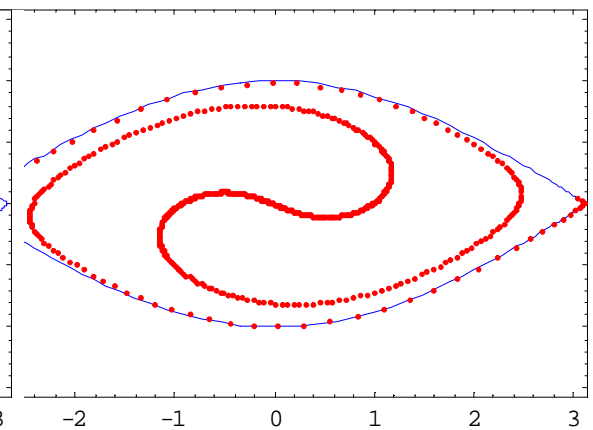
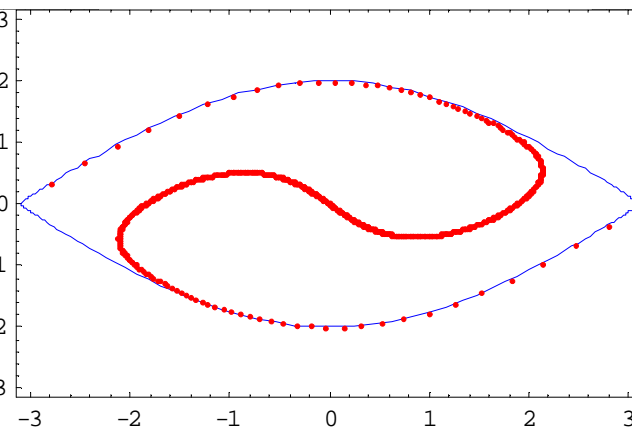
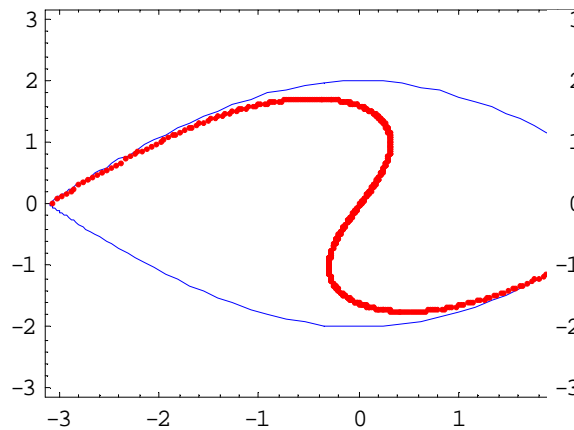
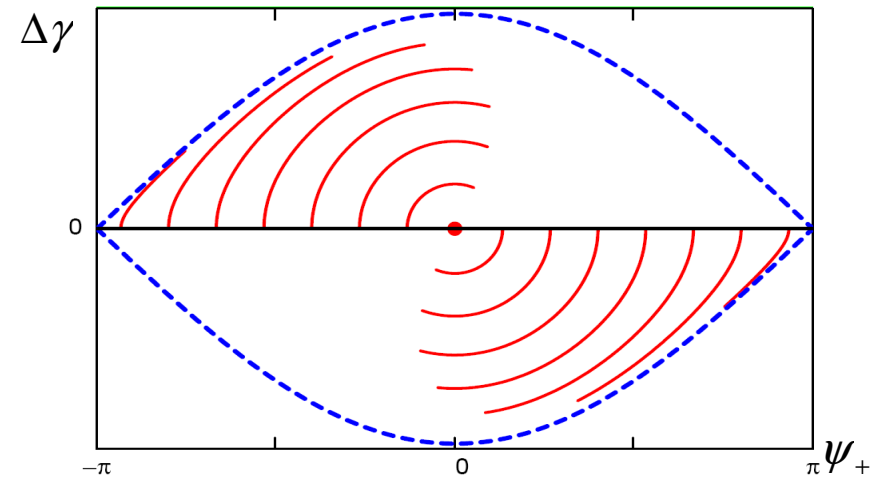
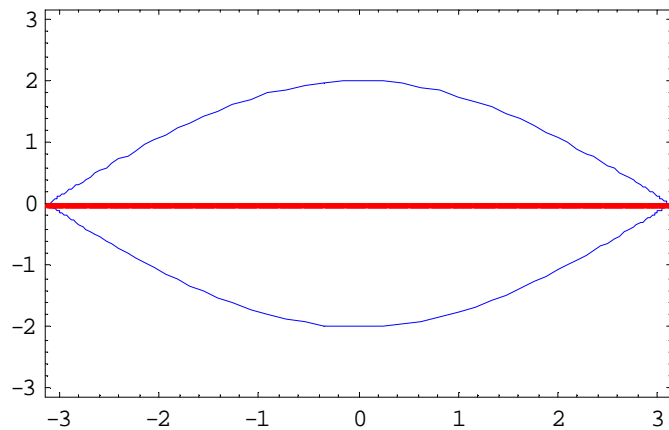




Energy Loss during Phase Space Motion



CHESS & LEPP



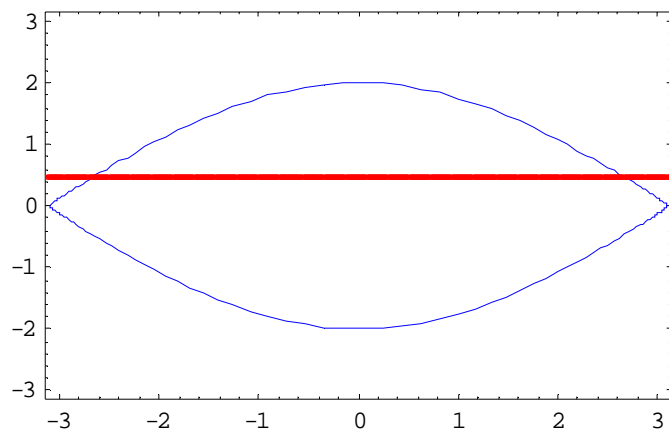
For a mono-energetic beam with $\Delta\gamma=0$, no net energy loss and beam filaments within the phase space separatrix.



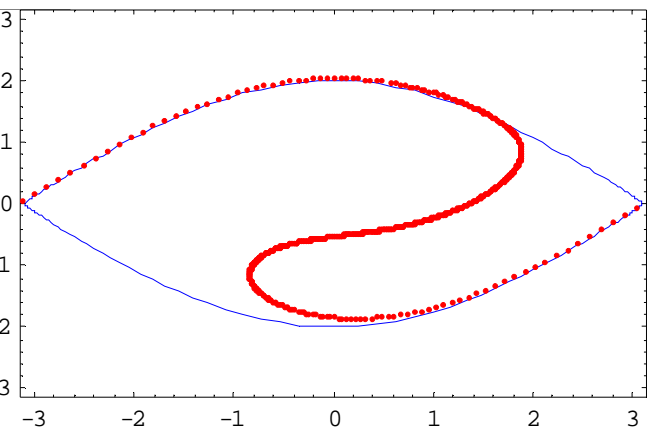
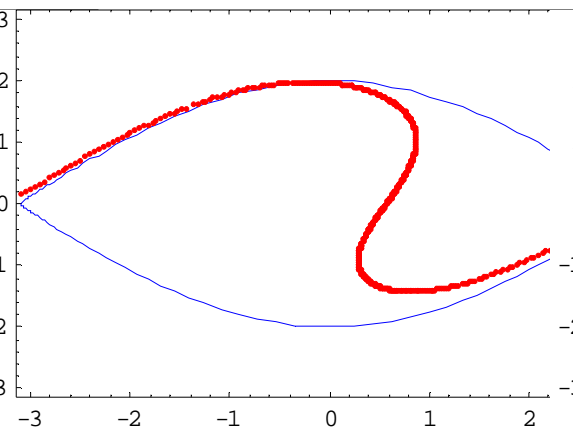
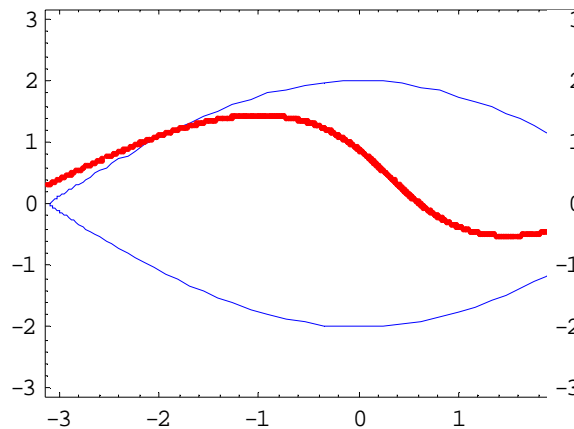
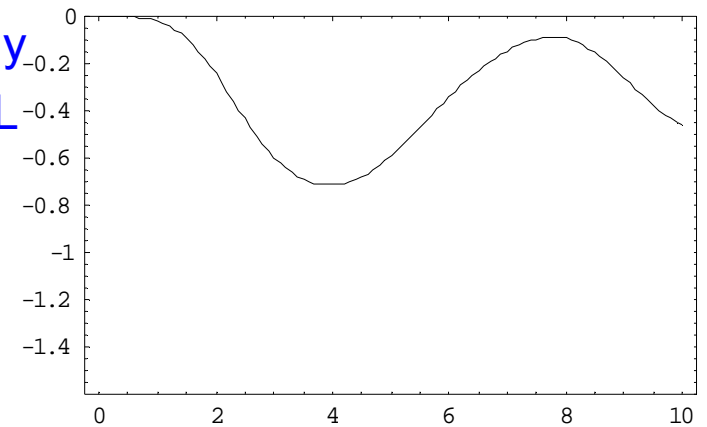
Energy Loss during Phase Space Motion



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Average energy
loss along FEL



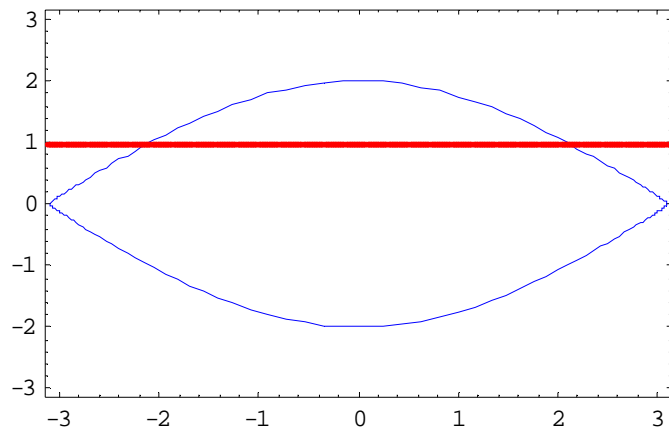
For a mono-energetic beam with $\Delta\gamma > 0$, net energy loss and beam filaments within the phase space separatrix.



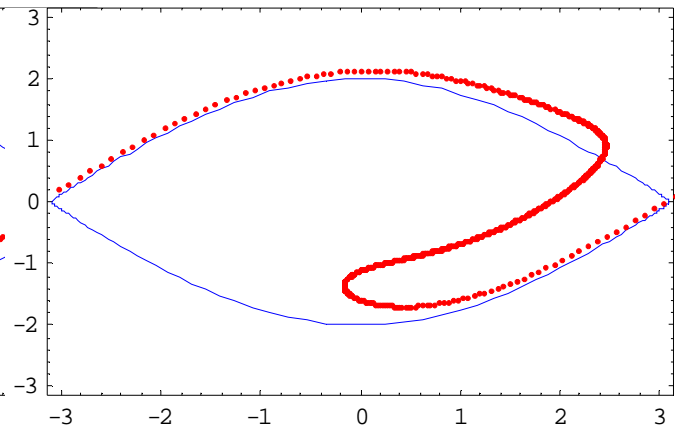
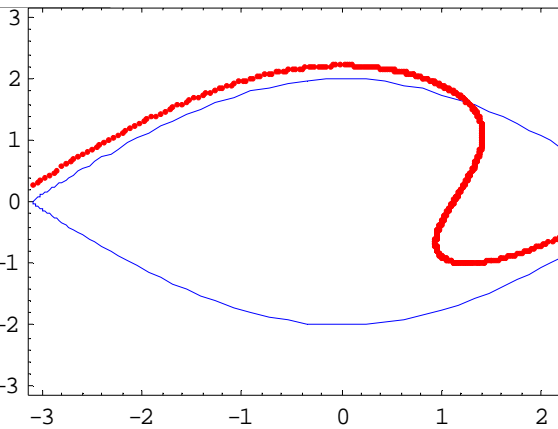
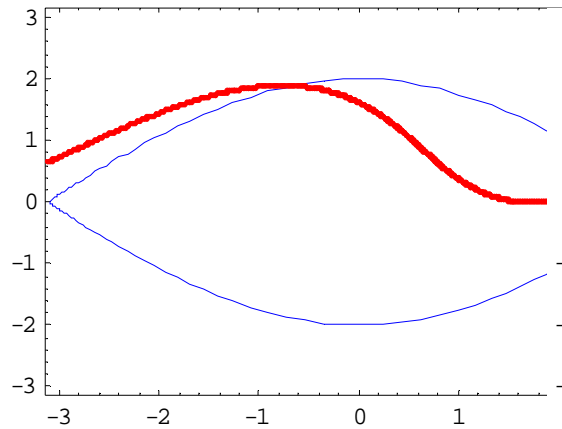
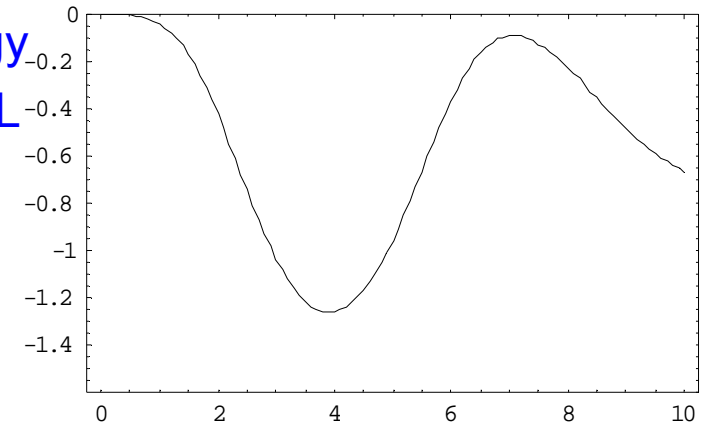
Energy Loss during Phase Space Motion



CHESS & LEPP



Average energy
loss along FEL



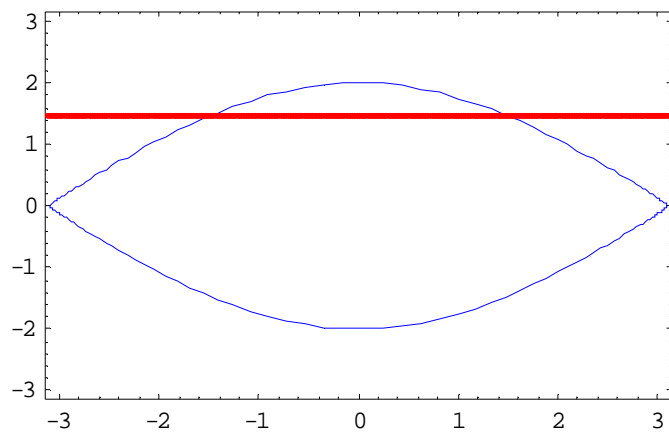
For a mono-energetic beam with $\Delta\gamma > 0$, net energy loss and beam filaments within the phase space separatrix.



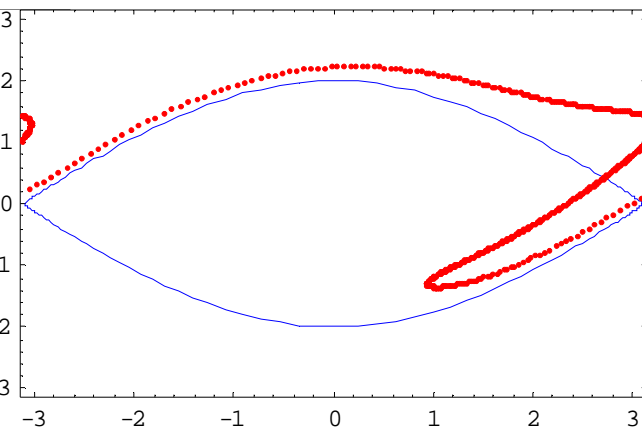
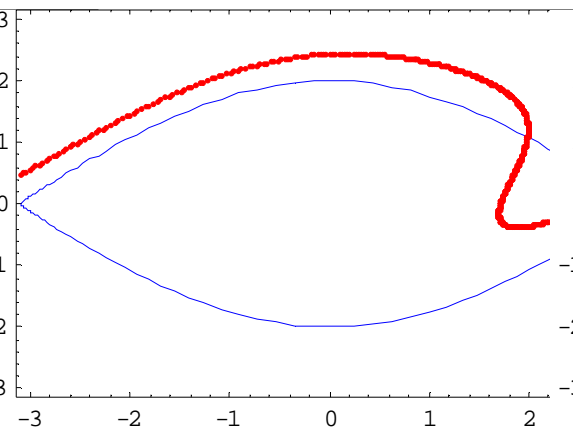
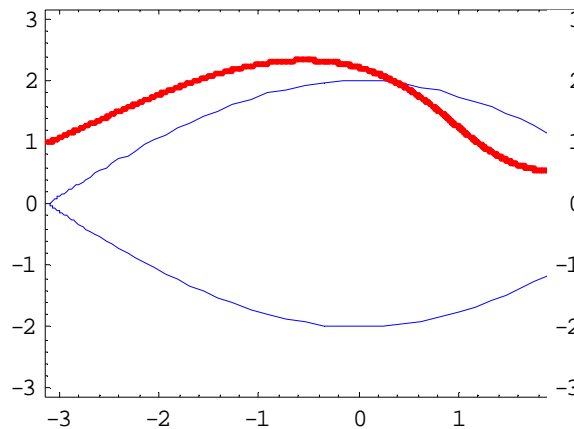
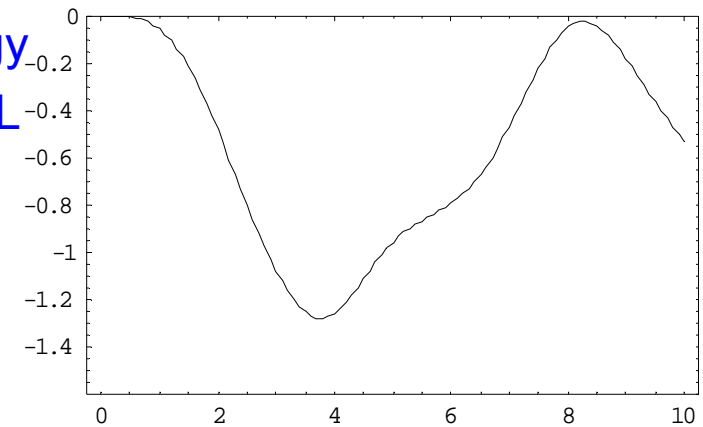
Energy Loss during Phase Space Motion



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Average energy
loss along FEL



For a mono-energetic beam with $\Delta\gamma > 0$, net energy loss and beam filaments within the phase space separatrix.



Change in Average Energy



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Wave number for small amplitudes: $\Omega = \frac{k_u}{\gamma} \sqrt{K_L K}$

$$\frac{d}{dz} \psi_+ = \frac{2k_u}{\gamma} \Delta\gamma + O^5\left(\frac{1}{\gamma}\right), \quad \frac{d}{dz} \Delta\gamma = -\frac{K_L k_u K}{2\gamma} \sin \psi_+ + O^3\left(\frac{1}{\gamma}\right)$$

$$\frac{d^2}{dz^2} \Delta\gamma = -\Omega^2 \Delta\gamma \cos \psi_+ + O^3\left(\frac{1}{\gamma}\right)$$

$$\frac{d^3}{dz^3} \Delta\gamma = \Omega^2 \left(\frac{2k_u}{\gamma} \Delta\gamma^2 \sin \psi_+ + \frac{K_L k_u K}{2\gamma} \sin \psi_+ \cos \psi_+ \right) + O^3\left(\frac{1}{\gamma}\right)$$

$$\left\langle \frac{d^4}{dz^4} \Delta\gamma \right\rangle_{\psi_+} = -\Omega^4 \Delta\gamma + O^3\left(\frac{1}{\gamma}\right)$$

Energy initially changes very slowly, only with z^4 .

How much energy can maximally be lost is determined by higher derivatives, i.e. by nonlinear terms in $\Delta\gamma$.

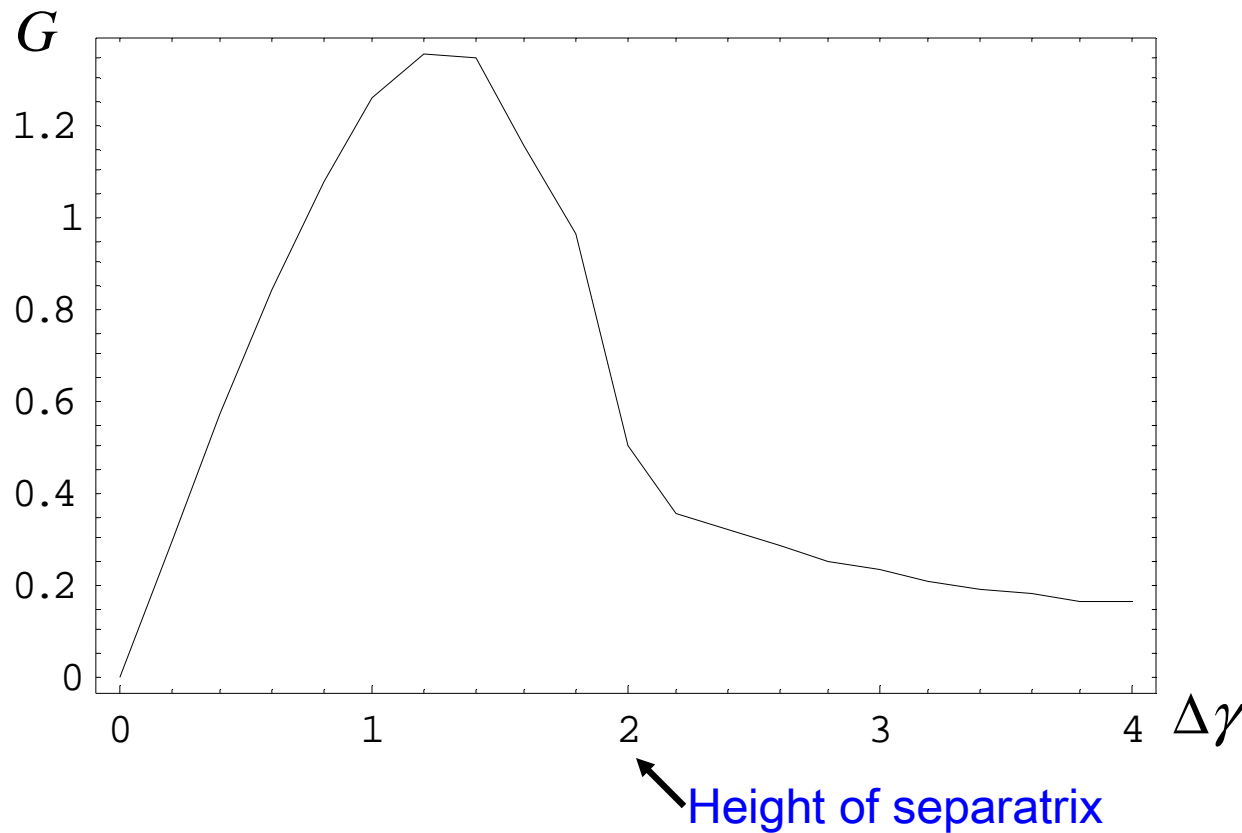
The saturation length of the FEL, i.e. when the maximal energy loss occurs, is also determined by nonlinear terms.



Gain of an FEL



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The Gain describes how much the field amplitude is amplified during one pass.
It is a function of $\Delta\gamma$.

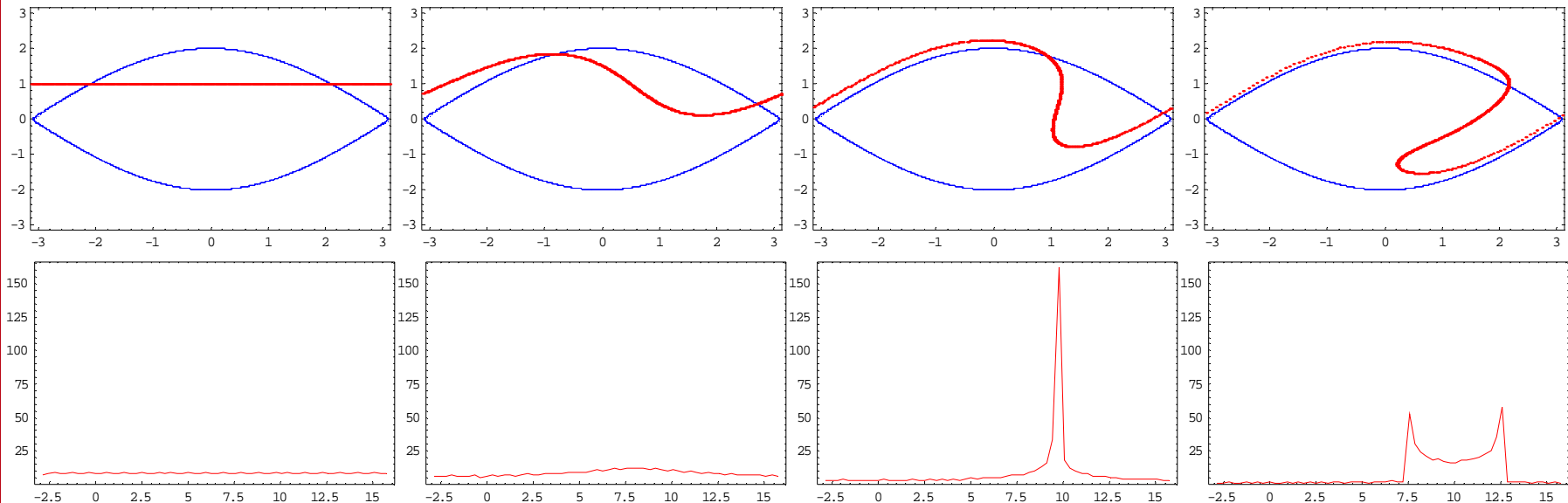


Microbunching



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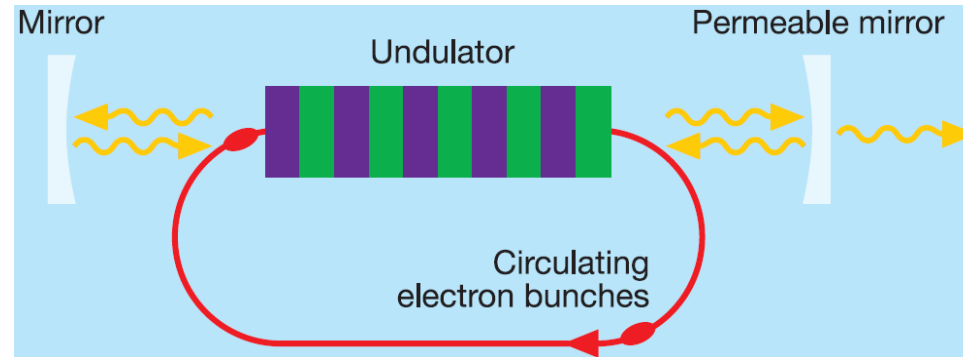
The FEL theory so far was for **weak amplification**. It has been considered how the field changes the electron distribution, but not how the electron distribution subsequently changes the field by radiation. This is the domain of **strong amplification** and needs a combined theory of radiation field and electron density.



- 1) Electron density becomes microbunched on the light wavelength! (Here by 20)
- 2) The microbunches radiate very strongly, as if they were one particle with N charges, radiating coherently N times the incoherent power !
- 3) The initial energy spread has to be very small to produce strong bunching.

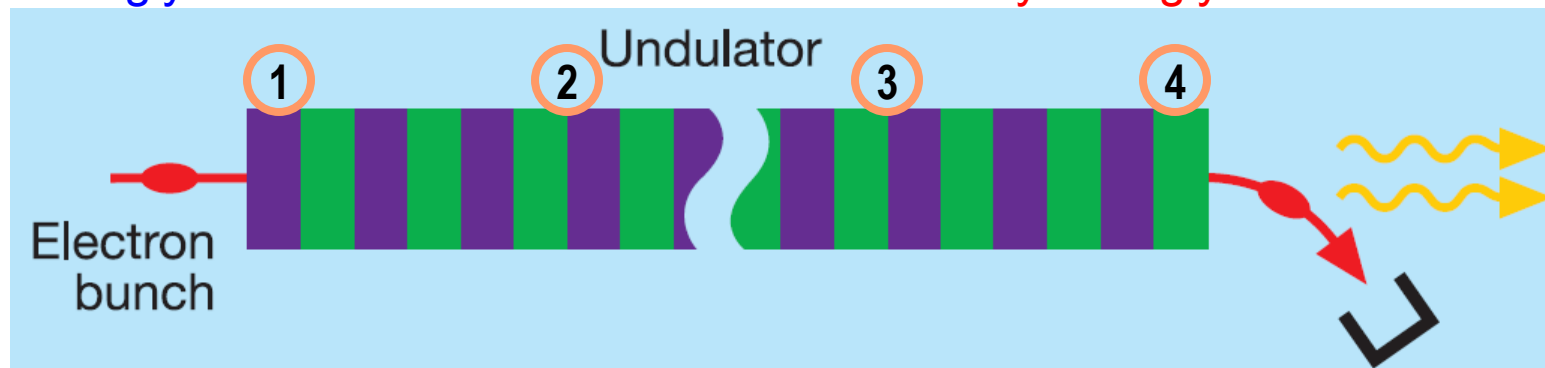
SASE FELs

Oscillator FEL: can be described well by the **weak amplification theory**.



SASE FEL: Self Amplification of Spontaneous Emission

- 1) Incoherent undulator radiation from each electron is produced
- 2) This radiation leads to weak microbunching
- 3) Microbunches radiate strongly and produce strong microbunching
- 4) Strongly microbunched beam **radiates extremely strongly!**



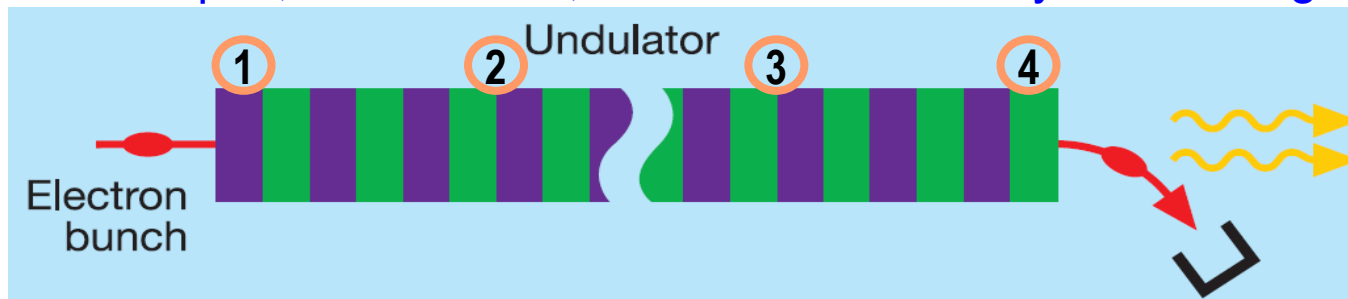


SASE FELs



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- 1) Undulators have to be very long (order 100m).
- 2) Current within an electron bunch has to be very large to produce enough field for bunching from initial incoherent radiation.
- 3) The phase of the initial incoherent radiation from (1) is determined by a statistical fluctuation of the density in the electron bunch. (A DC current would not radiate !)
The radiation process therefore start from noise.
- 4) The length has to be matched to saturation of the energy loss for maximum power.
- 5) The radiation power growth exponentially with length until saturation.
- 6) The start from noise lets the intensity fluctuate strongly, except at saturation.
- 7) The undulator has to have saturation length, where the power gets very large, destroying sample by the radiation from a single electron bunch.
- 8) SPRING8 / Japan, SLAC / USA, and DESY / Germany are building SASE FELs



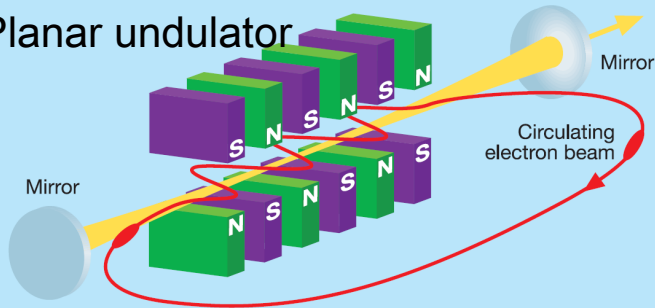


Higher Harmonics in an FEL



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Planar undulator



Only the fundamental wavelength $\lambda_L = \lambda_u \left(\frac{1}{\beta_z} - 1 \right)$ was discussed for an FEL.

The undulator radiate higher harmonics; can these Also be amplified by an E&M wave?

Change of particle energy due to the E&M wave (stimulated emission):

$$\frac{dE}{dt} = -qc \frac{E_{L0}K}{2\gamma} (\sin \Psi_+ + \sin \Psi_-), \quad \Psi_{\pm} = k_L [z - ct] + \phi_0 \pm k_u z$$

The energy transfer is large when $\frac{d}{dt} \Psi_+ = 0$ or $\frac{d}{dt} \Psi_- = 0 \Rightarrow \lambda_L = \lambda_u \left(\frac{1}{\beta_z} - 1 \right)$

But $z(t) = \bar{\beta}_z ct + \left(\frac{K}{2\gamma} \right)^2 \frac{1}{2k_u} \sin(2k_u \bar{\beta}_z ct) + O^4 \left(\frac{1}{\gamma} \right)$

$$\begin{aligned} \frac{dE}{dt} &= -qc \frac{E_{L0}K}{2\gamma} \{ \sin[\Psi_{1+} + (k_L + k_u) \hat{z} \sin(2k_u \bar{\beta}_z ct)] + \sin \Psi_- \} \\ &= -qc \frac{E_{L0}K}{2\gamma} \{ \sin[\Psi_{1+}] \cos[(k_L + k_u) \hat{z} \sin(2k_u \bar{\beta}_z ct)] + \cos[\Psi_{1+}] \sin[\dots] + \sin \Psi_- \} \\ &= -qc \frac{E_{L0}K}{2\gamma} \left\{ \sum_n A e^{\pm i \Psi_{1+}} e^{in 2k_u \bar{\beta}_z ct} + \sum_n A e^{\pm i \Psi_{1-}} e^{in 2k_u \bar{\beta}_z ct} \right\} \end{aligned}$$

Energy transfer can also be Large when one phase is constant: $k_L c (\bar{\beta}_z - 1) \pm k_u \bar{\beta}_z (1 + 2n) = 0 \Rightarrow \lambda_L = \frac{1}{n_{\text{odd}}} \lambda_u \left(\frac{1}{\beta_z} - 1 \right)$

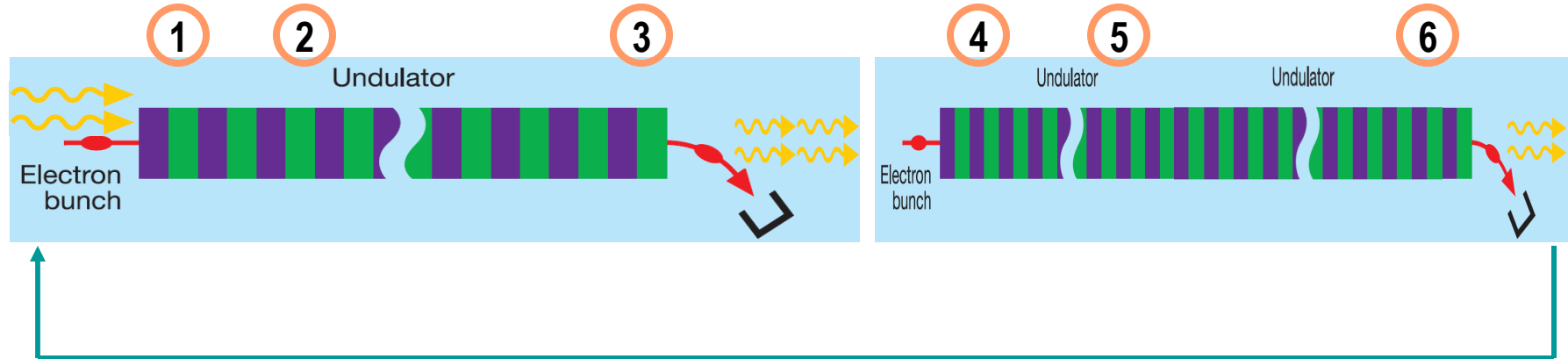


High Gain Harmonic Generation (HGHG)



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HGHG FELs avoid starting a SASE FEL from noise by seeding with a laser beam.



- 1) A beam from a quantum laser, often frequency multiplied by nonlinear media is sent into an undulator matched to the fundamental frequency of the e-beam.
- 2) Bunching of the e-beam in phase with the laser
- 3) Strong radiation of the bunches, including in higher harmonics, e.g. 3
- 4) This radiation is sent into an undulator with the 3rd harmonic as its fundamental
- 5) Bunching
- 6) Radiating

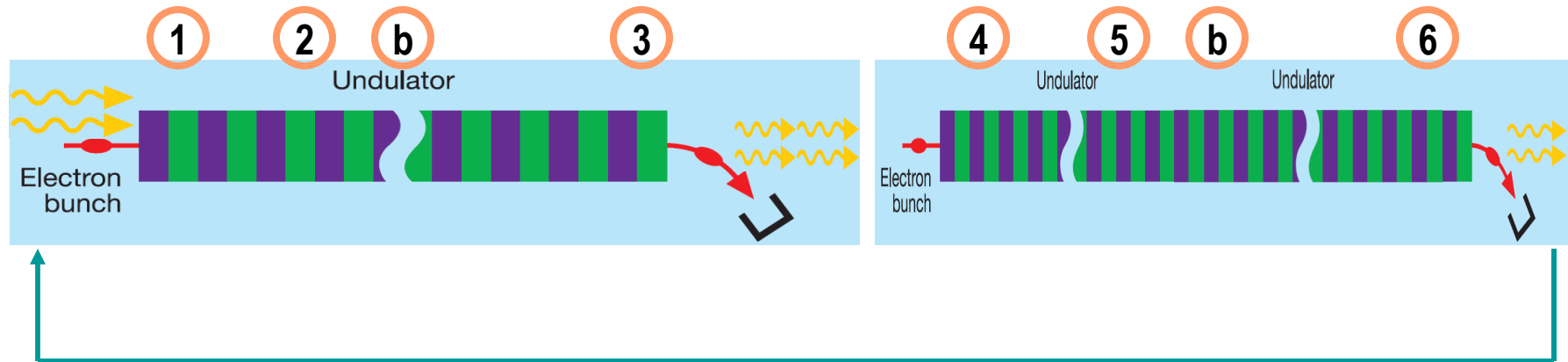


HGHG Modulator and Radiator

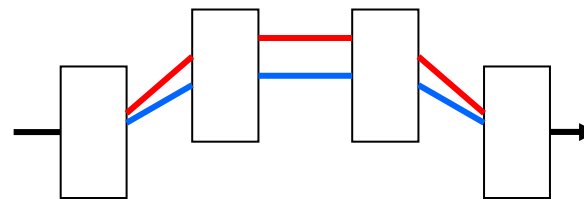
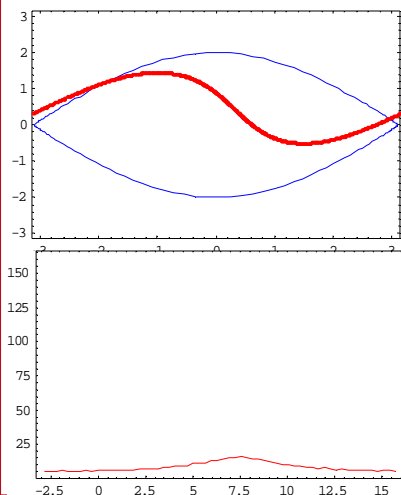


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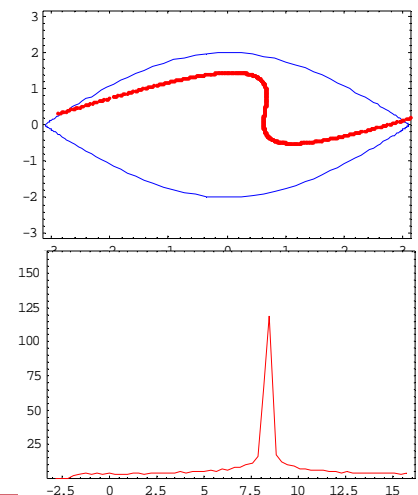
HGHG FELs avoid starting a SASE FEL from noise by seeding with a laser beam.



Note: In real projects, the microbunching would require too long undulators. Each undulator is therefore split in 2, one modulating the energy distribution (2), followed by a buncher section (b), the bunches then radiate, esp. harmonics (3).



Buncher: A chicane of bending magnets through which higher energetic particles can pass faster.





HGHG FELs



CHESS & LEPP

- 1) HGHG FELs have undulators with a total length significantly less than SASE FELs
- 2) Very small energy spread is needed.
- 3) So far only one stage of Modulator and Radiator has been tested.
- 4) BESSY / Germany will test a second stage.
- 5) Several Labs are proposing HGHG FELs for the 1nm soft x-ray regime with an approximately 2GeV e-beam with approximately 4 stages of HGHG, e.g. LBNL / USA, Madison / USA, BESSY / Germany, INFN / Italy, possibly STFC / UK.

