2 (\dot{\vec{p}}^2 - \frac{1}{c^2} \dot{\vec{E}}^2)$$





$$P = \frac{2r_e}{3mc} \gamma^2 (\dot{\vec{p}}^2 - \frac{1}{c^2} \dot{\vec{E}}^2)$$

Transverse acceleration:

Magnetic:
$$P = \frac{2r_e q^2}{3mc} \gamma^2 (\vec{v} \times \vec{B})^2 = \frac{2r_e q^2}{3mc} v^2 \gamma^2 B_{\perp}^2 = \frac{2r_e}{3m^3 c} \frac{p^4}{\rho^2} = \frac{2r_e c}{3} mc^2 \frac{(\beta\gamma)^4}{\rho^2}$$

Electric: $\vec{E} \perp \vec{v} \implies P = \frac{2r_e q^2}{3mc} \gamma^2 \vec{E}^2 = \frac{2r_e c}{3} mc^2 \frac{(\beta\gamma)^4}{\rho^2}$

Relative loss per turn: $\eta = \frac{2\pi\rho}{v} P \frac{1}{m\gamma c^2} = \frac{4\pi}{3} \frac{r_e}{\rho} (\beta\gamma)^3$ Example: $\rho = 60$ m, E = 5GeV $\rightarrow \eta = 2 \cdot 10^{-4}$

Longitudinal acceleration:

$$P = \frac{2r_e q^2}{3mc} \gamma^2 [\vec{E}^2 - \frac{1}{c^2} (\vec{E} \cdot \vec{v})^2] = \frac{2r_e q^2}{3mc} \vec{E}^2$$

Efficiency: $\eta = \frac{P}{q\vec{v}\cdot\vec{E}} = \frac{2}{3\beta} \frac{r_e |q\vec{E}|}{mc^2}$ Example: $|\vec{E}| = 30 \frac{\text{MeV}}{\text{m}} \to \eta = 10^{-13}$

84	Examples			S & LEPP
CESR @ CORNELL:	5.3GeV	ρ = 60m	η = 0.02%	
PETRA @ DESY:	23.5GeV	ρ = 195m	η = 0.6%	
LEP @ CERN:	105GeV	ho = 3000m	η = 3.4%	
HERA-e @ DESY:	27.5GeV	ρ = 608m	η = 0.3%	
HERA-p @ DESY:	920GeV	ho = 582m	η = 2.e-8	
For the same momentum, the energy loss scales with the third power of the mass, i.e. Electrons radiate 6.e9 times more than protons. Georg.Hoffstaetter@Cornell.edu Class Phys 488/688 Cornell University 02/01/2008				