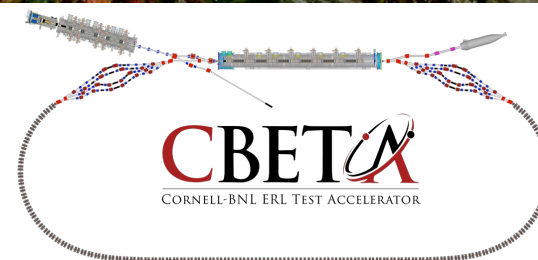
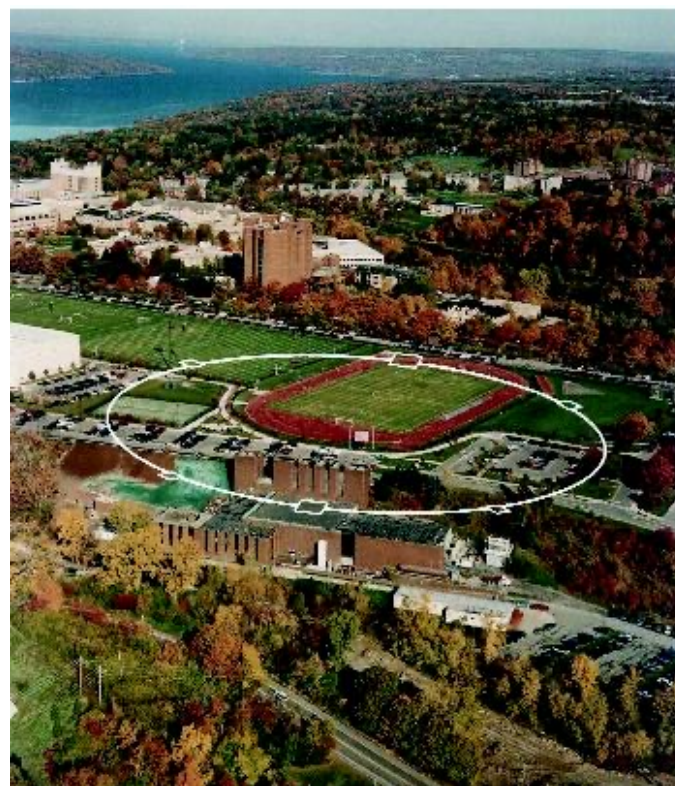


USPAS graduate Accelerator Physics

Content

1. Typical Particle Accelerators
2. Typical Accelerator Components
3. Linear Beam Optics (Circular & Straight)
4. Beam distributions
5. Nonlinear Beam Optics
6. Differential Algebra for Nonlinear Maps
7. Longitudinal Phase Space
8. Synchrotron Radiation
9. Polarization
10. Waveguides and Cavities

Accompanied by homework question and applied accelerator-optics design.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Class data

Dates:

Monday – Friday June 5 – 9, 2023, 9 - 12 and 2 - 5pm, HW after 7:30pm

Monday – Thursday June 12 – 15, 2023, 9 - 12 and 2 - 5pm, HW after 7:30pm

Friday: summary and final exam, 9am - noon

Expectation:

Collaborate on homework.

Consult TAs for homework and optics design.

Co-instructor for Bmad, design, and simulation: David Sagan dc16@cornell.edu

TAs: Ningdong Wang nw285@cornell.edu and Matthew Signorelli mgs255@cornell.edu

Lecture notes, homework, and homework solution will be on the class web page

<https://www.classe.cornell.edu/~hoff/LECTURES/23USPAS/>



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Accelerator optics and simulations

To practice what will be learned, accelerators will be simulated with the programs Tao and Bmad.

You all have accounts to run Bmad, which will be introduced Monday afternoon.

There will be Homework of analytical nature and design / optimization / simulation assignments.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Images are taken from many sources, including:

The Physics of Particle Accelerators, Klaus Wille, Oxford University Press, 2000, ISBN: 19 850549 3

Particle Accelerator Physics I, Helmut Wiedemann, Springer, 2nd edition, 1999, ISBN 3 540 64671 x

Teilchenbeschleuniger und Ionenoptik, Frank Hinterberger, 1997, Springer, ISBN 3 540 61238 6

Introduction to Ultraviolet and X-ray Free-Electron Lasers, Martin Dohlus, Peter Schmüser, Jörg Rossbach, Springer, 2008

Various web pages, 2003 – 2023



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Literature

Recommended as

Introduction

The Physics of Particle Accelerators: An Introduction, Klaus Wille, Oxford University Press

Wide selection of well explained topics

Particle Accelerator Physics, Helmut Wiedemann, Springer, (preferably 3rd edition)

Tremendous overview, with references for derivations and explanations

Handbook of Accelerator Physics and Engineering, Alexander Wu Cao, Maury Tigner, Hans Weise, Frank Zimmermann (3rd edition)



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

What is accelerator physics ?

Accelerator Physics has applications in particle accelerators for high energy physics or for x-ray science, in spectrometers, in electron microscopes, and in lithographic devices. These instruments have become so complex that an empirical approach to properties of the particle beams is by no means sufficient and a detailed theoretical understanding is necessary. This course will introduce into theoretical aspects of charged particle beams and into the technology used for their acceleration.

- Physics of beams
- Physics of non-neutral plasmas
- Physics of involved in the technology:
 - Superconductivity in magnets and radiofrequency (RF) devices
 - Surface physics in particle sources, vacuum technology, RF devices
 - Material science in collimators, beam dumps, superconducting materials



Cornell University



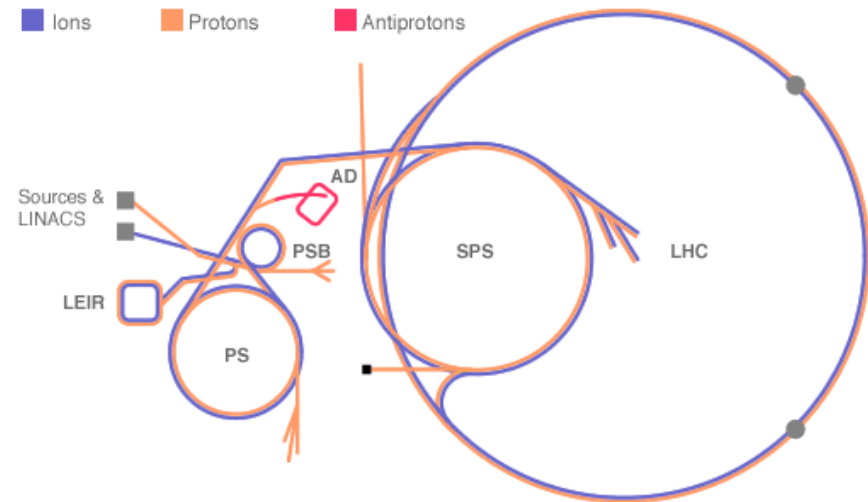
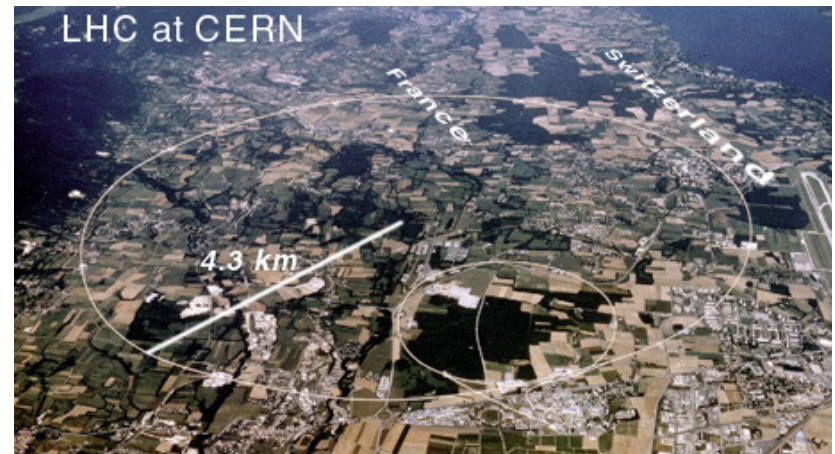
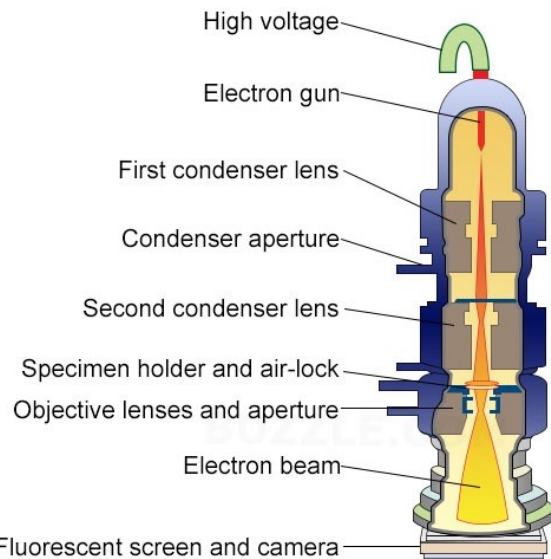
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Particle accelerators, large and small



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Why accelerator physics ?

- Industry

- Food & product safety
- Contraband detection
- Semiconductor fabrication
- Bridge safety

- Medicine

- Tumor detection and treatment.

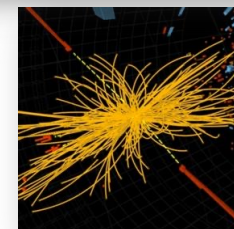
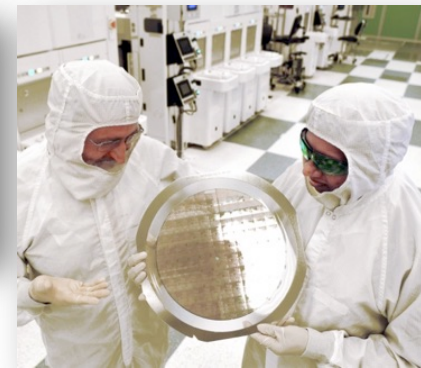
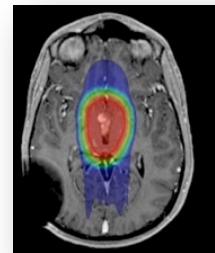
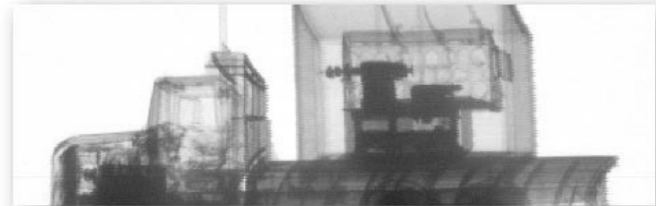
~30,000 industrial and medical accelerators are in use, with annual sales of \$3.5 B and 10% growth per year.

- Research

- X ray sources and colliders for nuclear & particle physics
- Electron microscopes

Since 1943, a Nobel Prize in **Physics** has been awarded to research benefiting from accelerators every 3 years.

Since 1997, the same has been true of **Chemistry**.

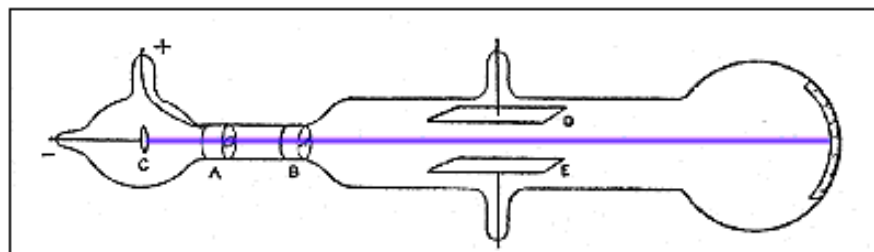


A short history of accelerators

- 1862: Maxwell theory of electromagnetism
- 1887: Hertz discovery of the electromagnetic wave
- 1886: Goldstein discovers positively charged rays (ion beams)
- 1894: Lenard extracts cathode rays (with a 2.65 μm Al Lenard window)
- 1897: JJ Thomson shows that cathode rays are particles since they followed the classical Lorentz force $m\vec{a} = e(\vec{E} + \vec{v} \times \vec{B})$ in an electromagnetic field
- 1926: GP Thomson shows that the electron is a wave
(1929-1930 in Cornell, NP in 1937)



NP 1905, Philipp E.A. von Lenard



NP 1906

Joseph J. Thomson



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

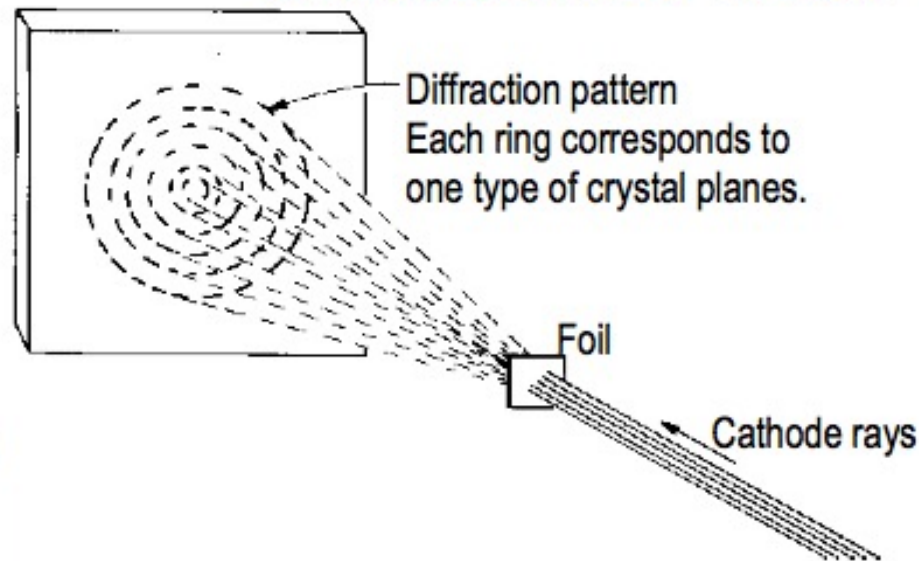
USPAS summer 2023

Discoveries with accelerated beams ...



George P. Thomson
(1892-1975)
1937 Nobel prize
Son of Joseph J. T.

In a powdered, microcrystalline substance there is always some crystal which has the correct angle for constructive interference $2d \cos \alpha = n\lambda$



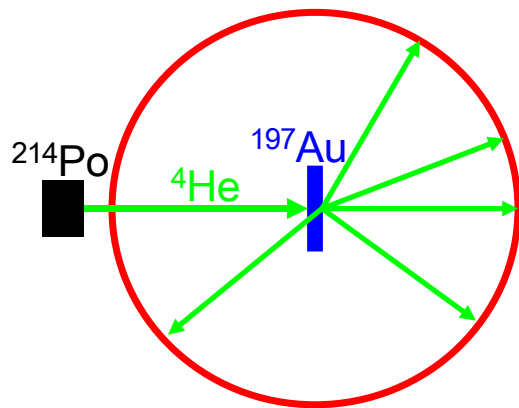
A magnetic field can change the rings, showing the the waves are associated with the electron charge.



Cornell

The need for higher energies

- 1911: Rutherford discovers the nucleus with 7.7 MeV ^4He from ^{214}Po alpha decay measuring the elastic crosssection of $^{197}\text{Au} + ^4\text{He} \mapsto ^{197}\text{Au} + ^4\text{He}$.



$$E = \frac{Z_1 e Z_2 e}{4\pi\epsilon_0 d} = Z_1 Z_2 m_e c^2 \frac{r_e}{d},$$

$$r_e = 2.8\text{fm}, \quad m_e c^2 = 0.511\text{MeV}$$

d = smallest approach for back scattering

- 1919: Rutherford produces first nuclear reactions with natural ^4He
 $^{14}\text{N} + ^4\text{He} \mapsto ^{17}\text{O} + \text{p}$
- 1921: Greinacher invents the cascade generator for several 100 keV
- Rutherford is convinced that several 10 MeV are in general needed for nuclear reactions. He therefore gave up the thought of accelerating particles.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

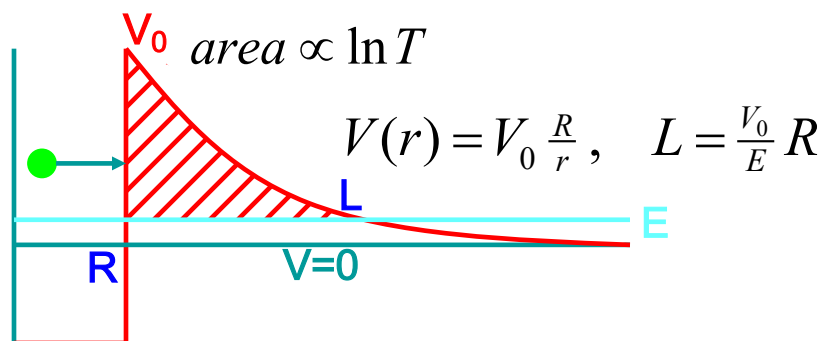
USPAS summer 2023

Tunneling allows low energies

- 1928: Explanation of alpha decay by Gamov as tunneling showed that several 100keV protons might suffice for nuclear reactions

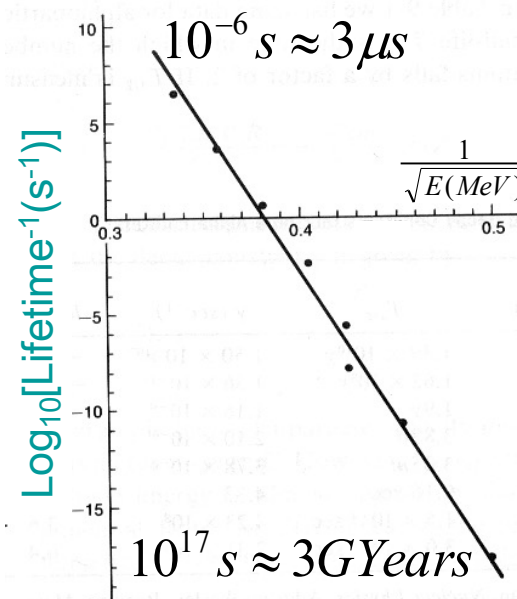
Schroedinger equation: $\frac{\partial^2}{\partial r^2} u(r) = \frac{2m}{\hbar^2} [V(r) - E] u(r), \quad T = \left| \frac{u(L)}{u(0)} \right|^2$

The transmission probability T for an alpha particle traveling from the inside towards the potential well that keeps the nucleus together determines the lifetime for alpha decay.



$$T \approx \exp\left[-2 \int_0^L \frac{\sqrt{2m[V(r)-E]}}{\hbar} dr\right]$$

$$\ln T \approx A - \frac{C}{\sqrt{E}}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

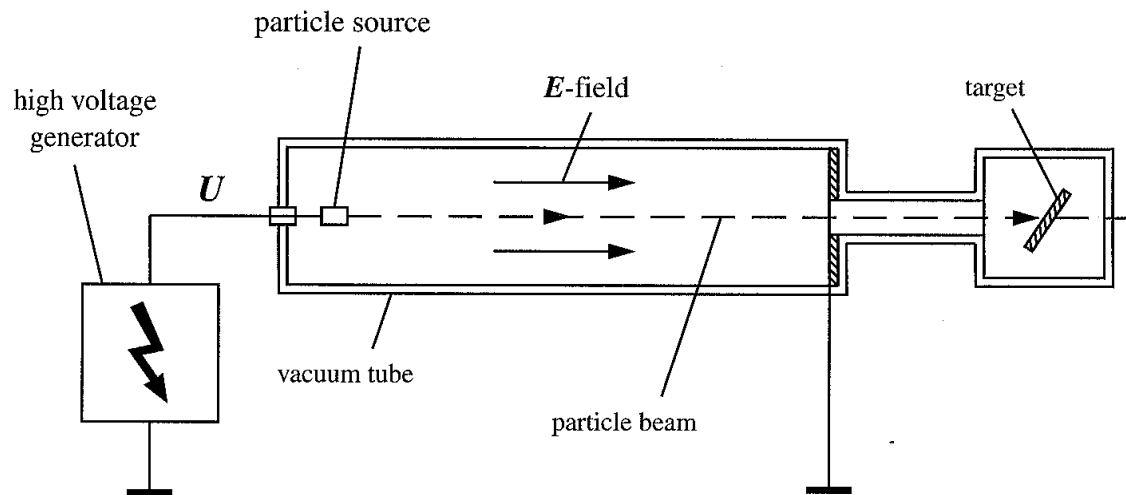
Graduate Accelerator Physics

USPAS summer 2023

Three historic lines of accelerators

Direct Voltage Accelerators

Resonant Accelerators Transformer Accelerator



Voltage 1MV
Charge Ze
Energy Z MeV

The energy limit is given by the maximum possible voltage. At the limiting voltage, electrons and ions are accelerated to such large energies that they hit the surface and produce new ions. An avalanche of charge carries causes a large current and therefore a breakdown of the voltage.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

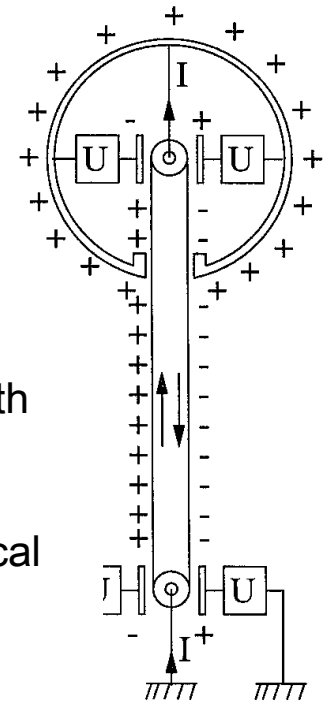
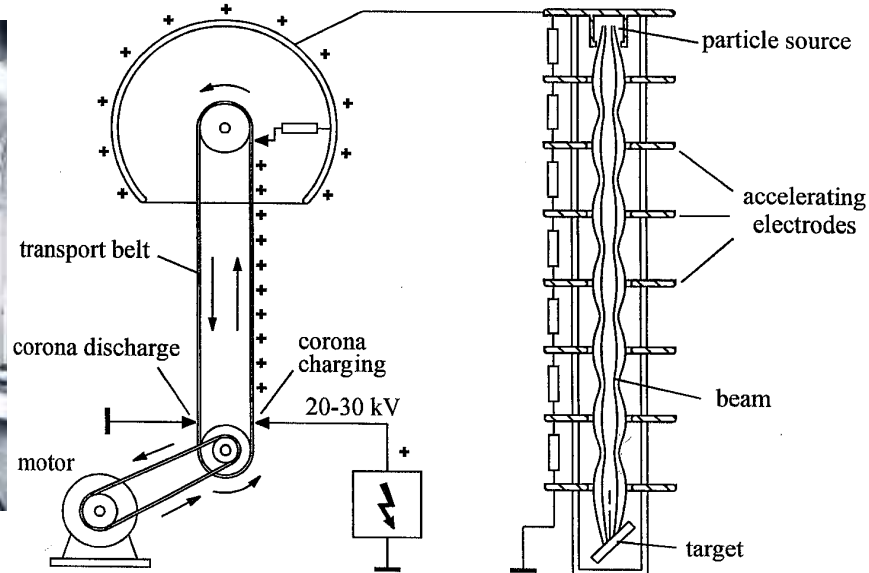
USPAS summer 2023

The Van de Graaff Accelerator

- 1930: van de Graaff builds the first 1.5MV high voltage generator



Van de
Graaff



- Today Peletrons (with chains) or Laddertron (with stripes) that are charged by influence are commercially available.
- Used as injectors, for electron cooling, for medical and technical n-source via $d + t \rightarrow n + \alpha$
- Up to 17.5 MV with insulating gas (1MPa SF_6)



Cornell University



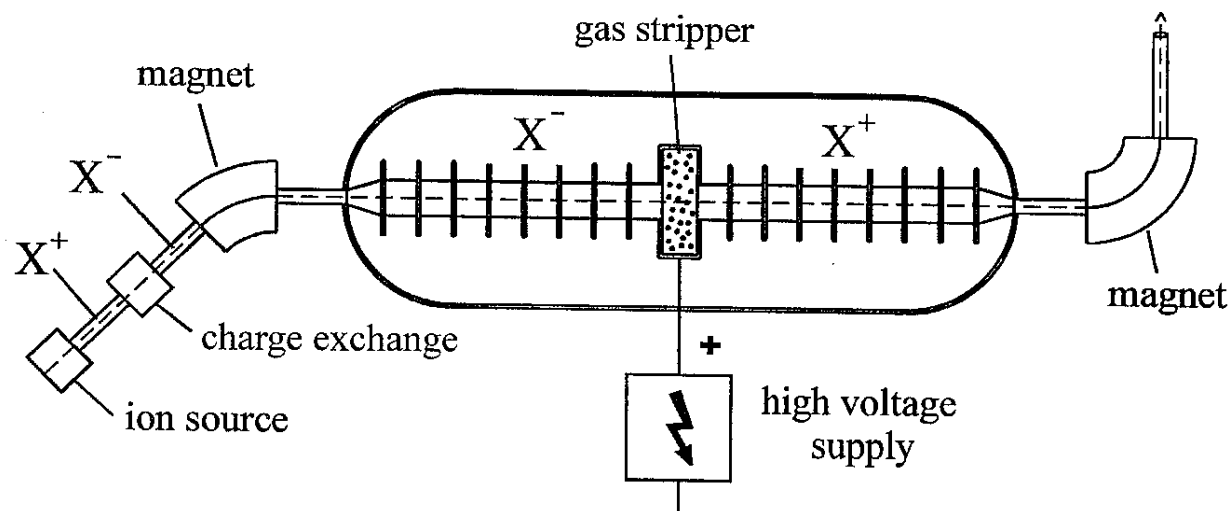
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

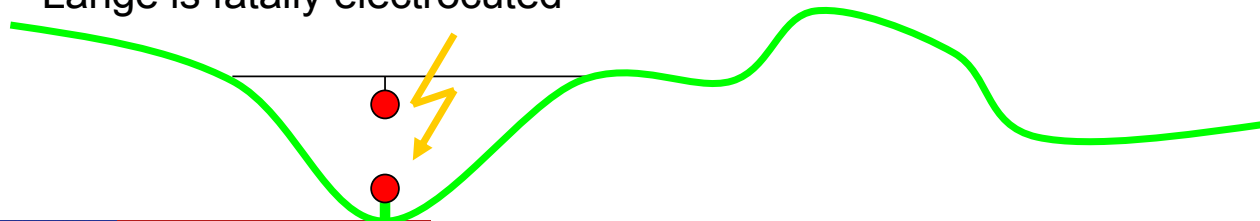
USPAS summer 2023

The Tandem (Van de Graaff) Accelerator

- Two Van de Graaffs, one + one -
- The Tandem Van de Graaff, highest energy 35MeV



- 1932: Brasch and Lange use potential from lightening, in the Swiss Alps, Lange is fatally electrocuted



Cornell University



Brookhaven
National Laboratory

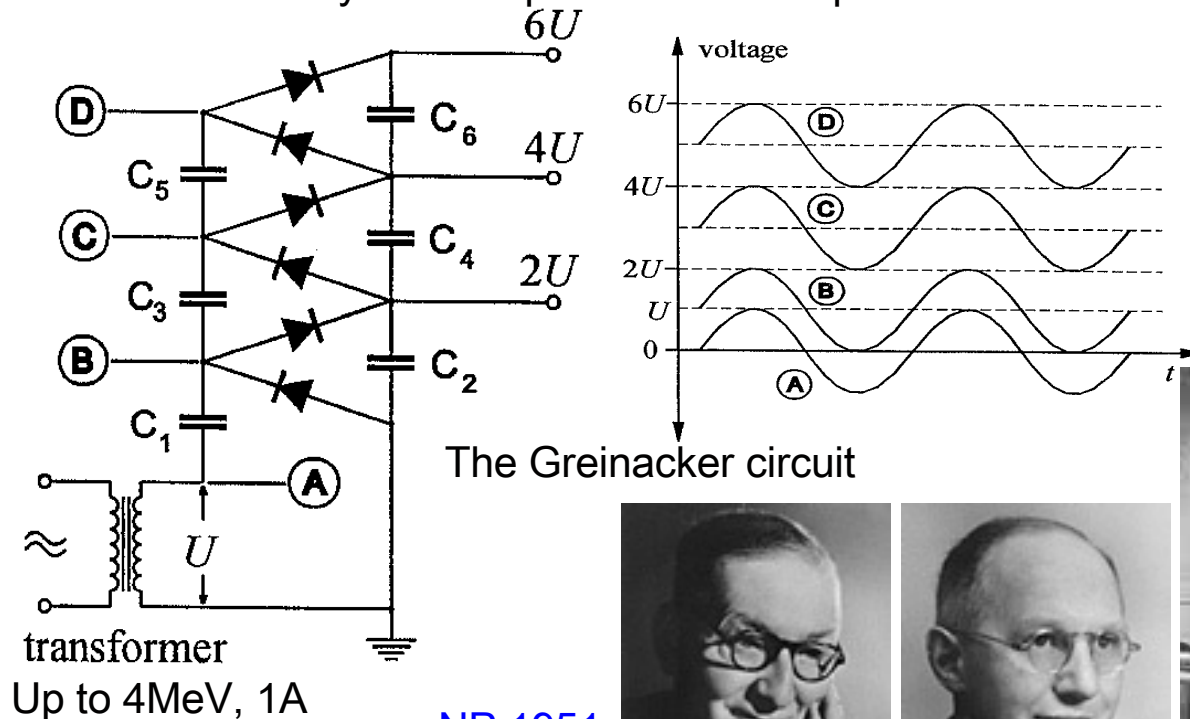
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

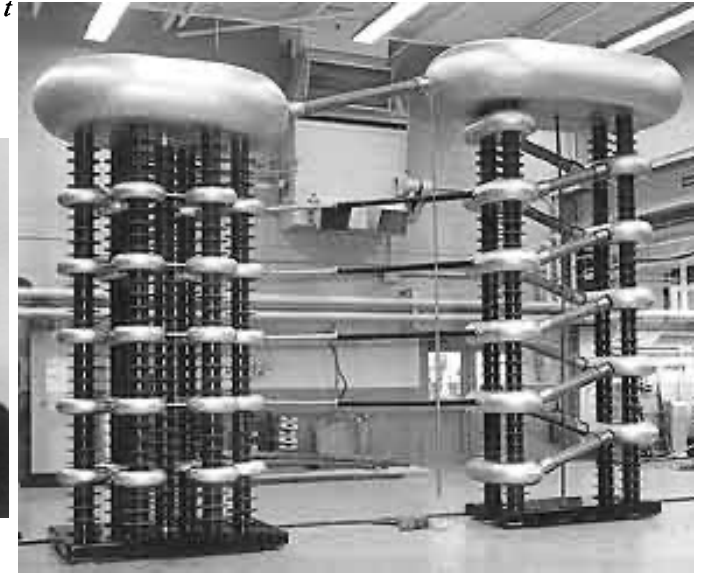
USPAS summer 2023

The Cockcroft-Walton Accelerator

1932: Cockcroft and Walton 1932: 700keV cascade generator (planned for 800keV)
and use initially 400keV protons for ${}^7\text{Li} + p \mapsto {}^4\text{He} + {}^4\text{He}$ and ${}^7\text{Li} + p \mapsto {}^7\text{Be} + n$



The Greinacker circuit



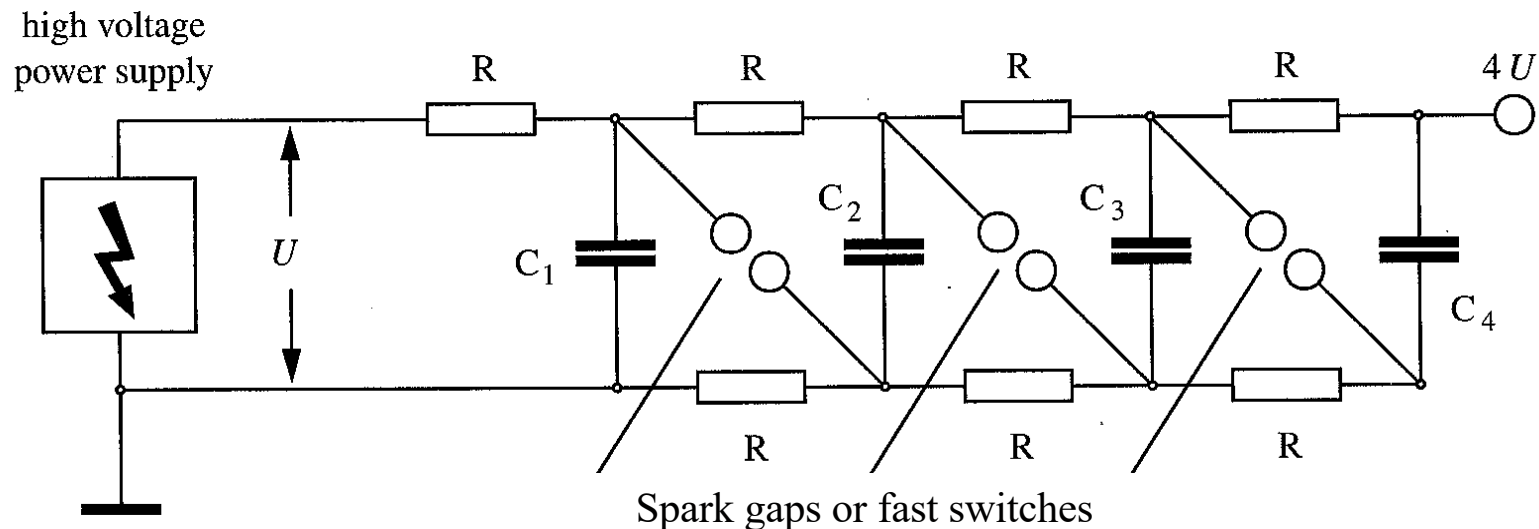
NP 1951

Sir John D Cockcrof

Ernest T S Walton

The Marx Generator

- 1932: Marx Generator achieves 6MV at General Electrics



After capacitors of around 2uF are filled to about 20kV, the spark gaps or switches close as fast as 40ns, allowing up to 500kA.

Today: The Z-machine (Physics Today July 2003) for z-pinch initial confinement fusion has 40TW for 100ns from 36 Marx generators



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

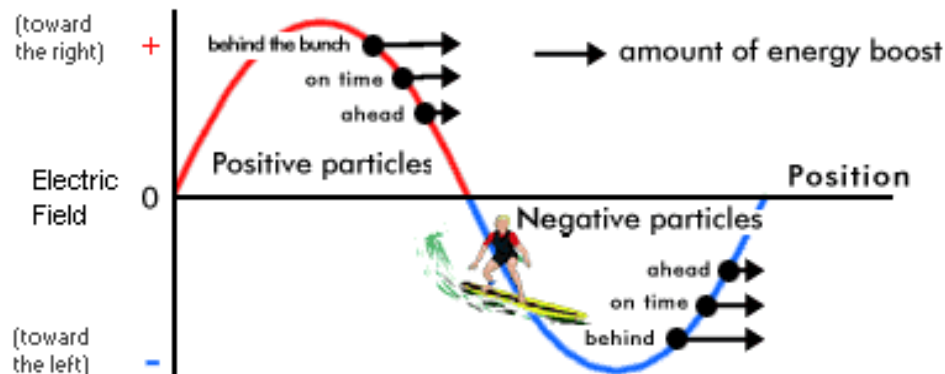
USPAS summer 2023

Three historic lines of accelerators

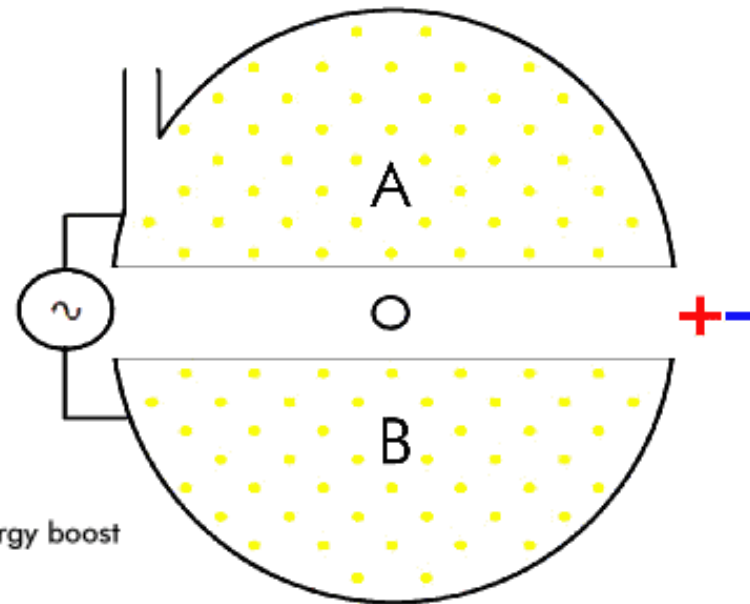
Direct Voltage Accelerators



Resonant Accelerators



Transformer Accelerator



Particles must have the correct phase relation to the accelerating voltage.



Cornell University



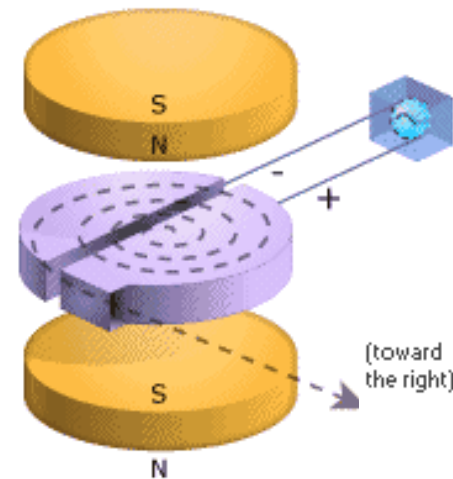
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

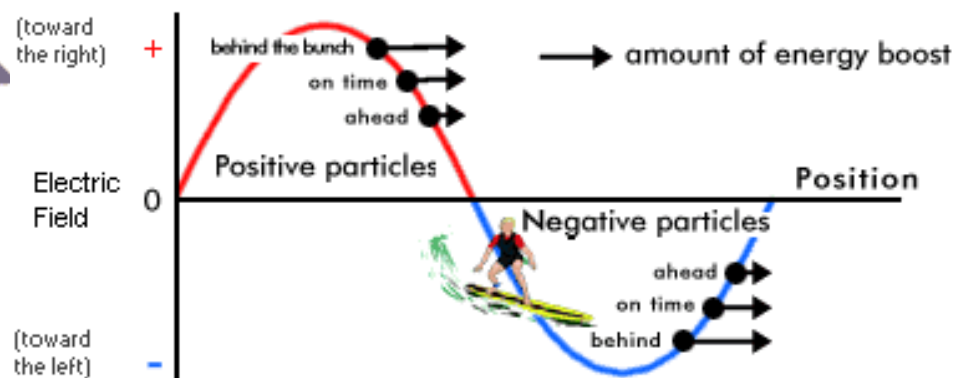
The Cyclotron



NP 1939

Ernest O Lawrence

- 1930: Lawrence proposes the Cyclotron
(before he develops a workable color TV screen)
- 1932: Lawrence and Livingston use a cyclotron for 1.25MeV protons and mention longitudinal (phase) focusing



- 1934: Livingston builds the first Cyclotron away from Berkely (2MeV protons) at Cornell (in room B54)



M Stanley Livingston



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Cyclotron frequency

$$F_r = m_0 \gamma \omega_z v = q v B_z$$

$$\omega_z = \frac{q}{m_0 \gamma} B_z = \text{const}$$

Condition: Non-relativistic particles.
Therefore, not for electrons.

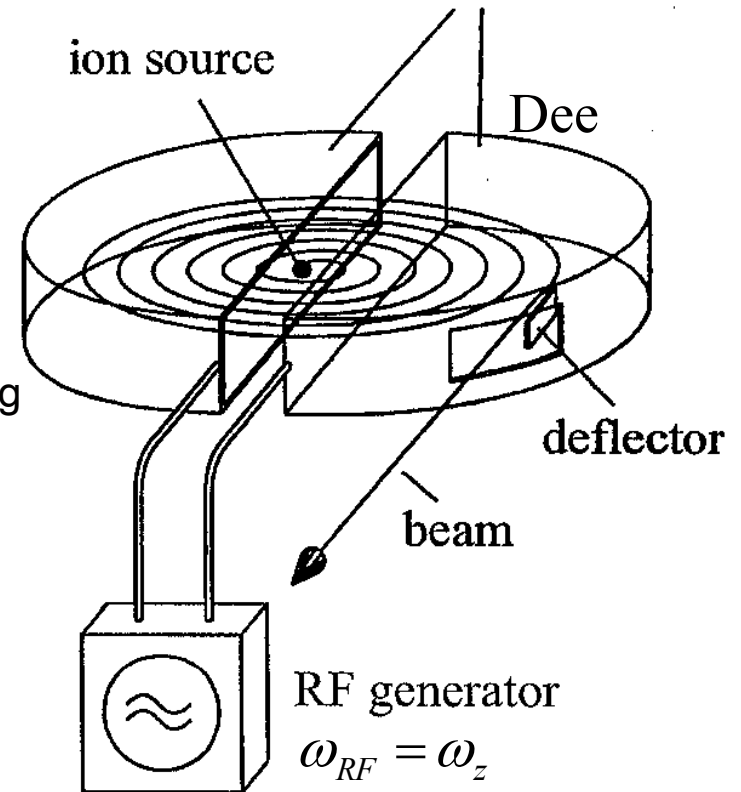
- The synchrocyclotron:
Acceleration of bunches with decreasing

$$\omega_z(E) = \frac{q}{m_0 \gamma(E)} B_z$$

- The isocyclotron with constant

$$\omega_z = \frac{q}{m_0 \gamma(E)} B_z(r(E))$$

Up to 600 MeV but
this vertically defocuses the beam



- 1938: Thomas proposes strong (transverse) focusing for a cyclotron



Cornell University



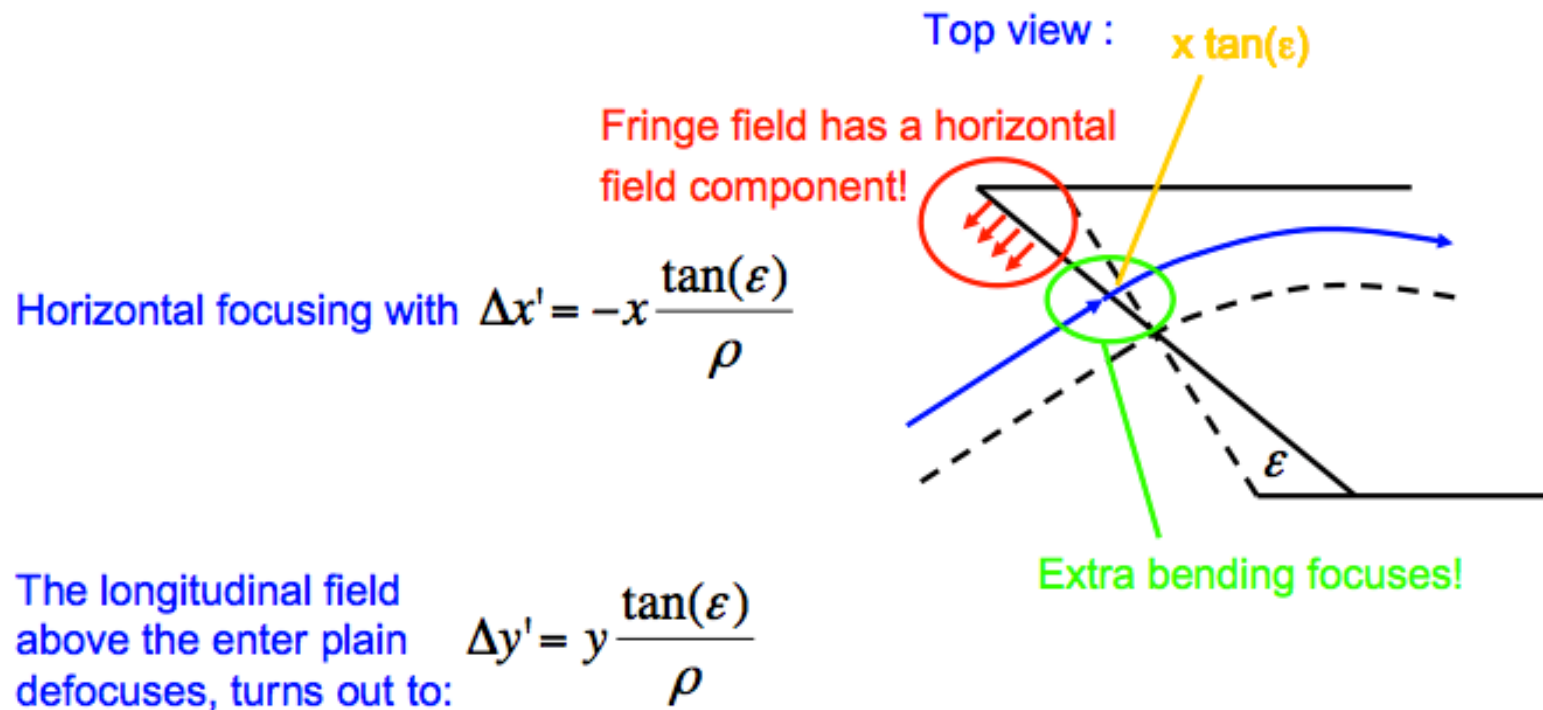
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Edge Focusing



Quadrupole effect: focusing in x and defocusing in y or
defocusing in x and focusing in y.



Cornell University



Brookhaven
National Laboratory

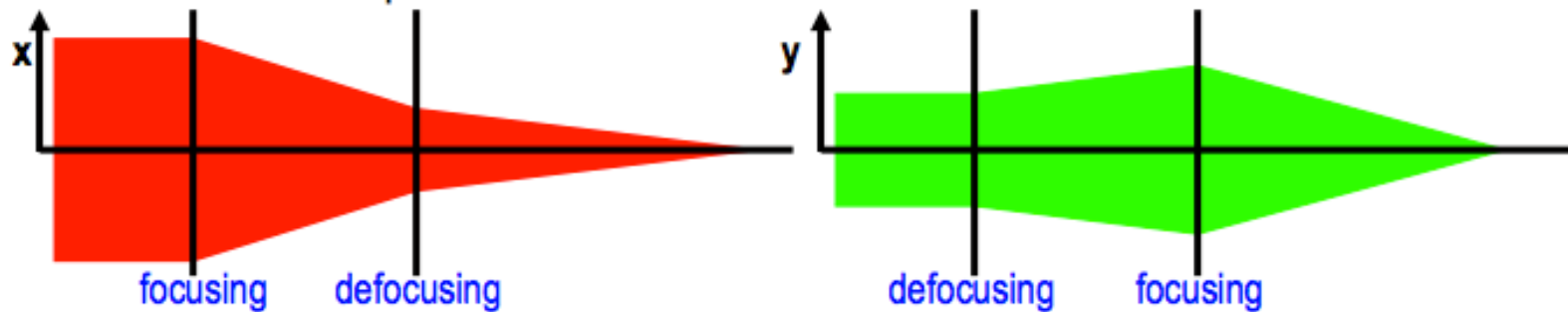
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Quadrupole Focusing

Transverse fields defocus in one plane if they focus in the other plane.
But two successive elements, one focusing the other defocusing,
can focus in both planes:



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

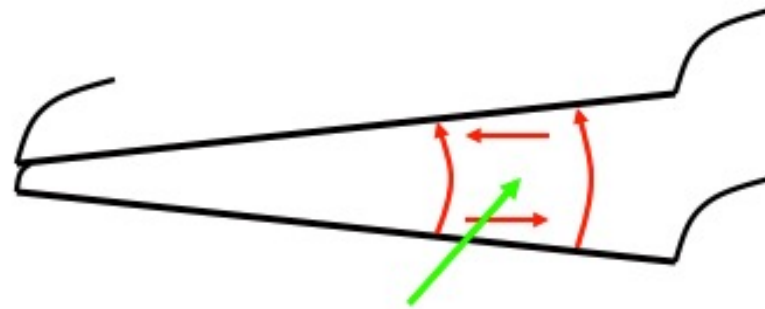
The Isocyclotron

- The isocyclotron with constant

$$\omega_z = \frac{q}{m_0 \gamma(E)} B_z(r(E))$$

Up to 600MeV but
this vertically defocuses the beam.

Edge focusing is therefore used.



First Medical Accelerators

- 1939: Lawrence uses 60' cyclotron for 9MeV protons, 19MeV deuterons, and 35MeV 4He. First tests of tumor therapy with neutrons via $d + t \rightarrow n + \alpha$
With 200-800keV d to get 10MeV neutrons.



Cornell University



Brookhaven
National Laboratory

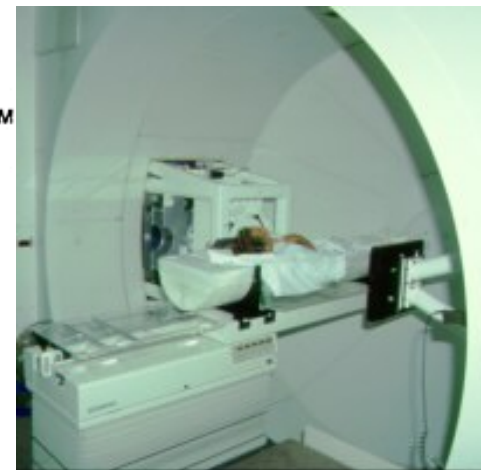
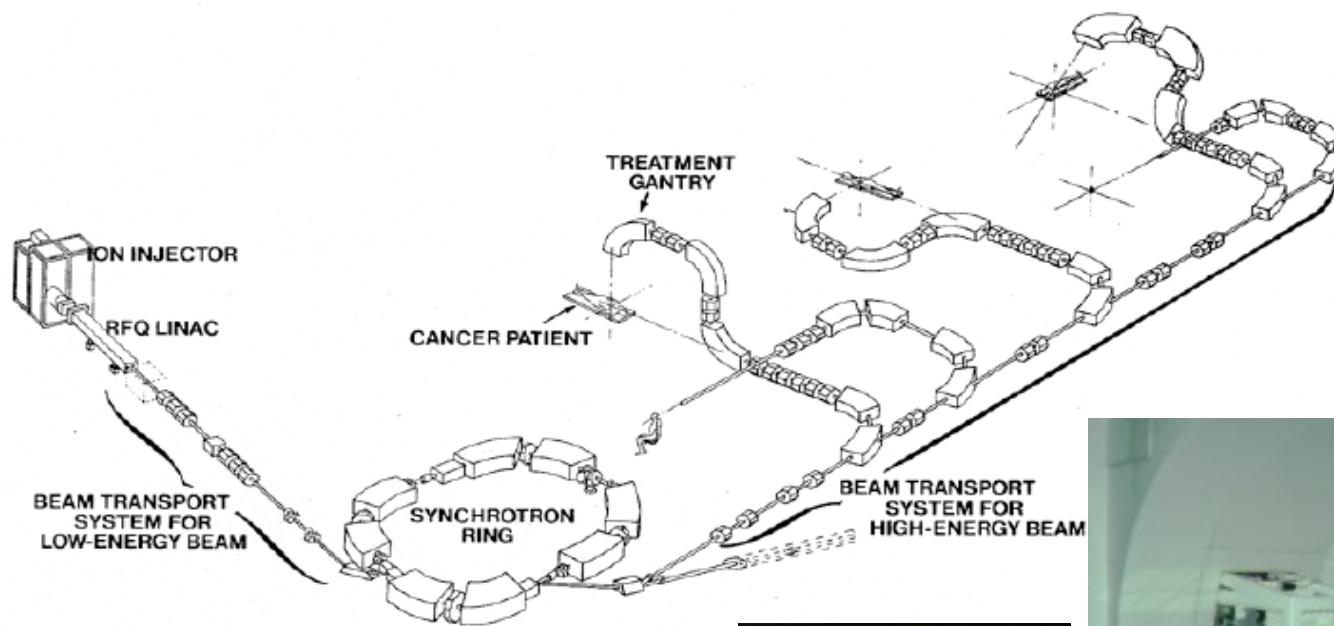
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Modern Nuclear Radiation Therapy

The Loma Linda proton therapy facility



Cornell University



Brookhaven
National Laboratory

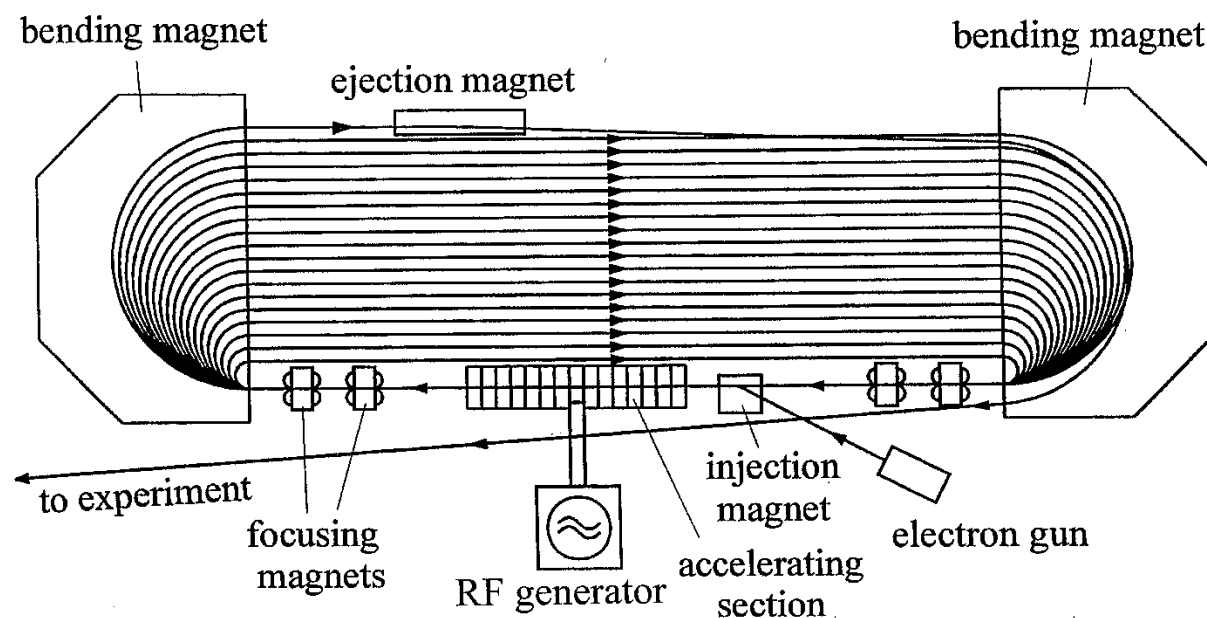
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Microtron

- Electrons are quickly relativistic and cannot be accelerated in a cyclotron.
- In a microtron the revolution frequency changes, but each electron misses an integer number of RF waves.



- Today: Used for medical applications with one magnet and 20MeV.
- Nuclear physics: MAMI designed for 820MeV as race track microtron.



Cornell University

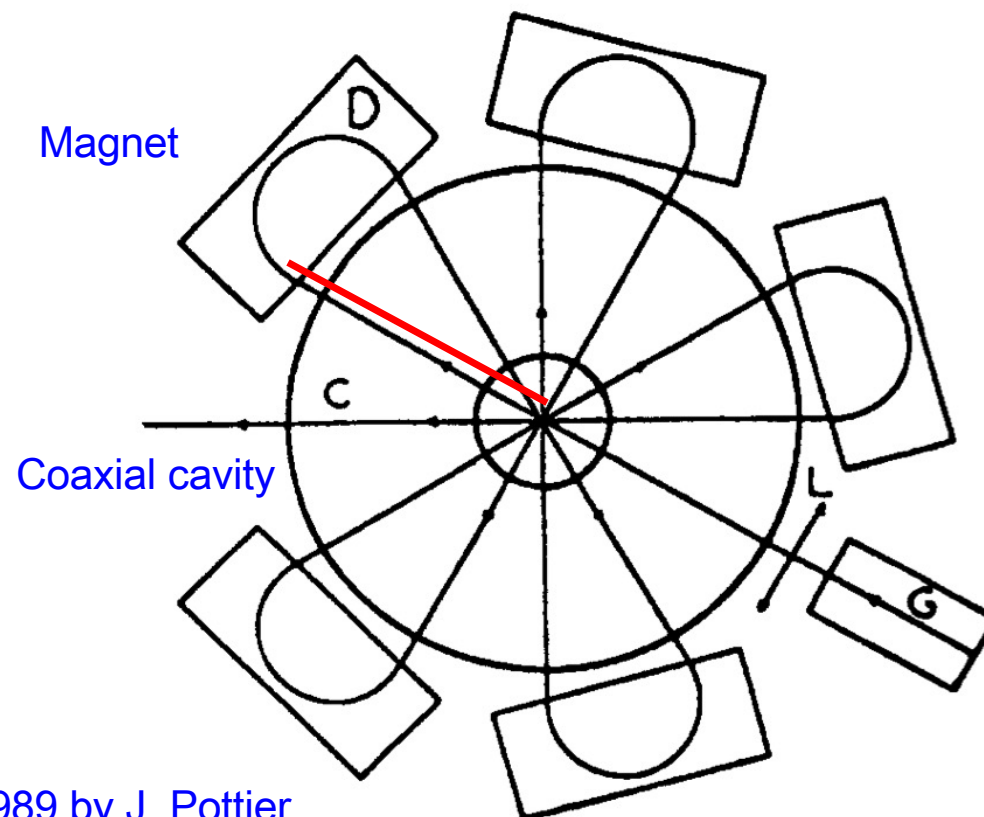


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Rhodotron



- Invented in 1989 by J. Pottier
(French, CEA)



Cornell University



Brookhaven
National Laboratory

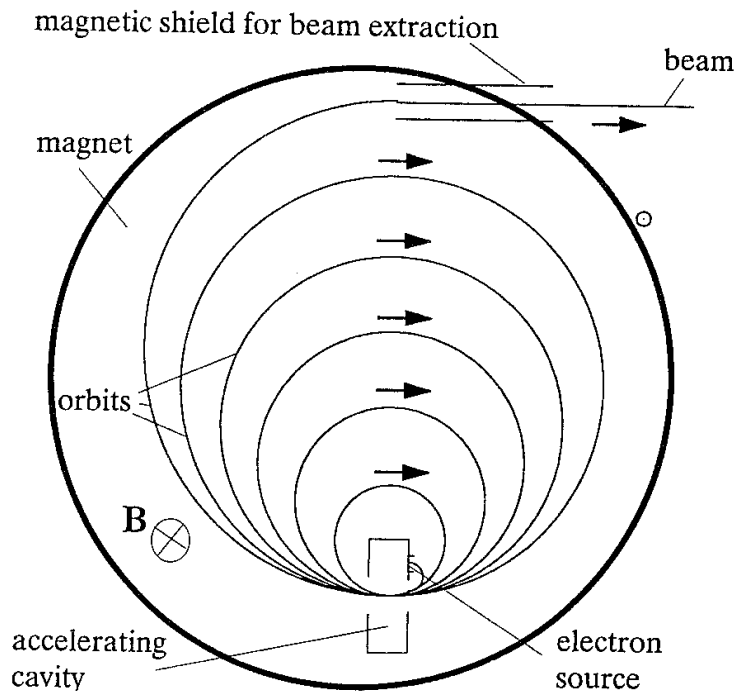
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

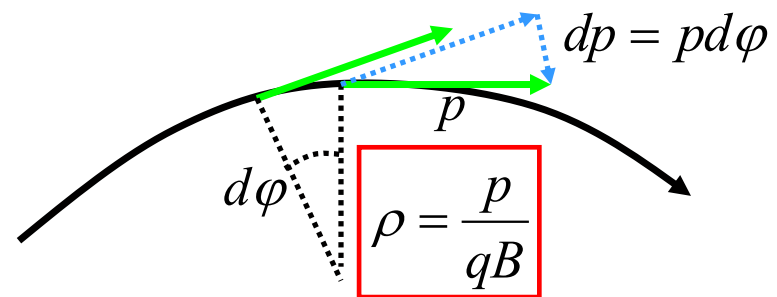
The Microtron Condition

- The extra time that each turn takes must be a multiple of the RF period.



$B=1\text{T}$, $n=1$, and $f_{\text{RF}}=3\text{GHz}$ leads to 4.78MeV
 This requires a small linear accelerator.

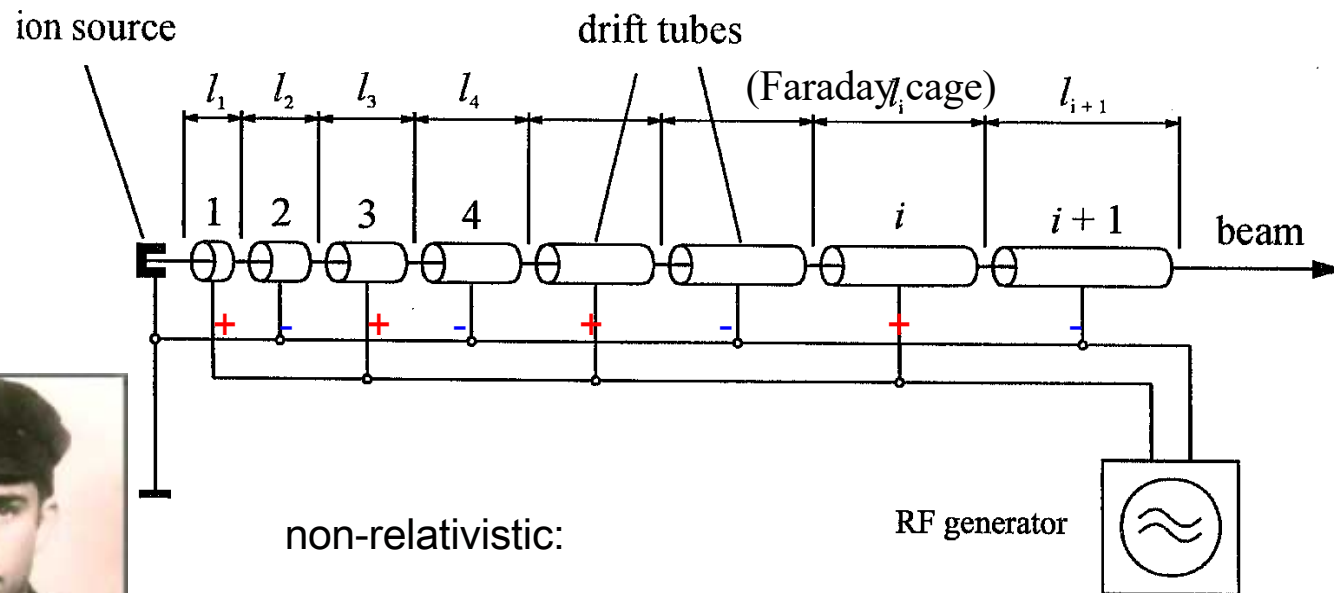
$$\frac{dp}{dt} = qvB \Rightarrow \rho = \frac{dl}{d\phi} = \frac{vdt}{dp/p} = \frac{p}{qB}$$



$$\begin{aligned} \Delta t &= 2\pi \left(\frac{\rho_{n+1}}{v_{n+1}} - \frac{\rho_n}{v_n} \right) \\ &= \frac{2\pi}{qB} (m_0 \gamma_{n+1} - m_0 \gamma_n) = \frac{2\pi}{qBc^2} \Delta K \end{aligned}$$

$$\Delta K = n \frac{qBc^2}{\omega_{\text{RF}}} \quad \text{for an integer } n$$

The Wideroe linear accelerator



Wideroe

non-relativistic:

$$K_n = nqU_{\max} \sin \psi_0 = \frac{1}{2} m v_n^2$$

$$l_n = \frac{1}{2} v_n T_{RF} = \frac{1}{2} \beta_n \lambda_{RF} \propto \sqrt{n}$$

Called the π or the $1/2\beta\lambda$ mode



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

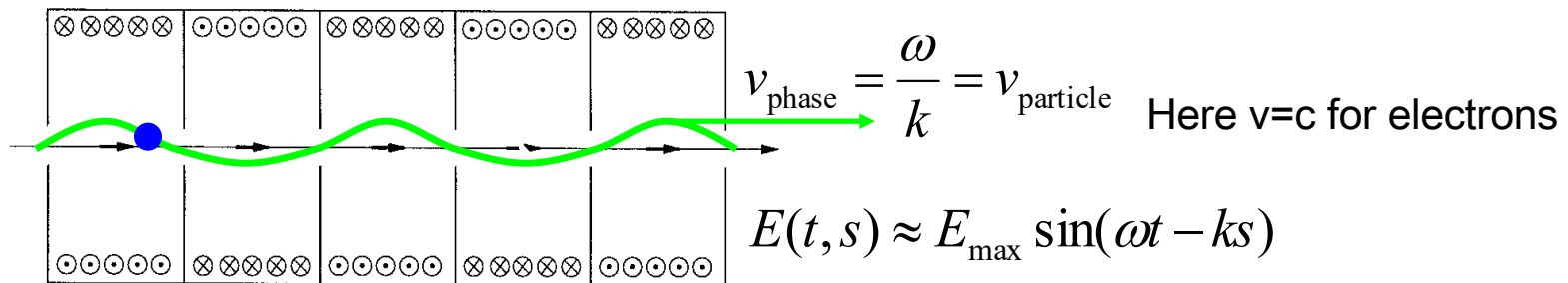
Graduate Accelerator Physics

USPAS summer 2023

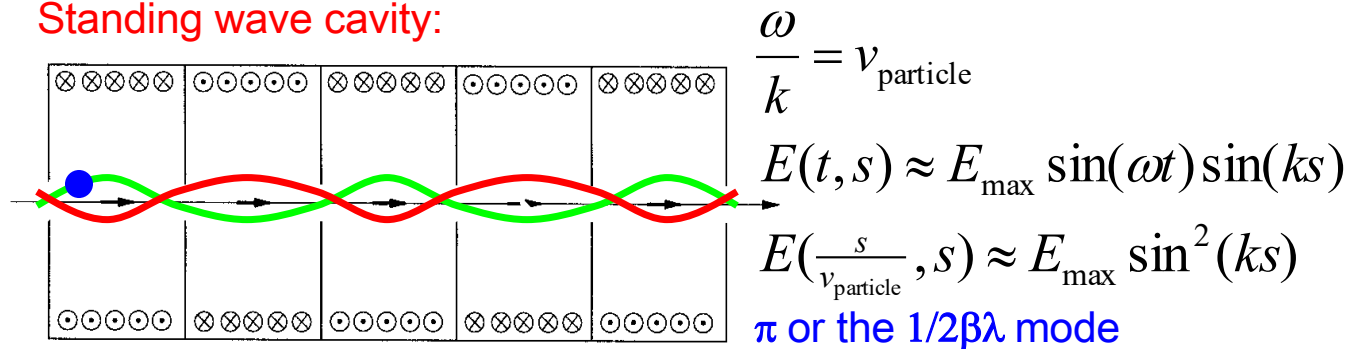
Accelerating Cavities

- 1933: J.W. Beams uses resonant cavities for acceleration

Traveling wave cavity:



Standing wave cavity:



Transit factor (for this example): $\langle E \rangle = \frac{1}{\lambda_{RF}} \int_0^{\lambda_{RF}} E\left(\frac{s}{v_{\text{particle}}}, s\right) ds \approx \frac{1}{2} E_{\text{max}}$



Cornell University



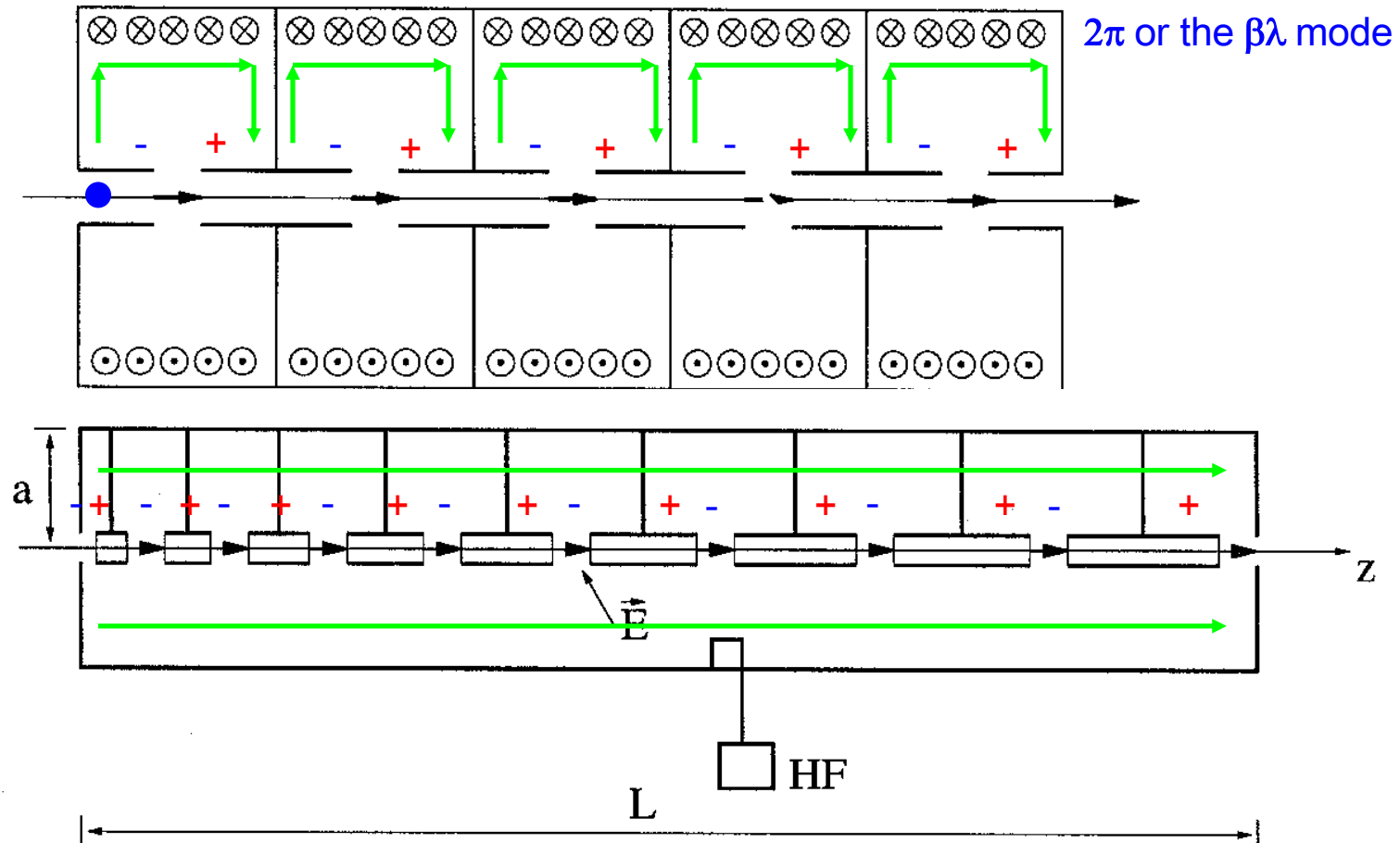
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Alvarez Linear Accelerator



Needs only one power input coupler and walls do not dissipate energy.



Cornell University



Brookhaven
National Laboratory

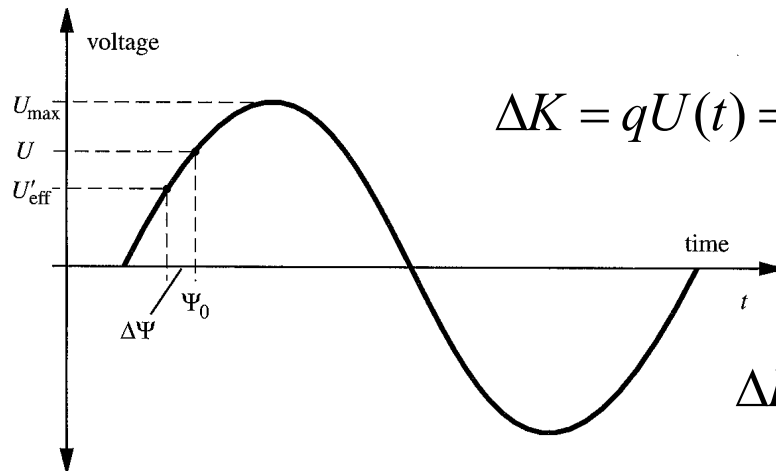
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Phase Focusing

- 1945: Veksler (UDSSR) and McMillan (USA) realize the importance of phase focusing



$$\Delta K = qU(t) = qU_{\max} \sin(\omega(t - t_0) + \psi_0)$$

Longitudinal position in the bunch:

$$\sigma = s - s_0 = -v_0(t - t_0)$$

$$\Delta K(\sigma) = qU_{\max} \sin\left(-\frac{\omega}{v_0}(s - s_0) + \psi_0\right)$$

$$\Delta K(0) > 0 \quad (\text{Acceleration})$$

$$\Delta K(\sigma) < \Delta K(0) \text{ for } \sigma > 0 \Rightarrow \frac{d}{d\sigma} \Delta K(\sigma) < 0 \quad (\text{Phase focusing})$$

$$\left. \begin{array}{l} qU(t) > 0 \\ q \frac{d}{dt} U(t) > 0 \end{array} \right\} \underline{\underline{\psi_0 \in (0, \frac{\pi}{2})}}$$

Phase focusing is required in any RF accelerator.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

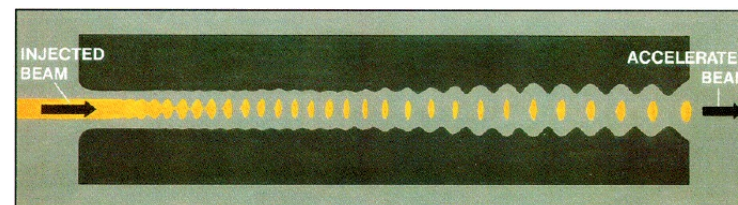
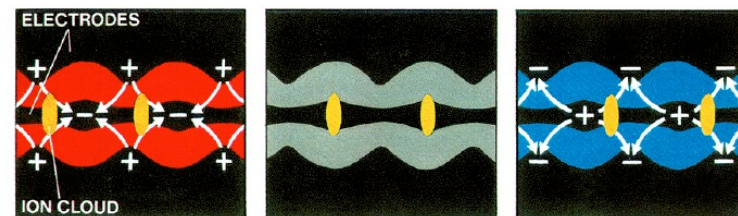
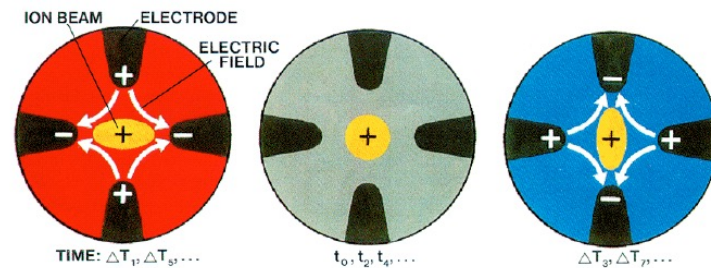
Graduate Accelerator Physics

USPAS summer 2023

Radio Frequency Quadrupoles



- 1970: Kapchinskii and Teplyakov invent the RFQ



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Three historic lines of accelerators

Direct Voltage Accelerators Resonant Accelerators Transformer Accelerator

- 1924: Wideroe invents the betatron
- 1940: Kerst and Serber build a betatron for 2.3MeV electrons and understand betatron (transverse) focusing (in 1942: 20MeV)

Betatron:

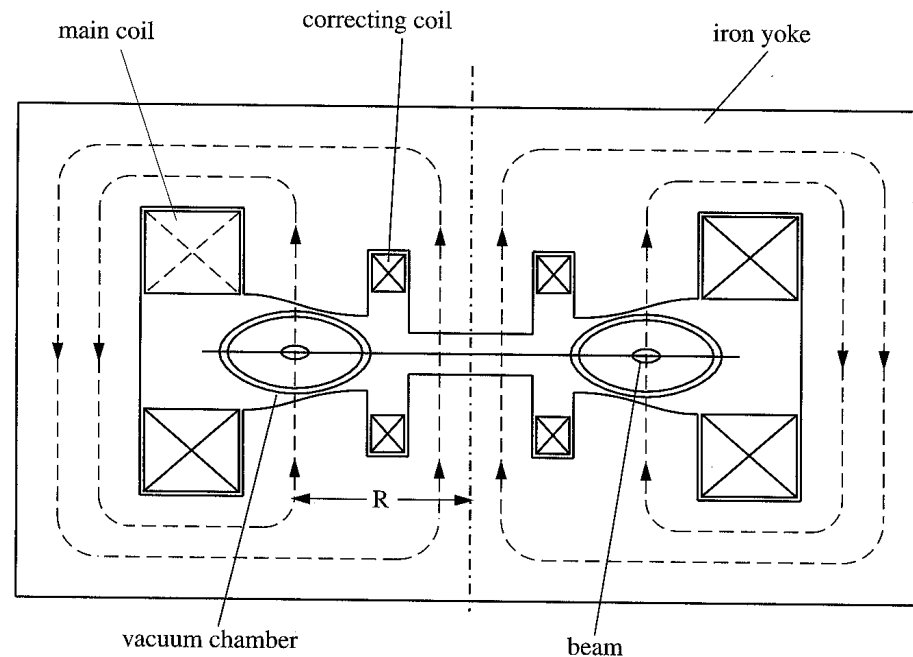
$R=\text{const}, B=B(t)$

Whereas for a cyclotron:

$R(t), B=\text{const}$

No acceleration section is needed since

$$\oint_{\partial A} \vec{E} \cdot d\vec{s} = - \iint_A \frac{d}{dt} \vec{B} \cdot d\vec{a}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Betatron Condition

Condition: $R = \frac{-p_\phi(t)}{qB_z(R,t)} = \text{const.}$ given $\oint_{\partial A} \vec{E} \cdot d\vec{s} = - \iint_A \frac{d}{dt} \vec{B} \cdot d\vec{a}$

$$E_\phi(R,t) = -\frac{1}{2\pi R} \int \frac{d}{dt} B_z(r,t) r dr d\phi = -\frac{R}{2} \left\langle \frac{d}{dt} B_z \right\rangle$$

$$\frac{d}{dt} p_\phi(t) = qE_\phi(R,t) = -q \frac{R}{2} \left\langle \frac{d}{dt} B_z \right\rangle$$

$$p_\phi(t) = p_\phi(0) - q \frac{R}{2} [\langle B_z \rangle(t) - \langle B_z \rangle(0)] = -RqB_z(R,t)$$

$$B_z(R,t) - B_z(R,0) = \frac{1}{2} [\langle B_z \rangle(t) - \langle B_z \rangle(0)]$$

Small deviations from this condition lead to transverse beam oscillations called **betatron oscillations** in all accelerators.

- Today: Betatrons with typically about 20MeV for medical applications



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Synchrotron

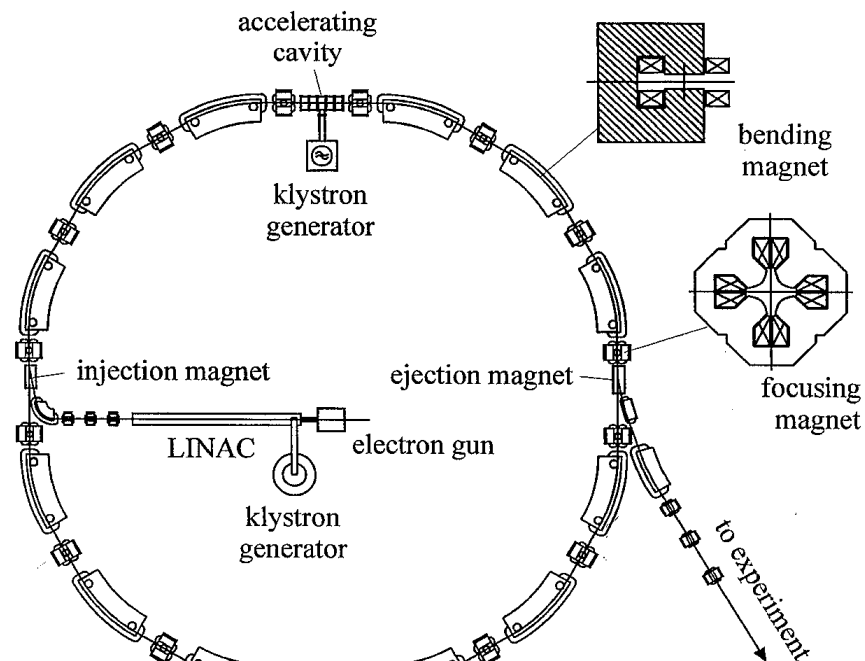
- 1945: Veksler (UDSSR) and McMillan (USA) invent the synchrotron
- 1946: Goward and Barnes build the first synchrotron (using a betatron magnet)
- 1949: Wilson et al. at Cornell are first to store beam in a synchrotron (later 300MeV, magnet of 80 Tons)
- 1949: McMillan builds a 320MeV electron synchrotron

- Many smaller magnets instead of one large magnet
- Only one acceleration section is needed, with

$$R = \frac{p(t)}{qB(R,t)} = \text{const.}$$

$$\omega = 2\pi \frac{v_{\text{particle}}}{L} n$$

for an integer n called the harmonic number



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

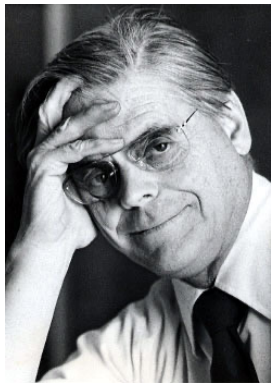
Graduate Accelerator Physics

USPAS summer 2023

Robert Wilson, Fermilab ...



Science Ed Center, FNAL (1990)



Robert R Wilson
USA 1914-2000



Wilson Hall, FNAL



Cornell University



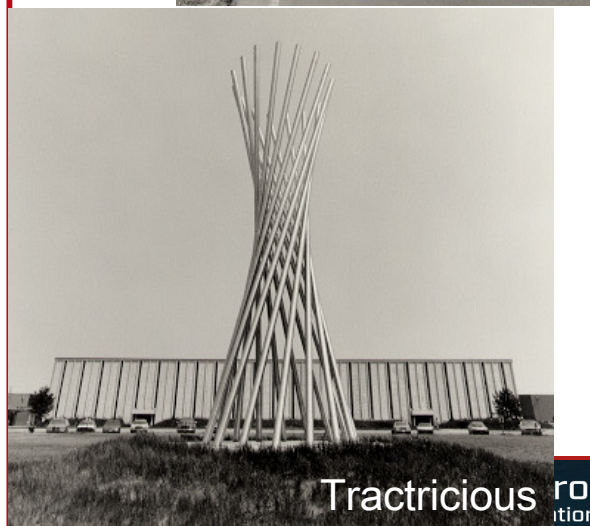
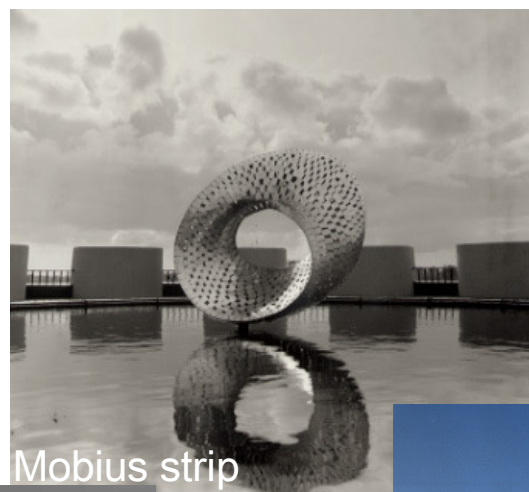
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

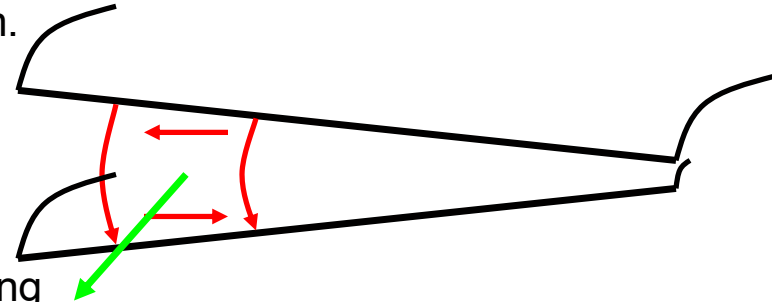
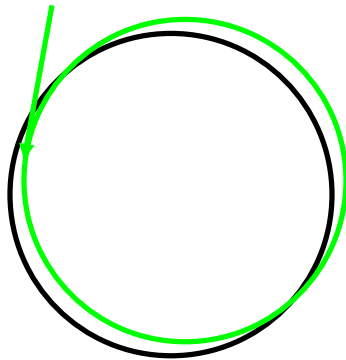
... art, architecture, and accelerators.



The weak focusing synchrotron

- 1952: Operation of the Cosmotron, 3.3 GeV proton synchrotron at Brookhaven: Beam pipe height: 15cm.

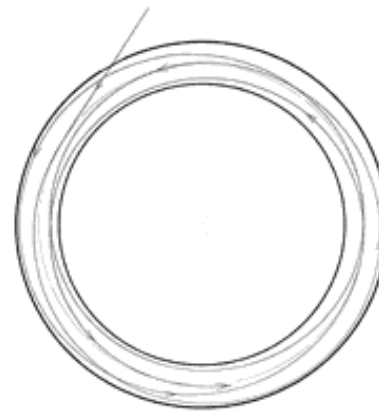
Natural ring focusing:



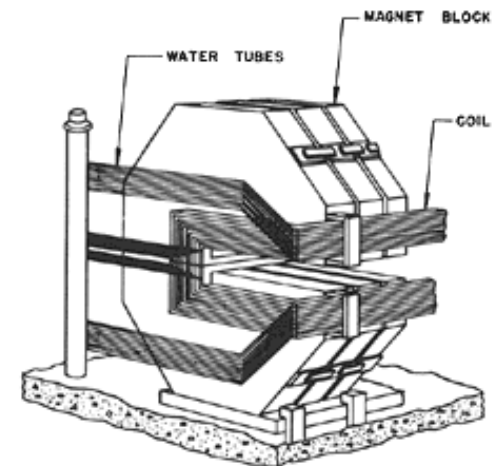
Vertical focusing
+ Horizontal defocusing + ring focusing
Focusing in both planes



The Cosmotron



Weak focusing accelerator



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

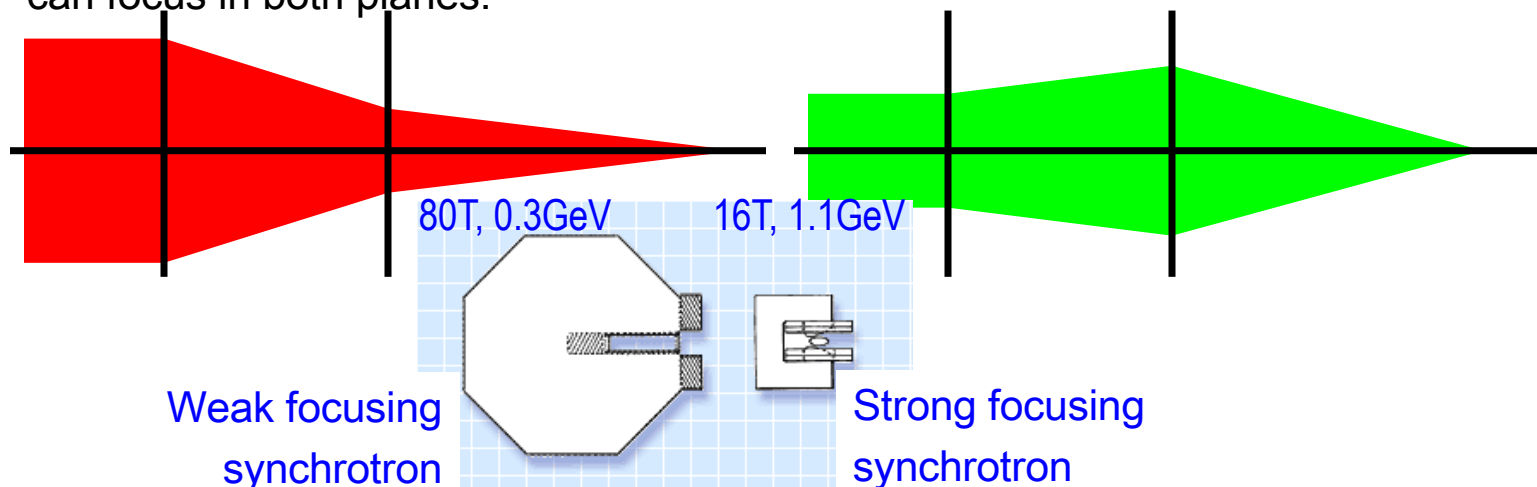
Graduate Accelerator Physics

USPAS summer 2023

Strong focusing synchrotrons

- 1952: Courant, Livingston, Snyder publish about strong focusing
- 1954: Wilson et al. build first synchrotron with strong focusing for 1.1MeV electrons at Cornell, 4cm beam pipe height, only 16 Tons of magnets.
- 1959: CERN builds the PS for 28GeV after proposing a 5GeV weak focusing accelerator for the same cost (still in use)

Transverse fields defocus in one plane if they focus in the other plane.
But two successive elements, one focusing the other defocusing,
can focus in both planes:



- Today: only strong focusing is used. Due to bad field quality at lower field excitations the injection energy is 20-500MeV from a linac or a microtron.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Limits of synchrotrons

$$\rho = \frac{p}{qB} \Rightarrow \text{The rings become too long}$$

Protons with $p = 20 \text{ TeV}/c$, $B = 6.8 \text{ T}$ would require a 87 km SSC tunnel

Protons with $p = 7 \text{ TeV}/c$, $B = 8.4 \text{ T}$ require CERN's 27 km LHC tunnel

$$P_{\text{radiation}} = \frac{c}{6\pi\epsilon_0} N \frac{q^2}{\rho^2} \gamma^4 \quad \Downarrow$$

Energy needed to compensate
Radiation becomes too large



Electron beam with $p = 0.1 \text{ TeV}/c$ in CERN's 27 km LEP tunnel radiated 20 MW
Each electron lost about 4 GeV per turn, requiring many RF accelerating sections.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

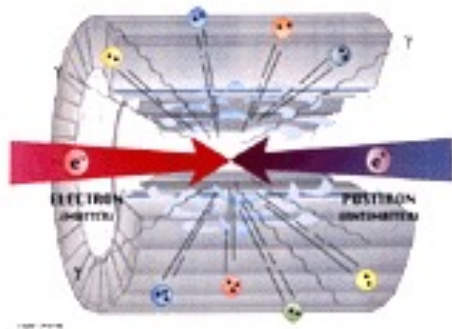
USPAS summer 2023

Colliding beam accelerators

- 1961: First storage ring for electrons and positrons (AdA) in Frascati for 250MeV
- 1972: SPEAR electron positron collider at 4GeV. Discovery of the J/Psi at 3.097GeV by Richter (SPEAR) and Ting (AGS) starts the November revolution and was essential for the quark model and chromodynamics.
- 1979: 5GeV electron positron collider CESR (designed for 8GeV)

Advantage:

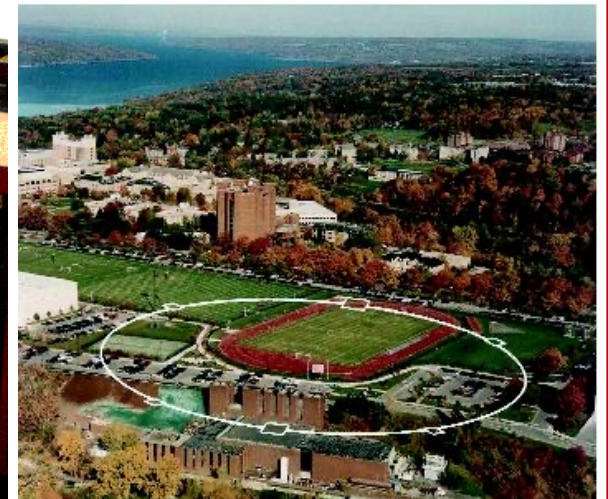
More center of mass energy



Drawback:

Less dense target

The beams therefore must be stored for a long time.



Cornell University



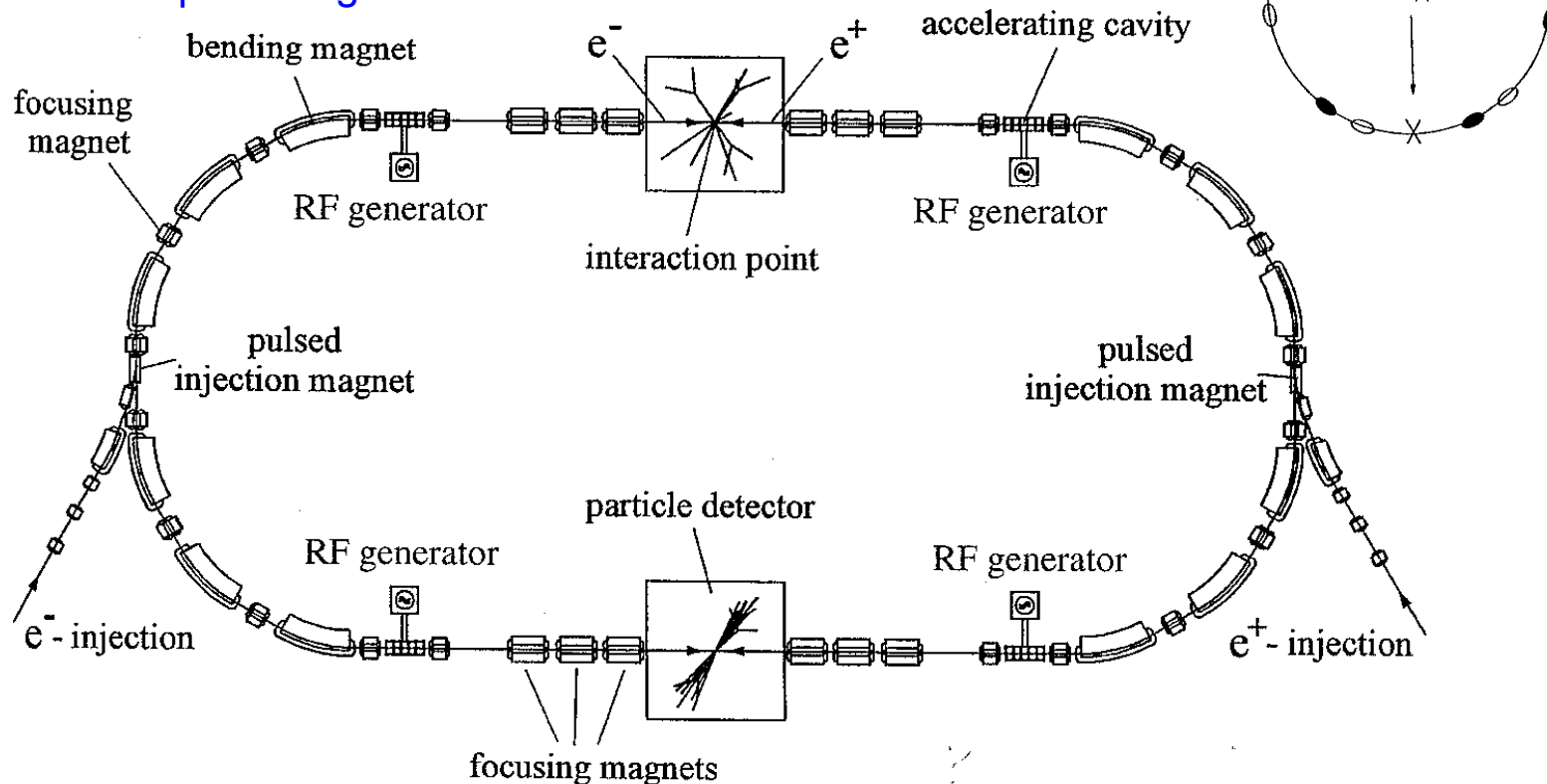
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Elements of a collider

- Saving one beam while injection another
- Avoiding collisions outside the detectors.
- Compensating the forces between e^+ and e^- beams



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Storage Rings

To avoid the loss of collision time during filling of a synchrotron, the beams in colliders must be stored for many millions of turns.

Challenges:

- Required vacuum of pressure below 10^{-7} Pa = 10^{-9} mbar, 3 orders of magnitude below that of other accelerators.
- Fields must be stable for a long time, often for hours.
- Field errors must be small, since their effect can add up over millions of turns.
- Even though a storage ring does not accelerate, it needs acceleration sections for phase focusing and to compensate energy loss due to the emission of radiation.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Further developments of Colliders

- 1981: Rubbia and van der Meer use stochastic cooling of anti-protons and discover W^+ , W^- and Z vector bosons of the weak interaction
- 1987: Start of the superconducting TEVATRON at FNAL
- 1989: Start of the 27km long LEP electron positron collider
- 1990: Start of the first asymmetric collider, electron (27.5GeV) proton (920GeV) in HERA at DESY
- 1998: Start of asymmetric two ring electron positron colliders KEK-B / PEP-II
- Today: 27km, 6.8 TeV proton collider LHC; Higgs discovery in 2012



NP 1984
Carlo Rubbia
Italy 1934 -



NP 1984
Simon van der Meer
Netherlands 1925 -



Cornell University



Brookhaven
National Laboratory

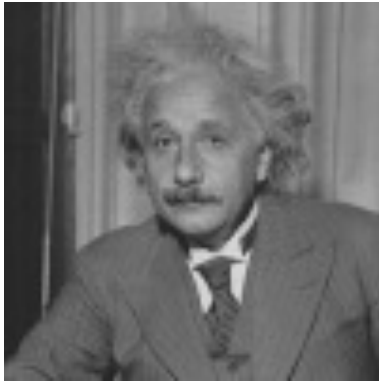
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Special Relativity

$$E = mc^2$$



Albert Einstein, 1879-1955
Nobel Prize, 1921
Time Magazine Man of the Century

Four-Vectors:

Quantities that transform according to the Lorentz transformation when viewed from a different inertial frame.

Examples:

$$X^\mu \in \{ct, x, y, z\}$$

$$P^\mu \in \{\frac{1}{c}E, p_x, p_y, p_z\}$$

$$\Phi^\mu \in \{\frac{1}{c}\phi, A_x, A_y, A_z\}$$

$$J^\mu \in \{c\rho, j_x, j_y, j_z\}$$

$$K^\mu \in \{\frac{1}{c}\omega, k_x, k_y, k_z\}$$

$$X^\mu \in \{ct, x, y, z\} \Rightarrow X^\mu X_\mu = (ct)^2 - \vec{x}^2 = \text{const.}$$

$$P^\mu \in \{\frac{1}{c}E, p_x, p_y, p_z\} \Rightarrow P^\mu P_\mu = \left(\frac{E}{c}\right)^2 - \vec{p}^2 = (m_0c)^2 = \text{const.}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

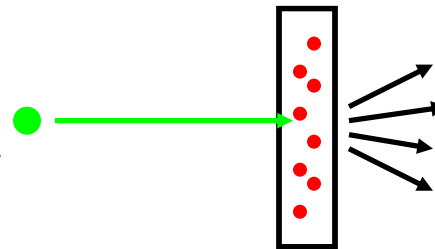
Graduate Accelerator Physics

USPAS summer 2023

Available Energy in Collisions

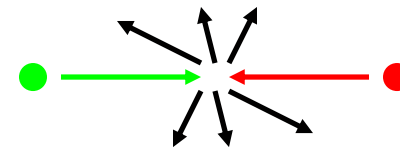
$$\begin{aligned}
 \frac{1}{c^2} E_{\text{cm}}^2 &= (P_1^\mu + P_2^\mu)_{\text{cm}} (P_{1\mu} + P_{2\mu})_{\text{cm}} \\
 &= (P_1^\mu + P_2^\mu)(P_{1\mu} + P_{2\mu}) \\
 &= \frac{1}{c^2} (E_1 + E_2)^2 - (p_{z1} - p_{z2})^2 \\
 &= 2\left(\frac{E_1 E_2}{c^2} + p_{z1} p_{z2}\right) + (m_{01} c)^2 + (m_{02} c)^2
 \end{aligned}$$

Operation of synchrotrons: fixed target experiments
where some energy is in the motion of the center of mass of the scattering products



$$E_1 \gg m_{01} c^2, m_{02} c^2; p_{z2} = 0; E_2 = m_{02} c^2 \Rightarrow E_{\text{cm}} = \sqrt{2E_1 m_{02} c^2}$$

Operation of colliders:
the detector is in the center of mass system



$$E_1 \gg m_{01} c^2; E_2 \gg m_{02} c^2 \Rightarrow E_{\text{cm}} = 2\sqrt{E_1 E_2}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Example: Fixed-target production of p-bar

- 1954: Operation of Bevatron, first proton synchrotron for 6.2 GeV, production of the anti-proton by Chamberlain and Segrè

$$p + p \mapsto p + p + p + \bar{p}$$

$$\frac{1}{c^2} E_{\text{cm}}^2 = 2 \left(\frac{E_1 E_2}{c^2} + p_{z1} p_{z2} \right) + (m_{01} c)^2 + (m_{02} c)^2$$

$$(4m_{p0}c)^2 < \frac{E_{\text{cm}}^2}{c^2} = 2 E_1 m_{p0} + (m_{p0}c)^2 + (m_{p0}c)^2$$

$$7m_{p0}c^2 < E_1$$

$$\underline{K_1 = E_1 - m_0 c^2} > \underline{6m_{p0}c^2 = 5.628 \text{ GeV}}$$



NP 1959

Emilio Gino Segrè

Italy 1905 – USA 1989



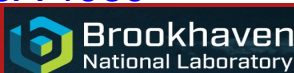
NP 1959

Owen Chamberlain

USA 1920 - 2006



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Example: production of c / c-bar states

- 1974: Observation of $c - \bar{c}$ resonances (J/Ψ) at $E_{cm} = 3095\text{MeV}$ at the e^+/e^- collider SPEAR

$$\frac{1}{c^2} E_{cm}^2 = 2\left(\frac{E_1 E_2}{c^2} + p_{z1} p_{z2}\right) + (m_{01}c)^2 + (m_{02}c)^2$$

Resonance in c/c-bar creation when $E_{cm} = 2m_{c0}c^2$ $E_1 = E_2 \Rightarrow E_{cm}^2 = 4E^2$

Energy per beam: $K = E - m_0c = \underline{1547\text{MeV}}$

Beam energy needed for an equivalent fixed target experiment: $\frac{E_{cm}^2}{c^2} = 2[Em + (mc)^2]$



NP 1976
Burton Richter
USA 1931 -

$$K = E - m_{0e}c^2 = \frac{E_{cm}^2}{2m_{0e}c^2} - 2m_{0e}c^2 = \underline{9.4\text{TeV}}$$



NP 1976
Samuel CC Ting
USA 1936 -

Rings for Synchrotron Radiation

- 1947: First detection of synchrotron light at General Electrics.
- 1952: First accurate measurement of synchrotron radiation power by Dale Corson with the Cornell 300MeV synchrotron.
- 1968: TANTALOS (U of Wisconsin), first dedicated storage ring for synchrotron radiation



Dale Corson
Cornell's 8th president
USA 1914 –



Cornell University



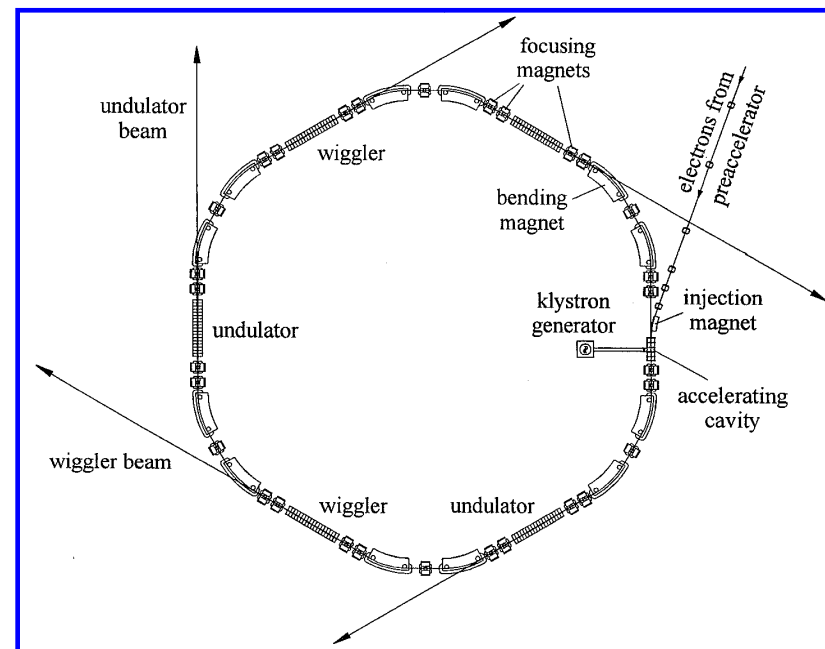
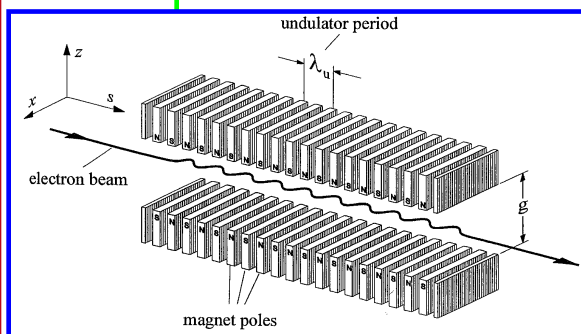
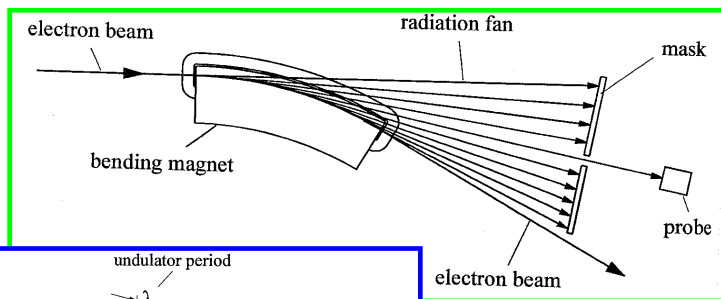
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

4 Generations of Light Sources

- 1st Generation (1970s): Many HEP rings are parasitically used for X-ray production
- 2nd Generation (1980s): Many dedicated X-ray sources (light sources)
- 3rd Generation (1990s): Several rings with dedicated radiation devices (wigglers and undulators)
- Today (4th Generation): Construction of Free Electron Lasers (FELs) driven by LINACs



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Macroscopic Fields in Accelerators

$$\frac{d}{dt} \vec{p} = q(\vec{E} + \vec{v} \times \vec{B})$$

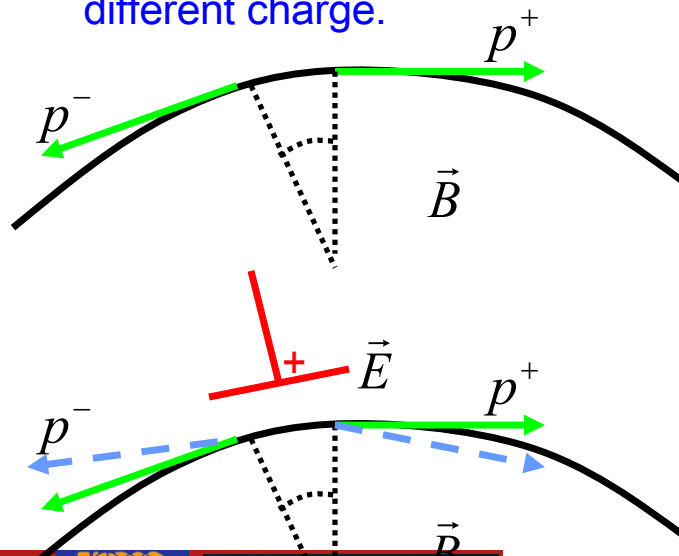
E has a similar effect as $v B$.

For relativistic particles $B = 1 \text{ T}$ has a similar effect as

$E = cB = 3 \cdot 10^8 \text{ V/m}$, such an

Electric field is beyond technical limits.

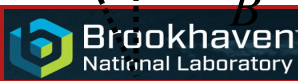
- Electric fields are only used for very low energies or
- For separating two counter rotating beams with different charge.



Electrostatic separators at CESR



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

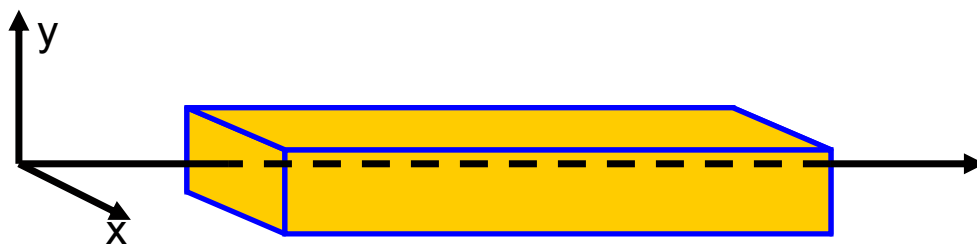
Magnetic Fields in Accelerators

Static magnetic fields: $\partial_t \vec{B} = 0$; $\vec{E} = 0$

Charge free space: $\vec{j} = 0$

$$\vec{\nabla} \times \vec{B} = \mu_0 (\vec{j} + \epsilon_0 \partial_t \vec{E}) = 0 \Rightarrow \vec{B} = -\vec{\nabla} \psi(\vec{r})$$

$$\vec{\nabla} \cdot \vec{B} = 0 \Rightarrow \vec{\nabla}^2 \psi(\vec{r}) = 0$$



$(x=0, y=0)$ is the beam's design curve

For finite fields on the design curve,
 Ψ can be power expanded in x and y :

$$\psi(x, y, z) = \sum_{n,m=0}^{\infty} b_{nm}(z) x^n y^m$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Surfaces of equal scalar magnetic potential



$$\vec{B}_{\perp}(\text{out}) = \vec{B}_{\perp}(\text{in})$$

$$\vec{H}_{\text{parallel}}(\text{out}) = \vec{H}_{\text{parallel}}(\text{in})$$

$$\vec{B}_{\text{parallel}}(\text{out}) = \frac{1}{\mu_r} \vec{B}_{\text{parallel}}(\text{in})$$

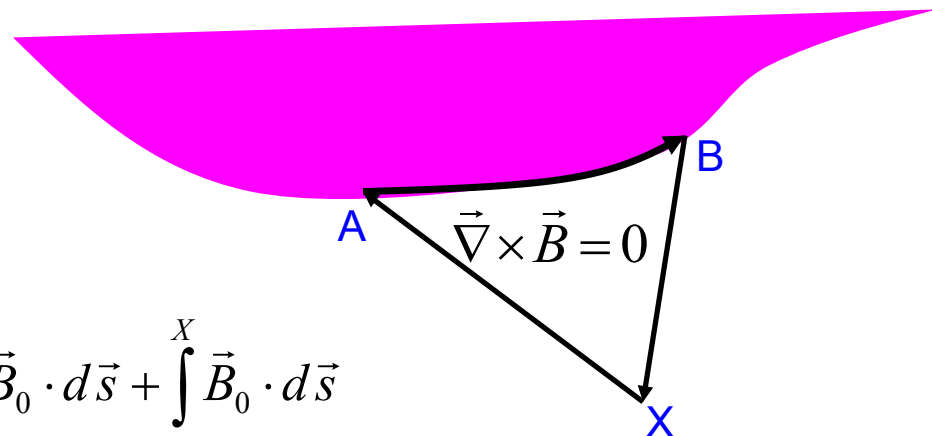
$$\vec{B}(\vec{r}) = -\vec{\nabla}\Psi(\vec{r})$$

$$0 = \oint \vec{B} \cdot d\vec{s} = \int_X^A \vec{B}_0 \cdot d\vec{s} + \int_A^B \vec{B}_0 \cdot d\vec{s} + \int_B^X \vec{B}_0 \cdot d\vec{s}$$

$$= \int_X^A \vec{B}_0 \cdot d\vec{s} + \frac{1}{\mu_r} \int_A^B \vec{B}_0 \cdot d\vec{s} + \int_B^X \vec{B}_0 \cdot d\vec{s}$$

$$\approx \int_X^A \vec{B}_0 \cdot d\vec{s} + \int_B^X \vec{B}_0 \cdot d\vec{s} = \Psi(A) - \Psi(B)$$

For large permeability, $H(\text{out})$ is perpendicular to the surface.



For highly permeable materials (like iron) surfaces have a constant potential.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Green's Theorem

$$\vec{\nabla}^2 \psi = 0$$

Green function:

$$\begin{aligned} \vec{\nabla}_0^2 G(\vec{r}, \vec{r}_0) &= \delta(\vec{r} - \vec{r}_0) & \psi(\vec{r}) &= \int_V \psi(\vec{r}_0) \delta(\vec{r} - \vec{r}_0) d^3 \vec{r}_0 \\ & & &= \int_V [\psi(\vec{r}_0) \vec{\nabla}_0^2 G - G \vec{\nabla}_0^2 \psi(\vec{r}_0)] d^3 \vec{r}_0 \\ & & &= \int_V \vec{\nabla}_0 [\psi(\vec{r}_0) \vec{\nabla}_0 G - G \vec{\nabla}_0 \psi(\vec{r}_0)] d^3 \vec{r}_0 \\ & & &= \int_{\partial V} [\psi(\vec{r}_0) \vec{\nabla}_0 G - G \vec{\nabla}_0 \psi(\vec{r}_0)] \cdot d^2 \vec{r}_0 \\ & & &= \int_{\partial V} [\psi(\vec{r}_0) \vec{\nabla}_0 G + \vec{B}(\vec{r}_0) G] \cdot d^2 \vec{r}_0 \end{aligned}$$

Knowledge of the field and the scalar magnetic potential on a closed surface inside a magnet determines the magnetic field for the complete volume which is enclosed.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Expansion of the scalar potential

If field data in a plane (for example the midplane of a cyclotron or of a beam line magnet) is known, the complete field is determined:

$$\psi(x, y, z) = \sum_{n=0}^{\infty} b_n(x, z) y^n \quad \Rightarrow \quad \vec{B}(x, 0, z) = - \begin{pmatrix} \partial_x b_0(x, z) \\ b_1(x, z) \\ \partial_z b_0(x, z) \end{pmatrix}$$

$$\begin{aligned} 0 = \vec{\nabla}^2 \psi &= \sum_{n=0}^{\infty} (\partial_x^2 + \partial_z^2) b_n y^n + \sum_{n=2}^{\infty} n(n-1) b_n y^{n-2} \\ &= \sum_{n=0}^{\infty} \left[(\partial_x^2 + \partial_z^2) b_n + (n+2)(n+1) b_{n+2} \right] y^n \end{aligned}$$

$$b_{n+2}(x, z) = - \frac{1}{(n+2)(n+1)} (\partial_x^2 + \partial_z^2) b_n(x, z)$$

Data of the magnetic field in the plane $y=0$ is used to determine $b_0(x, z)$ and $b_1(x, z)$.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Complex expansion of the potential

$$w = x + iy, \quad \bar{w} = x - iy$$

$$\partial_x = \partial_w + \partial_{\bar{w}}, \quad \partial_y = i\partial_w - i\partial_{\bar{w}} = i(\partial_w - \partial_{\bar{w}})$$

$$\underline{\vec{\nabla}^2 = \partial_x^2 + \partial_y^2 + \partial_z^2 = (\partial_w + \partial_{\bar{w}})^2 - (\partial_w - \partial_{\bar{w}})^2 + \partial_z^2 = 4\partial_w \partial_{\bar{w}} + \partial_z^2}$$

$$\psi = \text{Im} \left\{ \sum_{\nu, \lambda=0}^{\infty} a_{\nu\lambda}(z) \cdot (w\bar{w})^\lambda \bar{w}^\nu \right\}$$

$$\vec{\nabla}^2 \psi = \text{Im} \left\{ \sum_{\nu=0, \lambda=1}^{\infty} 4a_{\nu\lambda}(\lambda + \nu)\lambda(w\bar{w})^{\lambda-1} \bar{w}^\nu + \sum_{\nu=0, \lambda=0}^{\infty} a''_{\nu\lambda}(w\bar{w})^\lambda \bar{w}^\nu \right\}$$

$$= \text{Im} \left\{ \sum_{\nu, \lambda=0}^{\infty} [4(\lambda + 1 + \nu)(\lambda + 1)a_{\nu\lambda+1} + a''_{\nu\lambda}] (w\bar{w})^\lambda \bar{w}^\nu \right\} = 0$$

Iteration equation:
$$a_{\nu\lambda+1} = \frac{-1}{4(\lambda + 1 + \nu)(\lambda + 1)} a''_{\nu\lambda}, \quad a_{\nu 0} = \Psi_\nu(z)$$



Cornell University



The functions $\Psi_\nu(z)$ determine the complete field inside a magnet.

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

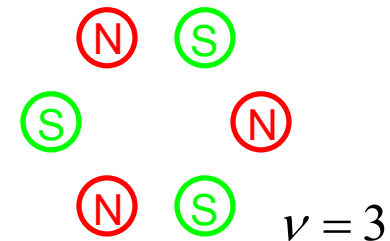
Multipole coefficients

$\Psi_\nu(z)$ are called the z-dependent multipole coefficients

$$\psi(x, y, z) = \text{Im} \left\{ \sum_{\nu, \lambda=0}^{\infty} \frac{(-1)^\lambda \nu!}{(\lambda + \nu)! \lambda!} \left(\frac{w \bar{w}}{4} \right)^\lambda \bar{w}^\nu \Psi_\nu^{[2\lambda]}(z) \right\}$$

$$\psi(r, \varphi, z) = \sum_{\nu, \lambda=0}^{\infty} \frac{(-1)^\lambda \nu!}{(\lambda + \nu)! \lambda!} \left(\frac{r}{2} \right)^{2\lambda} r^\nu \text{Im} \{ \Psi_\nu^{[2\lambda]}(z) e^{-i\nu\varphi} \}$$

The index ν describes C_ν Symmetry
around the z-axis \vec{e}_z
due to a sign change after $\Delta\varphi = \frac{\pi}{\nu}$



Cornell University



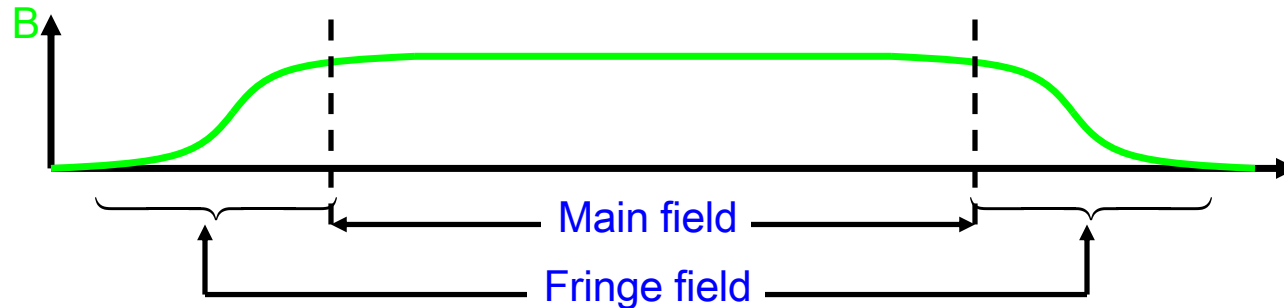
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Fringe Fields and Main fields



Only the fringe field region has terms with $\lambda \neq 0$ and $\partial_z^2 \psi \neq 0$

Main fields in accelerator physics: $\lambda = 0$, $\partial_z^2 \psi = 0$

$$\Psi_\nu = \begin{cases} e^{i\nu\vartheta_\nu} |\Psi_\nu| & \text{for } \nu \neq 0 \\ i |\Psi_0| & \text{for } \nu = 0 \end{cases}$$

$$\psi(r, \varphi) = \sum_{\nu=1}^{\infty} r^\nu |\Psi_\nu| \text{Im}\{e^{-i\nu(\varphi - \vartheta_\nu)}\} + |\Psi_0|$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Main-Field Potential

Main field potential:

$$\psi = |\Psi_0| - \sum_{\nu=1}^{\infty} r^{\nu} |\Psi_{\nu}| \sin[\nu(\varphi - \mathcal{G}_{\nu})]$$

The isolated multipole:

$$\psi = -r^{\nu} |\Psi_{\nu}| \sin(\nu\varphi)$$

Where the rotation \mathcal{G}_{ν} of the coordinate system is set to 0

The potentials produced by different multipole components Ψ_{ν} have

- a) Different rotation symmetry C_{ν}
- b) Different radial dependence r^{ν}



Cornell University

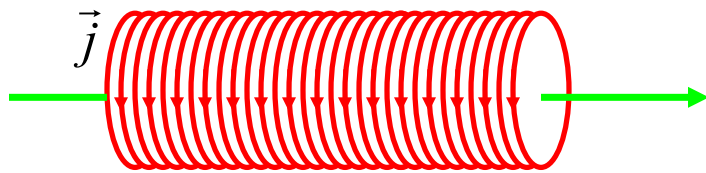


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Multipoles in Accelerators: $\nu=0$, Solenoids



$$m\gamma \begin{pmatrix} \ddot{x} \\ \ddot{y} \\ \ddot{z} \end{pmatrix} = q \begin{pmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{pmatrix} \times \begin{pmatrix} -\frac{x}{2} B'_z \\ -\frac{y}{2} B'_z \\ B_z \end{pmatrix}$$

\Downarrow

$$\begin{pmatrix} \ddot{x} \\ \ddot{y} \end{pmatrix} = \frac{qB_z}{m\gamma} \begin{pmatrix} \dot{y} \\ -\dot{x} \end{pmatrix} + \frac{qB'_z}{2m\gamma} \begin{pmatrix} y \\ -x \end{pmatrix}$$

\Downarrow

$$\ddot{w} = -i \frac{qB_z}{m\gamma} \dot{w} - i \frac{q\dot{B}_z}{2m\gamma} w$$

$$\psi = \Psi_0(z) - \frac{w\bar{w}}{4} \Psi_0''(z) \pm \dots$$

$$\vec{B} = \begin{pmatrix} \frac{x}{2} \Psi_0'' \\ \frac{y}{2} \Psi_0'' \\ -\Psi_0' \end{pmatrix} \Rightarrow \vec{\nabla} \cdot \vec{B} = 0$$

$$g = \frac{qB_z}{2m\gamma}, \quad w_0 = w e^{i \int_0^t g dt}$$

$$\ddot{w}_0 = (\ddot{w} + i2g\dot{w} + i\dot{g}w - g^2 w) e^{i \int_0^t g dt} = -g^2 w_0$$

$$\begin{aligned} \ddot{x}_0 &= -g^2 x_0 \\ \ddot{y}_0 &= -g^2 y_0 \end{aligned} \quad \begin{array}{l} \text{Focusing in a rotating} \\ \text{coordinate system} \end{array}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Strong vs. Solenoid Focusing

If the solenoids field was perpendicular to the particle' s motion,

its bending radius would be $\rho_z = \frac{p}{qB_z}$

$$\ddot{r} = -\left(\frac{qB_z}{2m\gamma}\right)^2 r = -\frac{qv_z}{m\gamma} B_z \frac{r}{4\rho_z}$$

Solenoid focusing is weak compared to the deflections created by a transverse magnetic field.

Transverse fields: $\vec{B} = B_x \vec{e}_x + B_y \vec{e}_y$

$$m\gamma \begin{pmatrix} \ddot{x} \\ \ddot{y} \\ \ddot{z} \end{pmatrix} = q \begin{pmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{pmatrix} \times \begin{pmatrix} B_x \\ B_y \\ 0 \end{pmatrix} \Rightarrow \begin{pmatrix} \ddot{x} \\ \ddot{y} \end{pmatrix} = \frac{qv_z}{m\gamma} \begin{pmatrix} -B_y \\ B_x \end{pmatrix}$$

Strong focusing

$$\ddot{x} = -\frac{q v_z}{m \gamma} \frac{\partial B_y}{\partial x} x$$

Weak focusing < Strong focusing by about magnet aperture / bending radius



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Natural Ring vs. Solenoid Focusing

Solenoid magnets are used in detectors for particle identification via $\rho = \frac{p}{qB}$

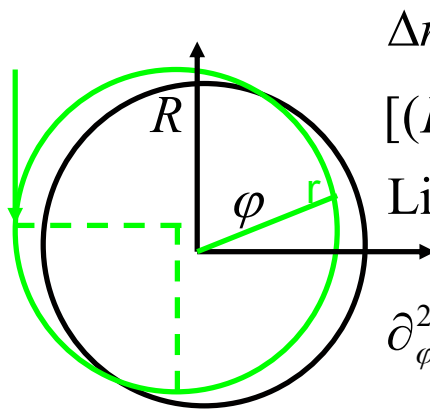
The solenoid's rotation $\dot{\phi} = -\frac{qB_z}{2m\gamma}$ of the beam is often compensated by a reversed solenoid called compensator.

Solenoid or Weak Focusing:

Solenoids are also used to focus low γ beams:

$$\ddot{w} = -\left(\frac{qB_z}{2m\gamma}\right)^2 w$$

Weak focusing from natural ring focusing:



$$\Delta r = r - R$$

$$[(R + \Delta r) \cos \varphi - \Delta x_0]^2 + [(R + \Delta r) \sin \varphi - \Delta y_0]^2 = R^2$$

$$\text{Linearization in } \Delta: \Delta r = (\cos \varphi \Delta x_0 + \sin \varphi \Delta y_0)$$

$$\partial_\varphi^2 \Delta r = -\Delta r \Rightarrow \Delta \ddot{r} = -\dot{\varphi}^2 \Delta r = -\left(\frac{v}{\rho}\right)^2 \Delta r = -\left(\frac{qB}{m\gamma}\right)^2 \Delta r$$

$$\frac{d^2 \Delta r}{ds^2} = \frac{d^2 \Delta r}{dt^2} \frac{1}{v^2} = \frac{1}{\rho^2} \Delta r$$

$$\text{Focusing strength: } \frac{1}{\rho^2}$$



Cornell University



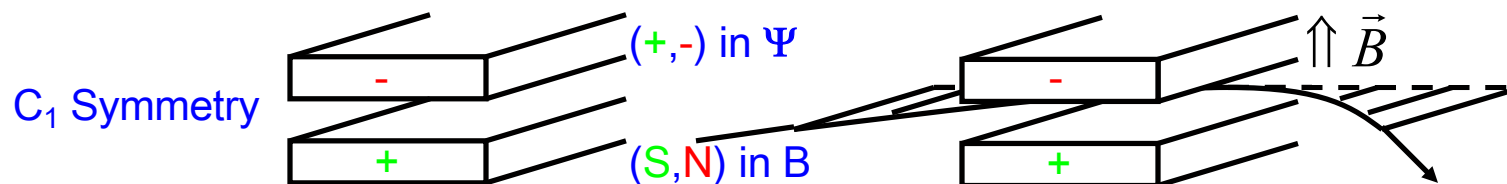
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

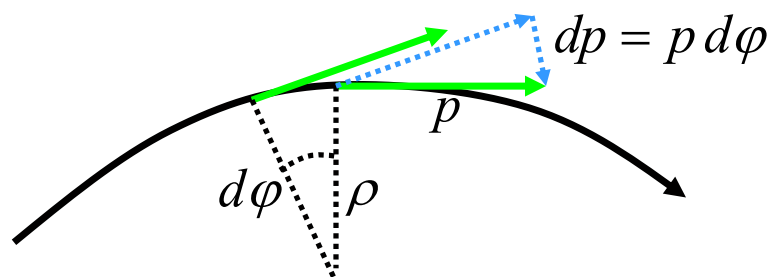
USPAS summer 2023

Multipoles in Accelerators: $\nu=1$, Dipoles

$$\psi = \Psi_1 \operatorname{Im}\{x - iy\} = -\Psi_1 \cdot y \Rightarrow \vec{B} = -\vec{\nabla} \psi = \Psi_1 \vec{e}_y \quad \text{Equipotential } y = \text{const.}$$



Dipole magnets are used for steering the beams direction



$$\frac{d\vec{p}}{dt} = q \vec{v} \times \vec{B} \Rightarrow \frac{dp}{dt} = qvB_{\perp} \Rightarrow \rho = \frac{dl}{d\phi} = \frac{vdt}{dp/p} = \frac{p}{qB_{\perp}}$$

Bending radius: $\rho = \frac{p}{qB}$



Cornell University



Brookhaven
National Laboratory

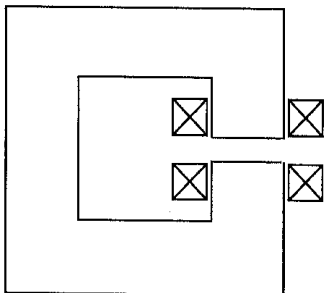
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

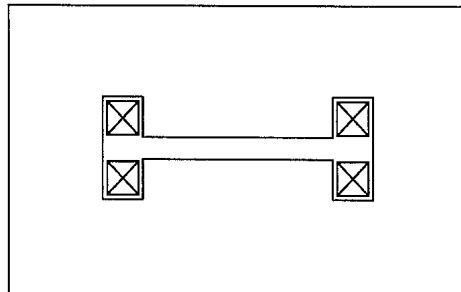
USPAS summer 2023

Types of iron-dominated Dipoles

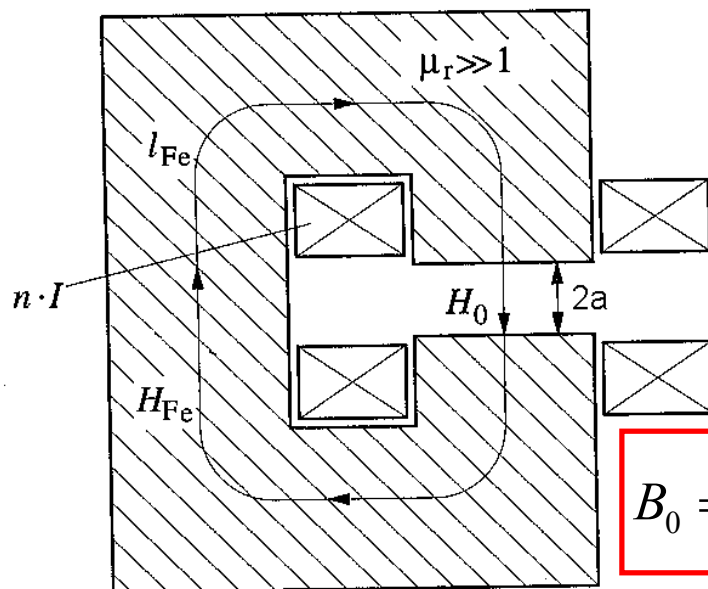
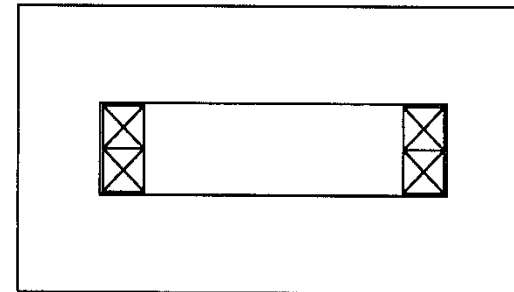
C-shape magnet:



H-shape magnet:



Window frame magnet:



$$\vec{B}_{\perp}(\text{out}) = \vec{B}_{\perp}(\text{in})$$

$$\vec{H}_{\perp}(\text{out}) = \mu_r \vec{H}_{\perp}(\text{in})$$

$$\begin{aligned} 2nI &= \oint \vec{H} \cdot d\vec{s} = H_{Fe} l_{Fe} + H_0 2a \\ &= \frac{1}{\mu_r} H_0 l_{Fe} + H_0 2a \approx H_0 2a \end{aligned}$$

$$B_0 = \mu_0 \frac{nI}{a}$$

Dipole strength: $\frac{1}{\rho} = \frac{q\mu_0}{p} \frac{nI}{a}$



Cornell University

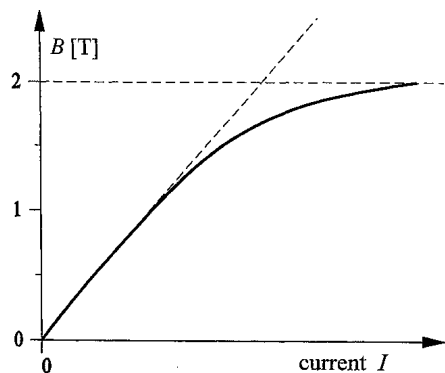


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Iron-dominated Dipoles Fields

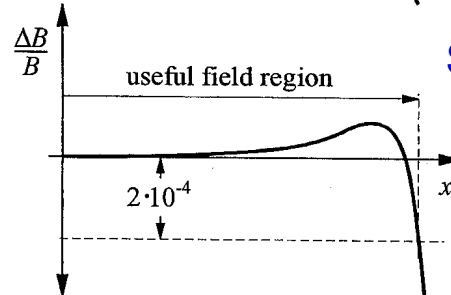
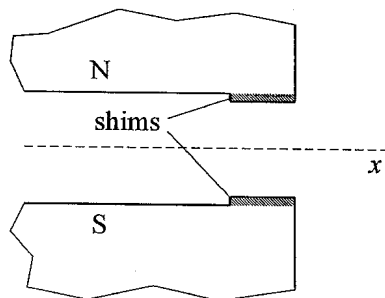
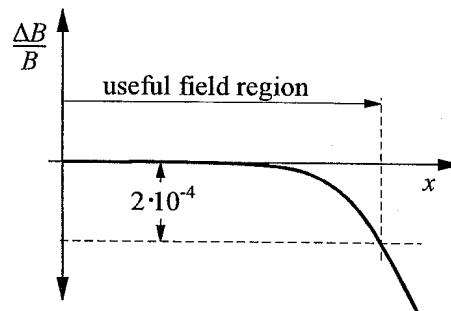
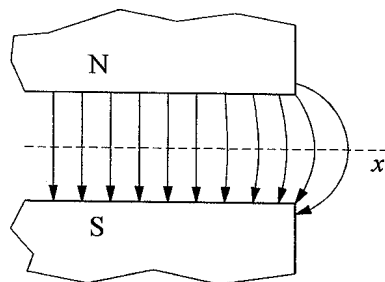


$B = 2 \text{ T}$: Typical limit, since the field becomes dominated by the coils, not the iron.

Limiting j for Cu is about 100 A/mm^2

$B < 1.5 \text{ T}$: Typically used region

$B < 1 \text{ T}$: Region in which $B_0 = \mu_0 \frac{nI}{a}$

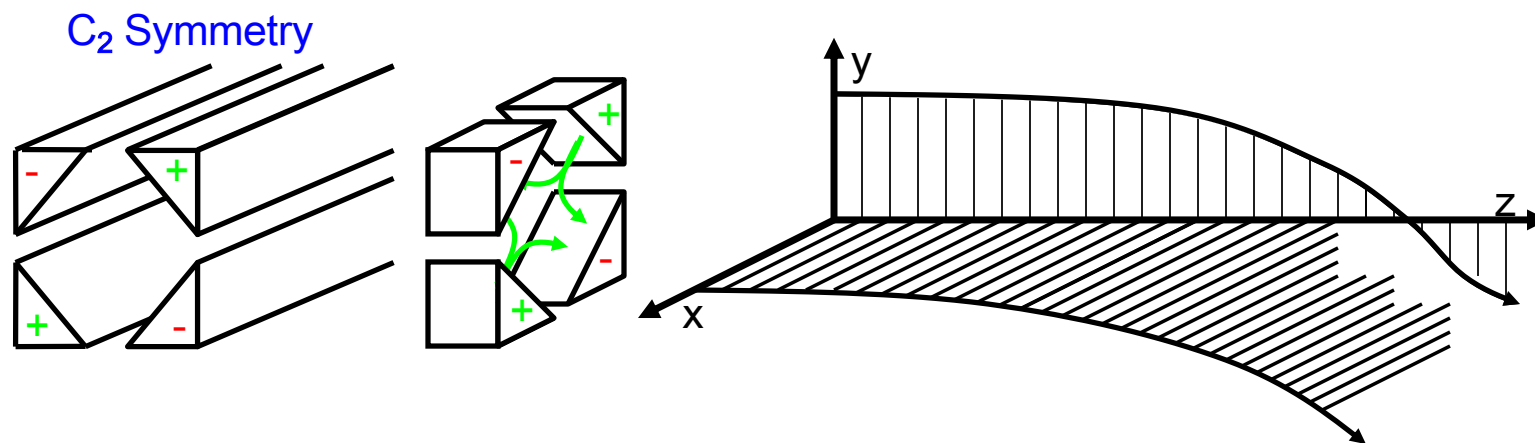


Shims reduce the space that is open to the beam, but they also reduce the fringe field region.



Multipoles in Accelerators: $\nu=2$, Quadrupoles

$$\psi = \Psi_2 \operatorname{Im}\{(x - iy)^2\} = -\Psi_2 \cdot 2xy \Rightarrow \vec{B} = -\vec{\nabla} \psi = \Psi_2 2 \begin{pmatrix} y \\ x \end{pmatrix}$$



In a **quadrupole** particles are focused in one plane and defocused in the other plane. Other modes of **strong focusing** are not possible.



Cornell University



Brookhaven
National Laboratory

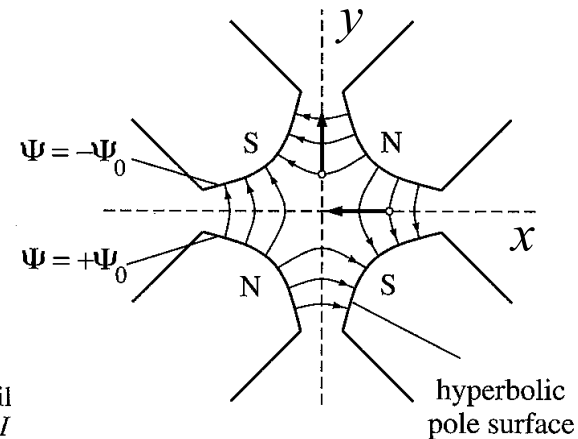
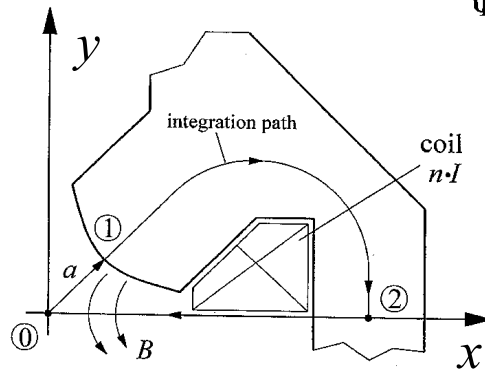
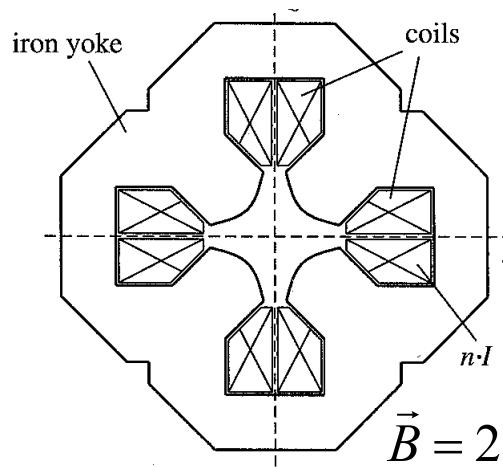
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Iron-dominated quadrupoles fields

$$\psi = -\Psi_2 \cdot 2xy \Rightarrow \text{Equipotential: } x = \frac{\text{const.}}{y}$$



$$\vec{B} = 2\Psi_2 \begin{pmatrix} y \\ x \end{pmatrix} \Rightarrow \vec{B}(0 \mapsto 1) = 2\Psi_2 r \vec{e}_r$$

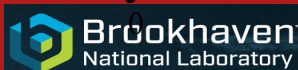
Quadrupole strength:

$$nI = \oint \vec{H} \cdot d\vec{s} \approx \int_a^a H_r dr = \Psi_2 \frac{a^2}{\mu_0}$$

$$k_1 = \frac{q}{p} \partial_x B_y \Big|_0 = \frac{q\mu_0}{p} \frac{2nI}{a^2}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Real Quadrupoles



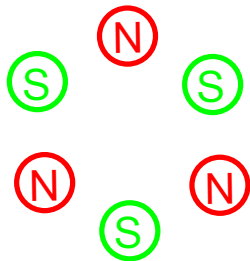
The coils show that this is an upright quadrupole not a rotated or skew quadrupole.



Multipoles in Accelerators: $\nu=3$, Sextupoles

$$\psi = \Psi_3 \operatorname{Im}\{(x-iy)^3\} = \Psi_3 \cdot (y^3 - 3x^2y) \Rightarrow \vec{B} = -\vec{\nabla} \psi = \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix}$$

C_3 Symmetry



i) Sextupole fields hardly influence the particles close to the center, where one can linearize in x and y .

ii) In linear approximation a by Δx shifted sextupole has a quadrupole field.

$$\vec{B} = -\vec{\nabla} \psi = \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix}$$

iii) When Δx depends on the energy, one can build an **energy dependent quadrupole**.

$$x \mapsto \Delta x + x$$

$$\vec{B} \approx \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix} + 6\Psi_3 \Delta x \begin{pmatrix} y \\ x \end{pmatrix} + O(\Delta x^2)$$



Cornell University



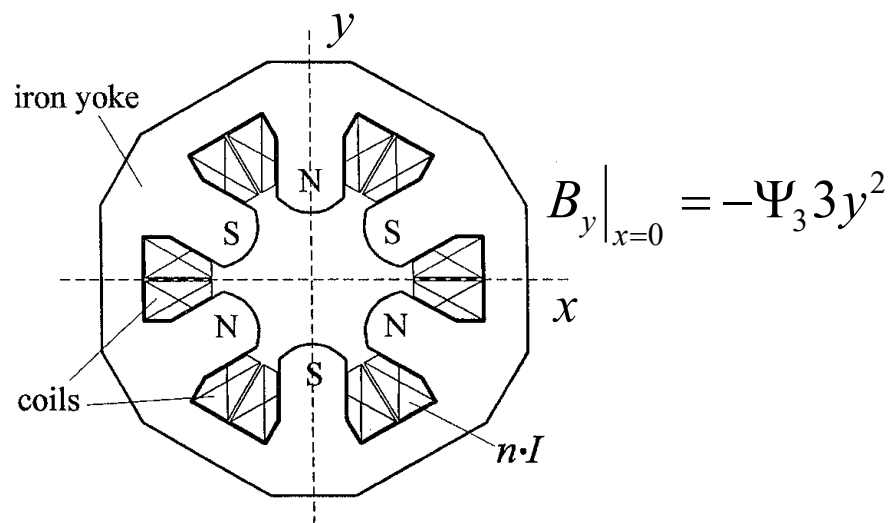
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

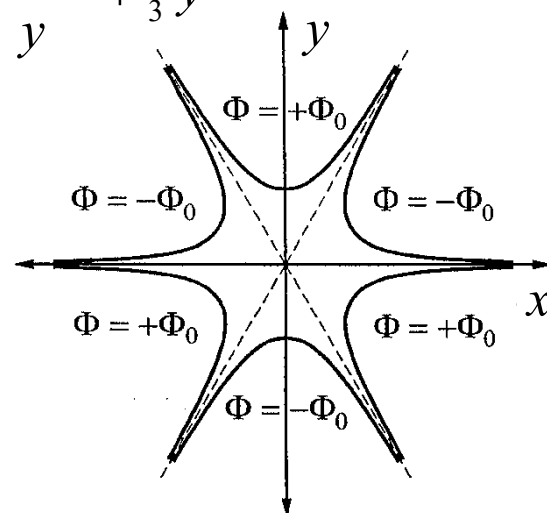
USPAS summer 2023

Iron-dominated sextupole fields

$$\psi = \Psi_2 \cdot (y^3 - 3x^2 y) \Rightarrow \text{Equipotential: } x = \sqrt{\frac{\text{const.}}{y} + \frac{1}{3} y^2}$$



$$nI = \oint \vec{H} \cdot d\vec{s} \approx \int_0^a H_r dr = \Psi_3 \frac{a^3}{\mu_0}$$



Quadrupole strength:

$$k_2 = \frac{q}{p} \partial_x^2 B_y|_0 = \frac{q\mu_0}{p} \frac{6nI}{a^3}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Real Sextupoles



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Higher-order multipoles

$$\psi = \Psi_n \operatorname{Im}\{(x - iy)^n\} = \Psi_n \cdot (\dots - i n x^{n-1} y) \Rightarrow \vec{B}(y=0) = \Psi_n n \begin{pmatrix} 0 \\ x^{n-1} \end{pmatrix}$$

Multipole strength: $k_n = \frac{q}{p} \partial_x^n B_y \big|_{x,y=0} = \frac{q}{p} \Psi_{n+1} (n+1)! \text{ units: } \frac{1}{\text{m}^{n+1}}$

p/q is also called **B_p** and used to describe the energy of multiply charge ions

Names: dipole, quadrupole, sextupole, octupole, decapole, duodecapole, ...

Higher order multipoles come from

- Field errors in magnets
- Magnetized materials
- From multipole magnets that compensate such erroneous fields
- To compensate nonlinear effects of other magnets
- To stabilize the motion of many particle systems
- To stabilize the nonlinear motion of individual particles



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

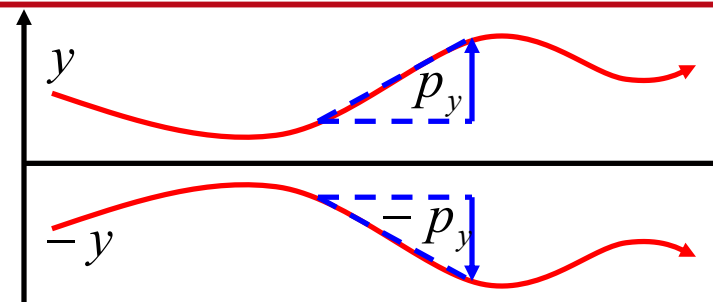
USPAS summer 2023

Midplane-symmetric motion

$$\vec{r}^{\oplus} = (x, -y, z)$$

$$\vec{p}^{\oplus} = (p_x, -p_y, p_z)$$

$$\frac{d}{dt} \vec{p} = \vec{F}(\vec{r}, \vec{p}) \Rightarrow \frac{d}{dt} \vec{p}^{\oplus} = \vec{F}(\vec{r}^{\oplus}, \vec{p}^{\oplus})$$



$$v_y B_z - v_z B_y = -v_y B_z(x, -y, z) - v_z B_y(x, -y, z) \quad B_x(x, -y, z) = -B_x(x, y, z)$$

$$v_z B_x - v_x B_z = -v_z B_x(x, -y, z) + v_x B_z(x, -y, z) \Rightarrow B_y(x, -y, z) = B_y(x, y, z)$$

$$v_x B_y - v_y B_x = v_x B_y(x, -y, z) + v_y B_x(x, -y, z) \quad B_z(x, -y, z) = -B_z(x, y, z)$$

$$\boxed{\psi(x, -y, z) = -\psi(x, y, z)}$$

$$\Psi_n \operatorname{Im} \left\{ e^{in\vartheta_n} (x + iy)^n \right\} = -\Psi_n \operatorname{Im} \left\{ e^{in\vartheta_n} (x - iy)^n \right\}$$

$$\Rightarrow \Psi_n \operatorname{Im} \left[e^{in\vartheta_n} 2 \operatorname{Re} \left\{ (x + iy)^n \right\} \right] = 0 \Rightarrow \boxed{\vartheta_n = 0}$$

The discussed multipoles

produce midplane symmetric motion. When the field is rotated by $\pi/2$,

i.e. $\vartheta_n = \pi/2n$, one speaks of a **skew multipole**.

Superconducting magnets

Above 2T the field from the bare coils dominate over the magnetization of the iron.

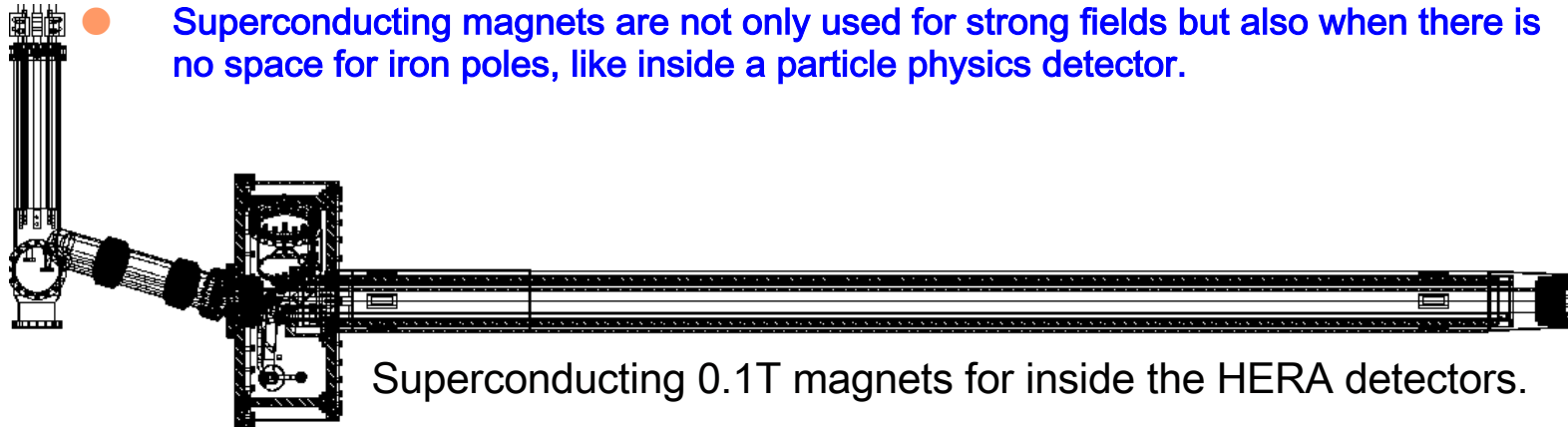
But Cu wires cannot create much field without iron poles:

5T at 5cm distance from a 3cm wire would require a current density of

$$j = \frac{I}{d^2} = \frac{1}{d^2} \frac{2\pi r B}{\mu_0} = 1389 \frac{\text{A}}{\text{mm}^2}$$

Cu can only support about 100A/mm².

- Superconducting cables routinely allow current densities of 1500A/mm² at 4.6 K and 6T. Materials used are usually Nb alloys, e.g. NbTi, Nb₃Ti or Nb₃Sn.
- Superconducting magnets are not only used for strong fields but also when there is no space for iron poles, like inside a particle physics detector.



Superconducting 0.1T magnets for inside the HERA detectors.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

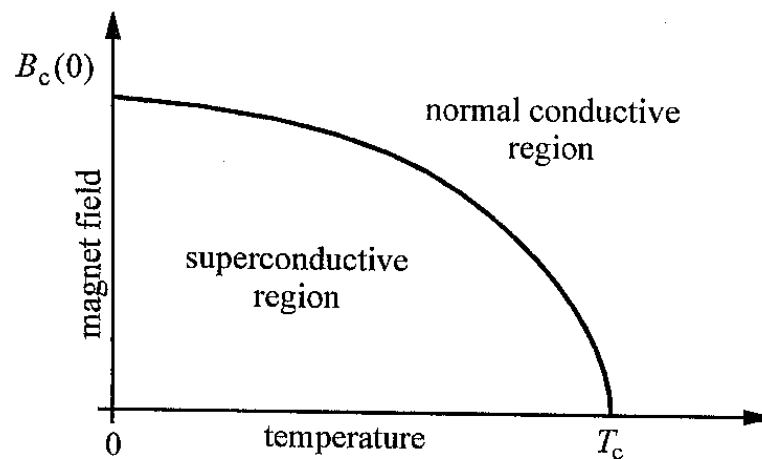
Graduate Accelerator Physics

USPAS summer 2023

Superconducting cables

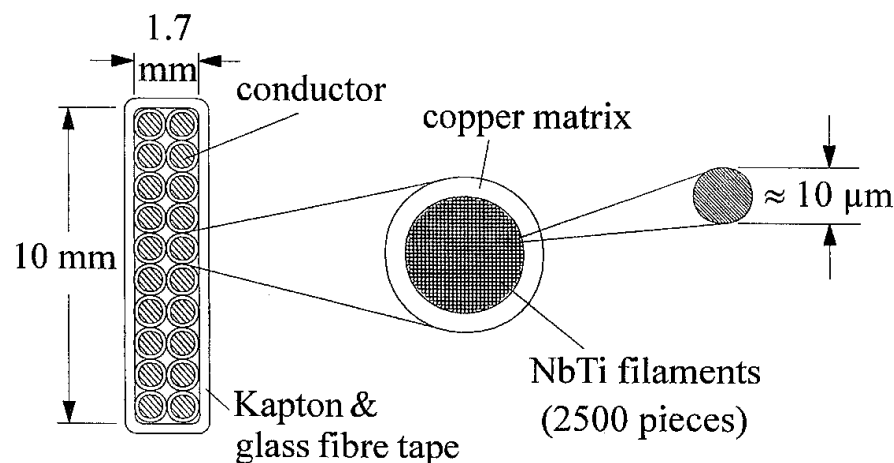
Problems:

- Superconductivity brakes down for too large fields
- Due to the Meissner-Ochsenfeld effect superconductivity current only flows on a thin surface layer.



Remedy:

- Superconducting cable consists of many very thin filaments (about $10\mu\text{m}$).



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Complex scalar magnetic potential of a wire

Straight wire at the origin: $\vec{\nabla} \times \vec{B} = \mu_0 \vec{j} \Rightarrow \vec{B}(r) = \frac{\mu_0 I}{2\pi r} \vec{e}_\phi = \frac{\mu_0 I}{2\pi r} \begin{pmatrix} -y \\ x \end{pmatrix}$

Wire at \vec{a} :

$$\vec{B}(x, y) = \frac{\mu_0 I}{2\pi (\vec{r} - \vec{a})^2} \begin{pmatrix} -[y - a_y] \\ x - a_x \end{pmatrix}$$

This can be represented by complex multipole coefficients Ψ_ν

$$\vec{B}(x, y) = -\vec{\nabla} \Psi \Rightarrow B_x + iB_y = -(\partial_x + i\partial_y)\psi = -2\partial_{\bar{w}}\psi$$

$$\begin{aligned} B_x + iB_y &= \frac{\mu_0 I}{2\pi} \frac{-i(w_a - w)}{(w_a - w)(\bar{w}_a - \bar{w})} = i \frac{\mu_0 I}{2\pi} \frac{-\frac{w_a}{a^2}}{1 - \frac{\bar{w}w_a}{a^2}} \\ &= i \frac{\mu_0 I}{2\pi} \partial_{\bar{w}} \ln(1 - \frac{\bar{w}w_a}{a^2}) = -2\partial_{\bar{w}} \operatorname{Im} \left\{ \frac{\mu_0 I}{2\pi} \ln(1 - \frac{\bar{w}w_a}{a^2}) \right\} \end{aligned}$$

$$\psi = \operatorname{Im} \left\{ \frac{\mu_0 I}{2\pi} \ln(1 - \frac{\bar{w}w_a}{a^2}) \right\} = -\operatorname{Im} \left\{ \frac{\mu_0 I}{2\pi} \sum_{\nu=1}^{\infty} \frac{1}{\nu} \left(\frac{w_a}{a^2} \right)^\nu \bar{w}^\nu \right\} \Rightarrow \Psi_\nu = \frac{\mu_0 I}{2\pi} \frac{1}{\nu} \frac{1}{a^\nu} e^{i\nu\phi_a}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Air-coil multipoles

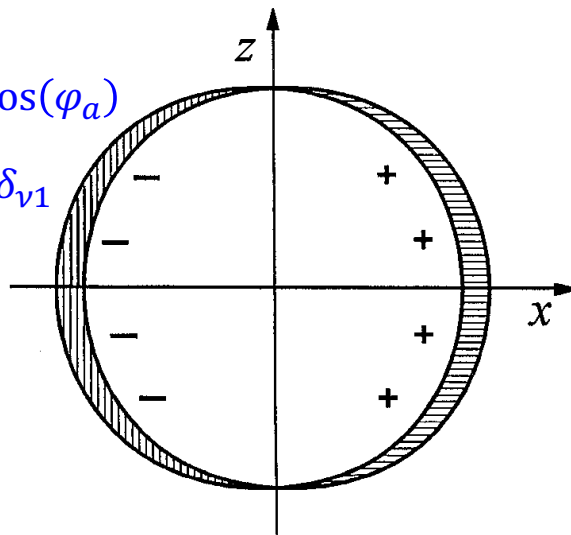
Creating a multipole be created by an arrangement of wires:

$$\Psi_\nu = \int_0^{2\pi} \frac{\mu_0}{2\pi} \frac{1}{\nu} \frac{1}{a^\nu} e^{i\nu\varphi_a} \frac{dI}{d\varphi_a} d\varphi_a$$

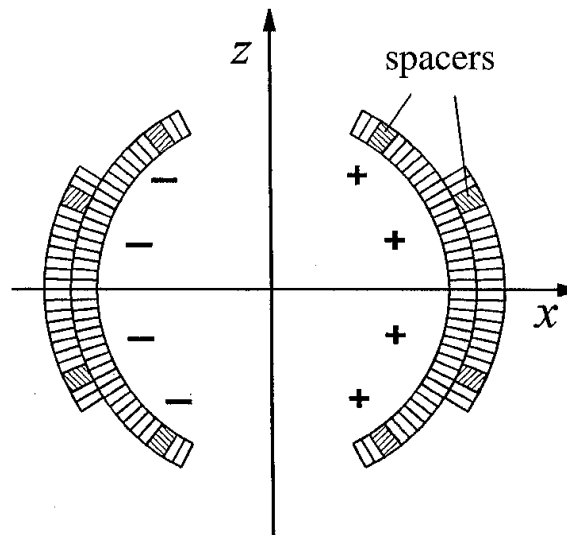
The Multipole coefficient Ψ_ν is the **Fourier coefficient** of the angular charge distribution $\frac{\mu_0}{\nu a^\nu} \frac{dI}{d\varphi_a}$.

$$\frac{dI}{d\varphi_a} \propto \cos(\varphi_a)$$

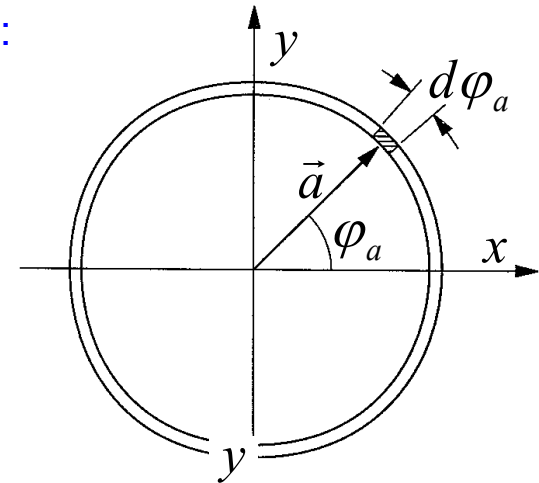
$$\Rightarrow \Psi_\nu \propto \delta_{\nu 1}$$



Ideal multipole



Approximate multipole



Cornell University

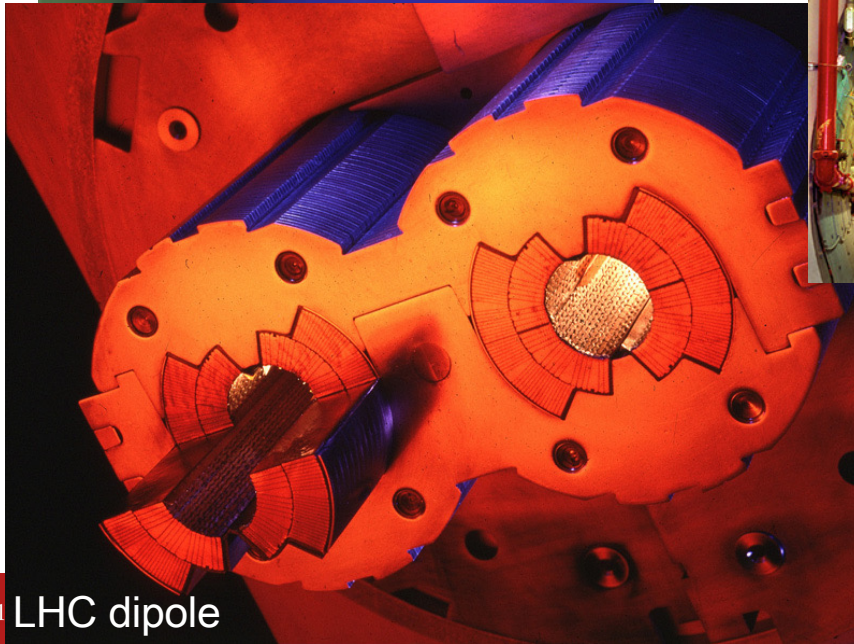


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Real Air-coil multipoles



LHC dipole



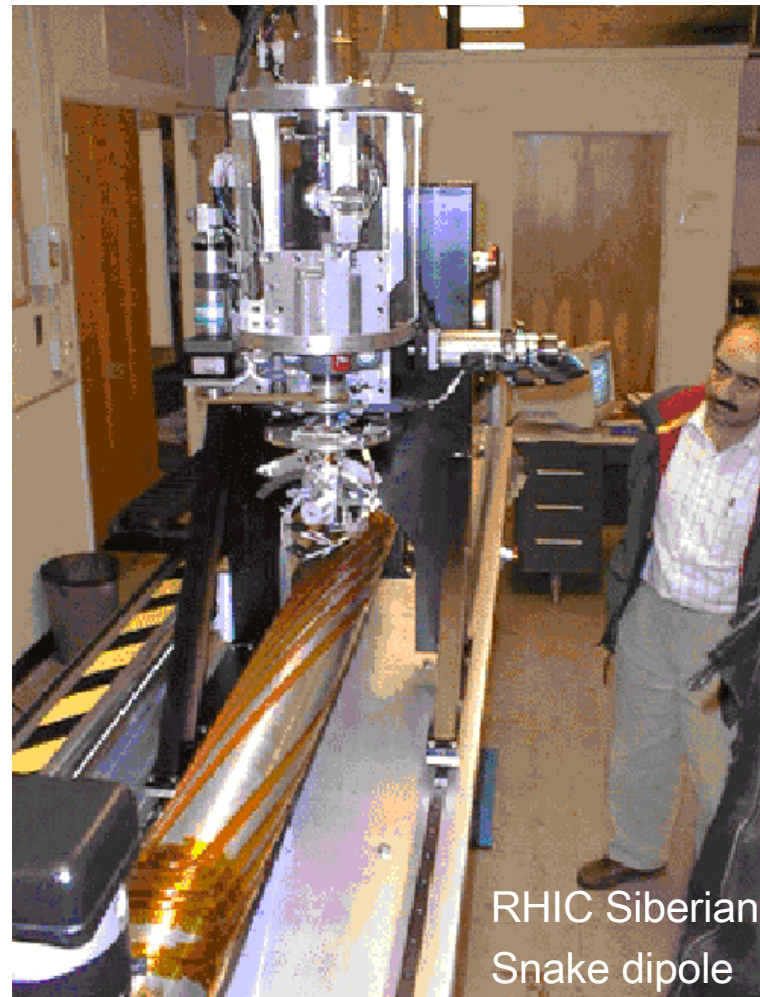
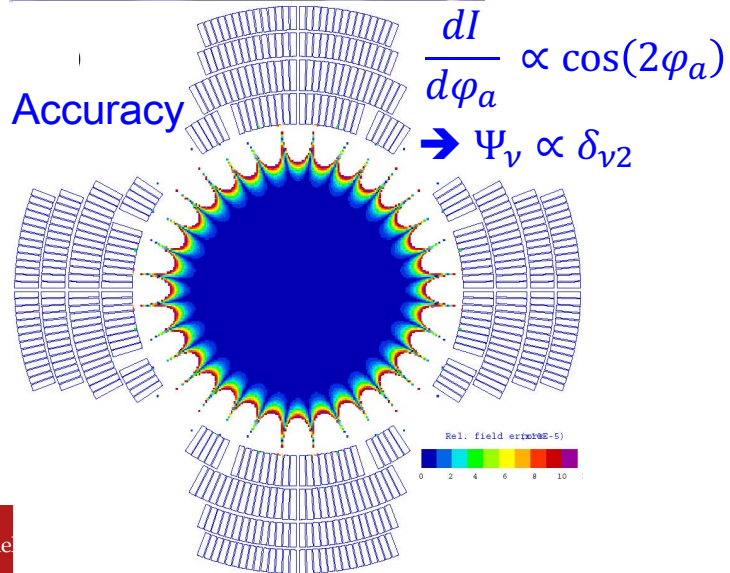
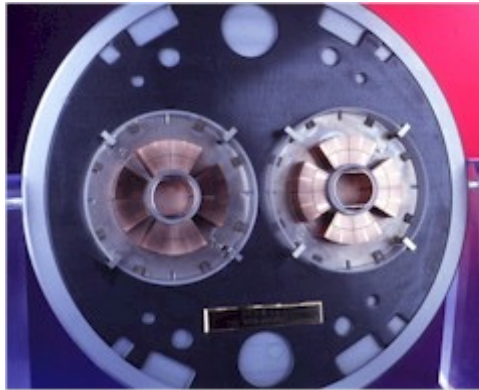
RHIC Tunnel



Cornell

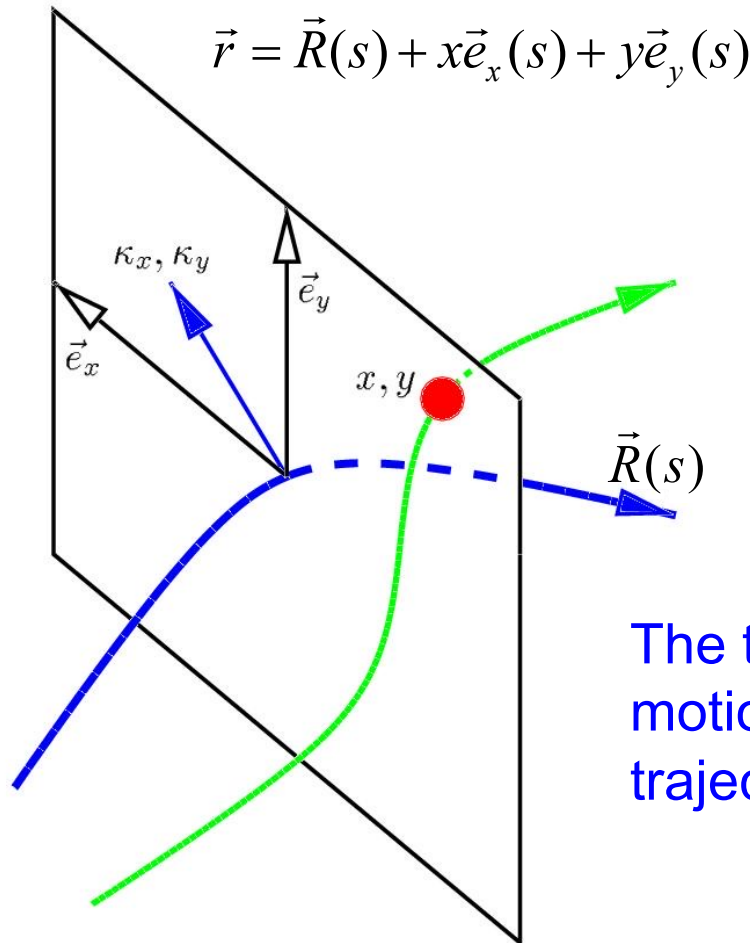
Special super-conducting Air-coil magnets

LHC double quadrupole



RHIC Siberian
Snake dipole

The comoving coordinate system



$$|d\vec{R}| = ds$$

$$\vec{e}_s \equiv \frac{d}{ds} \vec{R}(s)$$

The time dependence of a particle's motion is often not as interesting as the trajectory along the accelerator length "s".



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The 4-dimensional equation of motion

$$\frac{d^2}{dt^2} \vec{r} = \vec{f}_r(\vec{r}, \frac{d}{dt} \vec{r}, t)$$

3 dimensional ODE of 2nd order can be changed to a

6 dimensional ODE of 1st order:

$$\left. \begin{aligned} \frac{d}{dt} \vec{r} &= \frac{1}{m\gamma} \vec{p} = \frac{c}{\sqrt{p^2 - (mc)^2}} \vec{p} \\ \frac{d}{dt} \vec{p} &= \vec{F}(\vec{r}, \vec{p}, t) \end{aligned} \right\} \quad \frac{d}{dt} \vec{Z} = \vec{f}_Z(\vec{Z}, t), \quad \vec{Z} = (\vec{r}, \vec{p})$$

If the force does not depend on time, as in a typical beam line magnet, the energy is conserved so that one can reduce the dimension to 5. The equation of motion is then autonomous.

Furthermore, the time dependence is often not as interesting as the trajectory along the accelerator length “s”. Using “s” as the independent variable reduces the dimensions to 4. The equation of motion is then no longer autonomous.

$$\frac{d}{ds} \vec{Z} = \vec{f}_Z(\vec{Z}, s), \quad \vec{Z} = (x, y, p_x, p_y)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

6D equation of motion

Usually one prefers to compute the trajectory as a function of “s” along the accelerator even when the energy is not conserved, as when accelerating cavities are in the accelerator.

Then the energy “E” and the time “t” at which a particle arrives at the cavities are important. And the equations become 6 dimensional again:

$$\frac{d}{ds} \vec{z} = \vec{f}_z(\vec{z}, s), \quad \vec{z} = (x, y, p_x, p_y, -t, E)$$

But: $\vec{z} = (\vec{r}, \vec{p})$ is an especially suitable variable, since it is a phase space vector so that its equation of motion comes from a Hamiltonian, or by variation principle from a Lagrangian.

$$\delta \int [p_x \dot{x} + p_y \dot{y} + p_s \dot{s} - H(\vec{r}, \vec{p}, t)] dt = 0 \Rightarrow \text{Hamiltonian motion}$$

$$\delta \int [p_x x' + p_y y' - H t' + p_s (x, y, p_x, p_y, t, H)] ds = 0 \Rightarrow \text{Hamiltonian motion}$$

The new canonical coordinates are: $\vec{z} = (x, y, p_x, p_y, -t, E)$ with $E = H$

The new Hamiltonian is:

$$K = -p_s(\vec{z}, s)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

6D phase space motion

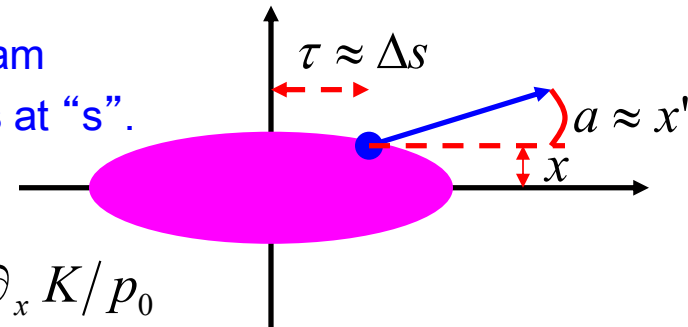
Using a reference momentum p_0 and a reference time t_0 :

$$\vec{z} = (x, a, y, b, \tau, \delta)$$

$$a = \frac{p_x}{p_0}, \quad b = \frac{p_y}{p_0}, \quad \delta = \frac{E - E_0}{E_0}, \quad \tau = (t_0 - t) \frac{c^2}{v_0} = (t_0 - t) \frac{E_0}{p_0}$$

Usually p_0 is the design momentum of the beam

And t_0 is the time at which the bunch center is at “s”.



$$\left. \begin{aligned} x' &= \partial_{p_x} K \\ p_x' &= -\partial_x K \end{aligned} \right\} \Rightarrow \left\{ \begin{aligned} x' &= \partial_a K / p_0, & a' &= -\partial_x K / p_0 \\ y' &= \partial_b K / p_0, & b' &= -\partial_y K / p_0 \end{aligned} \right.$$

$$-t' = \partial_E K \Rightarrow \tau' = \frac{c^2}{v_0} \partial_\delta K / E_0 = \partial_\delta K / p_0$$

$$E' = -\partial_{-t} K \Rightarrow \delta' = -\frac{1}{E_0} \partial_\tau K \frac{c^2}{v_0} = -\partial_\tau K / p_0$$

New Hamiltonian:

$$\tilde{H} = K / p_0$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

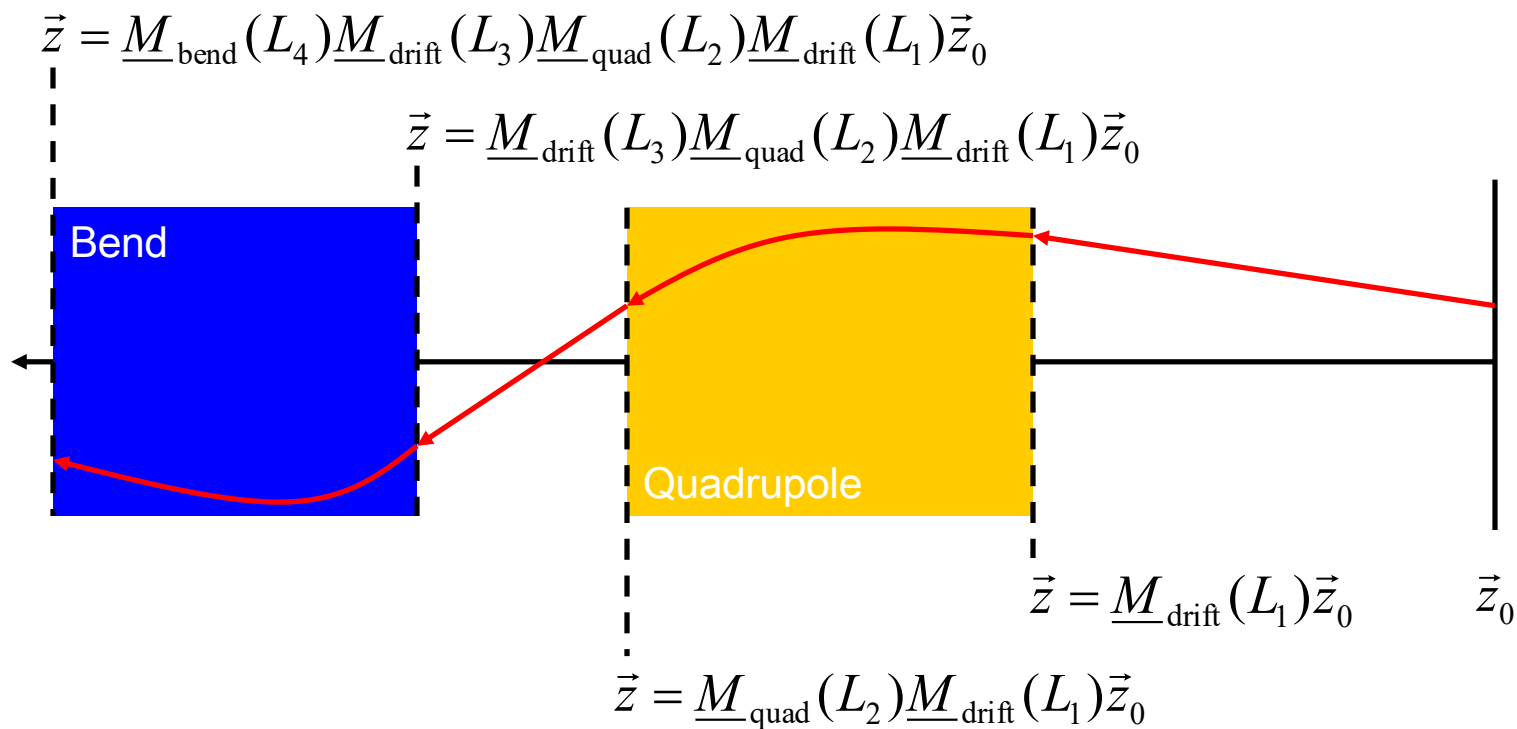
Graduate Accelerator Physics

USPAS summer 2023

The matrix solution of linear equations of motions

Linear equation of motion: $\vec{z}' = \underline{F}(s)\vec{z}$ $\rightarrow \vec{z}(s) = \underline{M}(s) \vec{z}_0$

Matrix solution of the starting condition $\vec{z}(0) = \vec{z}_0$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Simplest example: motion through an empty drift

$$\begin{pmatrix} x' \\ a' \\ y' \\ b' \\ \tau' \\ \delta' \end{pmatrix} = \begin{pmatrix} a \\ 0 \\ b \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

$$\ddot{x} = 0 \Rightarrow x'' = 0 \Rightarrow a = x', a' = 0$$

Linear solution:

$$x(s) = x_0 + x'_0 s$$

$$\begin{pmatrix} x \\ a \\ y \\ b \\ \tau \\ \delta \end{pmatrix} = \begin{pmatrix} x_0 + s a_0 \\ a \\ y_0 + s b_0 \\ b_0 \\ \tau_0 \\ \delta_0 \end{pmatrix} = \begin{pmatrix} 1 & s & \underline{0} & \underline{0} \\ 0 & 1 & \underline{0} & \underline{0} \\ \underline{0} & 1 & s & \underline{0} \\ \underline{0} & 0 & 1 & \underline{0} \\ \underline{0} & \underline{0} & 1 & 0 \\ \underline{0} & \underline{0} & 0 & 1 \end{pmatrix} \vec{z}_0$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Significance of the Hamiltonian

The equations of motion can be determined by one function:

$$\frac{d}{ds} x = \partial_{p_x} H(\vec{z}, s), \quad \frac{d}{ds} p_x = -\partial_x H(\vec{z}, s), \quad \dots$$

$$\frac{d}{ds} \vec{z} = \underline{J} \vec{\partial} H(\vec{z}, s) = \vec{F}(\vec{z}, s) \quad \text{with} \quad \underline{J} = \text{diag}(\underline{J}_2), \quad \underline{J}_2 = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$$

The force has a **Hamiltonian Jacobi Matrix**:

A linear force: $\vec{F}(\vec{z}, s) = \underline{F}(s) \cdot \vec{z}$

The **Jacobi Matrix** of a linear force: $\underline{F}(s)$

The general Jacobi Matrix : $F_{ij} = \partial_{z_j} F_i \quad \text{or} \quad \underline{F} = \left(\vec{\partial} \vec{F}^T \right)^T$

Hamiltonian Matrices: $\underline{F} \underline{J} + \underline{J} \underline{F}^T = 0$

Prove : $F_{ij} = \partial_{z_j} F_i = \partial_{z_j} J_{ik} \partial_{z_k} H = J_{ik} \partial_k \partial_j H \Rightarrow \underline{F} = \underline{J} \underline{D} \underline{H}$

$$\underline{F} \underline{J} + \underline{J} \underline{F}^T = \underline{J} \underline{D} \underline{J} \underline{H} + \underline{J} \underline{D}^T \underline{J}^T \underline{H} = 0$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Hamiltonian → Symplectic Flow

The flow of a Hamiltonian equation of motion has a **symplectic Jacobi Matrix**

The **flow** or **transport map**: $\vec{z}(s) = \vec{M}(s, \vec{z}_0)$

A linear flow: $\vec{z}(s) = \underline{M}(s) \cdot \vec{z}_0$

The Jacobi Matrix of a linear flow: $\underline{M}(s)$

The general **Jacobi Matrix** : $M_{ij} = \partial_{z_{0j}} M_i$ or $\underline{M} = \left(\vec{\partial}_0 \vec{M}^T \right)^T$

The **Symplectic Group** $SP(2N)$: $\underline{M} \underline{J} \underline{M}^T = \underline{J}$

$$\frac{d}{ds} \vec{z} = \frac{d}{ds} \vec{M}(s, \vec{z}_0) = \underline{J} \vec{\nabla} H = \vec{F} \quad \frac{d}{ds} M_{ij} = \partial_{z_{0j}} F_i(\vec{z}, s) = \partial_{z_{0j}} M_k \partial_{z_k} F_i(\vec{z}, s)$$

$$\underline{K} = \underline{M} \underline{J} \underline{M}^T \quad \frac{d}{ds} \underline{M}(s, \vec{z}_0) = \underline{F}(\vec{z}, s) \underline{M}(s, \vec{z}_0)$$

$$\frac{d}{ds} \underline{K} = \frac{d}{ds} \underline{M} \underline{J} \underline{M}^T + \underline{M} \underline{J} \frac{d}{ds} \underline{M}^T = \underline{F} \underline{M} \underline{J} \underline{M}^T + \underline{M} \underline{J} \underline{M}^T \underline{F}^T = \underline{F} \underline{K} + \underline{K} \underline{F}^T$$

$\underline{K} = \underline{J}$ is a solution. Since this is a linear ODE , $\underline{K} = \underline{J}$ is the unique solution.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Symplectic Flow → Hamiltonian

For every symplectic transport map there is a **Hamilton function**

The **flow or transport map**:

$$\vec{z}(s) = \vec{M}(s, \vec{z}_0)$$

Force vector:

$$\vec{h}(\vec{z}, s) = -\underline{J} \left[\frac{d}{ds} \vec{M}(s, \vec{z}_0) \right]_{\vec{z}_0 = \vec{M}^{-1}(\vec{z}, s)}$$

Since then:

$$\frac{d}{ds} \vec{z} = \underline{J} \vec{h}(\vec{z}, s)$$

There is a Hamilton function H with: $\vec{h} = \vec{\partial} H$

If and only if:

$$\partial_{z_j} h_i = \partial_{z_i} h_j \quad \Rightarrow \quad \underline{h} = \underline{h}^T$$

$$\underline{M} \underline{J} \underline{M}^T = \underline{J} \quad \Rightarrow \quad \begin{cases} \frac{d}{ds} \underline{M} \underline{J} \underline{M}^T = -\underline{M} \underline{J} \frac{d}{ds} \underline{M}^T \\ \underline{M}^{-1} = -\underline{J} \underline{M}^T \underline{J} \end{cases}$$

$$\vec{h} \circ \vec{M} = -\underline{J} \frac{d}{ds} \vec{M}$$

$$\underline{h}(\vec{M}) \underline{M} = -\underline{J} \frac{d}{ds} \underline{M}$$

$$\underline{h}(\vec{M}) = -\underline{J} \frac{d}{ds} \underline{M} \underline{M}^{-1} = \underline{J} \frac{d}{ds} \underline{M} \underline{J} \underline{M}^T \underline{J} = -\underline{J} \underline{M} \underline{J} \frac{d}{ds} \underline{M}^T \underline{J} = \underline{M}^{-T} \frac{d}{ds} \underline{M}^T \underline{J} = \underline{h}^T$$



Cornell University



Brookhaven
National Laboratory

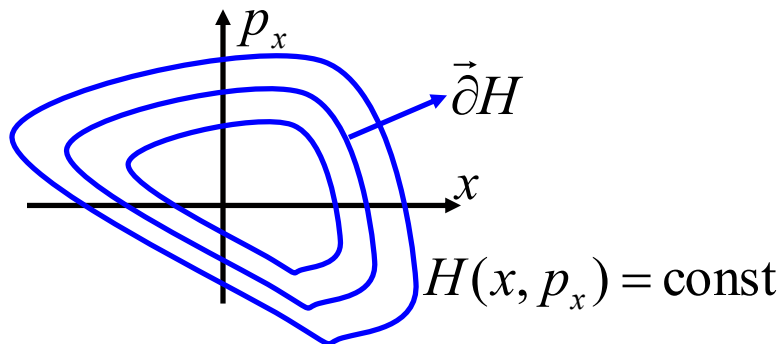
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

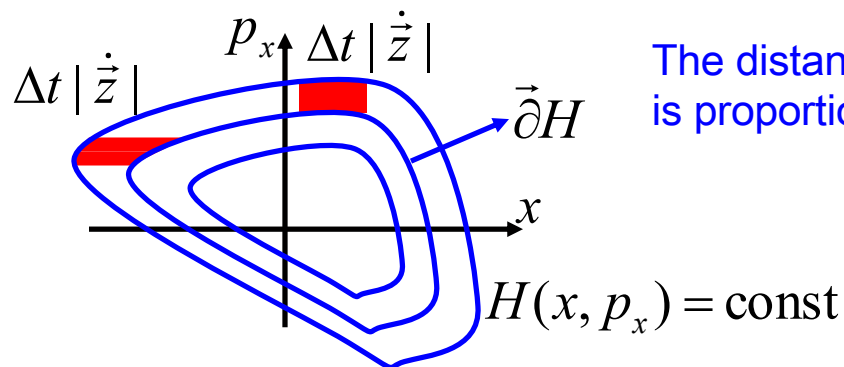
Phase space densities in 2D

- Phase space trajectories move on surfaces of constant energy



$$\frac{d}{ds} \vec{z} = \underline{J} \vec{\partial H} \Rightarrow \underline{\frac{d}{ds} \vec{z} \perp \vec{\partial H}}$$

- A phase space volume does not change when it is transported by Hamiltonian motion.



The distance d of lines with equal energy is proportional to $1 / |\vec{\partial H}| \propto |\dot{\vec{z}}|^{-1}$

$$d * \Delta t |\dot{\vec{z}}| = \text{const}$$



Cornell University



Brookhaven
National Laboratory

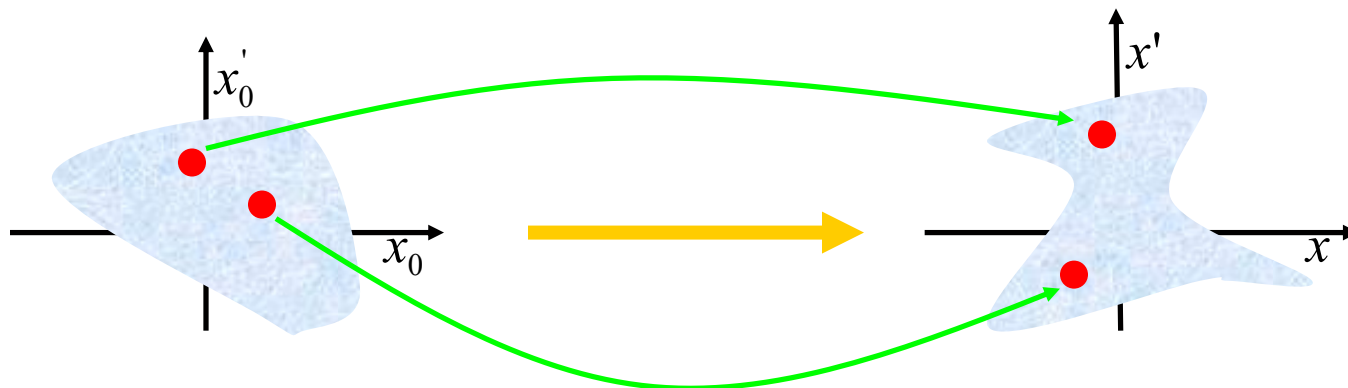
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Liouville's Theorem

- A phase space volume does not change when it is transported by Hamiltonian motion. $\vec{z}(s) = \underline{M}(s) \cdot \vec{z}_0$ with $\det[\underline{M}(s)] = +1$



$$\text{Volume} = V = \iint_V d^n \vec{z} = \iint_{V_0} \left| \frac{\partial \vec{z}}{\partial \vec{z}_0} \right| d^n \vec{z}_0 = \iint_{V_0} |\underline{M}| d^n \vec{z}_0 = \iint_{V_0} d^n \vec{z}_0 = V_0$$

$$\text{Hamiltonian Motion} \longrightarrow V = V_0$$

But Hamiltonian requires symplecticity, which is much more than just $\det[\underline{M}(s)] = +1$



Cornell University



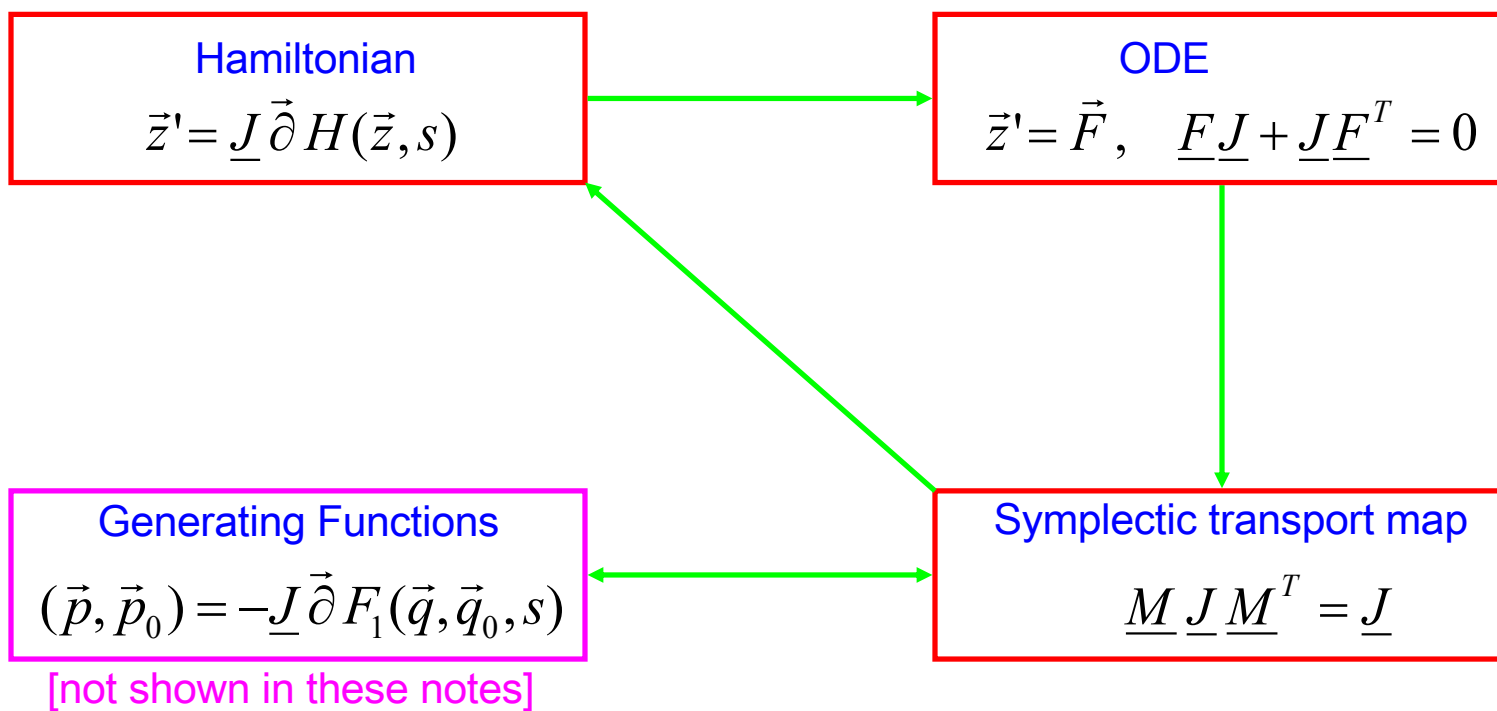
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Symplectic representations



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Eigenvalues of symplectic matrices

For matrices with real coefficients:

If there is an eigenvector and eigenvalue: $\underline{M}\vec{v}_i = \lambda_i \vec{v}_i$

then the complex conjugates are also eigenvector and eigenvalue: $\underline{M}\vec{v}_i^* = \lambda_i^* \vec{v}_i^*$

For symplectic matrices:

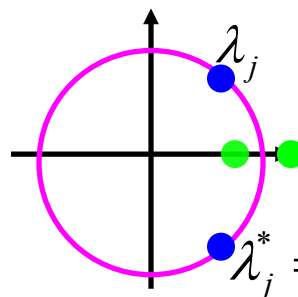
If there are eigenvectors and eigenvalues: $\underline{M}\vec{v}_i = \lambda_i \vec{v}_i$ with $\underline{J} = \underline{M}^T \underline{J} \underline{M}$

then $\vec{v}_i^T \underline{J} \vec{v}_j = \vec{v}_i^T \underline{M}^T \underline{J} \underline{M} \vec{v}_j = \lambda_i \lambda_j \vec{v}_i^T \underline{J} \vec{v}_j \Rightarrow \vec{v}_i^T \underline{J} \vec{v}_j (\lambda_i \lambda_j - 1) = 0$

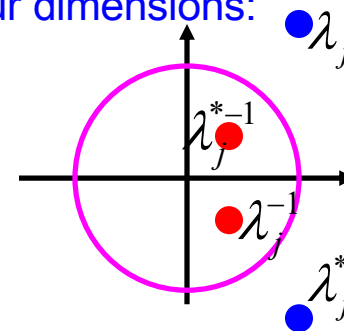
Therefore $\underline{J} \vec{v}_j$ is orthogonal to all eigenvectors with eigenvalues that are not $1/\lambda_j$. Since it cannot be orthogonal to all eigenvectors, there is at least one eigenvector with eigenvalue $1/\lambda_j$

Two dimensions: λ_j is eigenvalue
Then $1/\lambda_j$ and λ_j^* are eigenvalues

$$\lambda_2 = 1/\lambda_1 = \lambda_1^* \Rightarrow |\lambda_j| = 1$$



Four dimensions:



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Time of flight from symplecticity

$$\underline{M} = \begin{pmatrix} \underline{M}_4 & \vec{0} & \vec{D} \\ \vec{T}^T & 1 & M_{56} \\ \vec{0}^T & 0 & 1 \end{pmatrix} \text{ is in SU(6) and therefore } \underline{M} \underline{J} \underline{M}^T = \underline{J}$$

$$\begin{pmatrix} \underline{M}_4 \underline{J}_4 & -\vec{D} & \vec{0} \\ \vec{T}^T \underline{J}_4 & -M_{56} & 1 \\ \vec{0}^T & -1 & 0 \end{pmatrix} \begin{pmatrix} \underline{M}_4^T & \vec{T} & \vec{0} \\ \vec{0}^T & 1 & 0 \\ \vec{D}^T & M_{56} & 1 \end{pmatrix} = \begin{pmatrix} \underline{J}_4 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{pmatrix}$$

$$\begin{pmatrix} \underline{M}_4 \underline{J}_4 \underline{M}_4^T & \underline{M}_4 \underline{J}_4 \vec{T} - \vec{D} & \vec{0} \\ \vec{T}^T \underline{J}_4 \underline{M}_4^T + \vec{D}^T & 0 & 1 \\ \vec{0}^T & -1 & 0 \end{pmatrix} = \begin{pmatrix} \underline{J}_4 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{pmatrix}$$

$$\vec{T} = -\underline{J}_4 \underline{M}_4^{-1} \vec{D}$$

It is sufficient to compute the 4D map \underline{M}_4 , the Dispersion \vec{D} and the time of flight term M_{56}



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

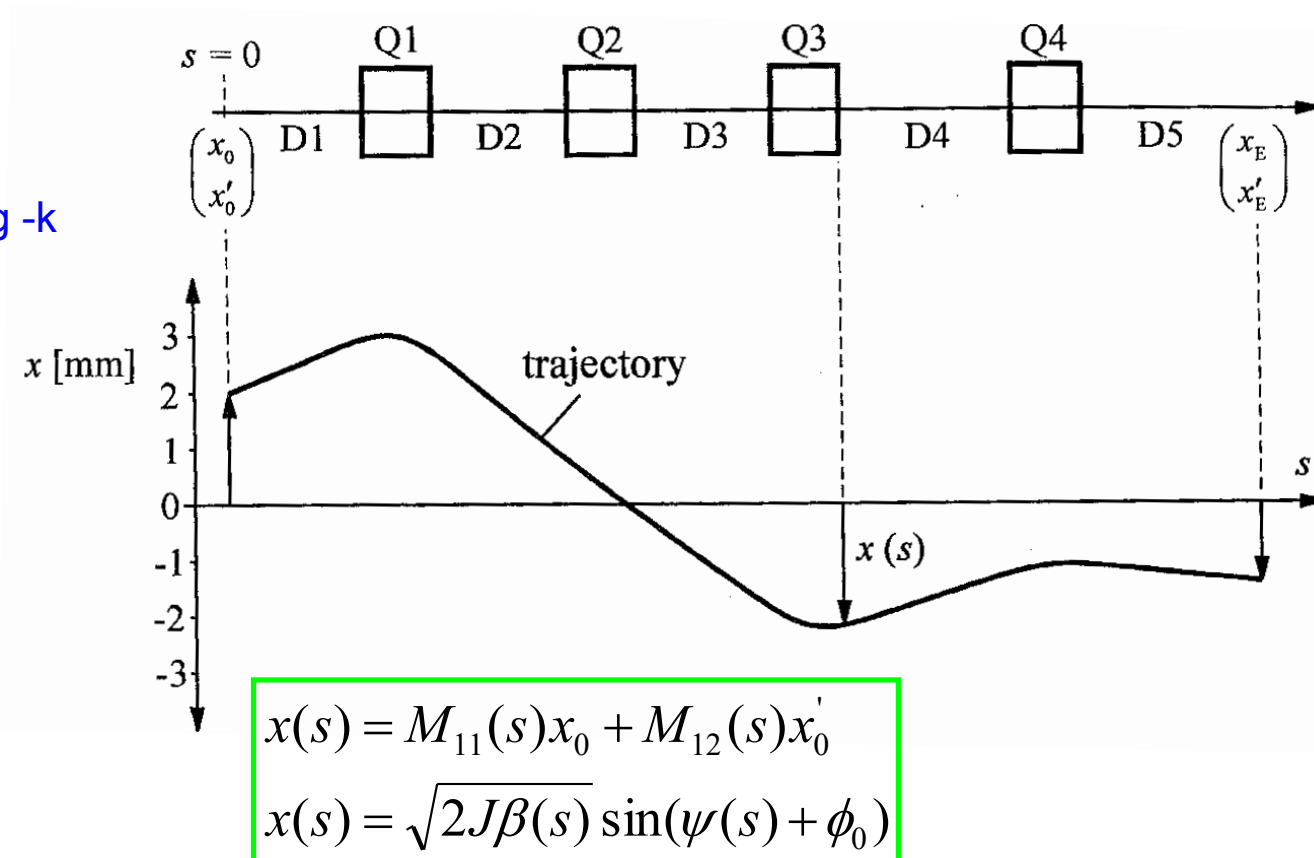
Betatron formalism for linear motion

$$x'' = -x K$$

$$y'' = y k$$

In y: quadrupole defocusing -k

$$\text{In x: } K = k + \frac{1}{\rho^2}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Twiss parameters

$$x'' = -k x$$

$$x(s) = \sqrt{2J\beta(s)} \sin(\psi(s) + \phi_0)$$

$$x'(s) = \sqrt{\frac{2J}{\beta}} [\beta\psi' \cos(\psi(s) + \phi_0) - \alpha \sin(\psi(s) + \phi_0)] \quad \text{with} \quad \alpha = -\frac{1}{2} \beta'$$

$$\begin{aligned} x''(s) &= \sqrt{\frac{2J}{\beta}} [(\beta\psi'' - 2\alpha\psi') \cos(\psi(s) + \phi_0) - (\alpha' + \frac{\alpha^2}{\beta} + \beta\psi'^2) \sin(\psi(s) + \phi_0)] \\ &= \sqrt{\frac{2J}{\beta}} [-k\beta \sin(\psi(s) + \phi_0)] \end{aligned}$$

$$\beta\psi'' - 2\alpha\psi' = \beta\psi'' + \beta'\psi' = (\beta\psi')' = 0 \quad \Rightarrow \quad \psi' = \frac{I}{\beta}$$

$$\alpha' + \gamma = k\beta \quad \text{with} \quad \underline{\gamma = \frac{I^2 + \alpha^2}{\beta}} \quad \text{Universal choice: } I=1!$$

$\alpha, \beta, \gamma, \psi$ are called
Twiss parameters.

$$\begin{aligned} \beta' &= -2\alpha \\ \alpha' &= k\beta - \gamma \\ \psi &= \int_0^s \frac{I}{\beta(s')} ds' \end{aligned}$$

What are the
initial conditions?



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The phase ellipse

Particles with a common J and different ϕ all lie on an ellipse in phase space:

$$\begin{pmatrix} x \\ x' \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi(s) + \phi_0) \\ \cos(\psi(s) + \phi_0) \end{pmatrix}$$

(Linear transform of a circle)

$$x_{\max} = \sqrt{2J\beta} \text{ at } x' = -\alpha\sqrt{\frac{2J}{\beta}}$$

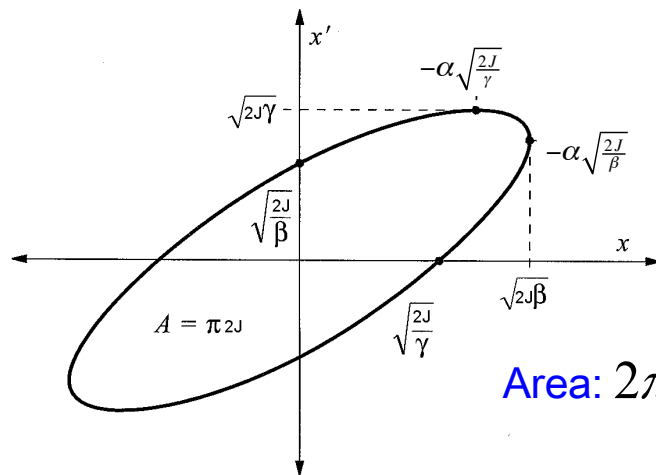
$$(x, x') \begin{pmatrix} \frac{1}{\sqrt{\beta}} & \frac{\alpha}{\sqrt{\beta}} \\ 0 & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} x \\ x' \end{pmatrix} = (x, x') \begin{pmatrix} \gamma & \alpha \\ \alpha & \beta \end{pmatrix} \begin{pmatrix} x \\ x' \end{pmatrix} = 2J$$

(Quadratic form)

$$\beta\gamma - \alpha^2 = I^2$$

$$\text{Area: } 2\pi J / I$$

$I=1$ is therefore a useful choice!



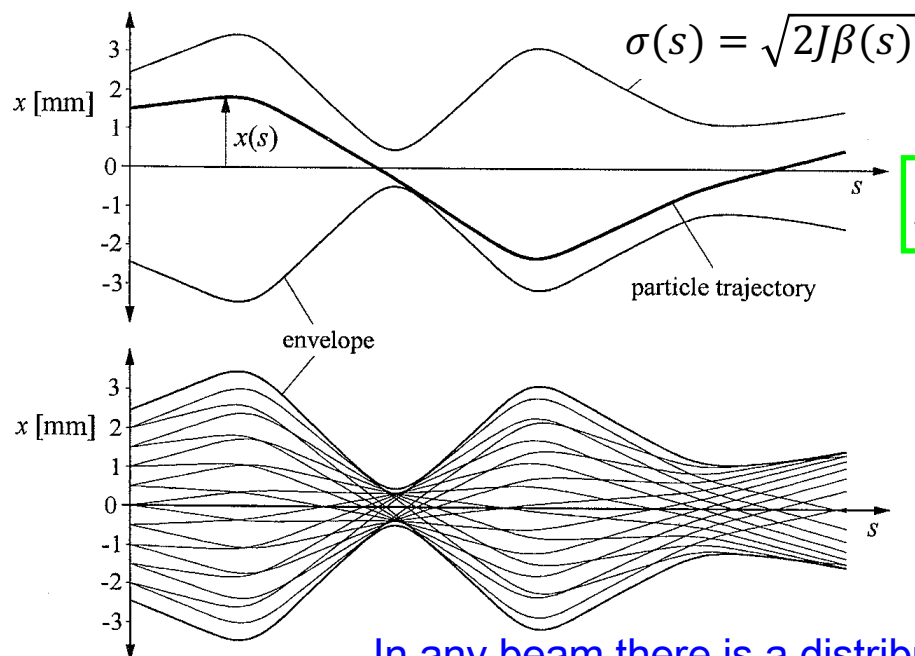
What β is for x , γ is for x'

$$x'_{\max} = \sqrt{2J\gamma} \text{ at } x = -\alpha\sqrt{\frac{2J}{\gamma}}$$

$$\text{Area: } 2\pi J \longrightarrow \int_0^{2\pi} \int_0^J dJ d\phi = 2\pi J = \iint dx dx'$$



The beam envelope



$$x(s) = \sqrt{2J\beta(s)} \sin(\psi(s) + \phi_0)$$

$$\sigma' = -\alpha \sqrt{\frac{2J}{\beta}}$$

$$\sigma'' = -(k\beta - \gamma) \sqrt{\frac{2J}{\beta}} - \alpha^2 \sqrt{\frac{2J}{\beta^3}}$$

$$\sigma'' = -(k\beta^2 - 1) \sqrt{\frac{2J}{\beta^3}}$$

In any beam there is a distribution of initial parameters. If the particles with the largest \$J\$ are distributed in \$\phi\$ over all angles, then the envelope of the beam is described by \$\sigma = \sqrt{2J_{max}\beta(s)}\$.

The initial conditions of \$\beta\$ and \$\alpha\$ are chosen so that this is approximately the case.

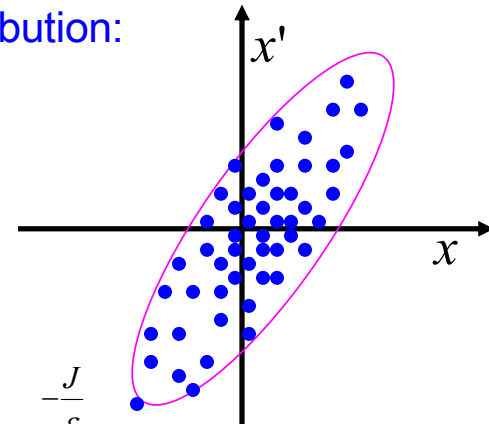
The envelope equation:
$$\sigma'' = -k\sigma + \frac{(2J)^2}{\sigma^3}$$

The phase space distribution

Often one can fit a Gauss distribution to the particle distribution:

$$\rho(x, x') = \frac{1}{2\pi\varepsilon} e^{-\frac{\gamma x^2 + 2\alpha xx' + \beta x'^2}{2\varepsilon}}$$

The equi-density lines are then ellipses. And one chooses the starting conditions for β and α according to these ellipses!



$$\begin{pmatrix} x \\ x' \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin \phi_0 \\ \cos \phi_0 \end{pmatrix}$$

$$\rho(J, \phi_0) = \frac{1}{2\pi\varepsilon} e^{-\frac{J}{\varepsilon}}$$

$$\langle 1 \rangle = \frac{1}{2\pi\varepsilon} \int_0^{2\pi} \int_0^\infty e^{-J/\varepsilon} dJ d\phi_0 = 1 \quad \text{Initial beam distribution} \longrightarrow \text{initial } \alpha, \beta, \gamma$$

$$\langle x^2 \rangle = \frac{1}{2\pi\varepsilon} \iint 2J\beta \sin^2 \phi_0 e^{-J/\varepsilon} dJ d\phi_0 = \varepsilon\beta \longrightarrow \langle x'^2 \rangle = \varepsilon\gamma$$

$$\langle xx' \rangle = -\frac{1}{2\pi\varepsilon} \iint 2J\alpha \sin \phi_0^2 e^{-J/\varepsilon} dJ d\phi_0 = -\varepsilon\alpha$$

$$\varepsilon = \sqrt{\langle x^2 \rangle \langle x'^2 \rangle - \langle xx' \rangle^2} \quad \text{is called the emittance.}$$



Cornell University



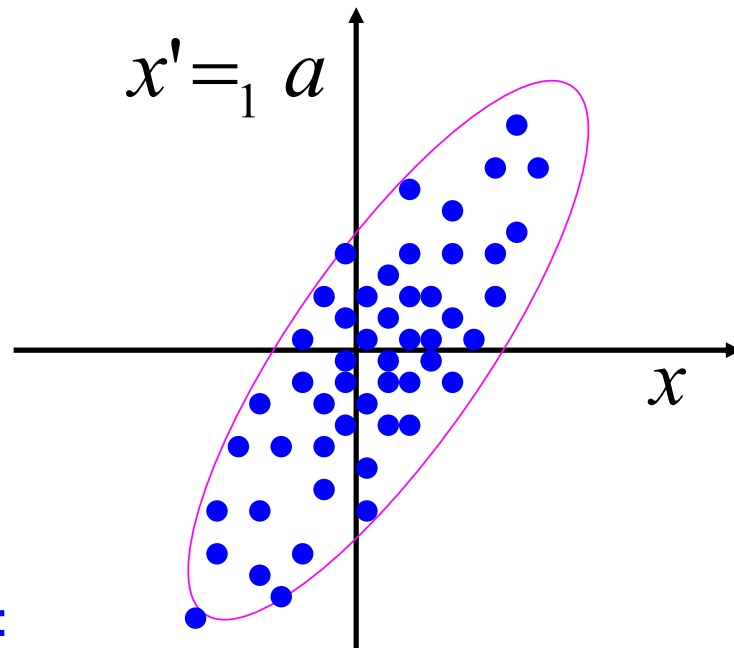
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The normalized emittance



with $a = p_x/p_0$
and the beam's reference
momentum p_0 .

Remarks:

- (1) The phase space area that a beam fills in (x, a) phase space shrinks during acceleration by the factor p_0/p . This area is the emittance ε .
- (2) The phase space area that a beam fills in (x, p_x) phase space is conserved. This area (divided by mc) is the normalized emittance ε_n .

$$\varepsilon = \varepsilon_n / \beta_r \gamma_r$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Invariant of motion

$$x(s) = \sqrt{2J\beta(s)} \sin(\psi(s) + \phi_0)$$

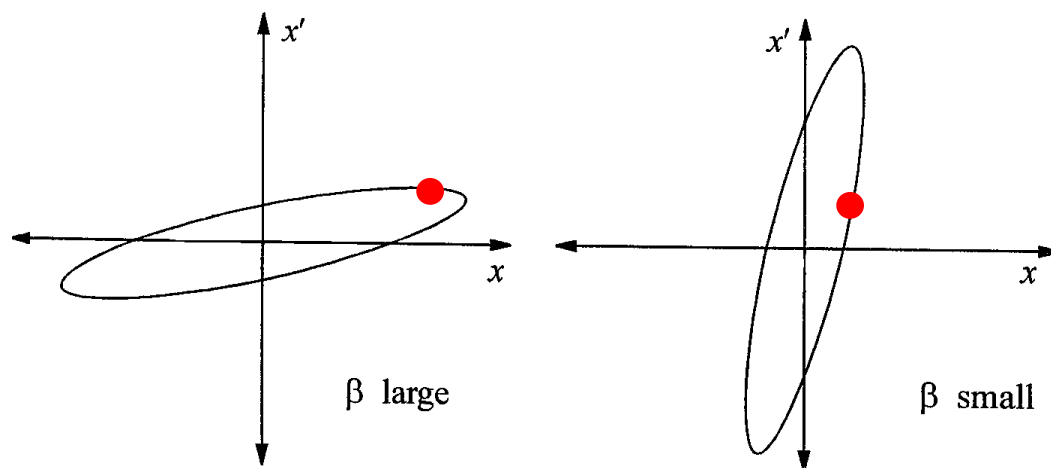
Where J and ϕ are given by the starting conditions x_0 and

$$\gamma x_0^2 + 2\alpha x_0 x'_0 + \beta x_0'^2 = 2J$$

Leads to the invariant of motion:

$$f(x, x', s) = \gamma(s)x^2 + 2\alpha(s)xx' + \beta(s)x'^2 \Rightarrow \frac{d}{ds} f = 0$$

It is called the **Courant-Snyder invariant**.



Twiss differential equation → usually too hard

$$\begin{aligned}\gamma &= \frac{1 + \alpha^2}{\beta} \\ \beta' &= -2\alpha \\ \alpha' &= k\beta - \gamma\end{aligned}$$

$$\beta'' = 2\gamma = 2 \frac{1 + \frac{1}{4}\beta'^2}{\beta} = \frac{d\beta'}{d\beta} \frac{d\beta}{ds}$$

$$\frac{\beta'}{1 + \frac{1}{4}\beta'^2} d\beta' = 2 \frac{d\beta}{\beta}$$

$$\log(1 + \frac{1}{4}\beta'^2) = \log(\beta / \beta_0)$$

$$\beta' = 2\sqrt{\beta / \beta_0 - 1}$$

$$\frac{d\beta}{2\sqrt{\beta / \beta_0 - 1}} = ds$$

$$\beta_0 \sqrt{\beta / \beta_0 - 1} = s - s_0$$

$$\beta(s) = \beta_0 \left(1 + \left(\frac{s - s_0}{\beta_0} \right)^2 \right)$$



Propagation of Twiss parameters

$$(x_0, x'_0) \begin{pmatrix} \gamma_0 & \alpha_0 \\ \alpha_0 & \beta_0 \end{pmatrix} \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix} = 2J$$

$$(x, x') \begin{pmatrix} \gamma & \alpha \\ \alpha & \beta \end{pmatrix} \begin{pmatrix} x \\ x' \end{pmatrix} = 2J = (x_0, x'_0) \underline{M}^T \begin{pmatrix} \gamma & \alpha \\ \alpha & \beta \end{pmatrix} \underline{M} \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix}$$

$$\begin{pmatrix} \gamma & \alpha \\ \alpha & \beta \end{pmatrix} = \underline{M}^{-T} \begin{pmatrix} \gamma_0 & \alpha_0 \\ \alpha_0 & \beta_0 \end{pmatrix} \underline{M}^{-1}$$

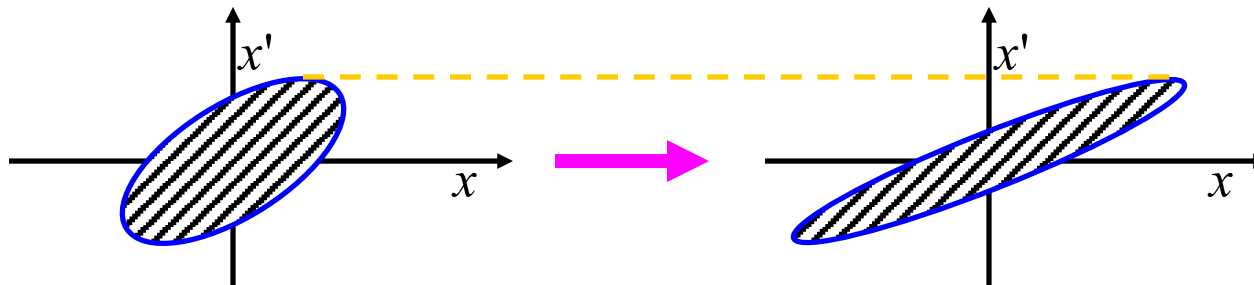
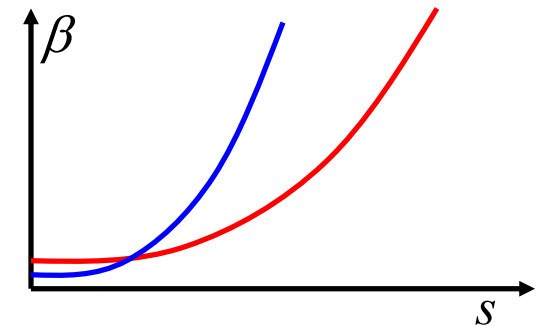
$$\begin{pmatrix} \beta & -\alpha \\ -\alpha & \gamma \end{pmatrix} = \underline{M} \begin{pmatrix} \beta_0 & -\alpha_0 \\ -\alpha_0 & \gamma_0 \end{pmatrix} \underline{M}^T$$



Twiss parameters in a drift

$$\begin{pmatrix} \beta & -\alpha \\ -\alpha & \gamma \end{pmatrix} = \begin{pmatrix} 1 & s \\ 0 & 1 \end{pmatrix} \begin{pmatrix} \beta_0 & -\alpha_0 \\ -\alpha_0 & \gamma_0 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ s & 1 \end{pmatrix} = \begin{pmatrix} \beta_0 - 2\alpha_0 s + \gamma_0 s^2 & \gamma_0 s - \alpha_0 \\ \gamma_0 s - \alpha_0 & \gamma_0 \end{pmatrix}$$

$$\beta = \beta_0^* \left[1 + \left(\frac{s}{\beta_0^*} \right)^2 \right] \quad \text{for} \quad \alpha_0^* = 0$$



Cornell University



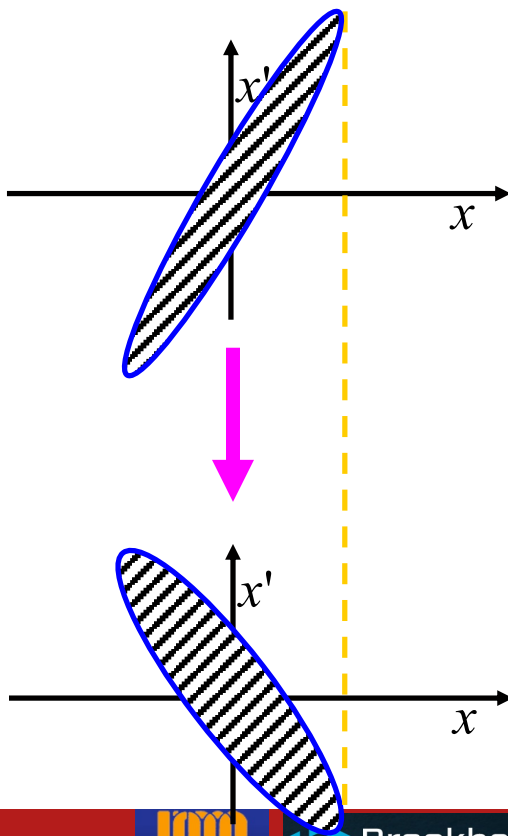
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

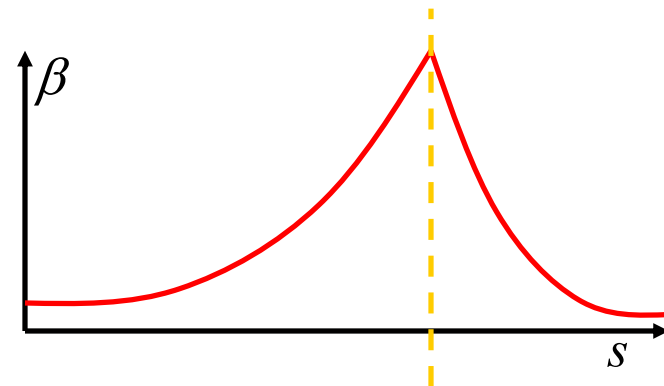
USPAS summer 2023

Twiss parameters in a thin quadrupole



$$\begin{pmatrix} \beta & -\alpha \\ -\alpha & \gamma \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ -k & 1 \end{pmatrix} \begin{pmatrix} \beta_0 & -\alpha_0 \\ -\alpha_0 & \gamma_0 \end{pmatrix} \begin{pmatrix} 1 & -k \\ 0 & 1 \end{pmatrix}$$

$$\alpha = \alpha_0 + k\beta_0$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

From Twiss parameter to Transfer Marix

$$\begin{pmatrix} x_0 \\ x'_0 \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta_0} & 0 \\ -\frac{\alpha_0}{\sqrt{\beta_0}} & \frac{1}{\sqrt{\beta_0}} \end{pmatrix} \begin{pmatrix} \sin(\phi_0) \\ \cos(\phi_0) \end{pmatrix}$$

$$\begin{pmatrix} x \\ x' \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi(s) + \phi_0) \\ \cos(\psi(s) + \phi_0) \end{pmatrix}$$

$$= \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \cos \psi(s) & \sin \psi(s) \\ -\sin \psi(s) & \cos \psi(s) \end{pmatrix} \begin{pmatrix} \sin \phi_0 \\ \cos \phi_0 \end{pmatrix}$$

$$\underline{M}(s) = \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \cos \psi(s) & \sin \psi(s) \\ -\sin \psi(s) & \cos \psi(s) \end{pmatrix} \begin{pmatrix} \frac{1}{\sqrt{\beta_0}} & 0 \\ \frac{\alpha_0}{\sqrt{\beta_0}} & \sqrt{\beta_0} \end{pmatrix}$$

$$= \begin{pmatrix} \sqrt{\frac{\beta}{\beta_0}} [\cos \psi + \alpha_0 \sin \psi] & \sqrt{\beta_0 \beta} \sin \psi \\ \sqrt{\frac{1}{\beta_0 \beta}} [(\alpha_0 - \alpha) \cos \psi - (1 + \alpha_0 \alpha) \sin \psi] & \sqrt{\frac{\beta_0}{\beta}} [\cos \psi - \alpha \sin \psi] \end{pmatrix}$$



Cornell University



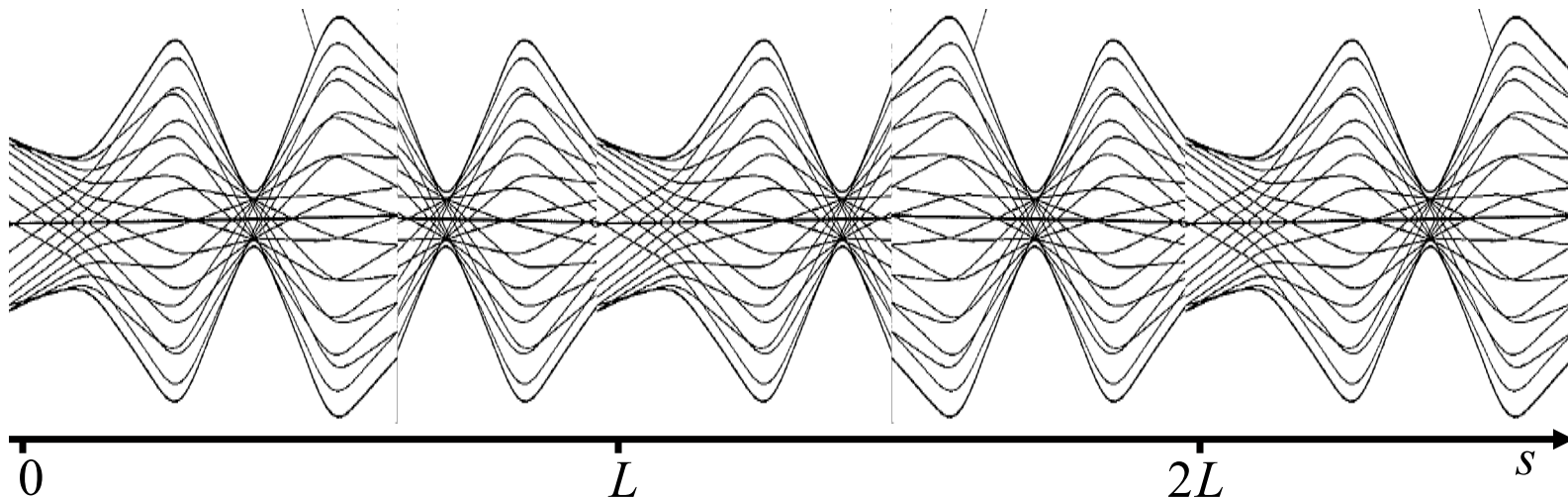
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Periodic solutions in a periodic accelerator



$$\vec{z}(s) = \underline{M}(s,0)\vec{z}(0)$$

$$\vec{z}(L) = \underline{M}(L,0)\vec{z}(0)$$

$$\vec{z}(s+L) = \underline{M}_0(s)\vec{z}(s) \quad , \quad \underline{M}_0 = \underline{M}(s+L,s)$$

$$\vec{z}(s+nL) = \underline{M}_0^n(s)\vec{z}(s)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Periodic beta functions

If the particle distribution in a ring or any other periodic structure is stable, it is periodic from turn to turn.

$$\rho(x, x', s + L) = \rho(x, x', s)$$

To be matched to such a beam, the Twiss parameters α , β , γ must be the same after every turn.

$$\underline{M}(s, 0) = \begin{pmatrix} \sqrt{\frac{\beta}{\beta_0}} [\cos \psi + \alpha_0 \sin \psi] & \sqrt{\beta_0 \beta} \sin \psi \\ \sqrt{\frac{1}{\beta_0 \beta}} [(\alpha_0 - \alpha) \cos \psi - (1 + \alpha_0 \alpha) \sin \psi] & \sqrt{\frac{\beta_0}{\beta}} [\cos \psi - \alpha \sin \psi] \end{pmatrix}$$

$$\underline{M}_p(s) = \begin{pmatrix} \cos \mu + \alpha \sin \mu & \beta \sin \mu \\ -\gamma \sin \mu & \cos \mu - \alpha \sin \mu \end{pmatrix} = \underline{1} \cos \mu + \begin{pmatrix} \alpha & \beta \\ -\gamma & -\alpha \end{pmatrix} \sin \mu$$

$$\mu = \psi(s + L) - \psi(s)$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

One turn matrix to periodic Twiss parameters

The periodic Twiss parameters are the solution of a nonlinear differential equation with periodic boundary conditions:

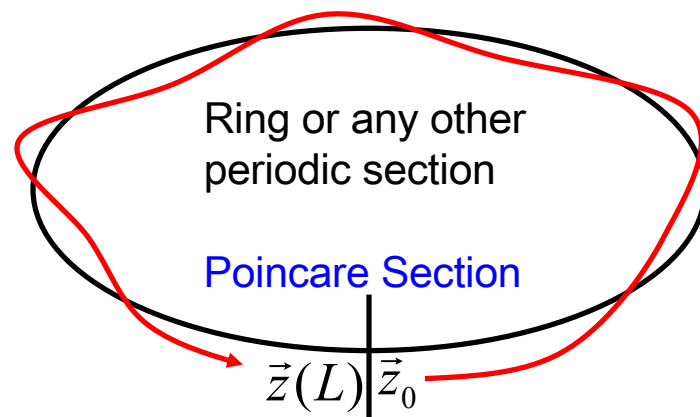
$$\begin{aligned} \beta' &= -2\alpha & \text{with } \beta(L) &= \beta(0) \\ \alpha' &= k\beta - \frac{1+\alpha^2}{\beta} & \text{with } \alpha(L) &= \alpha(0) \end{aligned}$$

$$\mu = \int_0^L \frac{1}{\beta(\hat{s})} d\hat{s}$$

Note: $\beta(s) > 0$

$$\underline{M}_0(s) = \underline{1} \cos \mu + \underline{\beta} \sin \mu ; \underline{\beta} = \begin{pmatrix} \alpha & \beta \\ -\gamma & -\alpha \end{pmatrix}$$

Stable beam motion and thus a periodic beta function can only exist when $|\text{Tr}[\underline{M}]| < 2$.



$$\cos \mu = \frac{1}{2} \text{Tr}[\underline{M}_0(s)]$$

$$\beta = \underline{M}_{0,12} \frac{1}{\sin \mu}$$

$$\alpha = (\underline{M}_{0,11} - \underline{M}_{0,22}) \frac{1}{2 \sin \mu}$$

$$\gamma = \frac{1+\alpha^2}{\beta}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The tune of a particle accelerator

The betatron phase advance per turn divided by 2π is called the **TUNE**.

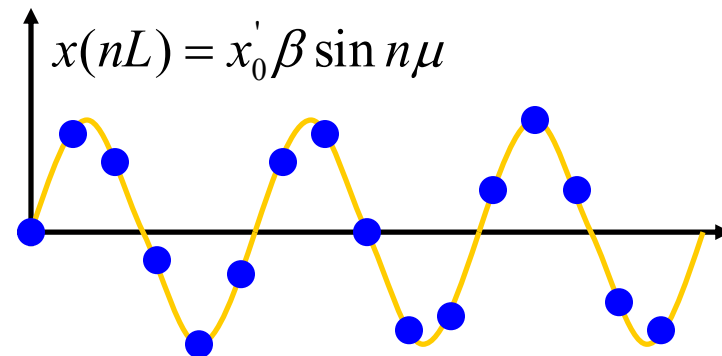
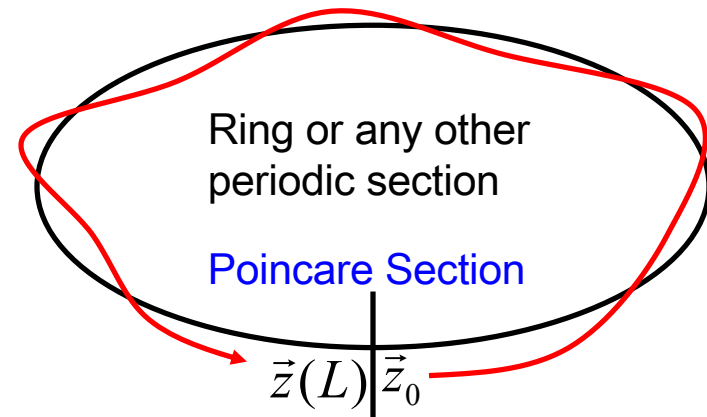
$$\mu = 2\pi\nu = \psi(s+L) - \psi(s)$$

It is a property of the ring and does not depend on the azimuth s .

$$\underline{M}_0(s) = \underline{1} \cos \mu + \begin{pmatrix} \alpha(s) & \beta(s) \\ -\gamma(s) & -\alpha(s) \end{pmatrix} \sin \mu$$

$$\begin{aligned} 2 \cos \underline{\mu}(s) &= \text{Tr}[\underline{M}_0(s)] = \text{Tr}[\underline{M}(s,0)\underline{M}_0(0)\underline{M}^{-1}(s,0)] \\ &= \text{Tr}[\underline{M}_0(0)] = 2 \cos \underline{\mu}(0) \end{aligned}$$

$$\underline{M}_0^n = \underline{1} \cos n\mu + \underline{\beta} \sin n\mu$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The drift

$$\begin{pmatrix} x' \\ a' \\ y' \\ b' \\ \tau' \\ \delta' \end{pmatrix} = \begin{pmatrix} a \\ 0 \\ b \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

Note that in nonlinear expansion $x' \neq a$
so that the drift does not have a linear
transport map even though $x(s) = x_0 + x'_0 s$
is completely linear.

$$\begin{pmatrix} x \\ a \\ y \\ b \\ \tau \\ \delta \end{pmatrix} = \begin{pmatrix} x_0 + s a_0 \\ a \\ y_0 + s b_0 \\ b_0 \\ \tau_0 \\ \delta_0 \end{pmatrix} = \begin{pmatrix} 1 & s & \underline{0} & \underline{0} \\ 0 & 1 & \underline{0} & \underline{0} \\ \underline{0} & 1 & s & \underline{0} \\ \underline{0} & 0 & 1 & \underline{0} \\ \underline{0} & \underline{0} & 1 & 0 \\ \underline{0} & \underline{0} & 0 & 1 \end{pmatrix} \vec{z}_0$$



The quadrupole

$$x'' = -x k$$

$$y'' = y k$$

$$\underline{M}_4 = \begin{pmatrix} \cos(\sqrt{k} s) & \frac{1}{\sqrt{k}} \sin(\sqrt{k} s) & \underline{0} \\ -\sqrt{k} \sin(\sqrt{k} s) & \cos(\sqrt{k} s) & \underline{0} \\ \underline{0} & \underline{0} & \cosh(\sqrt{k} s) & \frac{1}{\sqrt{k}} \sinh(\sqrt{k} s) \\ \underline{0} & \underline{0} & \sqrt{k} \sinh(\sqrt{k} s) & \cosh(\sqrt{k} s) \end{pmatrix}$$

As for a drift, the energy does not change, i.e. $\delta = \delta_0$.

The time of flight only depends on energy, i.e. $\tau = \tau_0 + M \delta$.

For $k < 0$ one has to take into account that

$$\cos(\sqrt{k} s) = \cosh(\sqrt{|k|} s), \quad \sin(\sqrt{k} s) = i \sinh(\sqrt{|k|} s)$$

$$\cosh(\sqrt{k} s) = \cos(\sqrt{|k|} s), \quad \sinh(\sqrt{k} s) = i \sin(\sqrt{|k|} s)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Variation of constants

$$\vec{z}' = \vec{f}(\vec{z}, s)$$

$$\vec{z}' = \underline{L}(s)\vec{z} + \Delta\vec{f}(\vec{z}, s) \quad \text{Field errors, nonlinear fields, etc can lead to } \Delta\vec{f}(\vec{z}, s)$$

$$\vec{z}'_H = \underline{L}(s)\vec{z}_H \Rightarrow \vec{z}_H(s) = \underline{M}(s)\vec{z}_{H0} \quad \text{with} \quad \underline{M}'(s)\vec{a} = \underline{L}(s)\underline{M}(s)\vec{a}$$

$$\vec{z}(s) = \underline{M}(s)\vec{a}(s) \Rightarrow \vec{z}'(s) = \underline{M}'(s)\vec{a} + \underline{M}(s)\vec{a}'(s) = \underline{L}(s)\vec{z} + \Delta\vec{f}(\vec{z}, s)$$

$$\vec{a}(s) = \vec{z}_0 + \int_0^s \underline{M}^{-1}(\hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s}$$

$$\vec{z}(s) = \underline{M}(s) \left\{ \vec{z}_0 + \int_0^s \underline{M}^{-1}(\hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s} \right\}$$

$$= \vec{z}_H(s) + \int_0^s \underline{M}(s - \hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s}$$

Perturbations are propagated
from s to s'



Cornell University



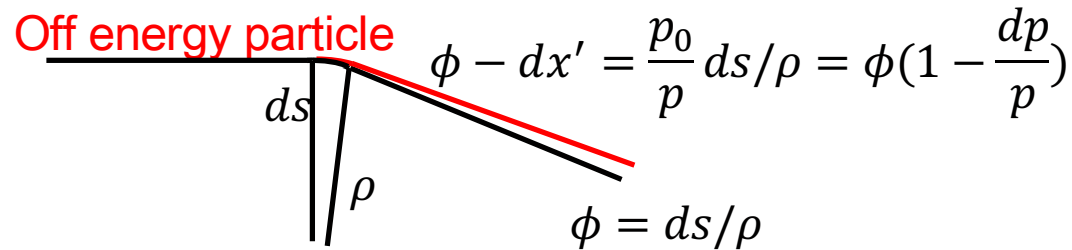
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The dipole equation of motion



$$\rightarrow x'' = \frac{1}{\rho} \frac{dp}{p} = \frac{1}{\rho} \frac{dp}{dE} \frac{E}{p} \delta = \frac{1}{\rho} \frac{1}{\beta^2} \delta$$

$$x'' + x \kappa^2 = \frac{\kappa}{\beta^2} \delta \quad \text{with} \quad \kappa = \frac{1}{\rho} \quad \text{or} \quad \begin{matrix} x' = a \\ a' = -\kappa^2 x + \frac{\kappa}{\beta^2} \delta \end{matrix}$$

$$x_H'' + x_H \kappa^2 = 0 \Rightarrow \begin{pmatrix} x \\ x' \end{pmatrix} = \begin{pmatrix} \cos(\kappa s) & \frac{1}{\kappa} \sin(\kappa s) \\ -\kappa \sin(\kappa s) & \cos(\kappa s) \end{pmatrix} \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix} = \underline{M}(s) \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix}$$

$$\begin{pmatrix} x \\ x' \end{pmatrix} - \underline{M}(s) \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix} = \int_0^s \underline{M}(s - \zeta) \begin{pmatrix} 0 \\ \frac{\kappa}{\beta^2} \delta \end{pmatrix} d\zeta = \int_0^s \begin{pmatrix} \frac{1}{\kappa} \sin(\kappa s) \\ \cos(\kappa s) \end{pmatrix} d\zeta \frac{\kappa}{\beta^2} \delta = \begin{pmatrix} \frac{1}{\kappa} (1 - \cos(\kappa s)) \\ \sin(\kappa s) \end{pmatrix} \frac{1}{\beta^2} \delta$$



Cornell University



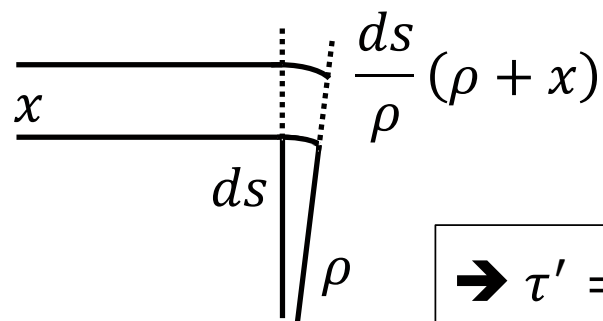
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The dipole transport matrix



$$\Rightarrow \tau' = -\frac{x}{\rho} \frac{1}{v} \frac{E_0}{p_0} = -\frac{x}{\rho} \frac{1}{\beta^2} = -\frac{\kappa}{\beta^2} x$$

$$\underline{M} = \begin{pmatrix} \cos(\kappa s) & \frac{1}{\kappa} \sin(\kappa s) & \underline{0} & 0 & \kappa^{-1}[1 - \cos(\kappa s)] \\ -\kappa \sin(\kappa s) & \cos(\kappa s) & \underline{0} & 0 & \sin(\kappa s) \\ & \underline{0} & 1 & s & \underline{0} \\ & & 0 & 1 & \underline{0} \\ -\sin(\kappa s) & \kappa^{-1}[\cos(\kappa s) - 1] & \underline{0} & 1 & \kappa^{-1}[\sin(\kappa s) - s\kappa] \\ 0 & 0 & \underline{0} & 0 & 1 \end{pmatrix}$$

(for $\beta = 1$)

$$M_{56} = -\kappa \int_0^s M_{16} ds$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The combined function bend

$$\underline{M}_6 = \begin{pmatrix} \underline{M}_x & \underline{0} & \vec{0} \vec{D} \\ \underline{0} & \underline{M}_y & \underline{0} \\ \underline{T} & \underline{0} & \underline{M}_\tau \end{pmatrix}$$

$$\underline{M}_x = \begin{pmatrix} \cos(\sqrt{K} s) & \frac{1}{\sqrt{K}} \sin(\sqrt{K} s) \\ -\sqrt{K} \sin(\sqrt{K} s) & \cos(\sqrt{K} s) \end{pmatrix}$$

$$\underline{M}_y = \begin{pmatrix} \cosh(\sqrt{k} s) & \frac{1}{\sqrt{k}} \sinh(\sqrt{k} s) \\ \sqrt{k} \sinh(\sqrt{k} s) & \cosh(\sqrt{k} s) \end{pmatrix}$$

$$\vec{D} = \begin{pmatrix} \frac{\kappa}{K} [1 - \cos(\sqrt{K} s)] \\ \frac{\kappa}{\sqrt{K}} \sin(\sqrt{K} s) \end{pmatrix}$$

$$\begin{aligned} x'' &= -x (\underbrace{\kappa^2 + k}_K) + \delta \kappa \\ y'' &= y k, \quad \tau' = -\kappa x \end{aligned}$$

Options:

- For $k > 0$:
focusing in x, defocusing in y.
- For $k < 0, K < 0$:
defocusing in x, focusing in y.
- For $k < 0, K > 0$:
weak focusing in both planes.

$$\underline{M}_\tau = \begin{pmatrix} 1 & -\kappa \int_0^s M_{16} ds \\ 0 & 1 \end{pmatrix}$$

\underline{T} from symplecticity



Cornell University



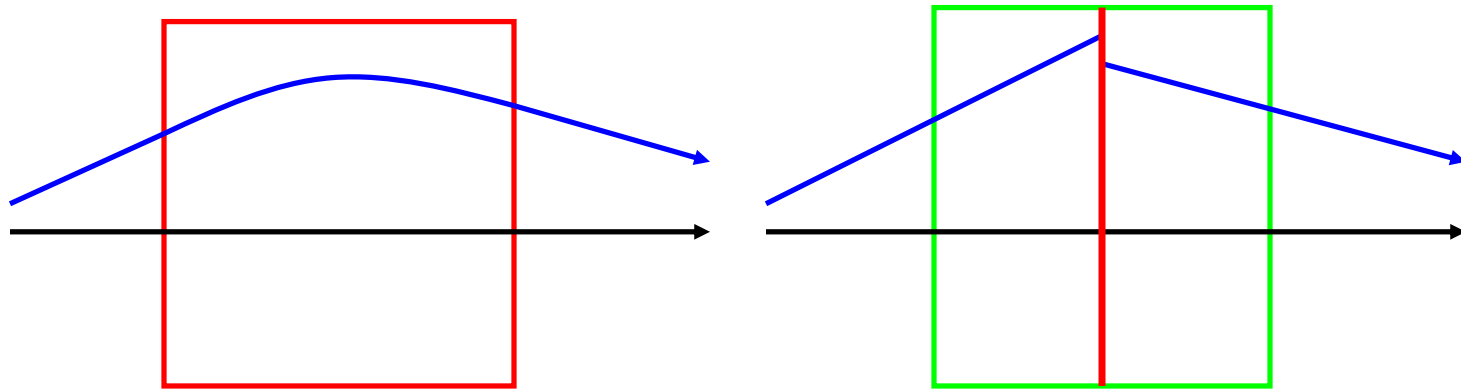
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Thin lens approximation



$$\vec{z}(s) = \underline{M}(s) \vec{z}_0 = \underline{D}(\frac{s}{2}) \underline{D}^{-1}(\frac{s}{2}) \underline{M}(s) \underline{D}^{-1}(\frac{s}{2}) \underline{D}(\frac{s}{2}) \vec{z}_0$$

Drift: $\underline{M}_{\text{drift}}^{\text{thin}}(s) = \underline{D}^{-1}(\frac{s}{2}) \underline{M}(s) \underline{D}^{-1}(\frac{s}{2}) = \underline{1}$



Cornell University



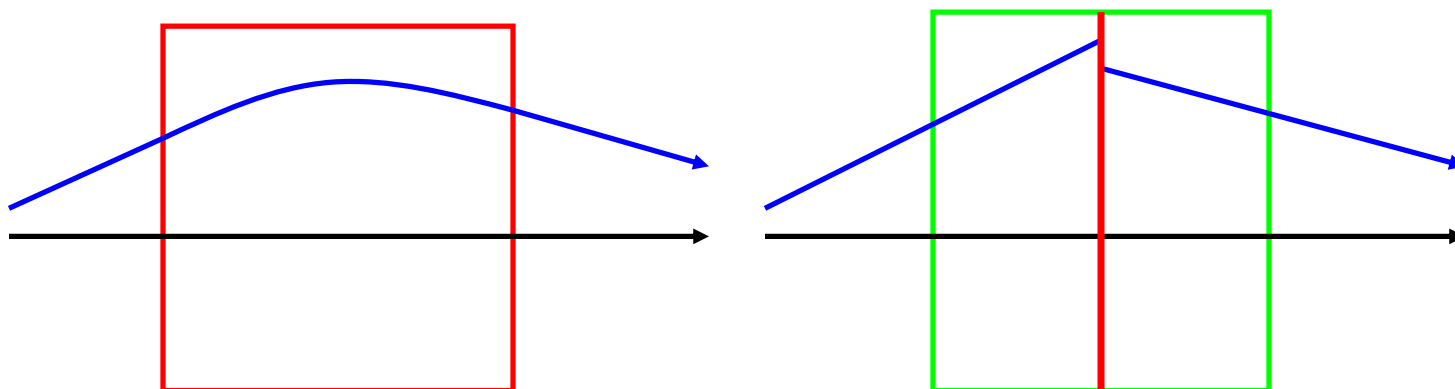
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The thin lens quadrupole



$$\underline{M}_{\text{quad},x}^{\text{thin}}(s) = \begin{pmatrix} 1 & -\frac{s}{2} \\ 0 & 1 \end{pmatrix} \begin{pmatrix} \cos(\sqrt{k}s) & \frac{1}{\sqrt{k}} \sin(\sqrt{k}s) \\ -\sqrt{k} \sin(\sqrt{k}s) & \cos(\sqrt{k}s) \end{pmatrix} \begin{pmatrix} 1 & -\frac{s}{2} \\ 0 & 1 \end{pmatrix}$$

$$\approx \begin{pmatrix} 1 & -\frac{s}{2} \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & s \\ -ks & 1 \end{pmatrix} \begin{pmatrix} 1 & -\frac{s}{2} \\ 0 & 1 \end{pmatrix} \approx \begin{pmatrix} 1 & -\frac{s}{2} \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & \frac{s}{2} \\ -ks & 1 + \frac{ks^2}{2} \end{pmatrix}$$

Weak magnet limit: $\sqrt{k}s \ll 1$

$$\underline{M}_{\text{quad},x}^{\text{thin}}(s) \approx \begin{pmatrix} 1 & 0 \\ -ks & 1 \end{pmatrix}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The thin lens dipole

$$\underline{M} = \begin{pmatrix} \cos(\kappa s) & \frac{1}{\kappa} \sin(\kappa s) & \underline{0} & 0 & \kappa^{-1}[1 - \cos(\kappa s)] \\ -\kappa \sin(\kappa s) & \cos(\kappa s) & \underline{0} & 0 & \sin(\kappa s) \\ & \underline{0} & 1 & s & \underline{0} \\ & & 0 & 1 & \underline{0} \\ -\sin(\kappa s) & \kappa^{-1}[\cos(\kappa s) - 1] & \underline{0} & & 1 & 0 \\ 0 & 0 & & & 0 & 1 \end{pmatrix}$$

Weak magnet limit: $\kappa s \ll 1$

$$\underline{M}_{\text{bend}, x\tau}^{\text{thin}}(s) = \underline{D}\left(-\frac{s}{2}\right) \underline{M}_{\text{bend}, x\tau} \underline{D}\left(-\frac{s}{2}\right) \approx \begin{pmatrix} 1 & 0 & 0 & 0 \\ -\kappa^2 s & 1 & 0 & \kappa s \\ -\kappa s & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The thin lens combined function bend

$$\underline{M}_6 = \begin{pmatrix} \underline{M}_x & \underline{0} & \vec{0} \vec{D} \\ \underline{0} & \underline{M}_y & \underline{0} \\ \underline{T} & \underline{0} & \underline{1} \end{pmatrix}$$

Weak magnet limit: $\kappa s \ll 1$

$$\underline{M}_x = \begin{pmatrix} \cos(\sqrt{K} s) & \frac{1}{\sqrt{K}} \sin(\sqrt{K} s) \\ -\sqrt{K} \sin(\sqrt{K} s) & \cos(\sqrt{K} s) \end{pmatrix}$$

$$\underline{M}_x^{\text{thin}} = \begin{pmatrix} 1 & 0 \\ -K s & 1 \end{pmatrix}$$

$$\underline{M}_y = \begin{pmatrix} \cosh(\sqrt{k} s) & \frac{1}{\sqrt{k}} \sinh(\sqrt{k} s) \\ \sqrt{k} \sinh(\sqrt{k} s) & \cosh(\sqrt{k} s) \end{pmatrix}$$

$$\underline{M}_y^{\text{thin}} = \begin{pmatrix} 1 & 0 \\ k s & 1 \end{pmatrix}$$

$$\vec{D} = \begin{pmatrix} \frac{\kappa}{K} [1 - \cos(\sqrt{K} s)] \\ \frac{\kappa}{\sqrt{K}} \sin(\sqrt{K} s) \end{pmatrix}$$

$$\vec{D} = \begin{pmatrix} 0 \\ \kappa s \end{pmatrix}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Edge focusing

Horizontal focusing with $\Delta x' = -x \frac{\tan(\varepsilon)}{\rho}$

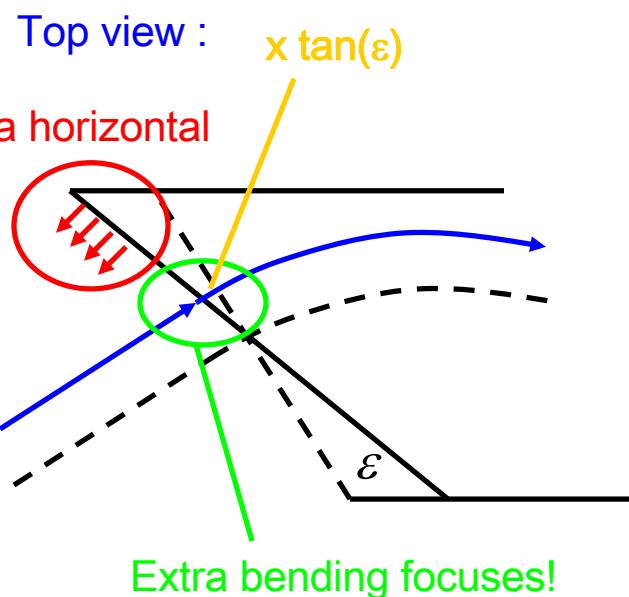
$$B_x = \partial_y B_s \Big|_{y=0} y \tan(\varepsilon) = \partial_s B_y \Big|_{y=0} y \tan(\varepsilon)$$

$$y'' = \frac{q}{p} \partial_s B_y \Big|_{y=0} y \tan(\varepsilon)$$

$$\Delta y' = \int y'' ds = \frac{q}{p} B_y y \tan(\varepsilon) = y \frac{\tan(\varepsilon)}{\rho}$$

Quadrupole effect with

$$kl = \frac{\tan(\varepsilon)}{\rho}$$



$$\vec{Z} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ -\frac{\tan(\varepsilon)}{\rho} & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & \frac{\tan(\varepsilon)}{\rho} & 1 \end{pmatrix} \vec{Z}_0$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Orbit distortions for a one-pass accelerator

$$x' = a$$

$$a' = -(\kappa^2 + k)x + \Delta f$$

The extra force can for example come from an erroneous dipole field or from a correction coil: $\Delta f = \frac{q}{p} \Delta B_y = \Delta \kappa$

Variation of constants: $\vec{z} = \underline{M}\vec{z}_0 + \Delta\vec{z}$ with $\Delta\vec{z} = \int_0^s \underline{M}(s - \hat{s}) \begin{pmatrix} 0 \\ \Delta\kappa(\hat{s}) \end{pmatrix} d\hat{s}$

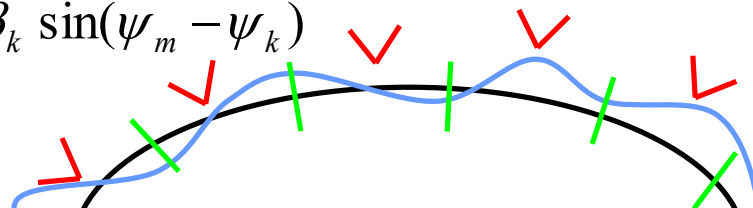
$$\Delta x(s) = \sum_k \Delta \mathcal{G}_k \sqrt{\beta(s)\beta_k} \sin(\psi(s) - \psi_k)$$



Orbit correction for a one-pass accelerator

When the closed orbit $x_{\text{co}}^{\text{old}}(s_m)$ is measured at beam position monitors (BPMs, index m) and is influenced by corrector magnets (index k), then the monitor readings before and after changing the kick angles created in the correctors by $\Delta\mathcal{G}_k$ are related by

$$\begin{aligned} x_{\text{co}}^{\text{new}}(s_m) &= x_{\text{co}}^{\text{old}}(s_m) + \sum_k \Delta\mathcal{G}_k \sqrt{\beta_m \beta_k} \sin(\psi_m - \psi_k) \\ &= x_{\text{co}}^{\text{old}}(s_m) + \sum_k O_{mk} \Delta\mathcal{G}_k \end{aligned}$$



$$\vec{x}_{\text{co}}^{\text{new}} = \vec{x}_{\text{co}}^{\text{old}} + \underline{O} \Delta\vec{\mathcal{G}}$$

$$\Delta\vec{\mathcal{G}} = -\underline{O}^{-1} \vec{x}_{\text{co}}^{\text{old}} \Rightarrow \vec{x}_{\text{co}}^{\text{new}} = 0$$

It is often better not to try to correct the closed orbit at the the BPMs to zero in this way since

1. computation of the inverse can be numerically unstable, so that small errors in the old closed orbit measurement lead to a large error in the corrector coil settings.
2. A zero orbit at all BPMs can be a bad orbit in-between BPMs



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

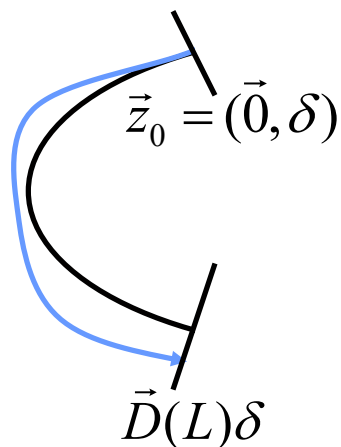
Dispersion of one-pass accelerators

$$x' = a$$

$$a' = -(\kappa^2 + k)x + \kappa\delta$$

$$\vec{z} = \underline{M}\vec{z}_0 + \int_0^s \underline{M}(s-\hat{s}) \begin{pmatrix} 0 \\ \delta\kappa(\hat{s}) \end{pmatrix} d\hat{s}$$

$$\Rightarrow \vec{D}(s) = \int_0^s \underline{M}(s-\hat{s}) \begin{pmatrix} 0 \\ \kappa(\hat{s}) \end{pmatrix} ds'$$



$$\Delta\kappa = \delta\kappa$$

$$D(s) = \sqrt{\beta(s)} \int_0^s \kappa(\hat{s}) \sqrt{\beta(\hat{s})} \sin(\psi(s) - \psi(\hat{s})) d\hat{s}$$

Alternatively, one can multiply the 6x6 matrices and take $D(s) = M_{16}(s)$



Cornell University



Brookhaven
National Laboratory

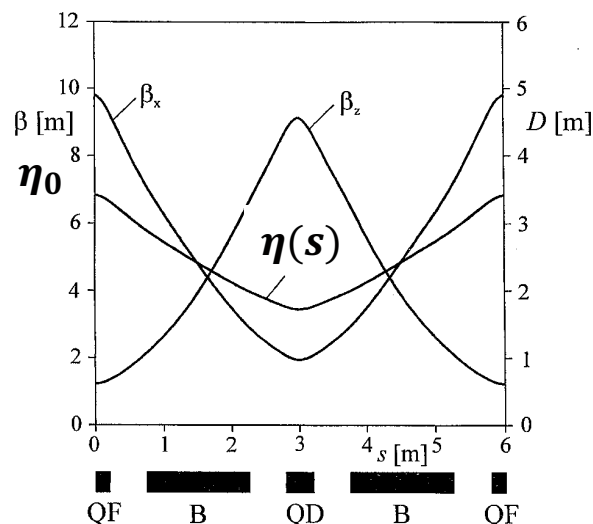
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Fodo Cells and periodic dispersion

Alternating gradients allow focusing in both transverse planes. Therefore, focusing and defocusing quadrupoles are usually alternated and interleaved with bending magnets.



The dispersion that starts with 0 is called $D(s)$, the dispersion that is periodic in a section is called $\eta(s)$.

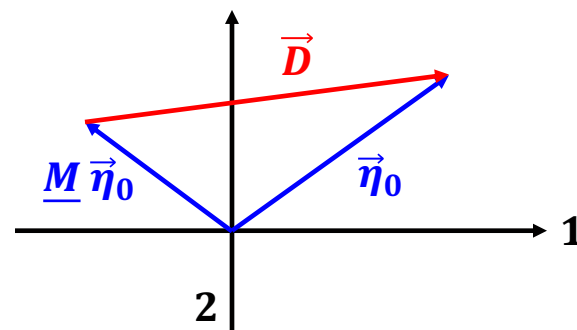
$$\underline{M}_0 = \underline{M}_{FoDo}^N$$

The periodic beta function and dispersion for each FODO is also periodic for an accelerator section that consists of many FODO cells. Often large sections of an accelerator consist of FODOs.

$$\vec{\eta}(s) = \begin{pmatrix} \eta(s) \\ \eta'(s) \end{pmatrix}, \vec{D}(s) = \begin{pmatrix} D(s) \\ D'(s) \end{pmatrix}$$

$$\vec{\eta}(s) = \underline{M}(s) \vec{\eta}_0 + \vec{D}(s)$$

After $\vec{\eta}_0 = \underline{M}(s) \vec{\eta}_0 + \vec{D} \Rightarrow \vec{\eta}_0 = (\underline{1} - \underline{M})^{-1} \vec{D}$
the FoDo:



Cornell University



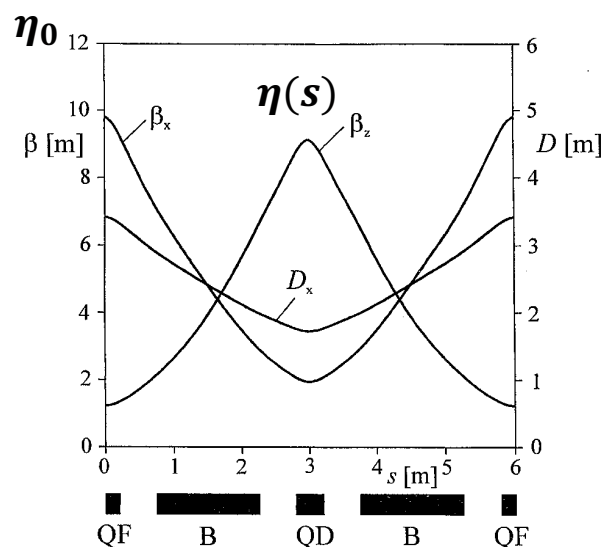
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

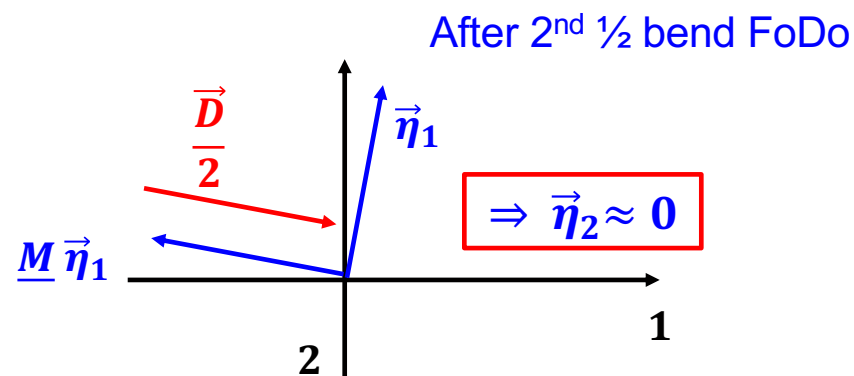
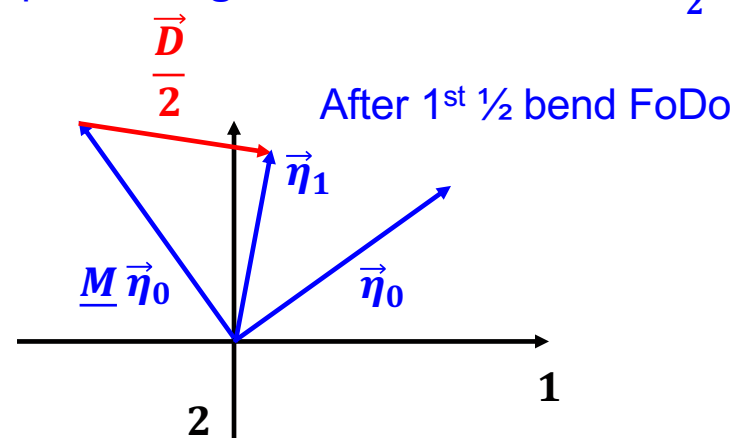
USPAS summer 2023

Dispersion suppression by missing bends

After an arc of periodic FoDo cells one would often like to suppress the dispersion to 0 while not changing the betas much from the periodic cells. This can be done by having one FoDo with bends reduced by a factor α followed by one reduced by $(1-\alpha)$.



Example: **90 degrees** FoDo cell and $\alpha = \frac{1}{2}$:



Cornell University



Brookhaven
National Laboratory

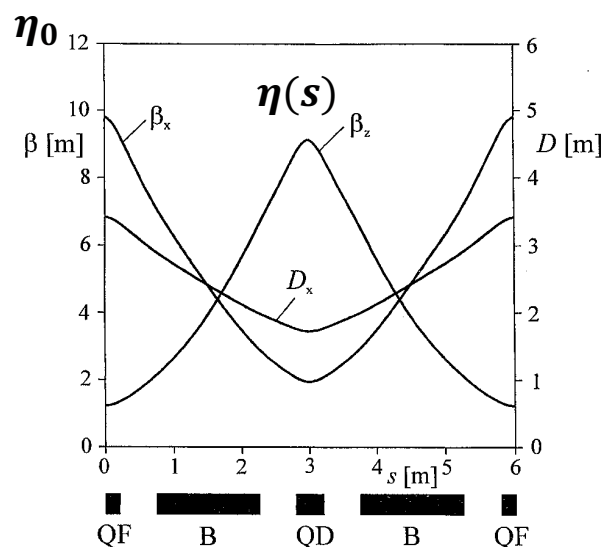
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

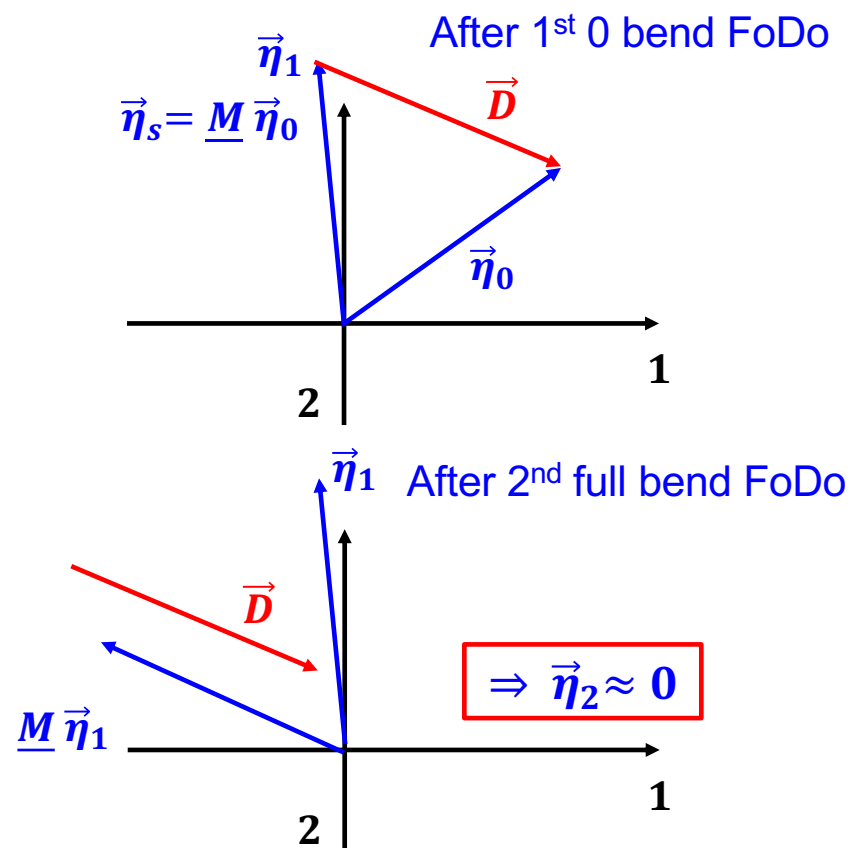
USPAS summer 2023

Dispersion suppression by missing bends

After an arc of periodic FoDo cells one would often like to suppress the dispersion to 0 while not changing the betas much from the periodic cells. This can be done by having one FoDo with bends reduced by a factor α followed by one reduced by $(1-\alpha)$.



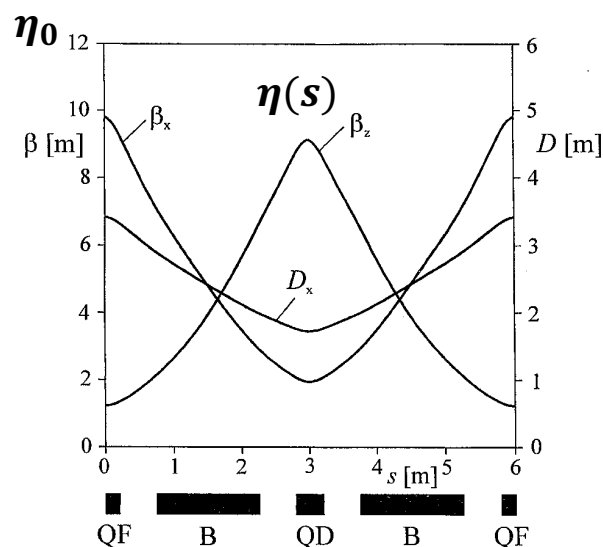
Example: 60 degrees FoDo cell and $\alpha = 0$:



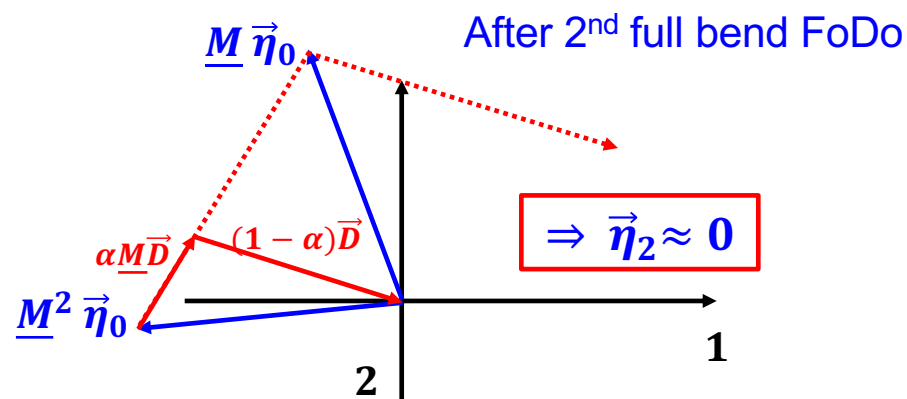
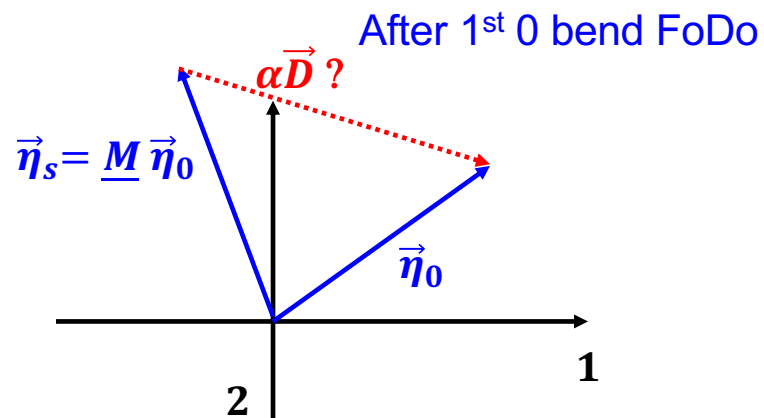
For every FoDo phase advance there is an α to make η 0.

Dispersion suppression by missing bends

After an arc of periodic FoDo cells one would often like to suppress the dispersion to 0 while not changing the betas much from the periodic cells. This can be done by having one FoDo with bends reduced by a factor α followed by one reduced by $(1-\alpha)$.



For any other FoDo phase advance: is there an α ?



For every FoDo phase advance there is an α to make η 0.

Closed orbit in periodic accelerators

$$x' = a$$

$$a' = -(\kappa^2 + k)x + \Delta f$$

The extra force can for example come from an erroneous dipole field or from a correction coil: $\Delta f = \frac{q}{p} \Delta B_y = \Delta \kappa$

Variation of constants: $\vec{z} = \underline{M} \vec{z}_0 + \Delta \vec{z}$ with $\Delta \vec{z} = \int_0^s \underline{M}(s - \hat{s}) \begin{pmatrix} 0 \\ \Delta \kappa(\hat{s}) \end{pmatrix} d\hat{s}$

For the periodic or closed orbit: $\vec{z}_{\text{co}} = \underline{M}_0 \vec{z}_{\text{co}} + \underline{M}_0 \int_0^L \underline{M}^{-1}(\hat{s}) \begin{pmatrix} 0 \\ \Delta \kappa(\hat{s}) \end{pmatrix} d\hat{s}$

$$\begin{aligned} \vec{z}_{\text{co}} &= [\underline{M}_0^{-1} - \underline{1}]^{-1} \int_0^L \underline{M}^{-1}(\hat{s}) \begin{pmatrix} 0 \\ \Delta \kappa(\hat{s}) \end{pmatrix} d\hat{s} \\ &= \frac{(\cos \mu - 1) \underline{1} + \sin \mu \underline{\beta}}{(\cos \mu - 1)^2 + \sin^2 \mu} \int_0^L \begin{pmatrix} -\sqrt{\beta \hat{\beta}} \sin \hat{\psi} \\ \sqrt{\frac{\hat{\beta}}{\beta}} [\cos \hat{\psi} + \alpha \sin \hat{\psi}] \end{pmatrix} \Delta \kappa(\hat{s}) d\hat{s} \end{aligned}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Closed orbit in periodic accelerators

$$\vec{z}_{co}(L) = \frac{\cos \mu \underline{1} + \sin \mu \underline{\beta} - \underline{1}}{2 - 2 \cos \mu} \int_0^L \begin{pmatrix} \sqrt{\beta \hat{\beta}} \sin(-\hat{\psi}) \\ \sqrt{\frac{\hat{\beta}}{\beta}} [\cos(-\hat{\psi}) - \alpha \sin(-\hat{\psi})] \end{pmatrix} \Delta \kappa(\hat{s}) d\hat{s}$$

$$\begin{aligned} x_{co}(L) &= \frac{1}{4 \sin^2 \frac{\mu}{2}} \int_0^L \sqrt{\beta \hat{\beta}} [\sin(\mu - \hat{\psi}) + \sin \hat{\psi}] \Delta \kappa(\hat{s}) d\hat{s} \\ &= \frac{1}{2 \sin \frac{\mu}{2}} \int_0^L \sqrt{\beta \hat{\beta}} \cos(\hat{\psi} - \frac{\mu}{2}) \Delta \kappa(\hat{s}) d\hat{s} \end{aligned}$$

$\sin\left(\left(\psi - \frac{\mu}{2}\right) + \frac{\mu}{2}\right) - \sin\left(\left(\psi - \frac{\mu}{2}\right) - \frac{\mu}{2}\right)$

$$\begin{aligned} x_{co}(s) &= \frac{1}{2 \sin \frac{\mu}{2}} \left[\int_s^L \sqrt{\beta \hat{\beta}} \cos(\hat{\psi} - \psi - \frac{\mu}{2}) \Delta \hat{\kappa} d\hat{s} + \int_0^s \sqrt{\beta \hat{\beta}} \cos(\hat{\psi} - \psi + \frac{\mu}{2}) \Delta \hat{\kappa} d\hat{s} \right] \\ &= \frac{\sqrt{\beta}}{2 \sin \frac{\mu}{2}} \int_0^L \sqrt{\hat{\beta}} \cos(|\hat{\psi} - \psi| - \frac{\mu}{2}) \Delta \hat{\kappa} d\hat{s} \end{aligned}$$



Cornell University

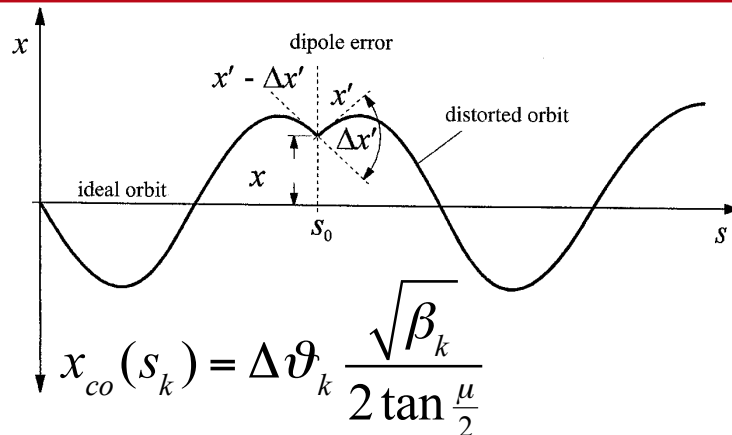


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Closed orbit for one kick



$$x_{co}(s) = \Delta\vartheta_k \frac{\sqrt{\beta\beta_k}}{2\sin\frac{\mu}{2}} \cos\left(\left|\psi - \psi_k\right| - \frac{\mu}{2}\right)$$

Free betatron oscillation

$$x'_{co}(s_k) - x'_{co}(s_k + L) = \Delta\vartheta_k \frac{-\sin(-\frac{\mu}{2}) + \sin(\frac{\mu}{2})}{2\sin\frac{\mu}{2}} = \Delta\vartheta_k$$

$$\longrightarrow x_{max} = \sqrt{2J\beta} \longrightarrow$$

$$x_{co}(s) = \sqrt{2J\beta} \sin(\psi + \varphi_0), \quad J = \frac{\Delta\vartheta_k^2 \beta_k}{8\sin^2\frac{\mu}{2}}$$

The oscillation amplitude J diverges when the tune ν is close to an integer.

$$s < s_k : \varphi_0 = \frac{\pi}{2} - \psi_k + \frac{\mu}{2} \quad , \quad s > s_k : \varphi_0 = \frac{\pi}{2} - \psi_k - \frac{\mu}{2} \quad \text{Phase jump by } \mu$$



Cornell University



Brookhaven
National Laboratory

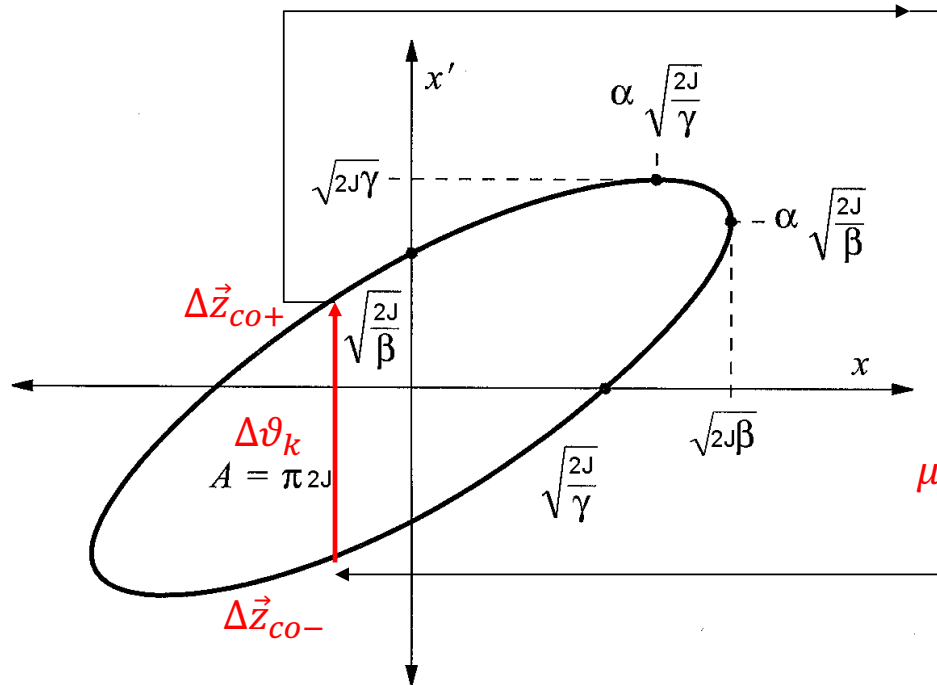
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Closed orbit for one kick

Free betatron oscillation



$$x_n(s) = \frac{\Delta\vartheta_k \sqrt{\beta\hat{\beta}}}{2\sin(\frac{\mu}{2})} \cos(\frac{\mu}{2} - |\psi - \psi_k|)$$

$$\begin{pmatrix} x_n \\ a_n \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} x \\ x' \end{pmatrix}$$

$$\begin{pmatrix} x_{n\pm} \\ a_{n\pm} \end{pmatrix} = \frac{\Delta\vartheta_k \sqrt{\hat{\beta}}}{\sin(\frac{\mu}{2})} \begin{pmatrix} \cos(\frac{\mu}{2}) \\ \pm \sin(\frac{\mu}{2}) \end{pmatrix}$$

$$\begin{pmatrix} x_{n\pm} \\ a_{n\pm} \end{pmatrix} = \frac{\Delta\vartheta_k \sqrt{\hat{\beta}}}{\sin(\frac{\mu}{2})} \left(\frac{\mu}{2} - |\psi - \psi_k| \right)$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Closed orbit correction in periodic accelerators

When the closed orbit $x_{\text{co}}^{\text{old}}(s_m)$ is measured at beam position monitors (BPMs, index m) and is influenced by corrector magnets (index k), then the monitor readings before and after changing the kick angles created in the correctors by $\Delta\mathcal{G}_k$ are related by

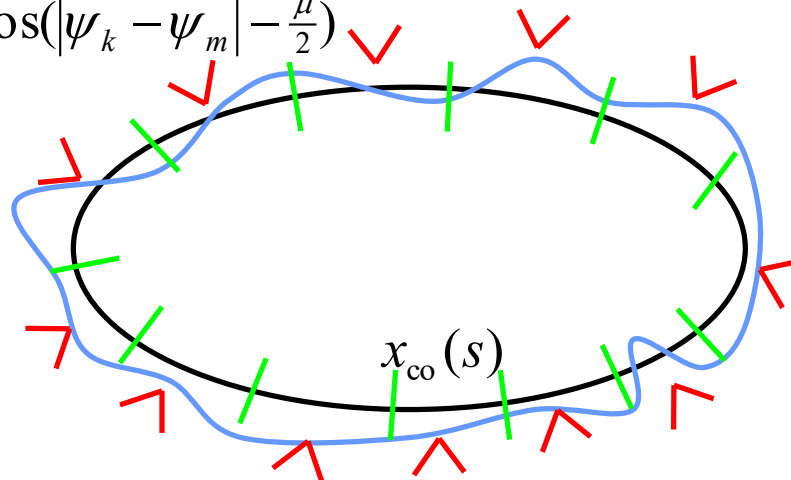
$$\begin{aligned} x_{\text{co}}^{\text{new}}(s_m) &= x_{\text{co}}^{\text{old}}(s_m) + \sum_k \Delta\mathcal{G}_k \frac{\sqrt{\beta_m\beta_k}}{2\sin\frac{\mu}{2}} \cos(|\psi_k - \psi_m| - \frac{\mu}{2}) \\ &= x_{\text{co}}^{\text{old}}(s_m) + \sum_k O_{mk} \Delta\mathcal{G}_k \end{aligned}$$

$$\vec{x}_{\text{co}}^{\text{new}} = \vec{x}_{\text{co}}^{\text{old}} + \underline{O} \Delta\vec{\mathcal{G}}$$

$$\Delta\vec{\mathcal{G}} = -\underline{O}^{-1} \vec{x}_{\text{co}}^{\text{old}} \Rightarrow \vec{x}_{\text{co}}^{\text{new}} = 0$$

It is often better not to try to correct the closed orbit at the the BPMs to zero in this way since

1. computation of the inverse can be numerically unstable, so that small errors in the old closed orbit measurement lead to a large error in the corrector coil settings.
2. A zero orbit at all BPMs can be a bad orbit inbetween BPMs



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Periodic dispersion

$$\begin{pmatrix} \underline{M}_{0x} \vec{z}_0 + \vec{D}(L) \delta \\ M_{56} \delta \\ \delta \end{pmatrix} = \begin{pmatrix} \underline{M}_{0x} & \vec{0} & \vec{D}(L) \\ \vec{T}^T & 1 & M_{56} \\ \vec{0}^T & 0 & 1 \end{pmatrix} \begin{pmatrix} \vec{z}_0 \\ 0 \\ \delta \end{pmatrix}$$

The periodic orbit for particles with relative energy deviation δ is

$$\vec{\eta}(0) = \underline{M}_0 \vec{\eta}(0) + \vec{D}(L) \quad \vec{\eta}(L) = \underline{M}_0 \vec{\eta}(0) + \vec{D}(L) \quad \text{with} \quad \vec{\eta}(s) = \vec{\eta}(0)$$

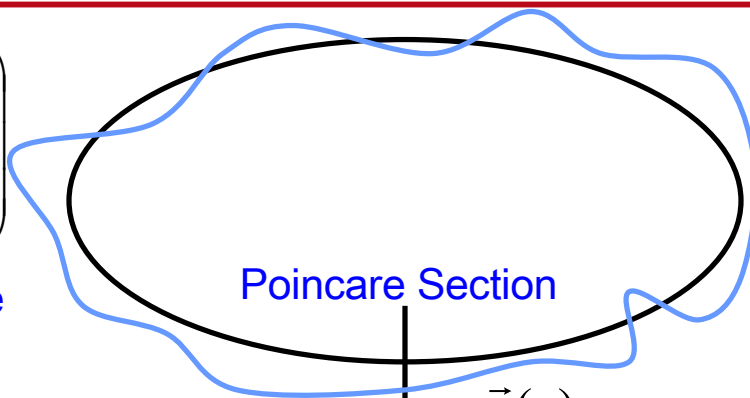
\Downarrow

$$\vec{\eta}(0) = [\underline{1} - \underline{M}_0(0)]^{-1} \vec{D}(L)$$

Particles with energy deviation δ oscillates around this periodic orbit.

$$\vec{z} = \vec{z}_\beta + \delta \vec{\eta}$$

$$\begin{aligned} \underline{\vec{z}}_\beta(L) + \delta \vec{\eta}(L) &= \vec{z}(L) = \underline{M}_0 \vec{z}(0) + \vec{D}(L) \delta = \underline{M}_0 [\vec{z}_\beta(0) + \delta \vec{\eta}(0)] + \vec{D}(L) \delta \\ &= \underline{M}_0 \underline{\vec{z}}_\beta(0) + \delta \vec{\eta}(L) \end{aligned}$$



Poincare Section



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

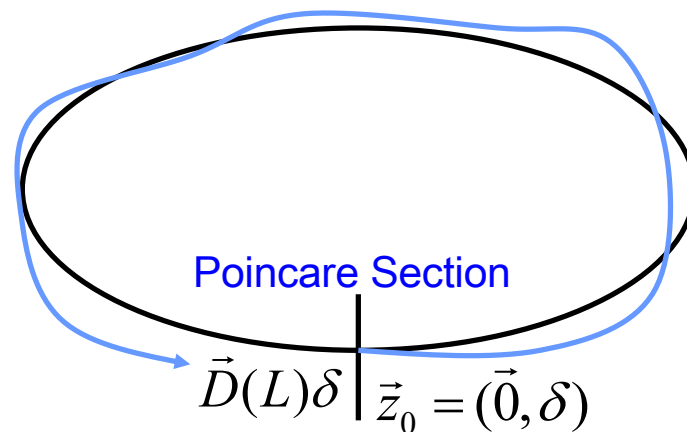
Periodic dispersion integral

$$x' = a$$

$$a' = -(\kappa^2 + k)x + \kappa\delta$$

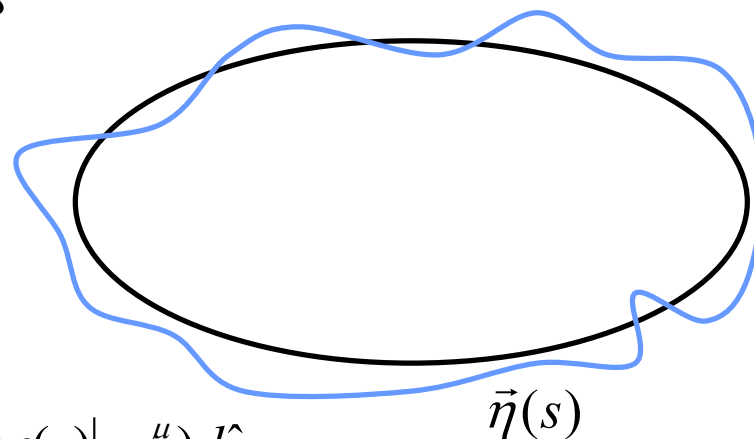
$$\vec{z} = \underline{M}\vec{z}_0 + \int_0^s \underline{M}(s - \hat{s}) \begin{pmatrix} 0 \\ \delta\kappa(\hat{s}) \end{pmatrix} d\hat{s}$$

$$\Rightarrow \vec{D}(L) = \int_0^L \underline{M}(L - \hat{s}) \begin{pmatrix} 0 \\ \kappa(\hat{s}) \end{pmatrix} ds'$$



$$\Delta\kappa = \delta\kappa$$

$$\eta(s) = \frac{\sqrt{\beta(s)}}{2\sin\frac{\mu}{2}} \oint \kappa(\hat{s}) \sqrt{\beta(\hat{s})} \cos(|\psi(\hat{s}) - \psi(s)| - \frac{\mu}{2}) d\hat{s}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Quadrupole errors in one-pass accelerators

$$\vec{z}' = \underline{L}(s) \vec{z} + \Delta \vec{f}(\vec{z}, s)$$

$$\vec{z}(s) = \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}(\vec{z}, \hat{s}) d\hat{s} \approx \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}(\vec{z}_H, \hat{s}) d\hat{s}$$

$$x'' = -(\kappa^2 + k)x - \Delta k(s)x \Rightarrow \begin{pmatrix} x' \\ a' \end{pmatrix} = \begin{pmatrix} a \\ -(\kappa^2 + k)x \end{pmatrix} - \begin{pmatrix} 0 & 0 \\ \Delta k(s) & 0 \end{pmatrix} \begin{pmatrix} x \\ a \end{pmatrix}$$

$$\vec{z}(s) = \left\{ \underline{M}(s) - \int_0^s \underline{M}(s, \hat{s}) \begin{pmatrix} 0 & 0 \\ \Delta k(\hat{s}) & 0 \end{pmatrix} \underline{M}(\hat{s}, 0) d\hat{s} \right\} \vec{z}_0$$

One quadrupole error:

$$\Delta \underline{M}(s, \hat{s}) = -\underline{M}(s, \hat{s}) \begin{pmatrix} 0 & 0 \\ \Delta kl(\hat{s}) & 0 \end{pmatrix}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Qadrupole errors and Twiss in one pass accelerators

$$\Delta \underline{M}(s, \hat{s}) = -\underline{M}(s, \hat{s}) \begin{pmatrix} 0 & 0 \\ \Delta kl(\hat{s}) & 0 \end{pmatrix} \quad \underline{M}(s) = \begin{pmatrix} \dots & \sqrt{\beta_0} \beta \sin \tilde{\psi} \\ \dots & \sqrt{\frac{\beta_0}{\beta}} [\cos \tilde{\psi} - \alpha \sin \tilde{\psi}] \end{pmatrix}$$

$$\Delta \underline{M}(s, \hat{s}) = -\Delta kl(\hat{s}) \begin{pmatrix} \sqrt{\hat{\beta}} \beta \sin \psi & 0 \\ \sqrt{\frac{\hat{\beta}}{\beta}} [\cos \psi - \alpha \sin \psi] & 0 \end{pmatrix}, \quad \psi = \psi(s) - \psi(\hat{s})$$

$$= \begin{pmatrix} \frac{\frac{1}{2} \Delta \beta [\cos \psi + \hat{\alpha} \sin \psi] + \Delta \psi \beta [\hat{\alpha} \cos \psi - \sin \psi]}{\sqrt{\hat{\beta}} \beta} & \sqrt{\hat{\beta}} \left(\frac{\frac{\Delta \beta}{2} \sin \psi + \Delta \psi \beta \cos \psi}{\sqrt{\beta}} \right) \\ \dots & \dots \end{pmatrix}$$

\downarrow (from top-left element) \downarrow (from top-right element)

$$\frac{1}{2} \Delta \beta \cos \psi + \frac{1}{2} \Delta \beta \frac{\sin^2 \psi}{\cos \psi} = \frac{1}{2} \Delta \beta \frac{1}{\cos \psi} = -\Delta kl(\hat{s}) \beta \hat{\beta} \sin \psi \quad \Delta \psi = -\frac{\Delta \beta}{2\beta} \tan \psi$$

$$\Delta \beta = -\Delta kl(\hat{s}) \beta \hat{\beta} \sin 2\psi$$

$$\Delta \psi = \Delta kl(\hat{s}) \hat{\beta} \frac{1}{2} (1 - \cos 2\psi)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Twiss changes in one-pass accelerators

$$\Delta\psi = \Delta kl_j \beta_j \sin^2(\psi - \psi_j)$$

→ More focusing always increases the tune

$$\frac{\Delta\beta}{\beta} = -\Delta kl_j \beta_j \sin(2[\psi - \psi_j])$$

→ Beta beat oscillates twice as fast as orbit.

Notice the self consistency: $\Delta\psi = \int_0^s \left(\frac{1}{\hat{\beta} + \Delta\hat{\beta}} - \frac{1}{\hat{\beta}} \right) d\hat{s} = - \int_0^s \frac{\Delta\hat{\beta}}{\hat{\beta}^2} d\hat{s} = - \int_0^s \frac{\Delta\hat{\beta}}{\hat{\beta}} d\hat{\psi}$

$$\Delta\alpha = -\frac{\Delta\beta'}{2} = \Delta kl_j \beta_j [\cos(2[\psi - \psi_j]) - \alpha \sin(2[\psi - \psi_j])]$$

$$\begin{pmatrix} \frac{\Delta\beta}{\beta} \\ \frac{\beta\Delta\alpha - \alpha\Delta\beta}{\beta} \end{pmatrix} = \Delta kl_j \beta_j \begin{pmatrix} \sin(2[\psi - \psi_j]) \\ \cos(2[\psi - \psi_j]) \end{pmatrix}$$

$$(\Delta kl_j \beta_j)^2 = \left[\left(\frac{\Delta\beta}{\beta} \right)^2 + \left(\Delta\alpha - \alpha \frac{\Delta\beta}{\beta} \right)^2 \right]$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Twiss correction in one-pass accelerators

$$\frac{\Delta\beta}{\beta} = -\sum_j \Delta k l_j \beta_j \sin(2[\psi - \psi_j]) \quad \Delta\psi = \sum_j \Delta k l_j \beta_j \frac{1}{2} [1 - \cos(2[\psi - \psi_j])]$$

When beta functions and betatron phases have been measured at many places, quadrupoles can be changed with these formulas to correct the Twiss errors.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Quadrupole errors and the tune

$$\begin{aligned}\cos(\mu + \Delta\mu) &= \frac{1}{2} \text{Tr}[M_0(s_j) + \Delta M_0(s_j)] \approx \cos \mu - \Delta\mu \sin \mu \\ &= \frac{1}{2} \text{Tr} \left[\begin{pmatrix} 1 & 0 \\ -\Delta k l_j & 1 \end{pmatrix} \begin{pmatrix} \cos \mu + \alpha_j \sin \mu & \beta_j \sin \mu \\ -\gamma_j \sin \mu & \cos \mu - \alpha_j \sin \mu \end{pmatrix} \right] \\ &= \cos \mu - \frac{1}{2} \Delta k l_j \beta_j \sin \mu\end{aligned}$$

$$\Delta\mu = \frac{1}{2} \Delta k l_j \beta_j$$

Oscillation frequencies can be measured relatively easily and accurately.

Measurement of beta function: Change k and measure tune.

Quadrupole errors and periodic beta function

One pass accelerators:

$$\Delta x(s) = \Delta \vartheta \sqrt{\beta \hat{\beta}} \sin(\psi - \hat{\psi}) \longrightarrow$$

Periodic accelerators:

$$\Delta x_{co}(s) = \frac{\Delta \vartheta \sqrt{\beta \hat{\beta}}}{\sin(\frac{\mu}{2})} \cos(|\psi - \hat{\psi}| - \frac{\mu}{2})$$

$$\frac{\Delta \beta}{\beta} = -\Delta k l \hat{\beta} \sin(2(\psi - \hat{\psi})) \longrightarrow$$

$$\frac{\Delta \beta}{\beta} = -\frac{\Delta k l \hat{\beta}}{2 \sin(\mu)} \cos(2|\psi - \hat{\psi}| - \mu)$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Energy dependent Twiss parameters

Natural Chromaticity: $\xi_x = \frac{1}{2\pi} \partial\mu_x/\partial\delta$, $\xi_y = \frac{1}{2\pi} \partial\mu_y/\partial\delta$

$$\Delta\mu_x = \frac{1}{2} \Delta kl \hat{\beta}_x \quad \longrightarrow \quad \partial\mu_x/\partial\delta = -\frac{1}{2} \int_0^L k(s) \beta_x(s) ds$$

$$\Delta\mu_y = \frac{1}{2} \Delta kl \hat{\beta}_y \quad \longrightarrow \quad \partial\mu_y/\partial\delta = \frac{1}{2} \int_0^L k(s) \beta_y(s) ds$$

The periodic beta functions will similarly depend on energy.



Cornell University



Georg.Hoffstaetter@Cornell.edu

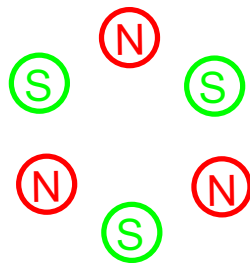
Graduate Accelerator Physics

USPAS summer 2023

Sextupoles (revisited)

$$\psi = \Psi_3 \operatorname{Im}\{(x - iy)^3\} = \Psi_3 \cdot (y^3 - 3x^2y) \Rightarrow \vec{B} = -\vec{\nabla} \psi = \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix}$$

C_3 Symmetry



i) Sextupole fields hardly influence the particles close to the center, where one can linearize in x and y .

ii) In linear approximation a by Δx shifted sextupole has a quadrupole field.

$$\vec{B} = -\vec{\nabla} \psi = \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix}$$

iii) When Δx depends on the energy, one can build an **energy dependent quadrupole**.

$$x \mapsto \Delta x + x$$

$$\vec{B} \approx \Psi_3 3 \begin{pmatrix} 2xy \\ x^2 - y^2 \end{pmatrix} + 6\Psi_3 \Delta x \begin{pmatrix} y \\ x \end{pmatrix} + O(\Delta x^2)$$

$$k_2 = \frac{q}{p} 3! \Psi_3 \Rightarrow k_1 = k_2 \Delta x$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Chromaticity and its correction

Chromaticity ξ = energy dependence of the tune

$$\nu(\delta) = \nu + \frac{\partial \nu}{\partial \delta} \delta + \dots$$

$$\xi = \frac{\partial \nu}{\partial \delta} \quad \text{with} \quad \nu = \frac{\mu}{2\pi}$$

Natural chromaticity ξ_0 = energy dependence of the tune due to quadrupoles only

$$\xi_{x0} = -\frac{1}{4\pi} \oint \beta_x(\hat{s}) k_1(\hat{s}) d\hat{s}$$

$$\xi_{y0} = \frac{1}{4\pi} \oint \beta_y(\hat{s}) k_1(\hat{s}) d\hat{s}$$

Particles with energy difference oscillate around the periodic dispersion leading to a quadrupole effect in sextupoles that also shifts the tune:

$$\xi_x = \frac{1}{4\pi} \oint \beta_x (-k_1 + \eta_x k_2) d\hat{s}$$

$$\xi_y = \frac{1}{4\pi} \oint \beta_y (k_1 - \eta_x k_2) d\hat{s}$$

Typically the chromaticity ξ is chosen to be slightly positive, between 0 and 3.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Chromatic beta beat minimization

Chromatic beta beat is mostly created by the strongest quadrupoles, it can be influenced by sextupoles, under the provision that these on average still correct the chromaticity.

$$\frac{d\beta}{d\delta} = \beta(k_1 l - k_2 l \cdot \eta) \hat{\beta} \sin(2|\psi - \hat{\psi}| - \mu)$$

Periodic accelerators:

$$\frac{d\beta}{d\delta} = \frac{\beta}{2 \sin \mu} (k_1 l - k_2 l \cdot \eta) \hat{\beta} \cos(2|\psi - \hat{\psi}| - \mu)$$

Sextupoles are used to compensate the chromaticity.

Several sextupoles are used to have their average compensate the chromaticity but have their regional variation compensate the chromatic beta beat on average and at critical sections, e.g. interaction points.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Single resonance model

$$x'' = -K x + \Delta f_x(x, y, s)$$

$$\frac{d}{d\vartheta} J = \sum_{n,m=-\infty}^{\infty} m H_{nm}(J) \sin(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$\frac{d}{d\vartheta} \varphi = \nu + \partial_J \sum_{n,m=-\infty}^{\infty} H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$

Strong deviation from: $J = J_0$, $\varphi = \nu \vartheta + \varphi_0$

Occur when there is coherence between the

perturbation and the phase space rotation: $n + m \frac{d}{ds} \varphi \approx 0$

Resonance condition: tune is rational $n + m \nu = 0$

On resonance the integral would increase indefinitely !

Neglecting all but the most important term

$$H(\varphi, J, \vartheta) \approx \nu J + H_{00}(J) + H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Sum and difference resonances

$n + m_x \nu_x + m_y \nu_y \approx 0$ means that oscillations in y can drive oscillations in x in

$$x'' = -K x + \Delta f_x(x, y, s)$$

The resonance term in the Hamiltonian then changes only slowly:

$$H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} + H_{00}(\vec{J}) + H_{n\vec{m}}(\vec{J}) \cos(n\vartheta + \vec{m} \cdot \vec{\varphi} + \Psi_{n\vec{m}}(\vec{J}))$$

$$\frac{d}{d\vartheta} \vec{\varphi} = \vec{\partial}_J H \quad , \quad \frac{d}{d\vartheta} \vec{J} = -\vec{\partial}_\varphi H$$

$$J = \vec{m} \cdot \vec{J}$$

$$J_\perp = m_x J_y - m_y J_x = \vec{m} \times \vec{J} \Rightarrow \frac{d}{d\vartheta} J_\perp = 0$$

Difference resonances lead to stable motion since:

$$n + |m_x| \nu_x - |m_y| \nu_y \approx 0 \Rightarrow |m_y| J_x + |m_x| J_y = \text{const}$$

Sum resonances lead to unstable motion since:

$$n + |m_x| \nu_x + |m_y| \nu_y \approx 0 \Rightarrow |m_y| J_x - |m_x| J_y = \text{const}$$

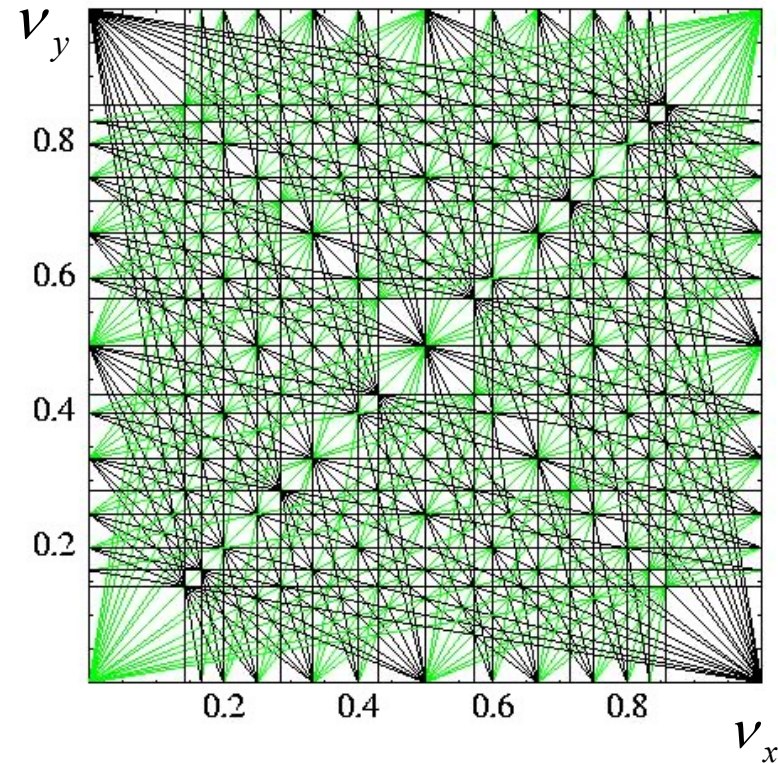
Common reasons for working points

(1) Avoid resonances $n + m_x \nu_x + m_y \nu_y \approx 0$

(2) +/- Colliders: be above a half integer to squeeze the beam size $\frac{\Delta\beta}{\beta} = -\frac{\Delta k\beta}{2 \tan \mu}$

(3) Polarized beams: close to integer

(4) Where the loss rate is smallest



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

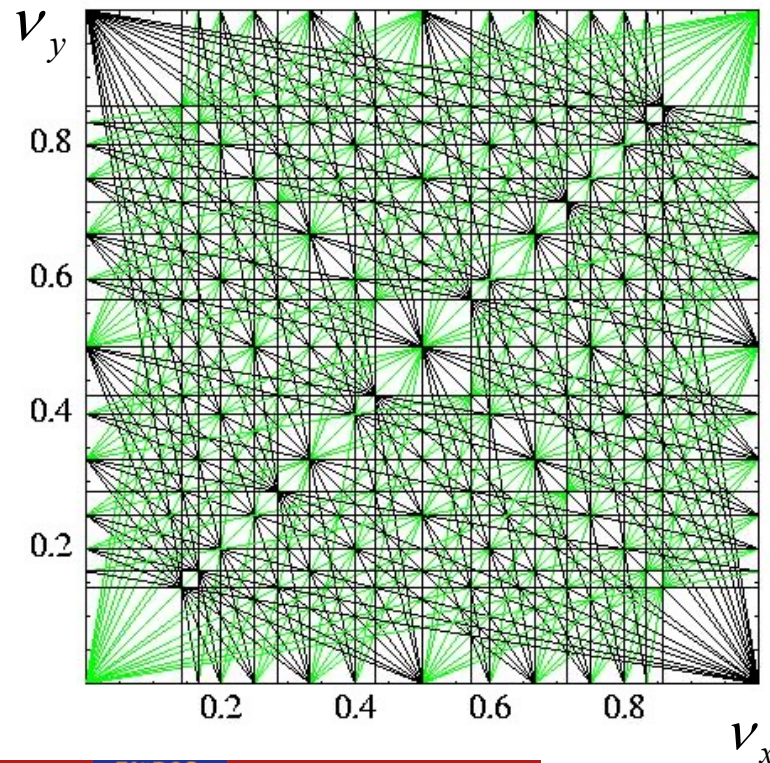
Graduate Accelerator Physics

USPAS summer 2023

Resonance diagram and choosing the tune

$n + m_x \nu_x + m_y \nu_y \approx 0$ means that oscillations in y can drive oscillations in x in

$$x'' = -K x + \Delta f_x(x, y, s)$$



All these resonances have to be avoided by their respective resonance width.

The position of an accelerator in the tune plane is called its Working Point.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Perturbations

$$\begin{pmatrix} x' \\ a' \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ -K & 0 \end{pmatrix} \begin{pmatrix} x \\ a \end{pmatrix} + \begin{pmatrix} 0 \\ \Delta f \end{pmatrix}$$

$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix} = \sqrt{2J} \underline{\beta} \vec{S}$$

This would be a solution with constant J and ϕ when $\Delta f=0$.


Variation of constants:

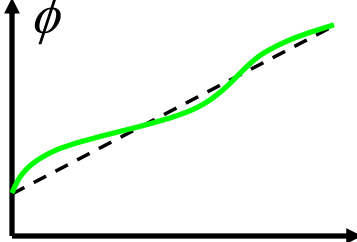
$$\frac{J'}{\sqrt{2J}} \underline{\beta} \vec{S} + \sqrt{2J} \phi_0' \begin{pmatrix} 0 & \sqrt{\beta} \\ -\frac{1}{\sqrt{\beta}} & -\frac{\alpha}{\sqrt{\beta}} \end{pmatrix} \vec{S} = \begin{pmatrix} 0 \\ \Delta f \end{pmatrix}$$

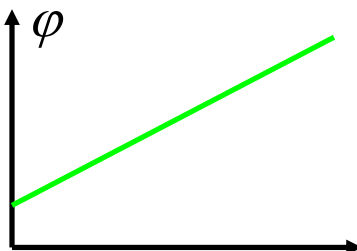
$$\frac{J'}{\sqrt{2J}} \vec{S} + \sqrt{2J} \phi_0' \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \vec{S} = \underline{\beta}^{-1} \begin{pmatrix} 0 \\ \Delta f \end{pmatrix} \quad \text{with} \quad \underline{\beta}^{-1} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix}$$

$$\frac{J'}{\sqrt{2J}} = \cos(\psi + \phi_0) \sqrt{\beta} \Delta f \quad , \quad \sqrt{2J} \phi_0' = -\sin(\psi + \phi_0) \sqrt{\beta} \Delta f$$

Simplification of linear motion

$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix} \Rightarrow \begin{matrix} J' = 0 \\ \phi_0' = 0 \end{matrix}$$


$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin \phi \\ \cos \phi \end{pmatrix} \Rightarrow \begin{matrix} J' = 0 \\ \phi' = \frac{1}{\beta} \end{matrix}$$


$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi - \mu \frac{s}{L} + \phi) \\ \cos(\psi - \mu \frac{s}{L} + \phi) \end{pmatrix} \Rightarrow \begin{matrix} J' = 0 \\ \phi' = \mu \frac{1}{L} \end{matrix}$$


$$\tilde{\psi} = \psi - \mu \frac{s}{L} \Rightarrow \tilde{\psi}(s + L) = \tilde{\psi}(s)$$

Corresponds to Floquet's Theorem



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Quasi-periodic perturbation

$$\begin{aligned} \downarrow J' &= \cos(\psi + \phi_0) \sqrt{2J\beta} \Delta f \quad , \quad \phi_0' = -\sin(\psi + \phi_0) \sqrt{\frac{\beta}{2J}} \Delta f \\ \tilde{\psi} &= \psi - \mu \frac{s}{L} \quad , \quad \varphi = \mu \frac{s}{L} + \phi_0 \\ \rightarrow J' &= \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \quad , \quad \varphi' = \mu \frac{1}{L} - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \end{aligned}$$

New independent variable $\mathcal{G} = 2\pi \frac{s}{L}$

$$\frac{d}{d\mathcal{G}} J = \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\mathcal{G}} \varphi = \nu - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

$$\Delta f(x) = \Delta f(\sqrt{2J\beta} \sin(\tilde{\psi} + \varphi))$$

The perturbations are 2π periodic in \mathcal{G} and in φ

φ is approximately $\varphi \approx \nu \cdot \mathcal{G}$

For irrational ν , the perturbations are **quasi-periodic**.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Tune shift with amplitude

$$\frac{d}{d\mathcal{G}} J = \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\mathcal{G}} \varphi = \nu - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

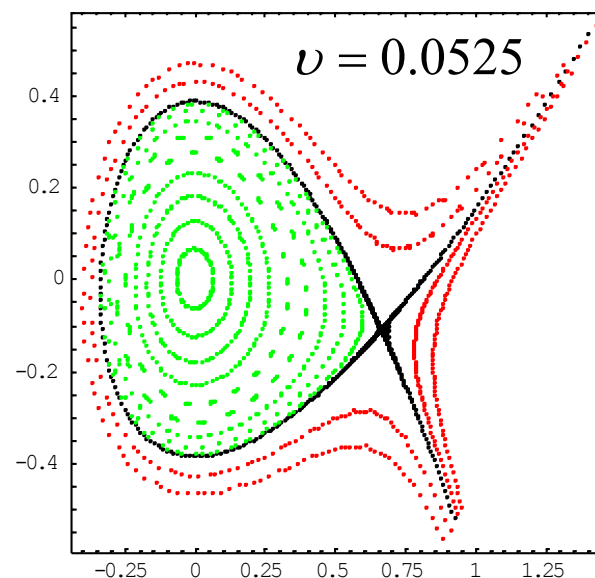
$$\frac{d}{d\mathcal{G}} \varphi = \partial_J H \quad , \quad \frac{d}{d\mathcal{G}} J = -\partial_\phi H \quad , \quad H(\varphi, J, \mathcal{G}) = \nu J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x}$$

The motion remains Hamiltonian in the perturbed coordinates !

If there is a part in $\partial_J H$ that does not depend on $\varphi, s \Rightarrow$ **Tune shift**
The effect of other terms tends to average out.

$$\varphi(\mathcal{G}) - \varphi_0 \approx \mathcal{G} \cdot \partial_J \langle H \rangle_{\varphi, \mathcal{G}}(J)$$

$$\nu(J) = \nu + \partial_J \langle \Delta H \rangle_{\varphi, \mathcal{G}}(J)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Shift examples

$H(\varphi, J) = \nu \cdot J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x} \quad , \quad \Delta \nu(J) = \partial_J \langle \Delta H \rangle_{\varphi, \vartheta}$
 Quadrupole: $\Delta f = -\Delta k x$

$$\Delta H = \frac{L}{2\pi} \Delta k \frac{1}{2} x^2 = \frac{L}{2\pi} \Delta k J \beta \sin^2(\tilde{\psi} + \varphi)$$

$\langle \Delta H \rangle_{\varphi, \vartheta} = \frac{1}{2\pi} \int_0^{2\pi} \Delta k \beta d\vartheta L \frac{J}{4\pi} = \int_0^L \Delta k \beta ds \frac{J}{4\pi} \Rightarrow \Delta \nu = \frac{1}{4\pi} \oint \Delta k \beta ds$
 Sextupole: $\Delta f = -k_2 \frac{1}{2} x^2$

$$\Delta H = \frac{L}{2\pi} k_2 \frac{1}{3!} x^3 = \frac{L}{2\pi} k_2 \frac{1}{3!} \sqrt{2J\beta}^3 \sin^3(\tilde{\psi} + \varphi)$$

$$\langle \Delta H \rangle_{\varphi, \vartheta} = 0 \Rightarrow \Delta \nu = 0$$

Octupole: $\Delta f = -k_3 \frac{1}{3!} x^3$

$$\Delta H = \frac{L}{2\pi} k_3 \frac{1}{4!} x^4 = \frac{L}{2\pi} k_3 \frac{1}{3!} (J\beta)^2 \sin^4(\tilde{\psi} + \varphi)$$

$$\langle \Delta H \rangle_{\varphi, \vartheta} = \frac{J^2}{3!2\pi} \oint k_3 \beta^2 ds \left\langle \frac{1}{2^4} (e^{i\varphi} - e^{-i\varphi})^4 \right\rangle_{\varphi} \Rightarrow \Delta \nu = J \frac{1}{16\pi} \oint k_3 \beta^2 ds$$



Nonlinear resonances

$$\frac{d}{d\vartheta} J = \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi = \nu - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} \varphi = \partial_J H \quad , \quad \frac{d}{d\vartheta} J = -\partial_\varphi H \quad , \quad H(\varphi, J, \vartheta) = \nu \cdot J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x}$$

The effect of the perturbation is especially strong when

$$\cos(\tilde{\psi} + \varphi) \sqrt{\beta} \Delta f \quad \text{or} \quad \sin(\tilde{\psi} + \varphi) \sqrt{\beta} \Delta f$$

has contributions that hardly change, i.e. the change of

$\sqrt{\beta(\vartheta)} \Delta f(x(\vartheta), \vartheta)$ is in resonance with the rotation angle $\varphi(\vartheta)$.

Periodicity allows Fourier expansion:

$$H(\varphi, J, \vartheta) = \sum_{n,m=-\infty}^{\infty} \hat{H}_{nm}(J) e^{i[n\vartheta+m\varphi]} = \sum_{n,m=-\infty}^{\infty} H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$H_{00}(J) = \langle H(\varphi, J, s) \rangle_{\varphi, s} \Rightarrow \text{Tune shift}$$

$$\text{Choosing: } \Psi_{00}(J) = 0$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Nonlinear motion

Sextupoles cause nonlinear dynamics, which can be chaotic and unstable.

$$\begin{pmatrix} x_{n+1} \\ x'_{n+1} \end{pmatrix} = \underline{M}_0 \left[\begin{pmatrix} x_n \\ x'_n \end{pmatrix} - \frac{k_2 l_s}{2} \begin{pmatrix} 0 \\ x_n^2 \end{pmatrix} \right] \quad \begin{pmatrix} x_n \\ x'_n \end{pmatrix} = \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \hat{x}_n \\ \hat{x}'_n \end{pmatrix}$$

$$\begin{pmatrix} \hat{x}_{n+1} \\ \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[\begin{pmatrix} \hat{x}_n \\ \hat{x}'_n \end{pmatrix} - \frac{k_2 l_s}{2} \sqrt{\beta} \begin{pmatrix} 0 \\ \beta \hat{x}_n^2 \end{pmatrix} \right]$$

$$\begin{pmatrix} \hat{x}_f \\ \hat{x}'_f \end{pmatrix} = \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \begin{pmatrix} 1 - \cos \mu & \sin \mu \\ -\sin \mu & 1 - \cos \mu \end{pmatrix}^{-1} \begin{pmatrix} 0 \\ \hat{x}_f^2 \end{pmatrix} = \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \frac{1}{2 \sin \frac{\mu}{2}} \begin{pmatrix} -\cos \frac{\mu}{2} \\ \sin \frac{\mu}{2} \end{pmatrix} \hat{x}_f^2$$

$$\left. \begin{aligned} \hat{x}_f &= -\frac{4}{k_2 l_s} \beta^{-\frac{3}{2}} \tan \frac{\mu}{2} \\ \hat{x}'_f &= \frac{4}{k_2 l_s} \beta^{-\frac{3}{2}} \tan^2 \frac{\mu}{2} \end{aligned} \right\} \hat{x} = \hat{x}_f + \Delta \hat{x} \quad J_f = \frac{1}{2} (\hat{x}_f^2 + \hat{x}'_f^2) = \frac{1}{2 \beta^3} \left(\frac{4}{k_2 l_s} \frac{\tan \frac{\mu}{2}}{\cos \frac{\mu}{2}} \right)^2$$

$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[\begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 - 4 \tan \frac{\mu}{2} \Delta \hat{x}_n \end{pmatrix} \right]$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Dynamic aperture (e.g., close to integer tunes)

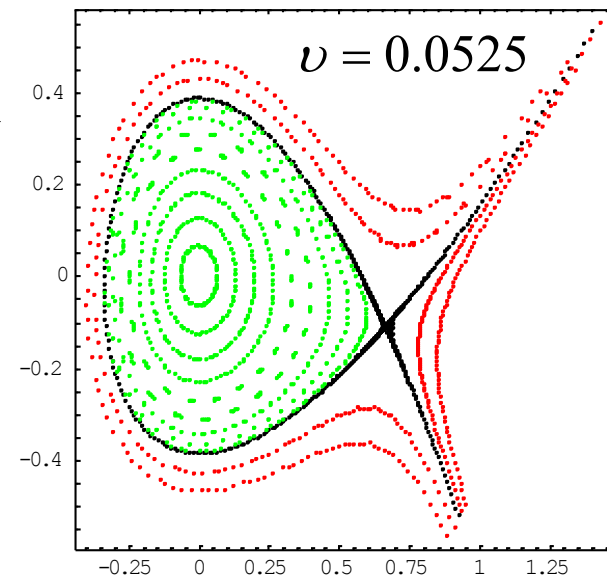
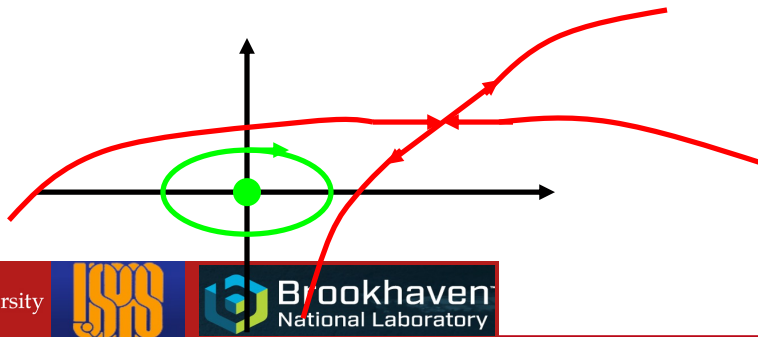
$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[\begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 - 4 \tan \frac{\mu}{2} \Delta \hat{x}_n \end{pmatrix} \right]$$

$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu + 4 \sin \mu \tan \frac{\mu}{2} & \sin \mu \\ -\sin \mu + 4 \cos \mu \tan \frac{\mu}{2} & \cos \mu \end{pmatrix} \left[\begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 \end{pmatrix} \right]$$

Example of one sextupole:

$$\frac{1}{2} \text{Tr}[\underline{M}] = 1 - 2 \sin^2 \frac{\mu}{2} + 4 \sin^2 \frac{\mu}{2} = 1 + 2 \sin^2 \frac{\mu}{2} \geq 1$$

The additional fixed point is unstable !



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

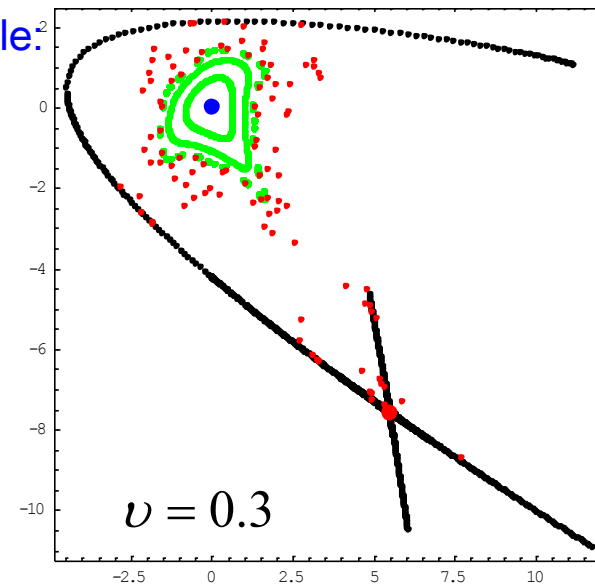
Sextupole Aperture

If the chromaticity is corrected by a single sextupole:

$$\xi_x = \xi_{0x} + \frac{1}{4\pi} \beta_x \eta_x k_2 l \approx 0$$

$$J_f = \frac{1}{2\beta^3} \left(\frac{4}{k_2 l_s} \frac{\tan \frac{\mu}{2}}{\cos^2 \frac{\mu}{2}} \right)^2 \approx \frac{1}{2\beta} \left(\frac{\eta}{\xi_0 \pi} \frac{\sin \frac{\mu}{2}}{\cos^2 \frac{\mu}{2}} \right)^2$$

Often the dynamic aperture is much smaller than the fixed point indicates !



When many sextupoles are used:

$$\xi_{0x} + \frac{N}{4\pi} \beta_x \eta_x k_2 l \approx 0$$

The sum of all k_2^2 is then reduced to about $\sum (k_2 l \beta)^2 \approx N (k_2 l \beta)^2 \approx \frac{1}{N} \left(\frac{4\pi}{\eta} \xi_0 \right)^2$

The dynamic aperture is therefore greatly increased when distributed sextupoles are used.



Cornell University

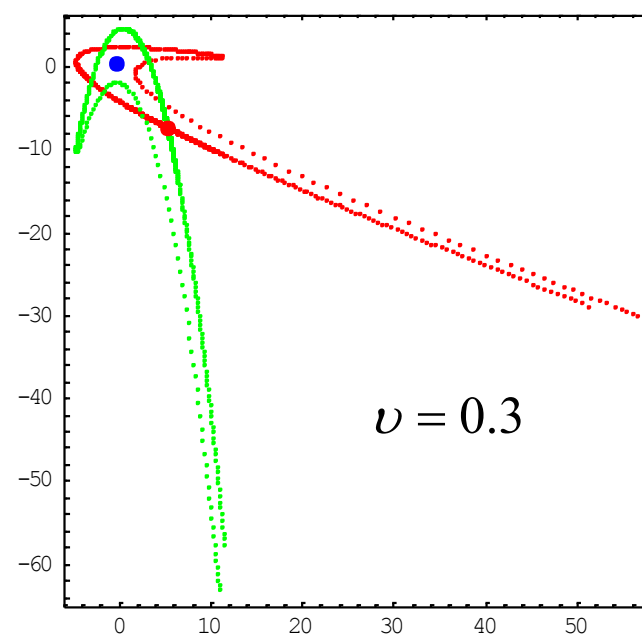
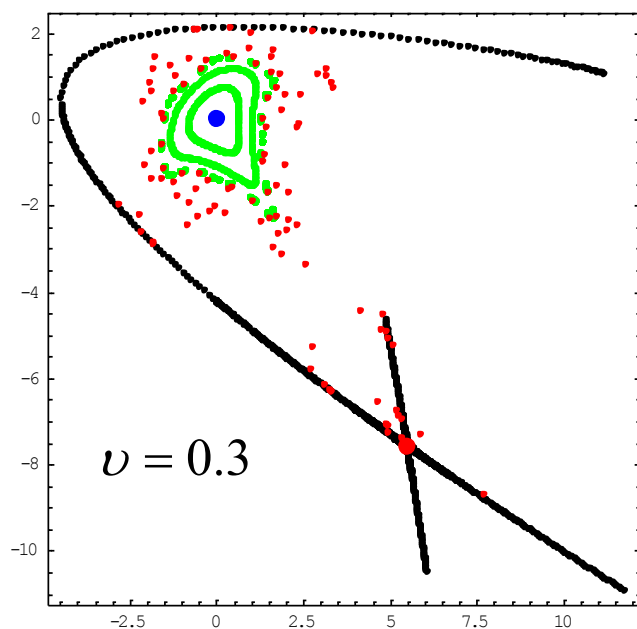


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Sextupole extraction



Due to the narrow region of unstable trajectories, sextupoles are used for slow particle extraction at a tune of $1/3$.

The intersection of **stable** and **unstable** manifolds is a certain indication of chaos.



Cornell University



Brookhaven
National Laboratory

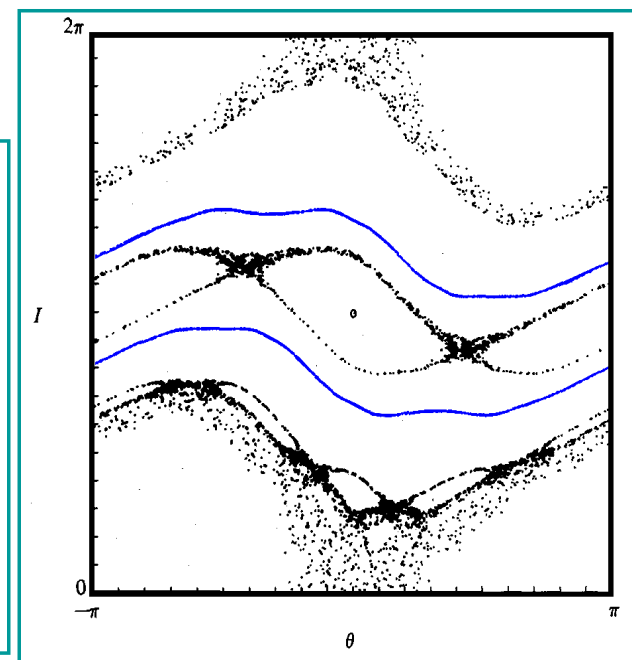
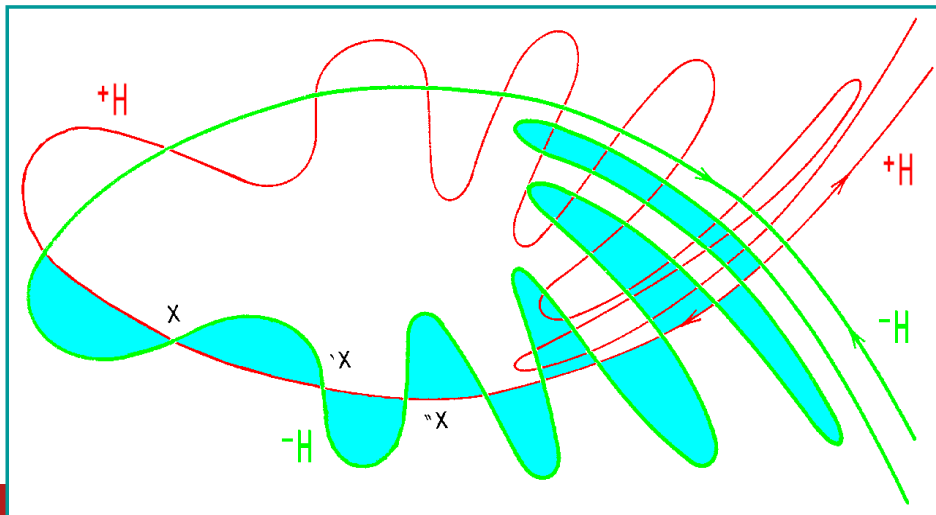
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Homoclinic points

- At instable fixed points, there is a **stable** and an **instable** invariant curve.
- Intersections of these curves (**homoclinic points**) lead to **chaos**.



Cornell University



Brookhaven
National Laboratory

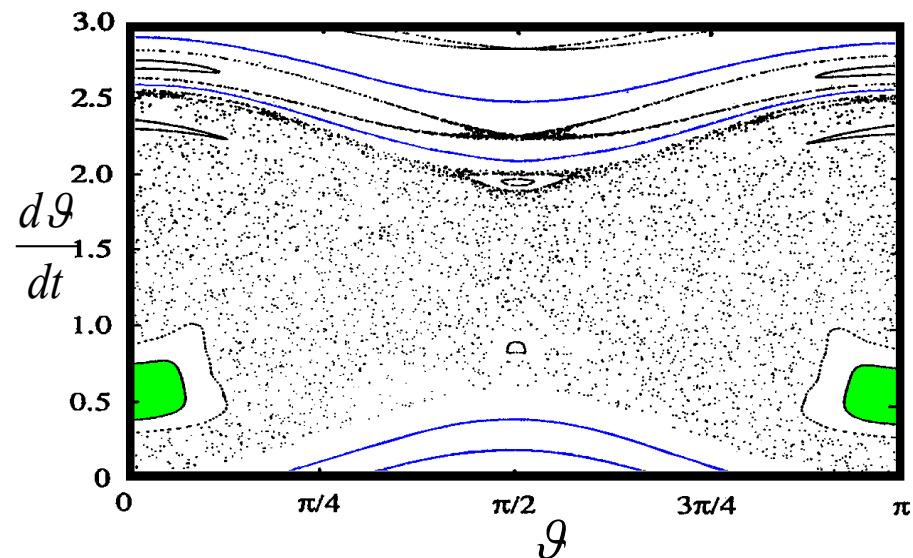
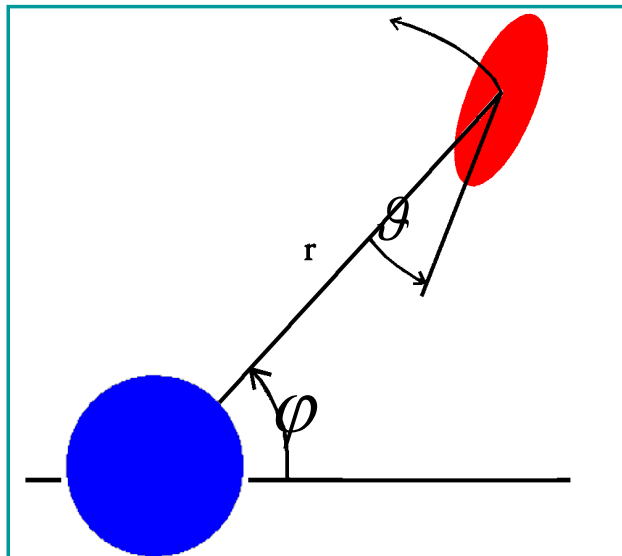
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Hyperion: rotation around the vertical

$$\frac{d^2(\vartheta + \varphi(t))}{dt^2} = -\alpha \left(\frac{a}{r(t)} \right)^3 \sin 2\vartheta$$



- On the path from Rotation to Libration around the Spin-Orbit-Coupling is a strong chaotic region.



Cornell University

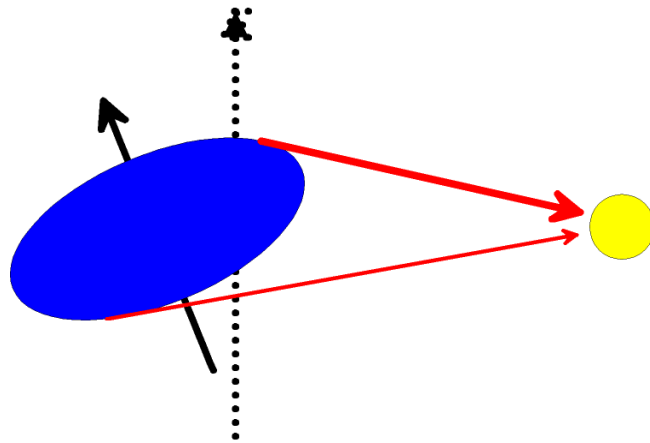


Georg.Hoffstaetter@Cornell.edu

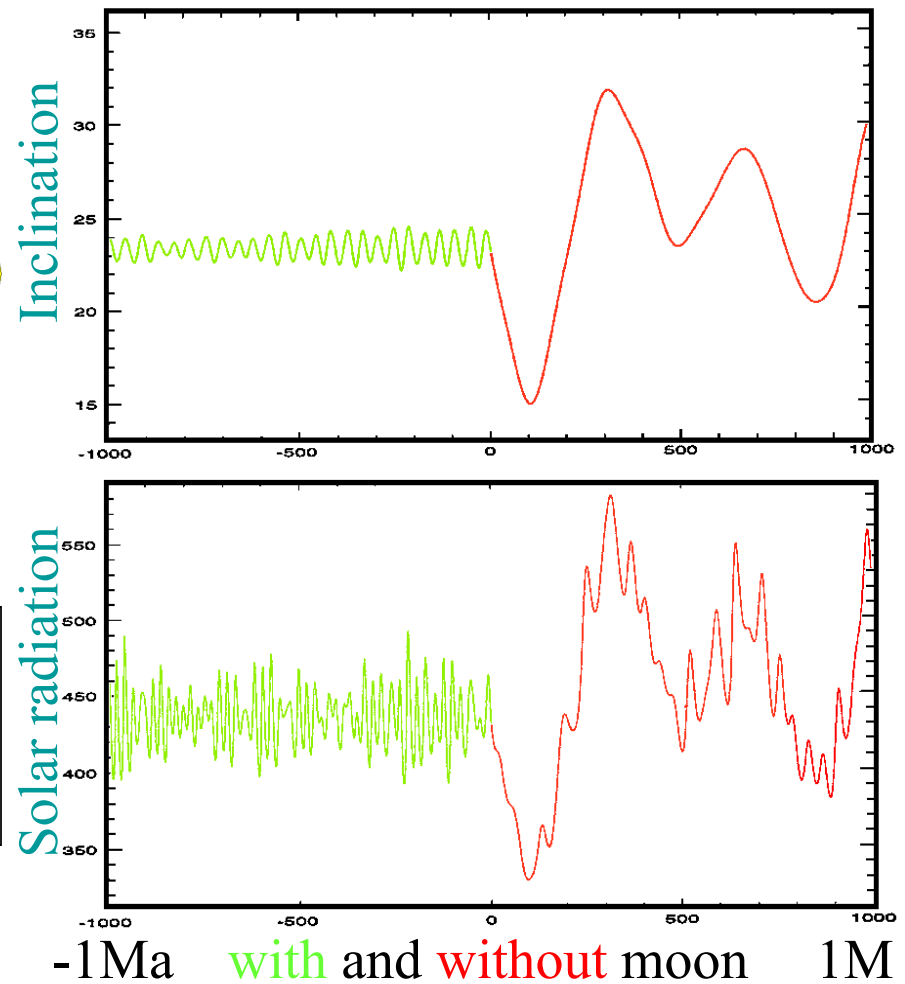
Graduate Accelerator Physics

USPAS summer 2023

Tilt of the earth



- Tidal forces from moon and sun cause a stabilization of the rotation axis.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

a

Single resonance model

$$\frac{d}{d\vartheta} J = \sum_{n,m=-\infty}^{\infty} m H_{nm}(J) \sin(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$\frac{d}{d\vartheta} \varphi = \nu + \partial_J \sum_{n,m=-\infty}^{\infty} H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$

Strong deviation from: $J = J_0$, $\varphi = \nu\vartheta + \varphi_0$

Occur when there is coherence between the

perturbation and the phase space rotation: $n + m \frac{d}{ds} \varphi \approx 0$

Resonance condition: tune is rational $n + m \nu = 0$

On resonance the integral would increase indefinitely !

Neglecting all but the most important term

$$H(\varphi, J, \vartheta) \approx \nu J + H_{00}(J) + H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Fixed points

$$\frac{d}{d\vartheta} J = mH_{nm}(J) \sin(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$\frac{d}{d\vartheta} \varphi = \nu + \Delta\nu(J) + \partial_J [H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))]$$

$$\Phi = \frac{1}{m} [n\vartheta + m\varphi + \Psi_{nm}(J)] , \quad \delta = \nu + \frac{n}{m}$$

$$\frac{d}{d\vartheta} J = mH_{nm}(J) \sin(m\Phi) , \quad \frac{d}{d\vartheta} \Phi = \delta + \Delta\nu(J) + H'_{nm}(J) \cos(m\Phi)$$

$$H(\Phi, J, \vartheta) \approx \delta J + H_{00}(J) + H_{nm}(J) \cos(m\Phi)$$

Fixed points: $\frac{d}{d\vartheta} J = mH_{nm}(J_f) \sin(m\Phi_f) = 0 \Rightarrow \Phi_f = \frac{k}{m} \pi$

If $\delta + \Delta\nu(J_f) \pm H'_{nm}(J_f) = 0$ has a solution.

$$\frac{d}{d\vartheta} \Delta J = \pm m^2 H_{nm}(J_f) \Delta\Phi , \quad \frac{d}{d\vartheta} \Delta\Phi = [\Delta\nu'(J_f) \pm H''_{nm}(J_f)] \Delta J$$

Stable fixed point for: $H_{nm}(J_f) [H''_{nm}(J_f) \pm \Delta\nu'(J_f)] < 0$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Third integer resonances

Sextupole: $\Delta f = -k_2 \frac{1}{2} x^2$

$$\begin{aligned}\Delta H &= \frac{L}{2\pi} k_2 \frac{1}{3!} x^3 = \frac{L}{2\pi} k_2 \frac{1}{3!} \sqrt{2J\beta}^3 \sin^3(\tilde{\psi} + \varphi) \\ &= \frac{L}{2\pi} k_2 \frac{1}{3!4} \sqrt{2J\beta}^3 [\sin(3[\tilde{\psi} + \varphi]) + 3\sin(\tilde{\psi} + \varphi)]\end{aligned}$$

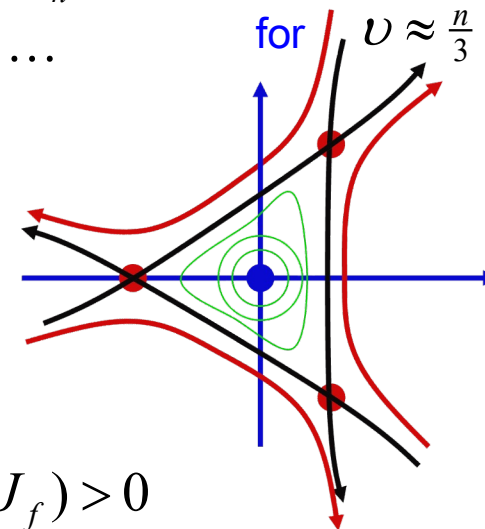
Simplification: one sextupole $k_2(\vartheta) = k_2 \delta(\vartheta) = k_2 \frac{1}{2\pi} \sum_{n=-\infty}^{\infty} \cos(n\vartheta)$

$$\Delta H = \frac{L}{2\pi} k_2 \frac{1}{3!4} \sqrt{2J\beta}^3 \frac{1}{2\pi} \cos(-n\vartheta + 3\varphi + 3\tilde{\psi} - \frac{\pi}{2}) + \dots$$

$$\Delta H \approx A_2 \sqrt{J}^3 \cos(3\Phi)$$

$$\left. \begin{aligned} \Phi_f &= 0, \frac{1}{3}\pi, \frac{2}{3}\pi, \dots \\ \delta \pm A_2 \frac{3}{2} \sqrt{J} &= 0 \end{aligned} \right\} \begin{aligned} \Phi_f &= \frac{1}{3}\pi, \pi, \frac{5}{3}\pi \\ &\text{for } \delta > 0 \end{aligned}$$

All these fixed points are instable since $H_{nm}(J_f) H''_{nm}(J_f) > 0$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Fourth integer resonances

Octupole: $\Delta f = -k_3 \frac{1}{3!} x^3$, $\Delta H = \frac{L}{2\pi} k_3 \frac{1}{4!} x^4 = \frac{L}{2\pi} k_3 \frac{1}{3!} J^2 \beta^2 \sin^4(\tilde{\psi} + \varphi)$
 $= \frac{L}{2\pi} k_3 \frac{1}{3!8} J^2 \beta^2 [\cos(4[\tilde{\psi} + \varphi]) - 4\cos(2[\tilde{\psi} + \varphi]) + 3]$

Simplification: one octupole $k_3(\mathcal{Q}) = k_3 \delta(\mathcal{Q}) = k_3 \frac{1}{2\pi} \sum_{n=-\infty}^{\infty} \cos(n\mathcal{Q})$

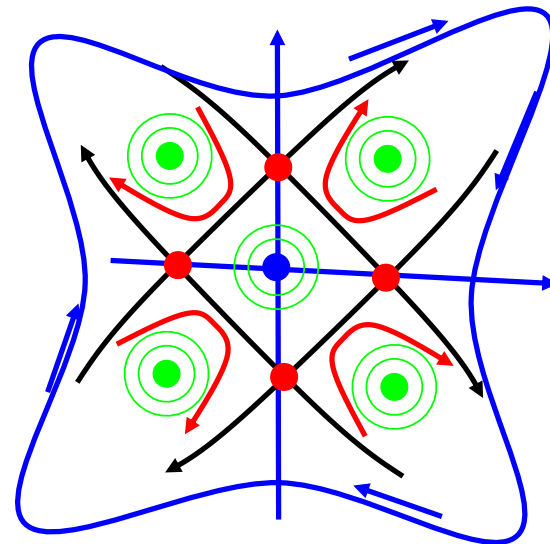
$\Delta H \approx A_3 J^2 [3 + \cos(4\Phi)]$ for $\nu \approx \frac{n}{4}$

$\Phi_f = 0, \frac{1}{4}\pi, \frac{2}{4}\pi, \dots$ Either 8 fixed points: $\delta < 0$

$\delta + A_3 2J(3 \pm 1) = 0$ or none for: $\delta > 0$

$H_{nm}(J_f)[H_{nm}''(J_f) \pm \Delta\nu'(J_f)] < 0$

Stability for $(2A_3 J)^2 [1 \pm 3] < 0$,
 i.e. for the 4 outer fixed points.



Cornell University



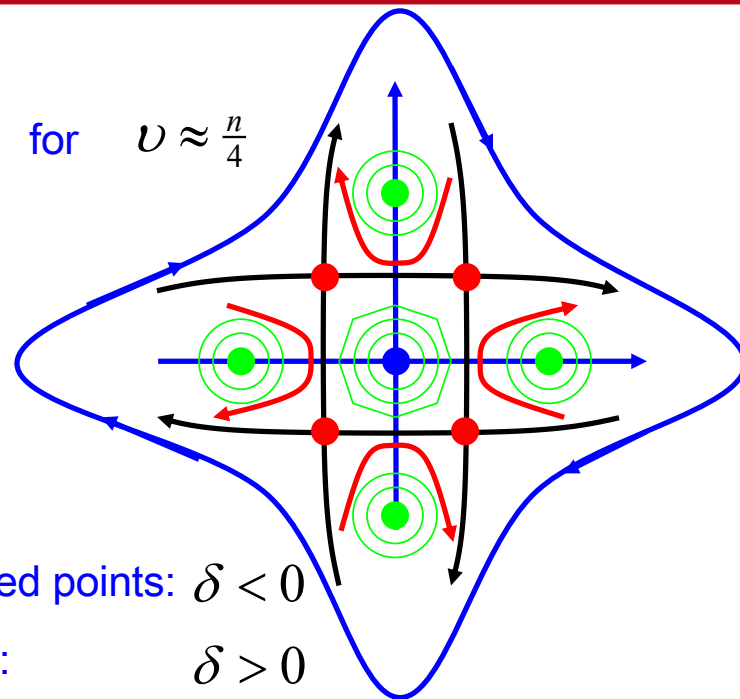
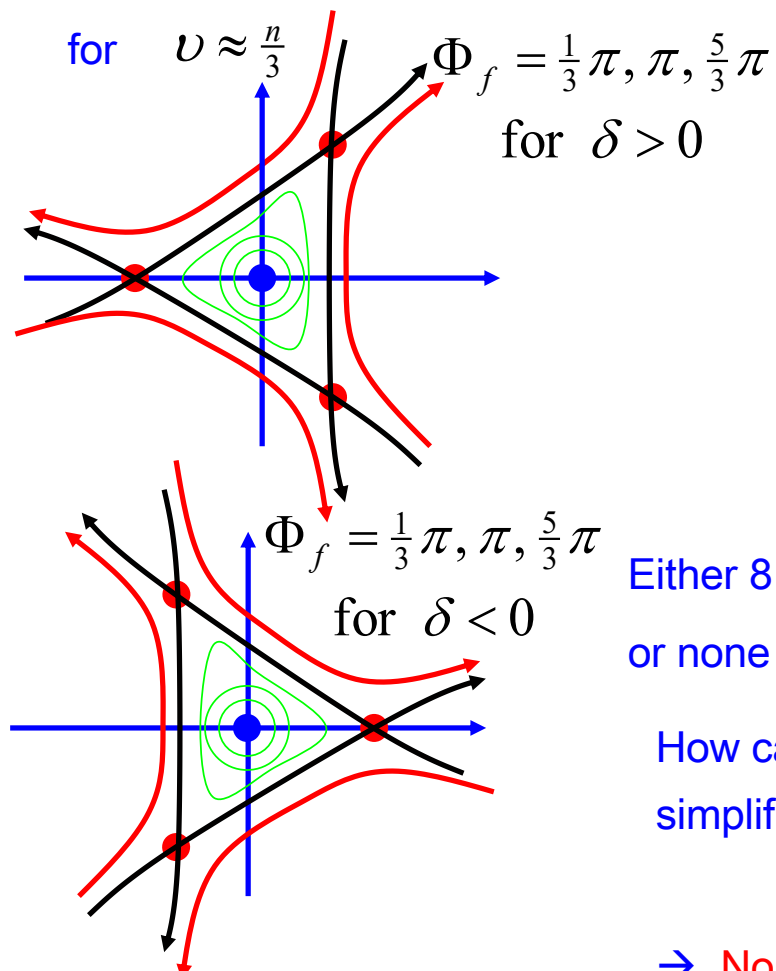
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Particle motion in the single resonance model



Either 8 fixed points: $\delta < 0$
or none for: $\delta > 0$

How can the motion inside the fixed points be simplified for a real accelerator ?

→ Normal Form Theory



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

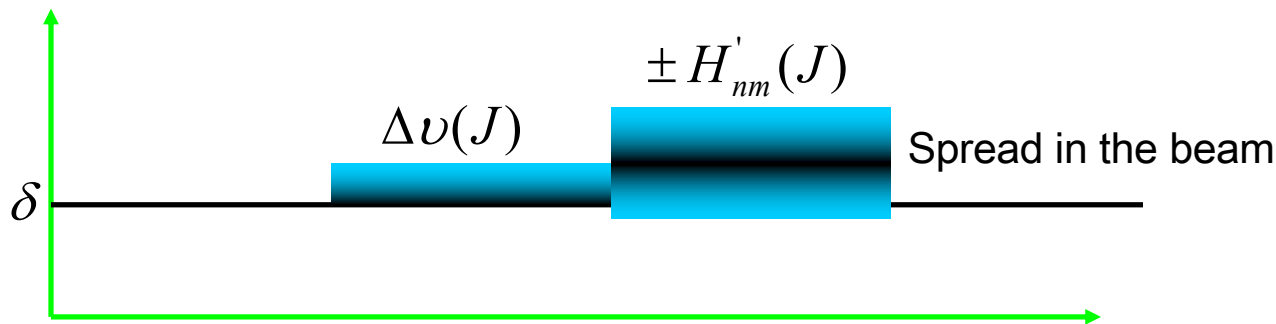
USPAS summer 2023

Resonance width (strength)

Fixed points: $\frac{d}{dg} J = mH_{nm}(J_f) \sin(m\Phi_f) = 0 \Rightarrow \Phi_f = \frac{k}{m} \pi$

If $\delta + \Delta\nu(J_f) \pm H'_{nm}(J_f) = 0$ has a solution.

δ has to avoid the region $\delta + \Delta\nu(J) \pm H'_{nm}(J) = 0$ for all particles.



Assuming that the tune shift and perturbation are monotonous in J :

This tune region has the width $\Delta_{nm} = 2 |H'_{nm}(J_{\max})|$ for strong resonances.

Δ_{nm} Is called **Resonance Width**, Resonance Strength, or Stop-Band Width



Cornell University



Brookhaven
National Laboratory

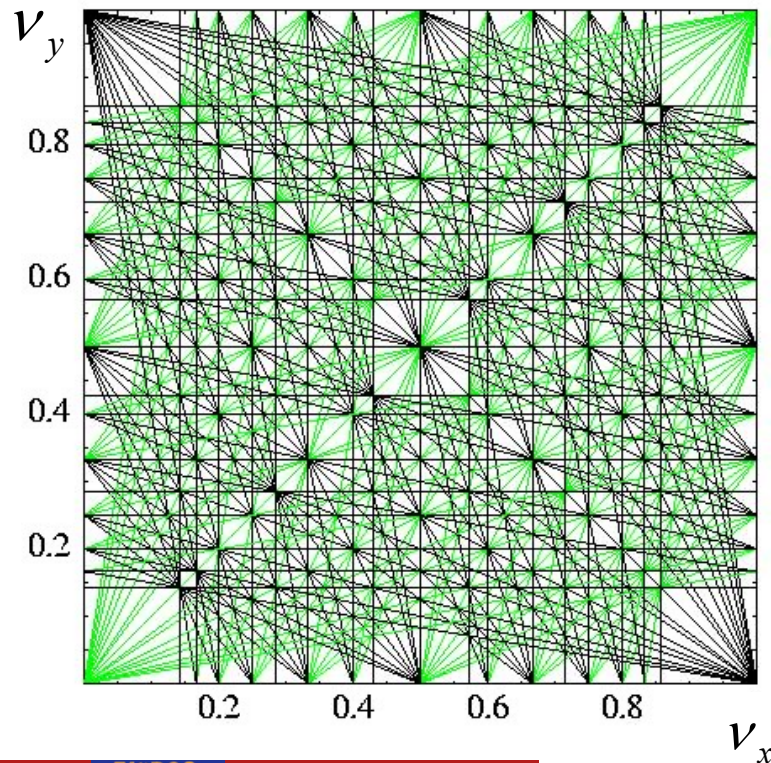
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Resonance width around resonance lines

$n + m_x \nu_x + m_y \nu_y \approx 0$ means that oscillations in y can drive oscillations in x in



All these resonances have to be avoided by their respective resonance width.

The position of an accelerator in the tune plane is called its Working Point.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Coupling resonances

$$\frac{d}{d\vartheta} J_x = \cos(\tilde{\psi}_x + \varphi_x) \sqrt{2J_x \beta_x} \Delta f_x \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi_x = \nu_x - \sin(\tilde{\psi}_x + \varphi_x) \sqrt{\frac{\beta_x}{2J_x}} \Delta f_x \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} J_y = \cos(\tilde{\psi}_y + \varphi_y) \sqrt{2J_y \beta_y} \Delta f_y \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi_y = \nu_y - \sin(\tilde{\psi}_y + \varphi_y) \sqrt{\frac{\beta_y}{2J_y}} \Delta f_y \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} \vec{\varphi} = \vec{\partial}_J H \quad , \quad \frac{d}{d\vartheta} \vec{J} = -\vec{\partial}_{\varphi} H \quad , \quad H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} - \frac{L}{2\pi} \int_0^{\vec{x}} \Delta \vec{f}(\hat{x}, s) d\hat{x}$$

The integral form can be chosen since it is path independent. This is due to the Hamiltonian nature of the force:

$$\Delta f_{x,y}(x, y, s) = -\partial_{x,y} \Delta H(x, y, s)$$

Single Resonance model for two dimensions means retaining only the amplitude dependent tune shift and one term in the two dimensional Fourier expansion:

$$H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} + H_{00}(\vec{J}) + H_{n\vec{m}}(\vec{J}) \cos(n\vartheta + m_x \varphi_x + m_y \varphi_y + \Psi_{n\vec{m}}(\vec{J}))$$

For $n + m_x \nu_x + m_y \nu_y \approx 0$

$$m_x \varphi_x + m_y \varphi_y = \vec{m} \cdot \vec{\varphi}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

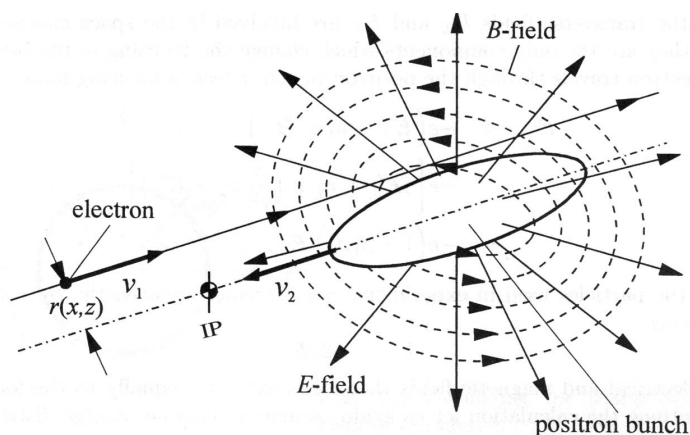
Graduate Accelerator Physics

USPAS summer 2023

Beam-beam tune shift

The force that acts from one beam to the other during collisions is focusing or defocusing in both planes for small distances.

For large distances it is very nonlinear.



The effects of E and B forces add.
Whereas they subtract for co-moving particles.

$$\Delta \nu_x^{(1)} = \frac{r_{cl}^{(1)} N_{cpb}^{(2)}}{2\pi} \frac{\beta_x^{(1)}}{\sigma_x^{(2)} (\sigma_x^{(2)} + \sigma_y^{(2)})}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Beam-beam force

$$\left. \begin{aligned} \rho_{\text{lab}} &= \frac{1}{2\pi\sigma_x\sigma_y} e^{-\left(\frac{x^2}{2\sigma_x^2} + \frac{y^2}{2\sigma_y^2}\right)} \rho_{\text{lab}z}(z + v^{(2)}t) \\ \vec{j}_{\text{lab}} &= \vec{\beta}c\rho_{\text{lab}z} \end{aligned} \right\} \left\{ \begin{aligned} \rho_{\text{rest}} &= \frac{1}{2\pi\sigma_x\sigma_y} e^{-\left(\frac{x^2}{2\sigma_x^2} + \frac{y^2}{2\sigma_y^2}\right)} \rho_z(z) \\ \vec{j}_{\text{rest}} &= 0 \end{aligned} \right.$$

$$E_x(x, y, z) = \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{2\pi\sigma_x\sigma_y} \int \frac{x-x_0}{(x-x_0)^2 + (y-y_0)^2} e^{-\left(\frac{x_0^2}{2\sigma_x^2} + \frac{y_0^2}{2\sigma_y^2}\right)} dx_0 dy_0$$

$$\approx \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{2\pi\sigma_x\sigma_y} \int \frac{u}{u^2 + v^2} e^{-\left(\frac{(x-u)^2}{2\sigma_x^2} + \frac{(y-v)^2}{2\sigma_y^2}\right)} dudv$$

$$= \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{2\pi\sigma_x\sigma_y} (\sigma_x^2 \partial_x + x) \int \frac{1}{u^2 + v^2} e^{-\left(\frac{(x-u)^2}{2\sigma_x^2} + \frac{(y-v)^2}{2\sigma_y^2}\right)} dudv$$

$$= \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{2\pi\sigma_x\sigma_y} (\sigma_x^2 \partial_x + x) \int \left[\int_0^\infty e^{-t(u^2 + v^2)} dt \right] e^{-\left(\frac{(x-u)^2}{2\sigma_x^2} + \frac{(y-v)^2}{2\sigma_y^2}\right)} dudv$$

$$= \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{2\pi\sigma_x\sigma_y} (\sigma_x^2 \partial_x + x) \int e^{-\left(u^2\left(\frac{1}{2\sigma_x^2} + t\right) - \frac{2xu}{2\sigma_x^2} + v^2\left(\frac{1}{2\sigma_y^2} + t\right) - \frac{2yv}{2\sigma_y^2}\right) - \left(\frac{x^2}{2\sigma_x^2} + \frac{y^2}{2\sigma_y^2}\right)} e^{-\left(\frac{x^2}{2\sigma_x^2} + \frac{y^2}{2\sigma_y^2}\right)} dudvdt$$



Beam-beam force

$$\begin{aligned}
 E_x(x, y, z) &\approx \frac{Q\rho_z}{4\pi\epsilon_0} \frac{1}{\pi\sigma_x\sigma_y} (\sigma_x^2 \partial_x + x) \int e^{-\left(u^2\left(\frac{1}{2\sigma_x^2}+t\right)-\frac{2xu}{2\sigma_x^2}+v^2\left(\frac{1}{2\sigma_y^2}+t\right)-\frac{2yv}{2\sigma_y^2}\right)} e^{-\left(\frac{x^2}{2\sigma_x^2}+\frac{y^2}{2\sigma_y^2}\right)} dudvdt \\
 &= \frac{Q\rho_z}{4\pi\epsilon_0} \frac{1}{\pi\sigma_x\sigma_y} (\sigma_x^2 \partial_x + x) \int e^{-\left(\tilde{u}^2\left(\frac{1}{2\sigma_x^2}+t\right)+\tilde{v}^2\left(\frac{1}{2\sigma_y^2}+t\right)\right)} e^{-\left(\frac{x^2}{2\sigma_x^2}-\frac{x^2}{4\sigma_x^4\left(\frac{1}{2\sigma_x^2}+t\right)}+\frac{y^2}{2\sigma_y^2}-\frac{y^2}{4\sigma_y^4\left(\frac{1}{2\sigma_y^2}+t\right)}\right)} dudvdt \\
 &= \frac{Q\rho_z}{4\pi\epsilon_0} 2(\sigma_x^2 \partial_x + x) \int_0^\infty \frac{e^{-\left(\frac{x^2}{2\sigma_x^2+\frac{1}{t}}+\frac{y^2}{2\sigma_y^2+\frac{1}{t}}\right)}}{\sqrt{2\sigma_x^2+\frac{1}{t}}\sqrt{2\sigma_y^2+\frac{1}{t}}} \frac{dt}{t} \\
 &= -\frac{Q\rho_z}{4\pi\epsilon_0} \partial_x \int_0^\infty \frac{e^{-\left(\frac{x^2}{2\sigma_x^2+q}+\frac{y^2}{2\sigma_y^2+q}\right)}}{\sqrt{2\sigma_x^2+q}\sqrt{2\sigma_y^2+q}} dq = -\partial_x U \\
 &\approx \frac{Q\rho_z}{2\pi\epsilon_0} x \int_0^\infty \frac{1}{\sqrt{2\sigma_x^2+q}^3 \sqrt{2\sigma_y^2+q}} dq = \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{\sigma_x(\sigma_x+\sigma_y)} x
 \end{aligned}$$



Beam-beam force

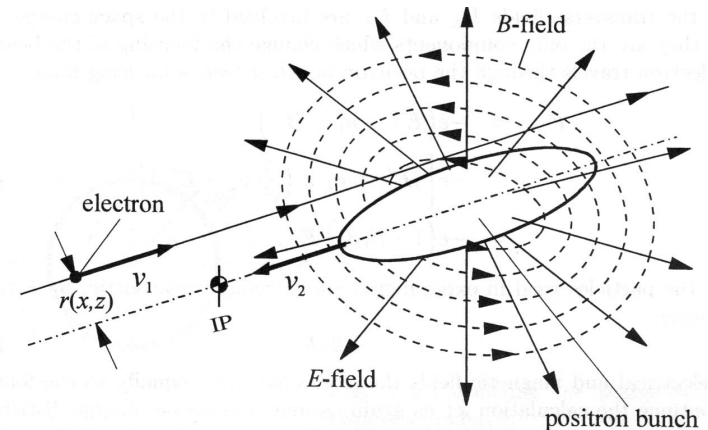
$$\left. \begin{aligned} \vec{E}(x, y, z) &\approx \frac{Q\rho_z}{2\pi\epsilon_0} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix} \\ \vec{B}(x, y, z) &= 0 \end{aligned} \right\} \left\{ \begin{aligned} \vec{E}_{\text{lab}}(x, y, z) &\approx \frac{Q\rho_{\text{lab}z}}{2\pi\epsilon_0} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix} \\ \vec{B}_{\text{lab}}(x, y, z) &= \frac{1}{c} \vec{\beta} \times \vec{E}_{\text{lab}}(x, y, z) \end{aligned} \right.$$

$$\Delta\vec{p}(x, y) = \int \vec{F}(x, y, v^{(1)}t) dt$$

$$\approx \frac{Q}{2\pi\epsilon_0} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix} \int \rho_{\text{lab}z} (v^{(1)}t + v^{(2)}t) dt (1 + \beta^{(1)}\beta^{(2)})$$

$$= \frac{Q}{2\pi\epsilon_0} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix} \frac{1 + \beta^{(1)}\beta^{(2)}}{v^{(1)} + v^{(2)}}$$

$$\approx \frac{Q}{2\pi\epsilon_0 c} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

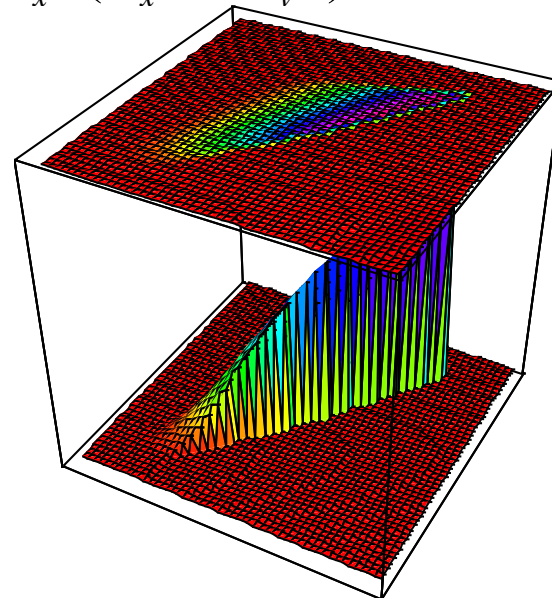
USPAS summer 2023

Beam-beam tune shift

$$\begin{pmatrix} \Delta x' \\ \Delta y' \end{pmatrix} = \frac{\Delta \vec{p}}{p} \approx \frac{q^{(1)} Q}{2\pi \varepsilon_0 p^{(1)} c} \frac{1}{\sigma_x + \sigma_y} \begin{pmatrix} \frac{x}{\sigma_x} \\ \frac{y}{\sigma_y} \end{pmatrix} = \begin{pmatrix} k_x x \\ k_y y \end{pmatrix}$$

$$\Delta \nu_x^{(1)} \approx \frac{q^{(1)} Q}{8\pi^2 \varepsilon_0 p^{(1)} c} \frac{\beta_x^{(1)}}{\sigma_x (\sigma_x + \sigma_y)} \approx \frac{r_{cl}^{(1)} N_{cpb}^{(2)}}{2\pi} \frac{\beta_x^{(1)}}{\sigma_x^{(2)} (\sigma_x^{(2)} + \sigma_y^{(2)})}$$

Tune spread over the beam,
amplitude dependent tune shift
and the tune shift cravat



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

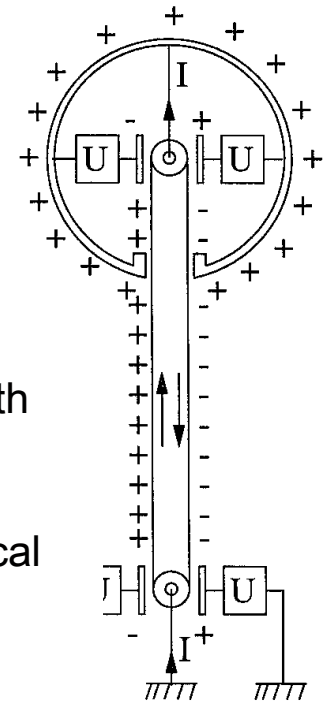
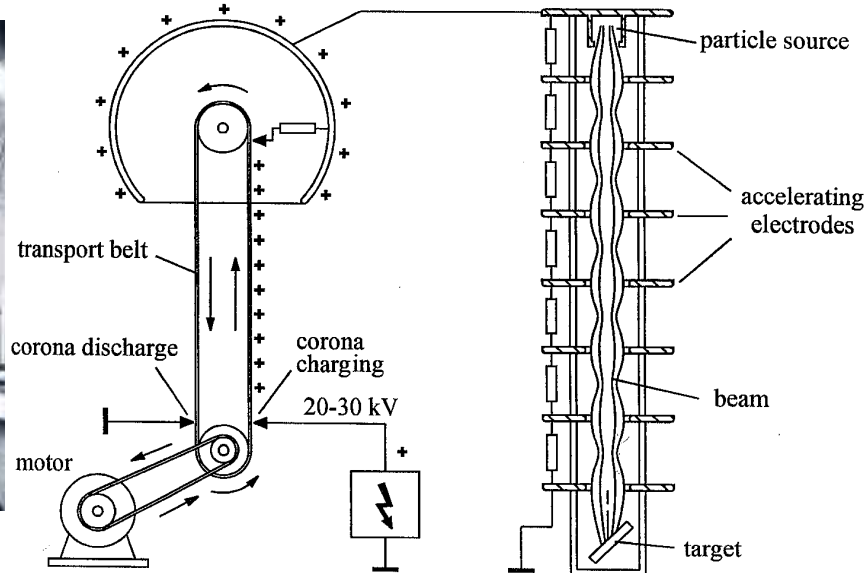
USPAS summer 2023

The Van de Graaff Accelerator

- 1930: van de Graaff builds the first 1.5MV high voltage generator



Van de
Graaff



- Today Pelletrons (with chains) or Laddertron (with stripes) that are charged by influence are commercially available.
- Used as injectors, for electron cooling, for medical and technical n-source via $d + t \rightarrow n + \alpha$
- Up to 17.5 MV with insulating gas (1MPa SF_6)



Cornell University



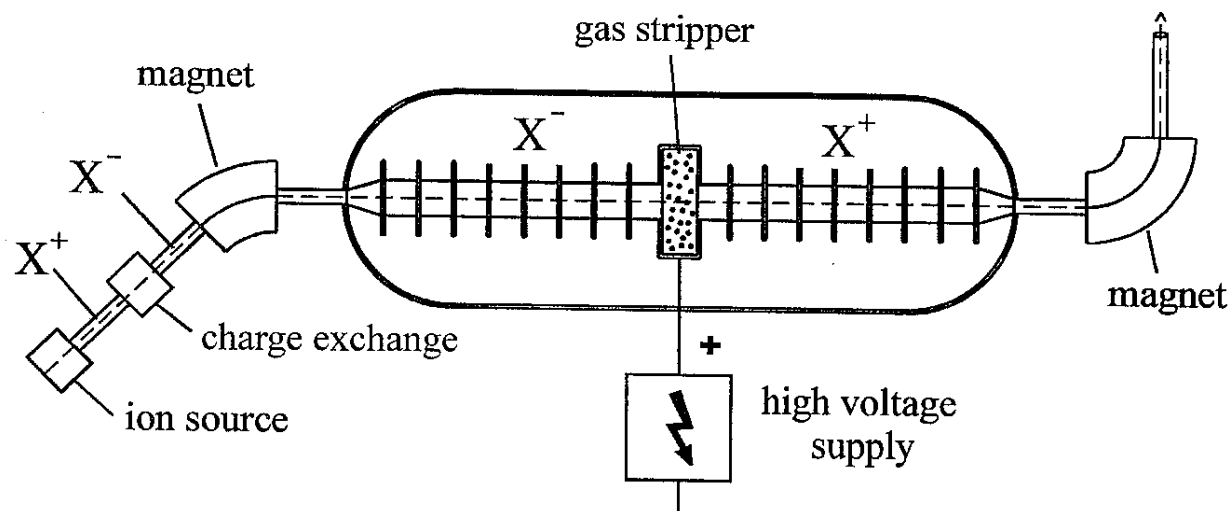
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

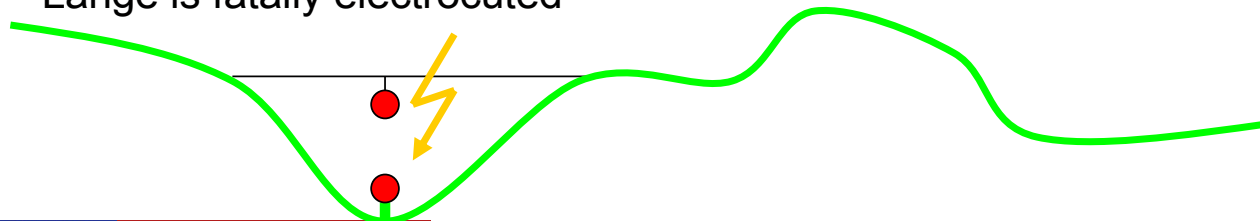
USPAS summer 2023

The Tandem (Van de Graaff) Accelerator

- Two Van de Graaffs, one + one -
- The Tandem Van de Graaff, highest energy 35MeV



- 1932: Brasch and Lange use potential from lightening, in the Swiss Alps, Lange is fatally electrocuted



Cornell University



Brookhaven
National Laboratory

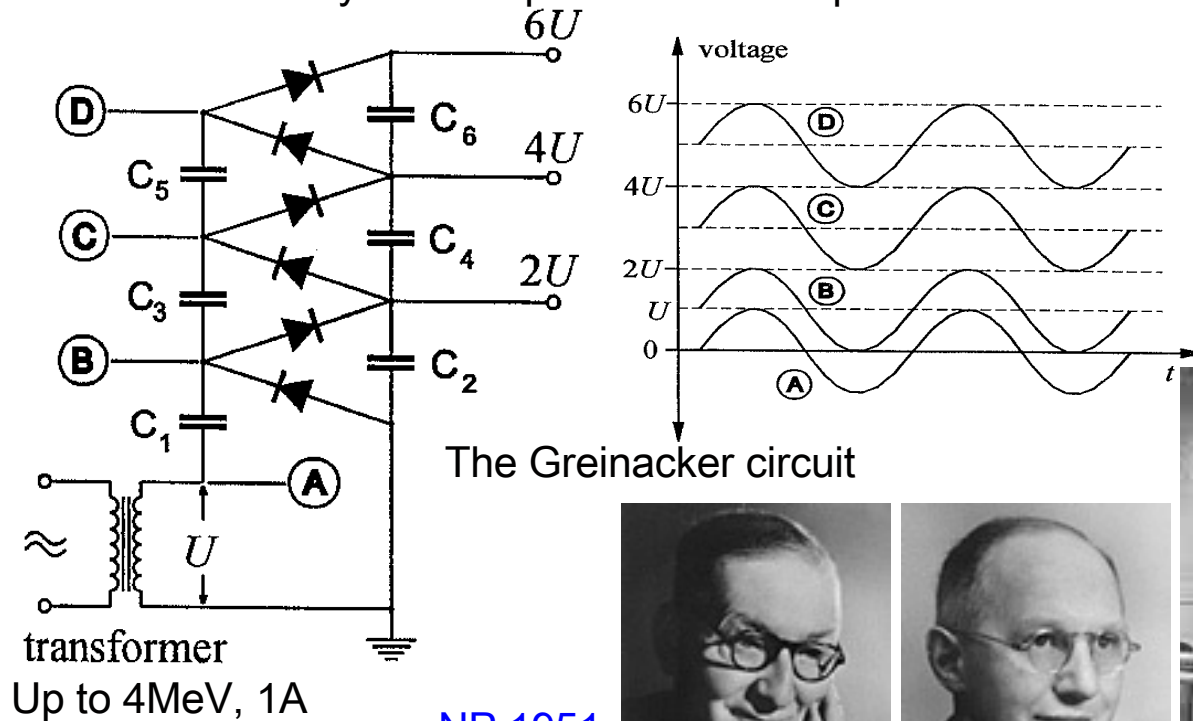
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

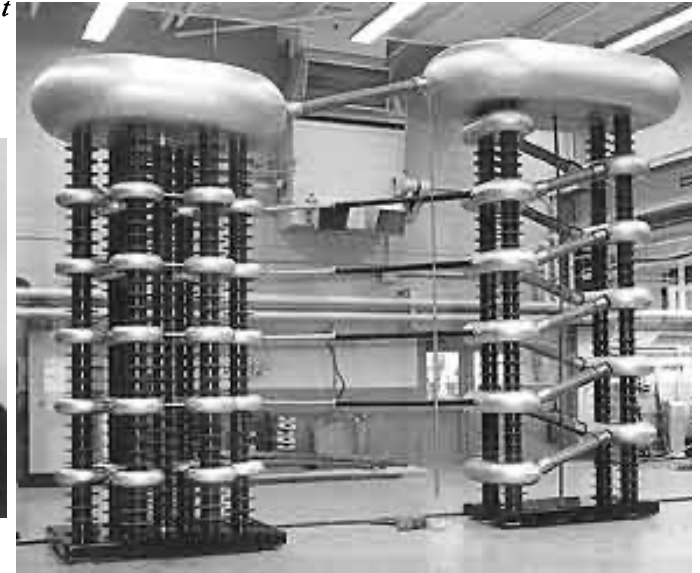
USPAS summer 2023

The Cockcroft-Walton Accelerator

1932: Cockcroft and Walton 1932: 700keV cascade generator (planned for 800keV)
and use initially 400keV protons for ${}^7\text{Li} + p \mapsto {}^4\text{He} + {}^4\text{He}$ and ${}^7\text{Li} + p \mapsto {}^7\text{Be} + n$



The Greinacker circuit



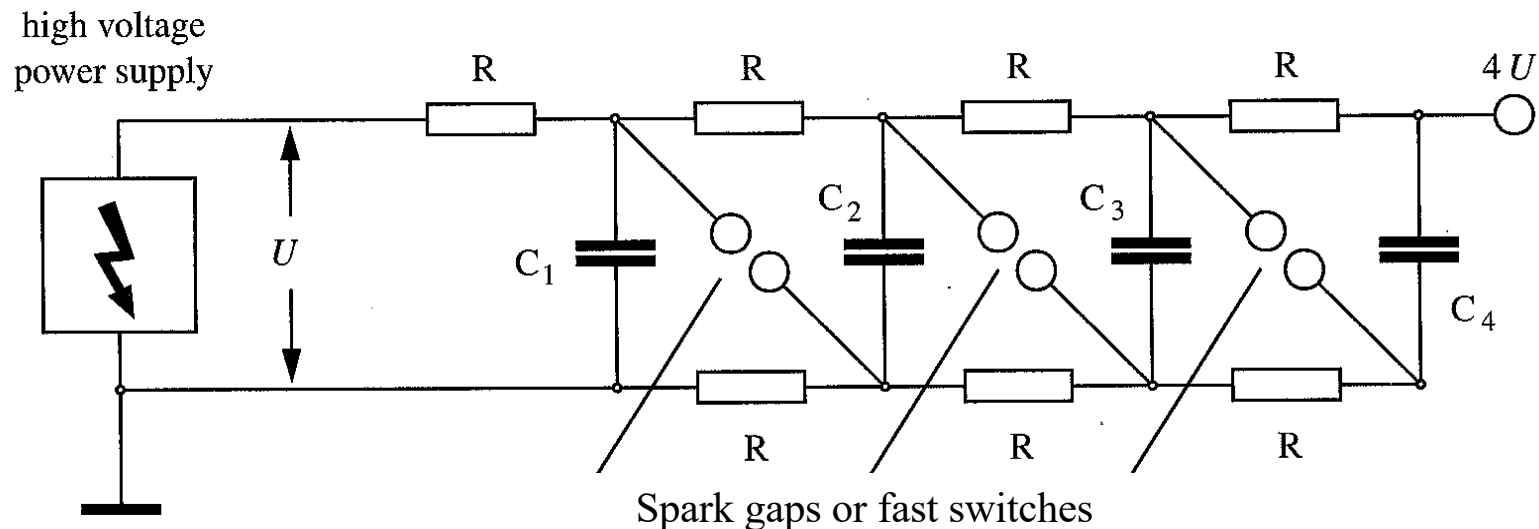
NP 1951

Sir John D Cockcrof

Ernest T S Walton

The Marx Generator

- 1932: Marx Generator achieves 6MV at General Electrics



After capacitors of around 2uF are filled to about 20kV, the spark gaps or switches close as fast as 40ns, allowing up to 500kA.

Today: The Z-machine (Physics Today July 2003) for z-pinch initial confinement fusion has 40TW for 100ns from 36 Marx generators



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

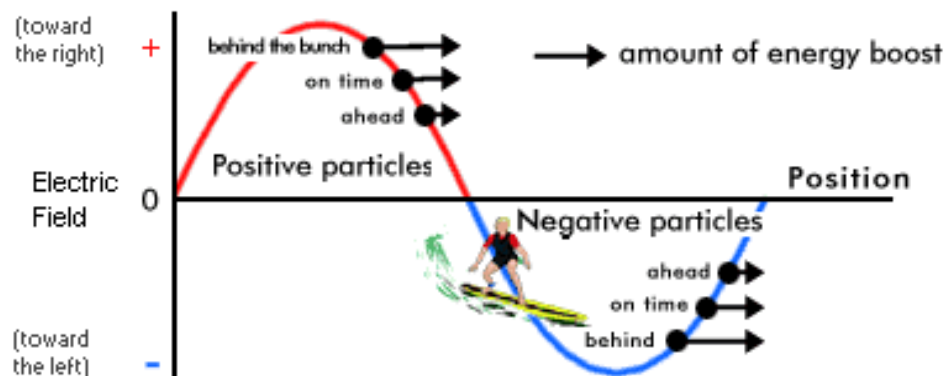
USPAS summer 2023

Three historic lines of accelerators

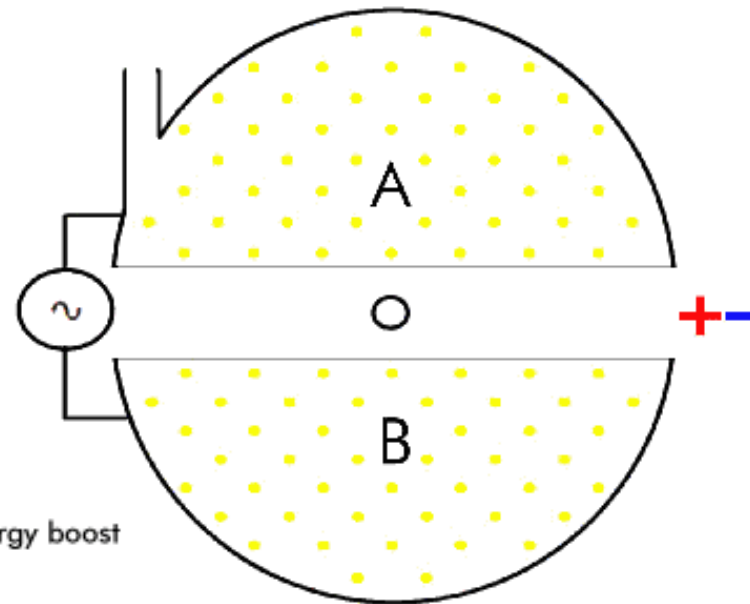
Direct Voltage Accelerators



Resonant Accelerators



Transformer Accelerator



Particles must have the correct phase relation to the accelerating voltage.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Cyclotron frequency

$$F_r = m_0 \gamma \omega_z v = q v B_z$$

$$\omega_z = \frac{q}{m_0 \gamma} B_z = \text{const}$$

Condition: Non-relativistic particles.
Therefore, not for electrons.

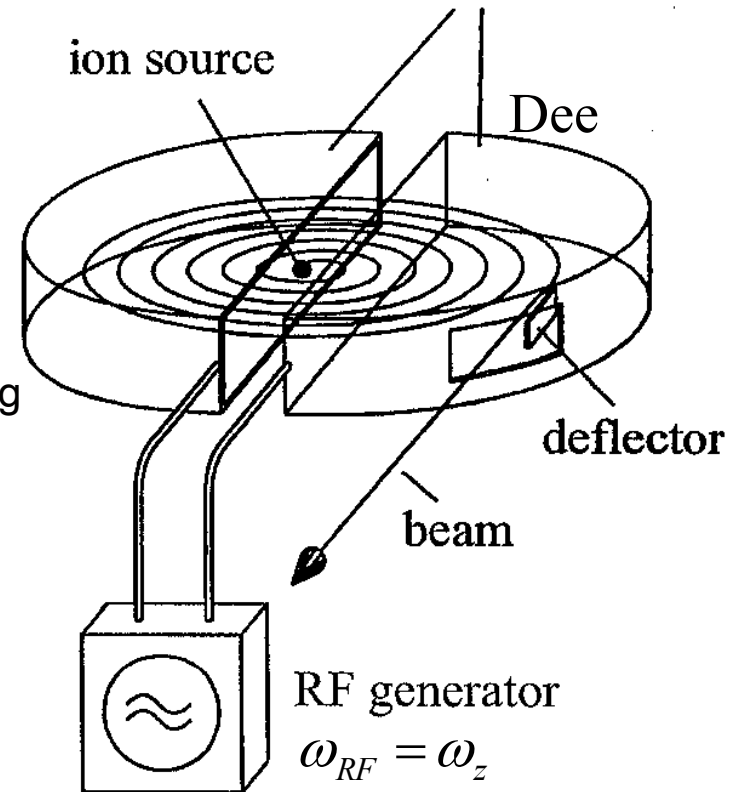
- The synchrocyclotron:
Acceleration of bunches with decreasing

$$\omega_z(E) = \frac{q}{m_0 \gamma(E)} B_z$$

- The isocyclotron with constant

$$\omega_z = \frac{q}{m_0 \gamma(E)} B_z(r(E))$$

Up to 600 MeV but
this vertically defocuses the beam



- 1938: Thomas proposes strong (transverse) focusing for a cyclotron



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

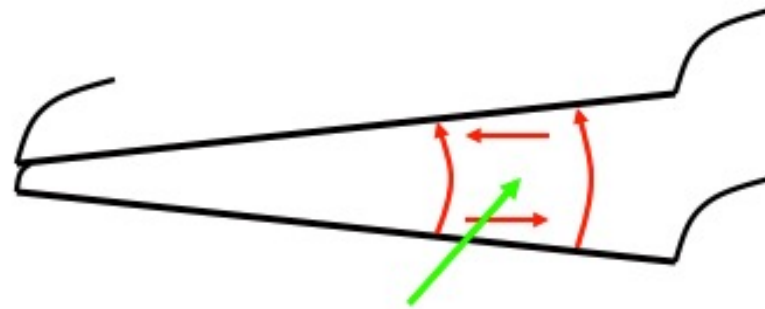
The Isocyclotron

- The isocyclotron with constant

$$\omega_z = \frac{q}{m_0 \gamma(E)} B_z(r(E))$$

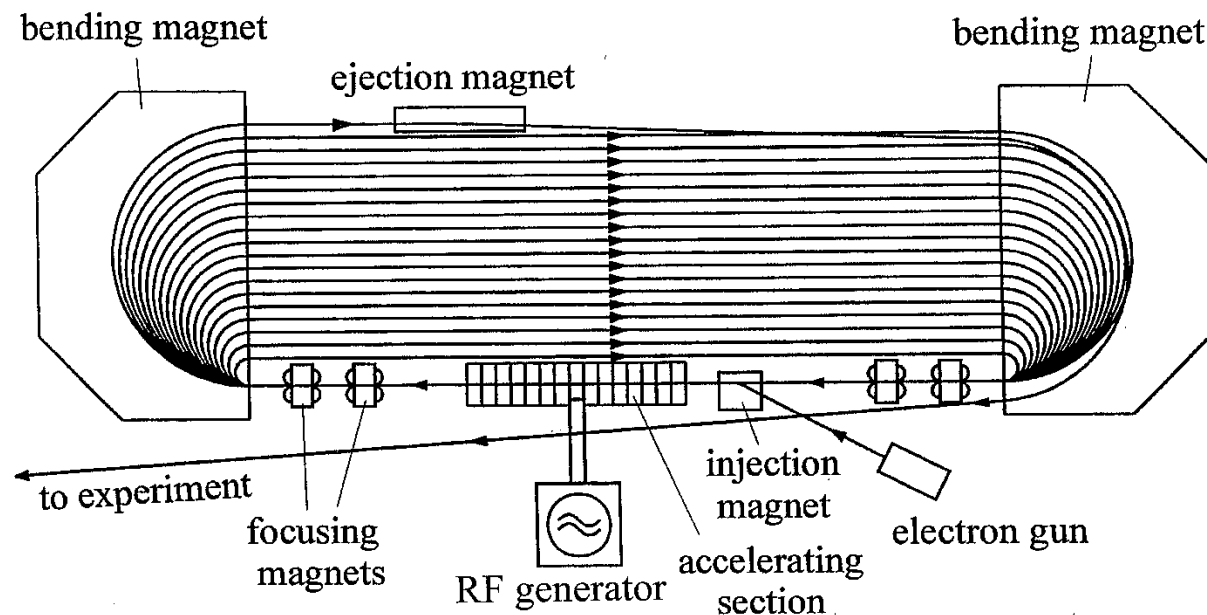
Up to 600MeV but
this vertically defocuses the beam.

Edge focusing is therefore used.



The Microtron

- Electrons are quickly relativistic and cannot be accelerated in a cyclotron.
- In a microtron the revolution frequency changes, but each electron misses an integer number of RF waves.



- Today: Used for medical applications with one magnet and 20MeV.
- Nuclear physics: MAMI designed for 820MeV as race track microtron.



Cornell University



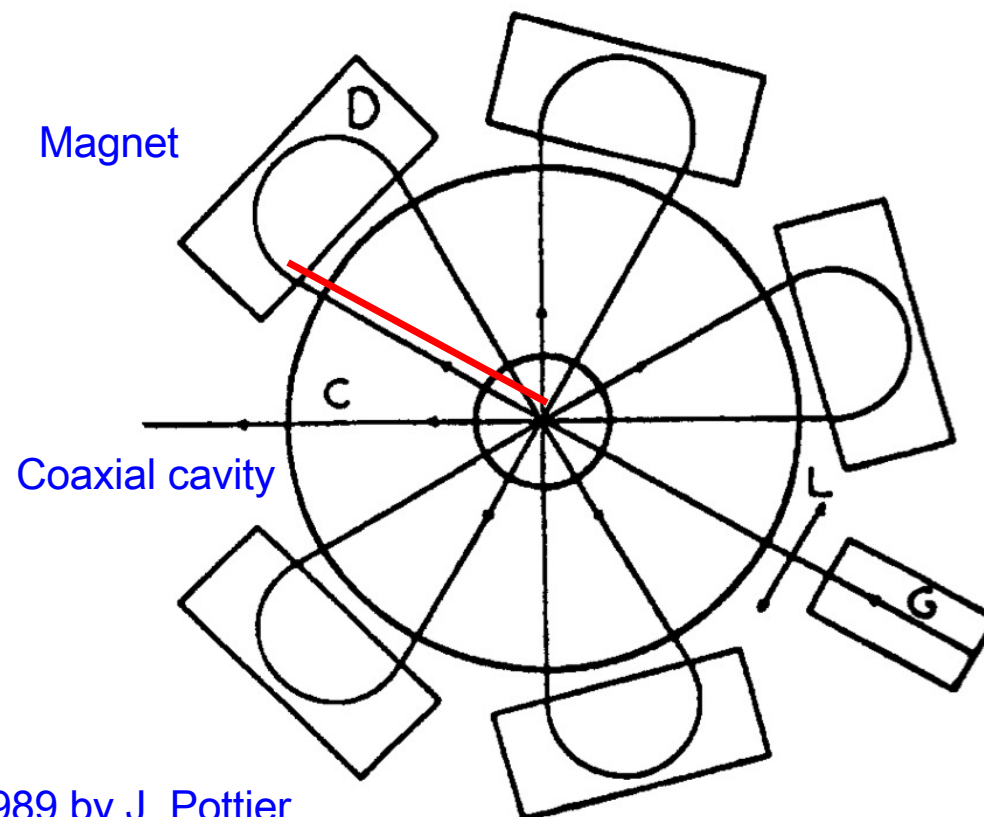
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Rhodotron



- Invented in 1989 by J. Pottier
(French, CEA)



Cornell University



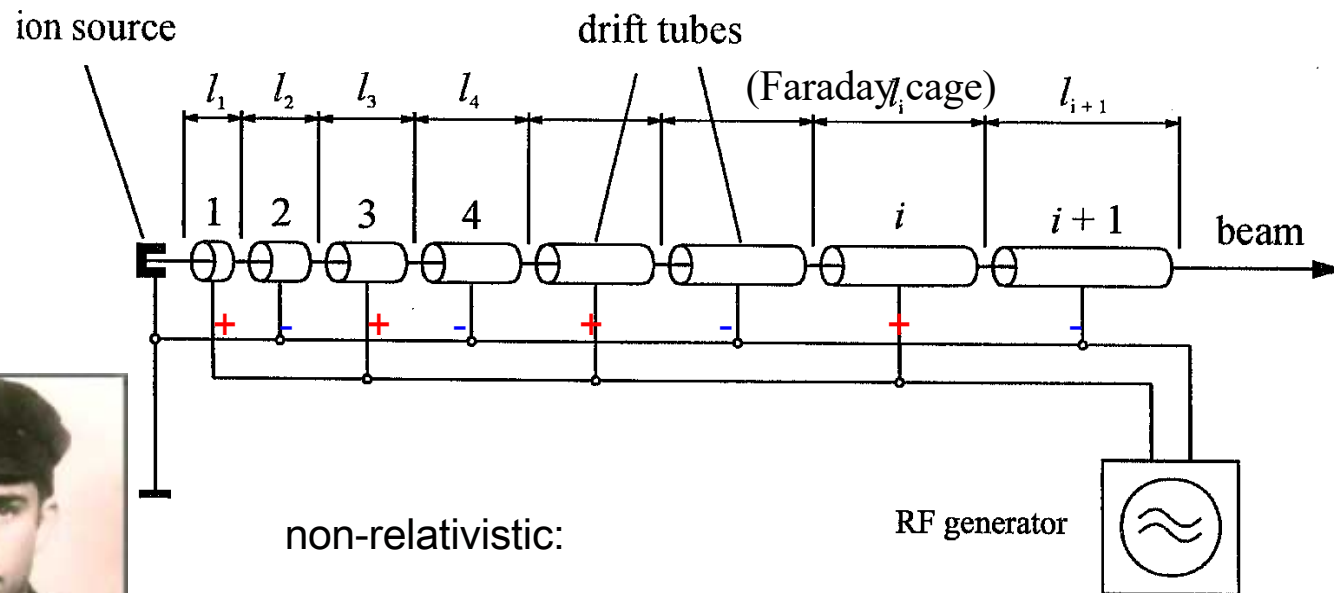
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Wideroe linear accelerator



Wideroe

non-relativistic:

$$K_n = nqU_{\max} \sin \psi_0 = \frac{1}{2} m v_n^2$$

$$l_n = \frac{1}{2} v_n T_{RF} = \frac{1}{2} \beta_n \lambda_{RF} \propto \sqrt{n}$$

Called the π or the $1/2\beta\lambda$ mode



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

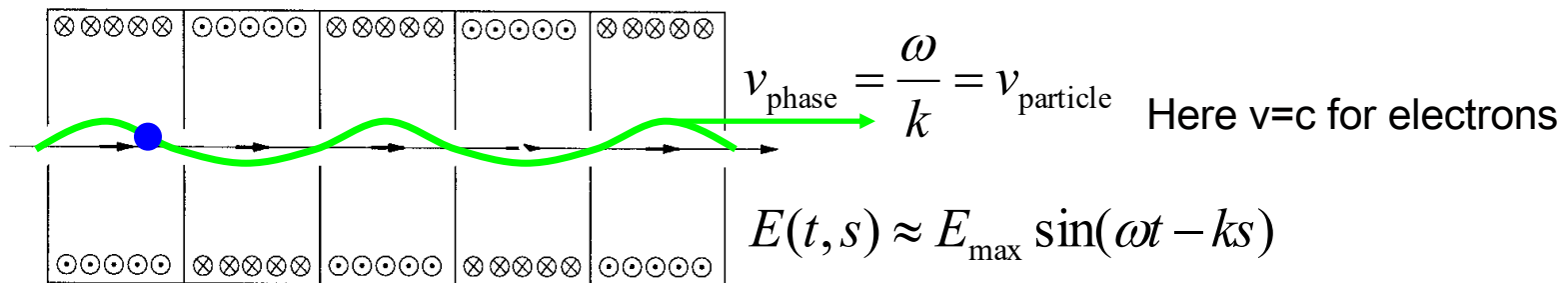
Graduate Accelerator Physics

USPAS summer 2023

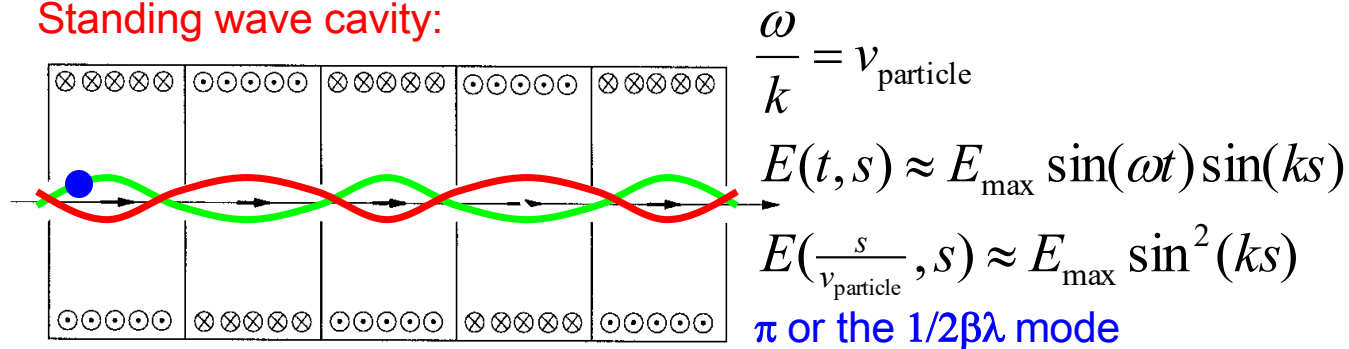
Accelerating Cavities

- 1933: J.W. Beams uses resonant cavities for acceleration

Traveling wave cavity:



Standing wave cavity:



Transit factor (for this example): $\langle E \rangle = \frac{1}{\lambda_{RF}} \int_0^{\lambda_{RF}} E\left(\frac{s}{v_{\text{particle}}}, s\right) ds \approx \frac{1}{2} E_{\text{max}}$



Cornell University



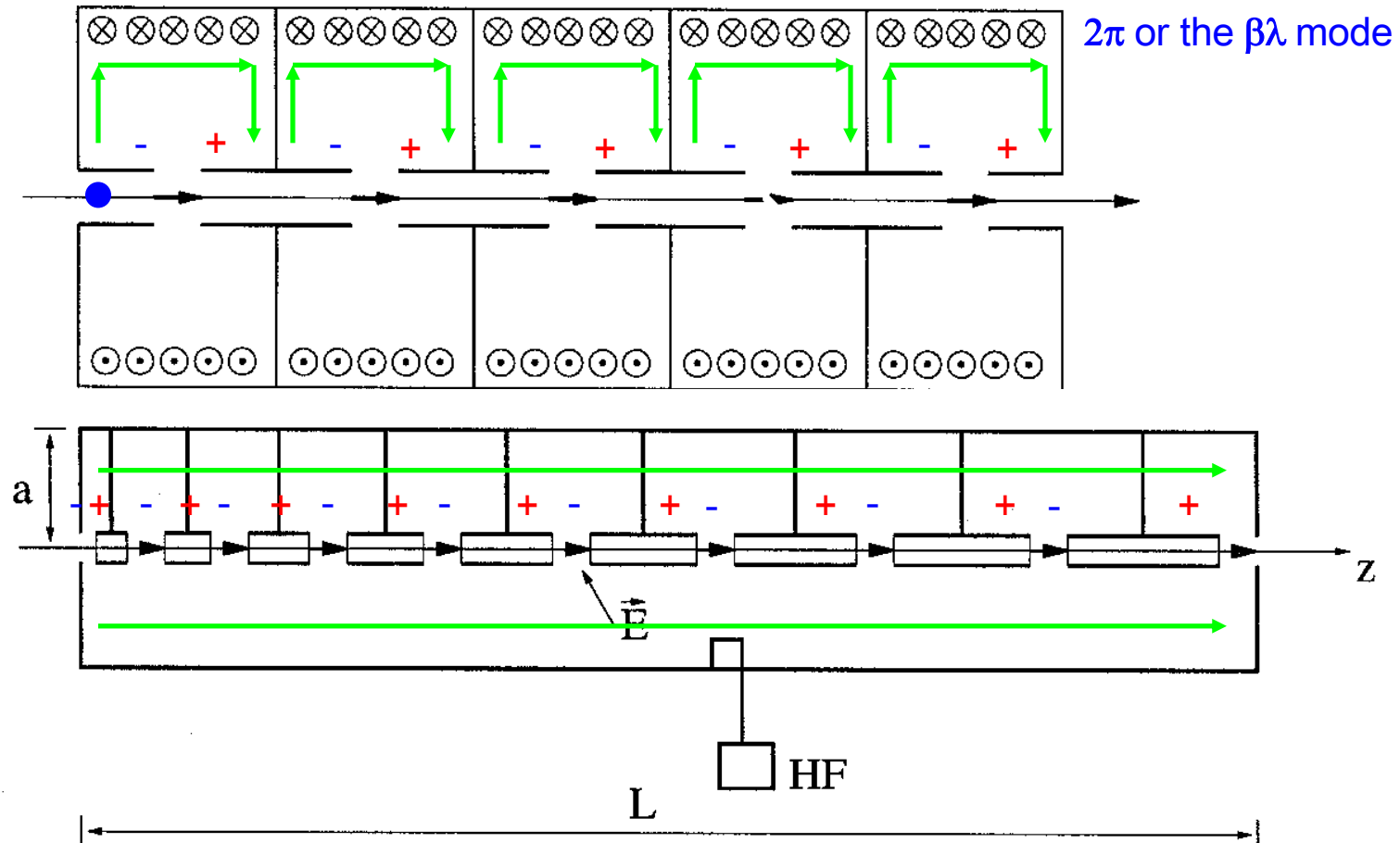
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Alvarez Linear Accelerator



Needs only one power input coupler and walls do not dissipate energy.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

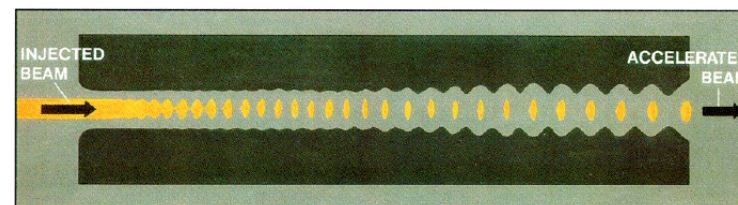
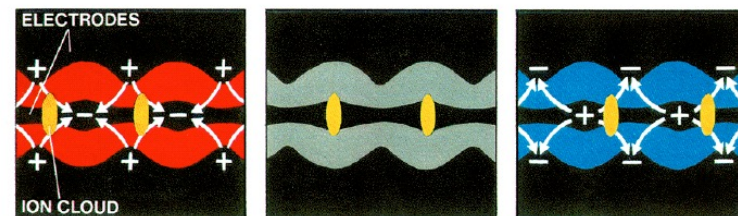
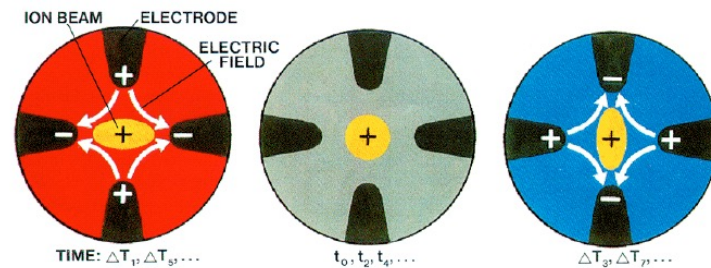
Graduate Accelerator Physics

USPAS summer 2023

Radio Frequency Quadrupoles



- 1970: Kapchinskii and Teplyakov invent the RFQ



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Three historic lines of accelerators

Direct Voltage Accelerators Resonant Accelerators Transformer Accelerator

- 1924: Wideroe invents the betatron
- 1940: Kerst and Serber build a betatron for 2.3MeV electrons and understand betatron (transverse) focusing (in 1942: 20MeV)

Betatron:

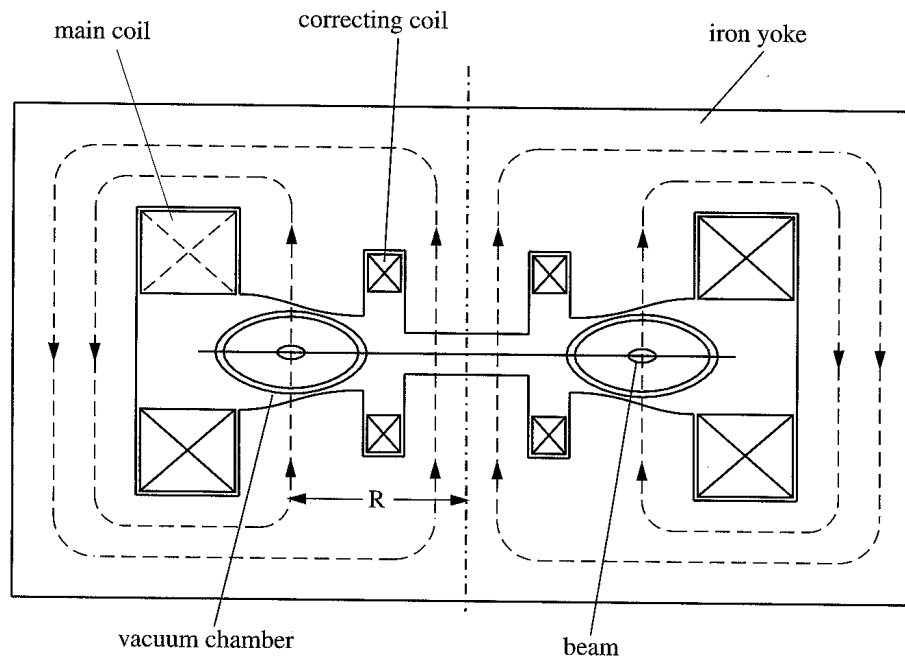
$R=\text{const}, B=B(t)$

Whereas for a cyclotron:

$R(t), B=\text{const}$

No acceleration section is needed since

$$\oint_{\partial A} \vec{E} \cdot d\vec{s} = - \iint_A \frac{d}{dt} \vec{B} \cdot d\vec{a}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Synchrotron

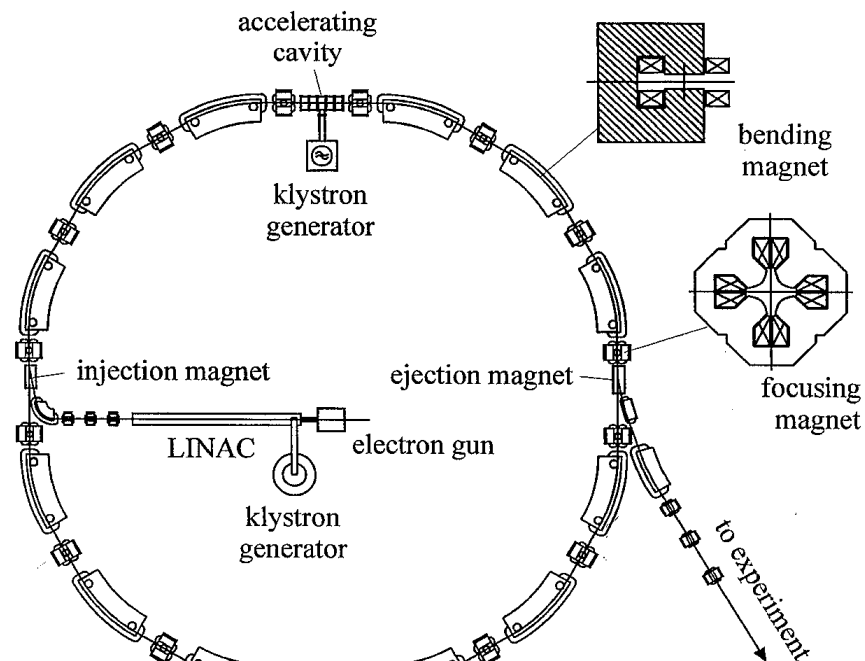
- 1945: Veksler (UDSSR) and McMillan (USA) invent the synchrotron
- 1946: Goward and Barnes build the first synchrotron (using a betatron magnet)
- 1949: Wilson et al. at Cornell are first to store beam in a synchrotron (later 300MeV, magnet of 80 Tons)
- 1949: McMillan builds a 320MeV electron synchrotron

- Many smaller magnets instead of one large magnet
- Only one acceleration section is needed, with

$$R = \frac{p(t)}{qB(R,t)} = \text{const.}$$

$$\omega = 2\pi \frac{v_{\text{particle}}}{L} n$$

for an integer n called the harmonic number



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

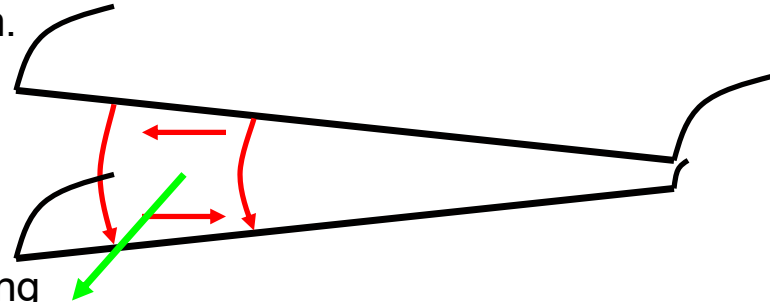
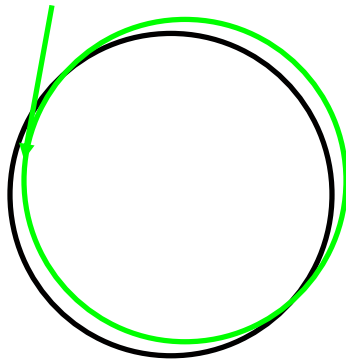
Graduate Accelerator Physics

USPAS summer 2023

The weak focusing synchrotron

- 1952: Operation of the Cosmotron, 3.3 GeV proton synchrotron at Brookhaven: Beam pipe height: 15cm.

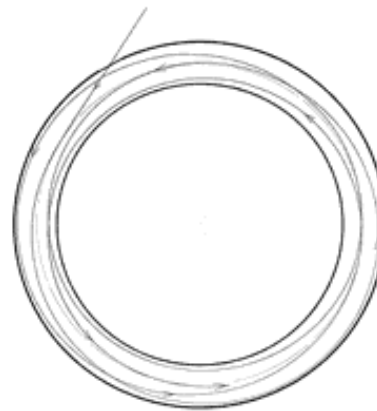
Natural ring focusing:



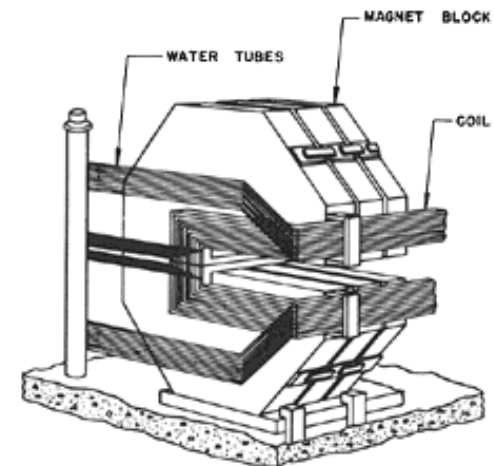
Vertical focusing
+ Horizontal defocusing + ring focusing
Focusing in both planes



The Cosmotron



Weak focusing accelerator



Cornell University



Brookhaven
National Laboratory

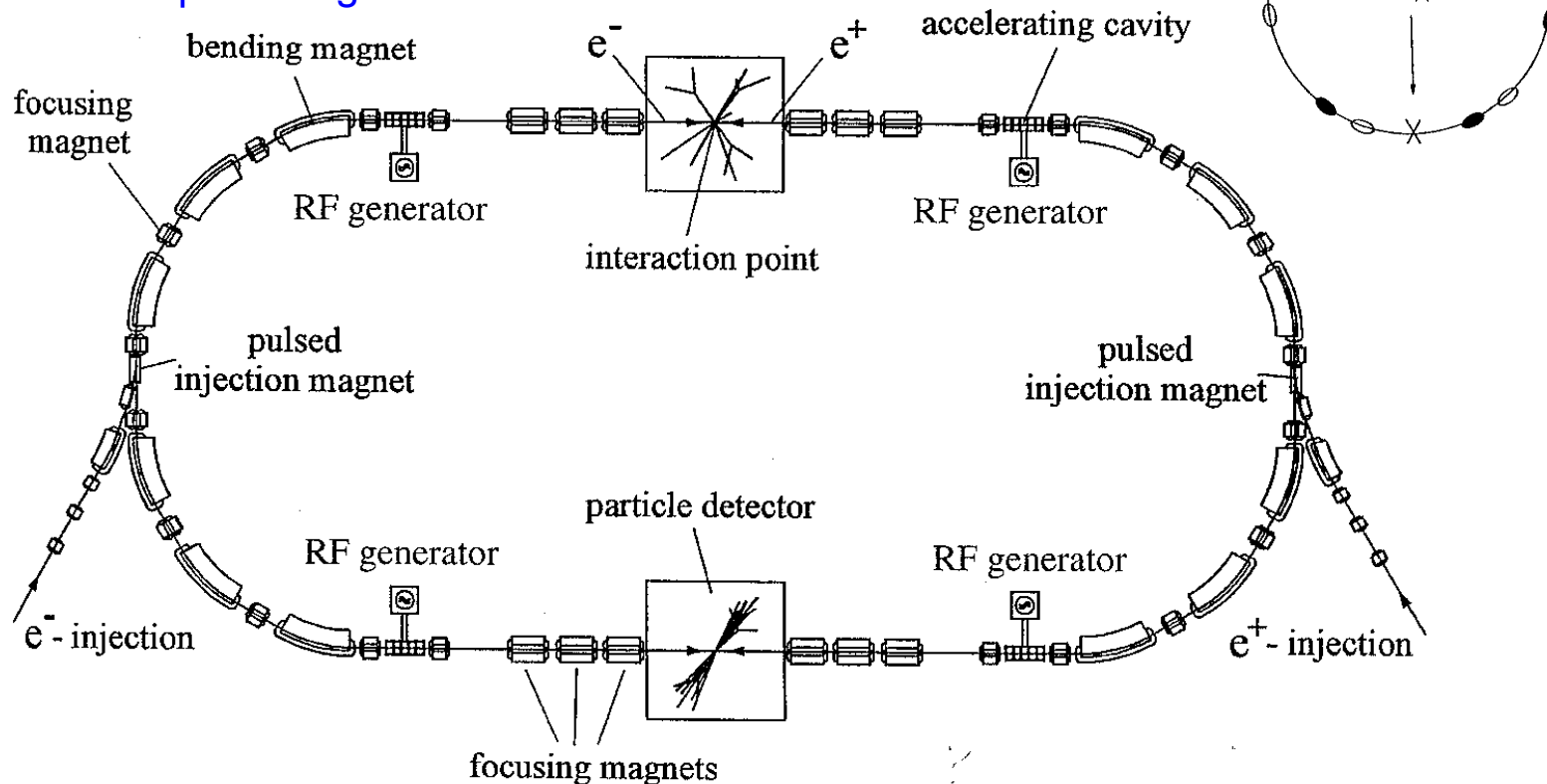
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Elements of a collider

- Saving one beam while injection another
- Avoiding collisions outside the detectors.
- Compensating the forces between e^+ and e^- beams



Cornell University



Brookhaven
National Laboratory

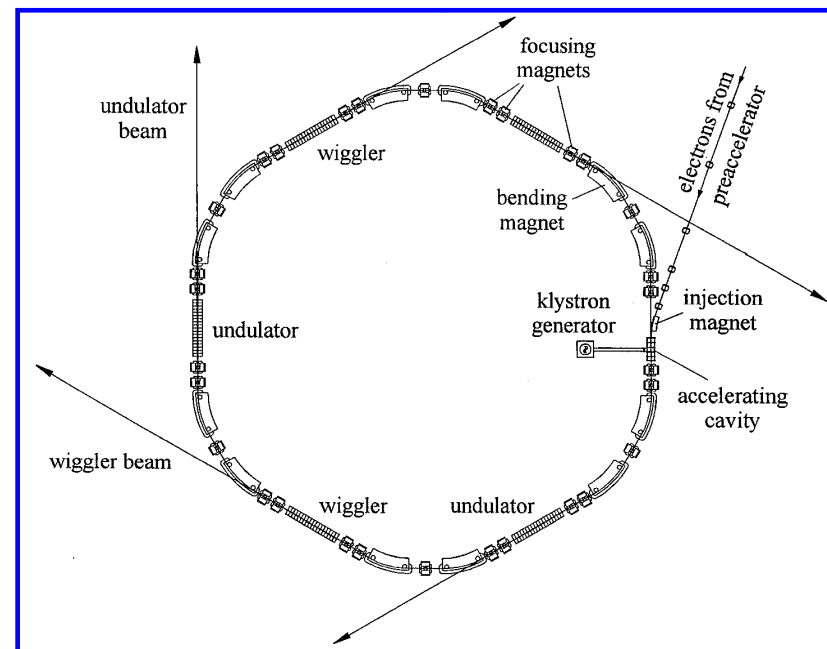
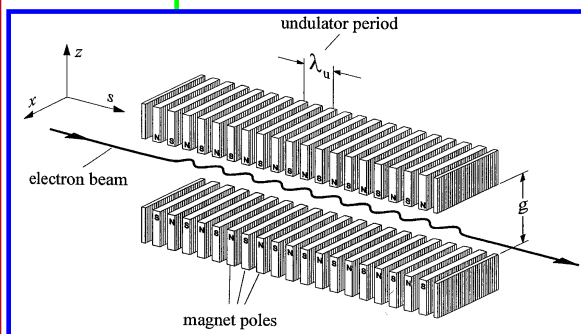
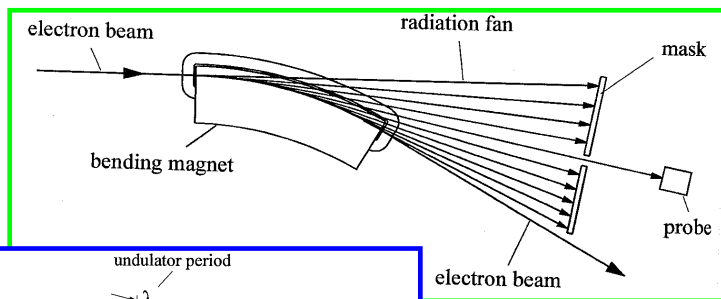
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

4 Generations of Light Sources

- 1st Generation (1970s): Many HEP rings are parasitically used for X-ray production
- 2nd Generation (1980s): Many dedicated X-ray sources (light sources)
- 3rd Generation (1990s): Several rings with dedicated radiation devices (wigglers and undulators)
- Today (4th Generation): Construction of Free Electron Lasers (FELs) driven by LINACs



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

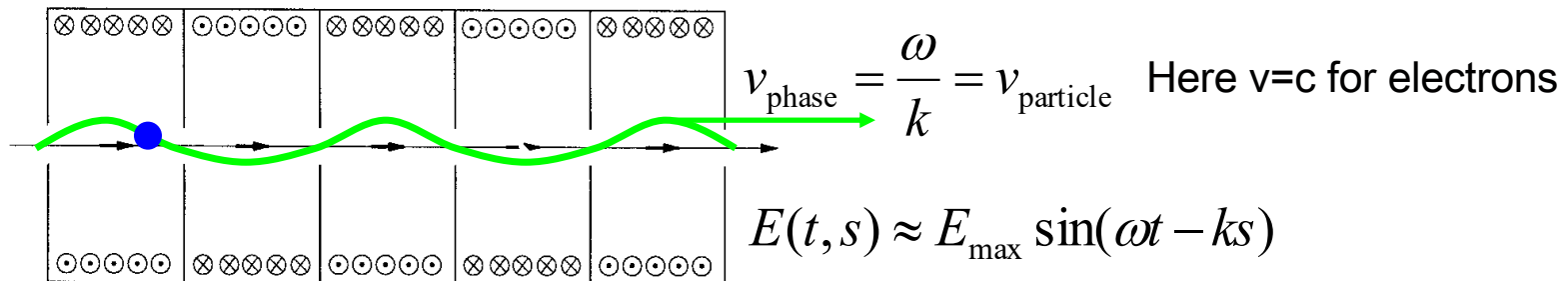
Graduate Accelerator Physics

USPAS summer 2023

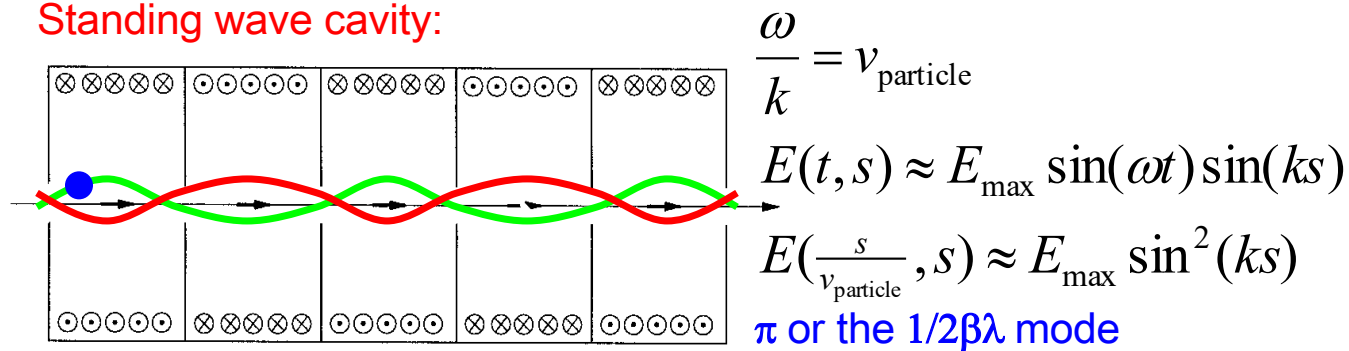
Accelerating cavities

- 1933: J.W. Beams uses resonant cavities for acceleration

Traveling wave cavity:



Standing wave cavity:



Transit factor (for this example): $\langle E \rangle = \frac{1}{\lambda_{RF}} \int_0^{\lambda_{RF}} E(\frac{s}{v_{\text{particle}}}, s) ds = \frac{1}{2} E_{\text{max}}$



Cornell University



Brookhaven
National Laboratory

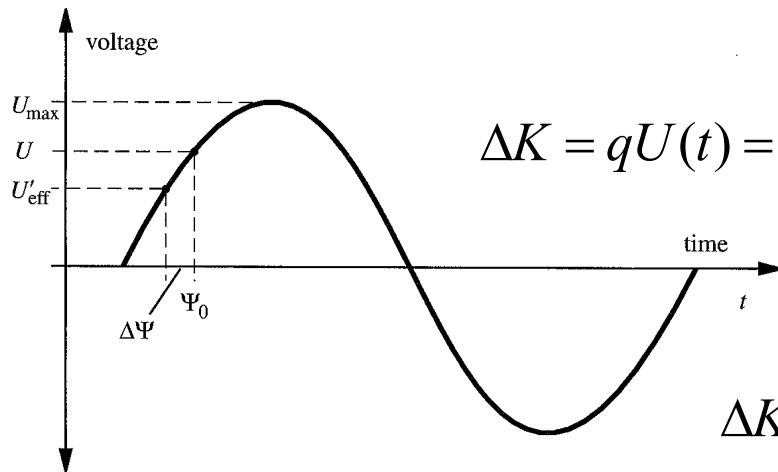
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Phase focusing

- 1945: Veksler (UDSSR) and McMillan (USA) realize the importance of phase focusing



$$\Delta K = qU(t) = qU_{\max} \sin(\omega(t - t_0) + \psi_0)$$

Longitudinal position in the bunch:

$$\sigma = s - s_0 = -v_0(t - t_0)$$

$$\Delta K(\sigma) = qU_{\max} \sin\left(-\frac{\omega}{v_0}(s - s_0) + \psi_0\right)$$

$$\Delta K(0) > 0 \quad (\text{Acceleration})$$

$$\Delta K(\sigma) < \Delta K(0) \text{ for } \sigma > 0 \Rightarrow \frac{d}{d\sigma} \Delta K(\sigma) < 0 \quad (\text{Phase focusing})$$

$$\left. \begin{array}{l} qU(t) > 0 \\ q \frac{d}{dt} U(t) > 0 \end{array} \right\} \underline{\underline{\psi_0 \in (0, \frac{\pi}{2})}}$$

Phase focusing is required in any RF accelerator.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Longitudinal phase space in a Linac

Other particles: $\frac{dE}{ds} = \hat{E} \cos \Phi$ Reference particle: $\frac{dE_0}{ds} = \hat{E} \cos \Phi_0$

$$\phi = \Phi - \Phi_0 = \omega(t - t_0)$$

$$\frac{d\delta}{ds} = \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \approx -\phi \frac{\hat{E}}{E_0} \sin \Phi_0$$

$$\frac{d\phi}{ds} = \omega \left(\frac{1}{v} - \frac{1}{v_0} \right) \approx \omega \left(\frac{1}{v_0 + \left. \frac{dv}{d\delta} \right|_0 \delta} - \frac{1}{v_0} \right) \approx -\omega \frac{\left. \frac{dv}{d\delta} \right|_0}{v_0^2} \delta = -\omega \frac{c^2}{v_0^3 \gamma_0^2} \delta$$

$$\frac{dv}{d\delta} = E \frac{dv}{dE} = \gamma \frac{dv}{d\gamma} = c\gamma \frac{d\sqrt{1 - \frac{1}{\gamma^2}}}{d\gamma} = \frac{c}{\gamma^2 \beta}$$

$$\frac{d^2\phi}{ds^2} \approx -\omega \frac{c^2}{v_0^3 \gamma_0^2} \frac{d\delta}{ds} \approx \frac{\hat{E}}{E_0} \sin \Phi_0 \omega \frac{c^2}{v_0^3 \gamma_0^2} \phi$$

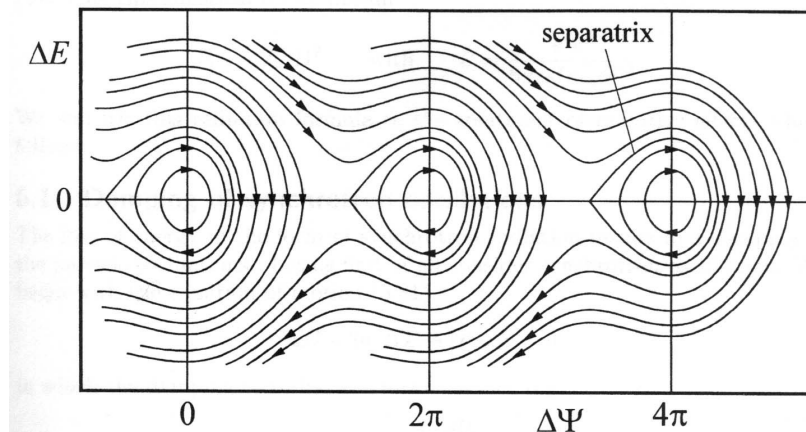
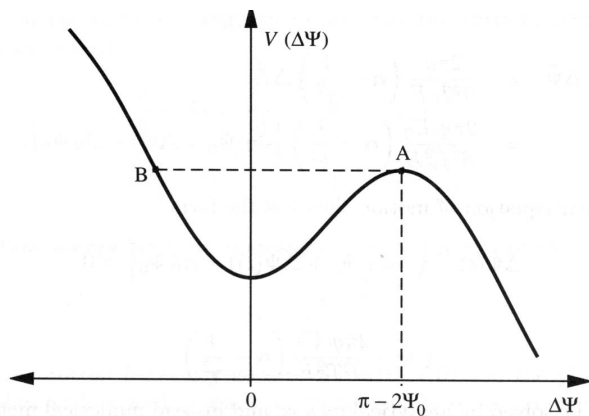
Stability for small phases when the factor on the right-hand side is negative.

For not very small phases one cannot linearize.

$$\frac{d\delta}{ds} = \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \quad \frac{d\phi}{ds} \approx -\omega \frac{c^2}{v_0^3 \gamma_0^2} \delta$$

$$H(\phi, \delta) = -\frac{q\bar{E}_s}{K_0} (\sin(\Phi_0 + \phi) - \phi \cos \Phi_0) - \omega \frac{c^2}{v_0^3 \gamma_0^2} \delta^2$$

$$\frac{d}{dt} \phi = \frac{\partial}{\partial \delta} H, \quad \frac{d}{dt} \delta = -\frac{\partial}{\partial \phi} H$$

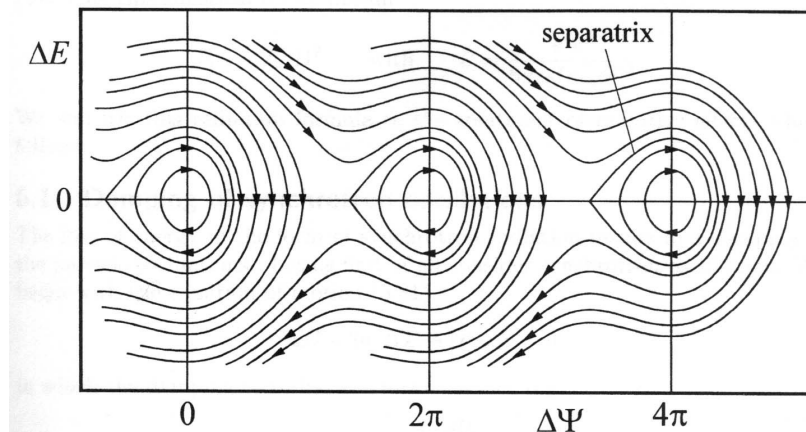
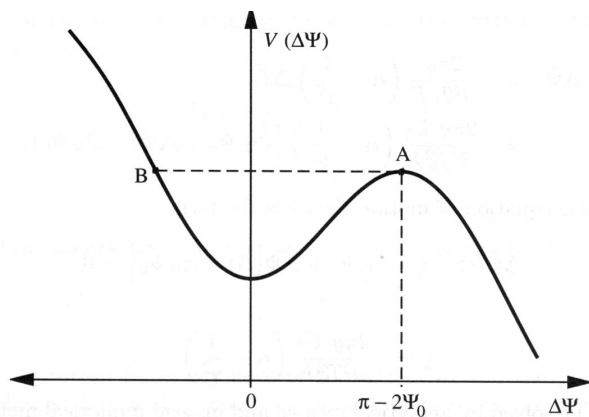


Cornell

$$\frac{d\delta}{ds} = \frac{q\overline{E}_s}{K_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \approx -\phi \frac{q\overline{E}_s}{K_0} \sin \Phi_0$$

$$\begin{aligned} \frac{d^2\phi}{ds^2} &\approx -\omega \frac{c^2}{v_0^3 \gamma_0^2} \frac{d\delta}{ds} = -\omega \frac{c^2}{v_0^3 \gamma_0^2} \frac{q\overline{E}_s}{K_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \\ &= -\frac{d}{d\phi} \omega \frac{c^2}{v_0^3 \gamma_0^2} \frac{q\overline{E}_s}{K_0} (\sin(\Phi_0 + \phi) - \phi \cos \Phi_0) \end{aligned}$$

Effective potential



Cornell

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS Summer 2023
Georg.Hoffstaetter@Cornell.edu

Longitudinal phase space in rings

Other particles: $\frac{dE}{ds} = \hat{E} \cos \Phi$

Reference particle: $\frac{dE_0}{ds} = \hat{E} \cos \Phi_0$

$$\phi = \Phi - \Phi_0 = \omega(t - t_0)$$

$$\frac{d\delta}{ds} = \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \approx -\phi \frac{\hat{E}}{E_0} \sin \Phi_0$$

Momentum compaction $\alpha = \frac{p}{L} \frac{dL}{dp}$, closed orbit length change with momentum.

$$\frac{d\phi}{ds} = \omega(t - t_0) = \omega \left(\frac{L}{v} - \frac{L_0}{v_0} \right) = \omega \left(\frac{L_0}{v} + \alpha \frac{L_0}{p_0 v_0} dp - \frac{L_0}{v_0} \right) = \frac{\omega L_0}{v_0} \left(\alpha \frac{dp}{p_0} - \frac{dv}{v_0} \right) = \frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \delta$$

Stability for small phases when the factor on the right-hand side is negative.

$$\frac{d^2\phi}{ds^2} = -\phi \frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \sin \phi_0$$

Transition energy and phase focusing

Stability for small phases when the factor on the right-hand side is negative.

$$\frac{d^2\phi}{ds^2} = -\phi \frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \sin\phi_0$$

For only natural ring focusing, $\rho \propto p \Rightarrow \alpha = 1$. For stronger focusing, $\alpha < 1$.

For small γ , $\alpha - \frac{1}{\gamma^2} < 0$ and $\phi_0 \in [0, \pi]$ leads to a stable phase.

For large γ , $\alpha - \frac{1}{\gamma^2} < 0$ and $\phi_0 \in [\pi, 2\pi]$ leads to a stable phase.

At the transition energy, one must jump the RF phase by π . This jump does not have to happen very quickly, because at that energy, phases hardly change because of

$$\frac{d^2\phi}{ds^2} \approx 0.$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

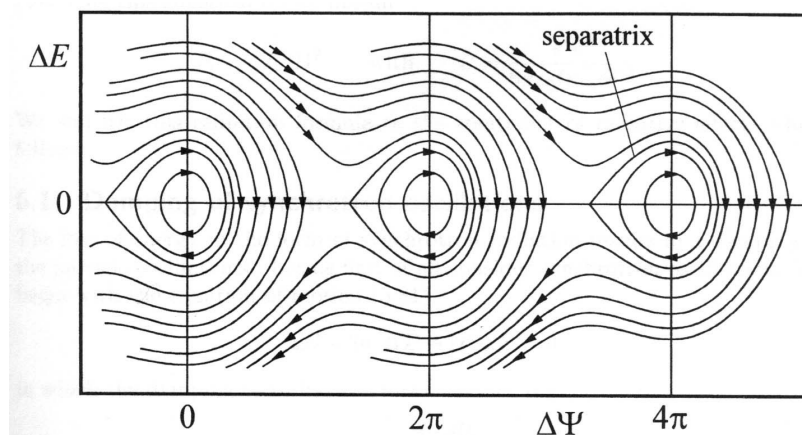
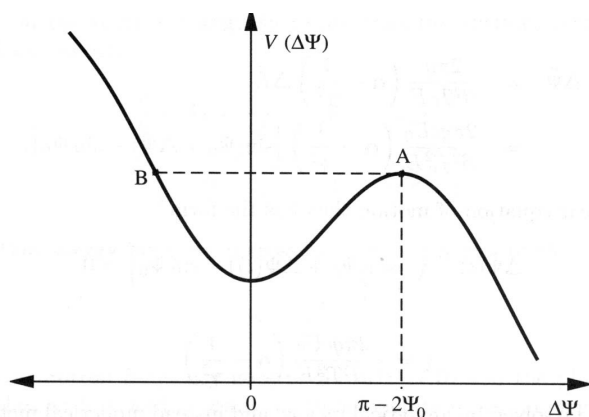
Accelerating longitudinal phase space

For not very small phases one cannot linearize.

$$\frac{d\delta}{ds} = \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \quad \frac{d\phi}{ds} = -\frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \delta$$

$$H(\phi, \delta) = -\frac{qE}{E} (\sin(\Phi_0 + \phi) - \phi \cos(\Phi_0)) - \frac{\omega L_0}{2 v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \delta^2$$

$$\frac{d}{dt} \phi = \frac{\partial}{\partial \delta} H, \quad \frac{d}{dt} \delta = -\frac{\partial}{\partial \phi} H$$



Cornell

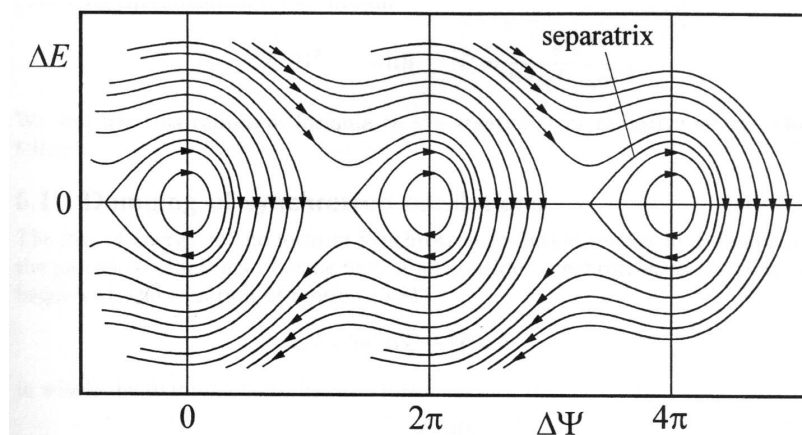
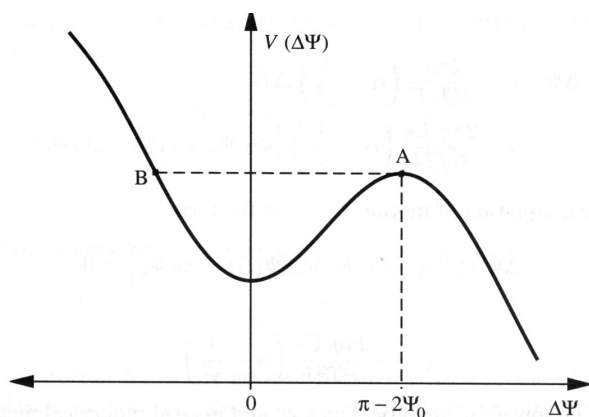
The effective accelerating potential

$$\frac{d\delta}{ds} = \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos \Phi_0) \quad \frac{d\phi}{ds} = -\frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \delta$$

$$\frac{d^2\phi}{ds^2} = -\frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \frac{\hat{E}}{E_0} (\cos(\Phi_0 + \phi) - \cos(\Phi_0))$$

$$\frac{d^2\phi}{ds^2} = -\frac{\partial}{\partial \phi} \frac{\omega L_0}{v_0 \beta_0^2} \left(\alpha - \frac{1}{\gamma^2} \right) \frac{\hat{E}}{E_0} (\sin(\Phi_0 + \phi) - \phi \cos(\Phi_0))$$

Effective potential



Cornell

Transport maps of cavities

(1) Linearization: $E_r(r, z, t) = -\frac{r}{2} \partial_z E_z(0, z, t) \Rightarrow \vec{\nabla} \cdot \vec{E} = 0$

$$B_\varphi(r, z, t) = \frac{1}{c^2} \frac{r}{2} \partial_t E_z(0, z, t) \Rightarrow \vec{\nabla} \times \vec{B} = \frac{1}{c^2} \partial_t \vec{E}$$

(2) Equation of motion:

$$a = \frac{p_x}{p_0}$$

$$a' = \frac{1}{p_0 v} (F_x - a F_z) = -\frac{q}{p_0 v} \left[\frac{r}{2} (\partial_z + \frac{v}{c^2} \partial_t) E_z + a E_z \right]$$

$$= -\frac{q}{p_0 v} \left[\frac{r}{2} \left(\frac{d}{dz} - \frac{1}{v} \left(1 - \frac{v^2}{c^2} \right) \partial_t \right) E_z + a E_z \right] \approx -\frac{1}{p_0} \left[r \frac{1}{2} p_0'' + a p_0' \right]$$

$$u = r \sqrt{p} \quad \text{p denotes } p_0 \text{ for simplicity}$$

$$u' = a \sqrt{p} + r \sqrt{p} \frac{p'}{2p}$$

Focusing !

$$u'' \approx -\frac{1}{\sqrt{p}} \left(r \frac{1}{2} p'' + a p' \right) + a \sqrt{p} \frac{p'}{p} + r \left(\frac{p''}{2\sqrt{p}} - \frac{p'}{4\sqrt{p}^3} \right) = -u \left(\frac{p'}{2p} \right)^2$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Transport maps of cavities

(3) Average focusing over one period with relatively little energy change:

$$u'' \approx -u \frac{\Delta^2/4}{p^2}, \quad \Delta = \sqrt{\langle p'^2 \rangle}$$

(4) Continuous energy change:

$$p' \approx \Omega, \quad \Omega = \langle p' \rangle$$

$$\frac{d^2}{dp^2} u \approx \frac{1}{\Omega^2} u'' \approx -u \frac{(\Delta/\Omega)^2}{4p^2}$$

$$\frac{d^2}{dp^2} (r \sqrt{p}) = \frac{d^2}{dp^2} r \sqrt{p} + \frac{d}{dp} r \frac{1}{\sqrt{p}} - r \frac{1}{4\sqrt{p}^3} \approx -r \frac{(\Delta/\Omega)^2}{4\sqrt{p}^3}$$

$$\frac{d^2}{dp^2} r + \frac{d}{dp} r \frac{1}{p} \approx -r \frac{(\Delta/\Omega)^2 - 1}{4p^2} = -r \frac{\varepsilon^2}{p^2}$$

$$r(p) = \eta(-\ln(p)) \quad \Rightarrow \quad \frac{d^2}{dp^2} r = \frac{1}{p^2} \eta' + \frac{1}{p^2} \eta'' = \frac{1}{p^2} \eta' - \eta \frac{\varepsilon^2}{p^2}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Transport maps of traveling wave cavities

$$\eta'' = -\varepsilon^2 \eta, \quad \eta(\xi) = A \cos(\varepsilon \xi) - B \sin(\varepsilon \xi)$$

$$\begin{pmatrix} r \\ a \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & \frac{p'}{p} \end{pmatrix} \begin{pmatrix} \cos(\varepsilon \ln(p)) & \sin(\varepsilon \ln(p)) \\ -\sin(\varepsilon \ln(p)) & \cos(\varepsilon \ln(p)) \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix}$$

$$\begin{pmatrix} r \\ a \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & \frac{p'}{p} \end{pmatrix} \begin{pmatrix} \cos(\varepsilon \ln(\frac{p}{p_i})) & \sin(\varepsilon \ln(\frac{p}{p_i})) \\ -\sin(\varepsilon \ln(\frac{p}{p_i})) & \cos(\varepsilon \ln(\frac{p}{p_i})) \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & \frac{p_i}{p'} \end{pmatrix} \begin{pmatrix} r_i \\ a_i \end{pmatrix}$$

$$E_z = \sum_{n=0}^{\infty} g_n \cos(n \frac{2\pi}{L} z + \alpha_n) \cos(\omega t - kz + \varphi_0)$$

$$= \sum_{n=0}^{\infty} g_n \cos(n \frac{2\pi}{L} z + \alpha_n) \cos(\varphi_0), \quad \alpha_0 = 0$$

$$\langle p' \rangle = g_0 \cos(\varphi_0), \quad \langle p'^2 \rangle = \left(g_0^2 + \frac{1}{2} \sum_{n=1}^{\infty} g_n^2 \right) \cos^2(\varphi_0) \Rightarrow \varepsilon = \frac{1}{2} \sqrt{\frac{1}{2} \sum_{n=1}^{\infty} \frac{g_n^2}{g_0^2}}$$



Transport maps of standing wave cavities

$$\begin{pmatrix} r \\ a \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & \frac{p'}{p} \end{pmatrix} \begin{pmatrix} \cos(\varepsilon \ln(\frac{p}{p_i})) & \sin(\varepsilon \ln(\frac{p}{p_i})) \\ -\sin(\varepsilon \ln(\frac{p}{p_i})) & \cos(\varepsilon \ln(\frac{p}{p_i})) \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & \frac{p_i}{p'} \end{pmatrix} \begin{pmatrix} r_i \\ a_i \end{pmatrix}$$

$$E_z = \sum_{n=-\infty}^{\infty} g_n e^{in\frac{2\pi}{L}z} \cos(kz) \cos(\omega t + \varphi_0) , \quad \omega t = kz , k = m \frac{\pi}{L} , \quad g_{-n} = g_n^*$$

$$= \frac{1}{4} \sum_{n=-\infty}^{\infty} g_n [e^{i(n+m)\frac{2\pi}{L}z + \varphi_0} + e^{i(n-m)\frac{2\pi}{L}z - \varphi_0} + e^{in\frac{2\pi}{L}z + \varphi_0} + e^{in\frac{2\pi}{L}z - \varphi_0}]$$

$$= \frac{1}{4} \sum_{n=-\infty}^{\infty} [g_{n-m} e^{i\varphi_0} + g_{n+m} e^{-i\varphi_0} + 2g_n \cos(\varphi_0)] e^{in\frac{2\pi}{L}z} = \sum_{n=-\infty}^{\infty} f_n e^{in\frac{2\pi}{L}z}$$

$$\left. \begin{aligned} \langle p' \rangle &= f_0 , \\ \langle p'^2 \rangle &= \sum_{n=0}^{\infty} |f_n|^2 \end{aligned} \right\} \varepsilon = \frac{1}{2} \sqrt{\sum_{n=0}^{\infty} \left| \frac{f_n}{f_0} \right|^2 - 1}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Phase space preservation in cavities

Average focusing over one period with relatively little energy change:

$$\begin{pmatrix} r \\ a \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 \\ 0 & \frac{p'}{p} \end{pmatrix} \begin{pmatrix} \cos(\varepsilon \ln(\frac{p}{p_i})) & \sin(\varepsilon \ln(\frac{p}{p_i})) \\ -\sin(\varepsilon \ln(\frac{p}{p_i})) & \cos(\varepsilon \ln(\frac{p}{p_i})) \end{pmatrix}}_{\underline{M}} \begin{pmatrix} 1 & 0 \\ 0 & \frac{p_i}{p'} \end{pmatrix} \begin{pmatrix} r_i \\ a_i \end{pmatrix}$$

$$\det(\underline{M}) = \frac{p_i}{p}$$

Because the determinant is not 1, the phase space volume is no longer conserved but changes by p_0/p .

A new propagation and definition of Twiss parameters is therefore needed:

$$r = \sqrt{2J \frac{1}{\beta_r \gamma_r}} \beta \sin(\psi + \phi_0)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Twiss parameters in accelerating cavities

$$\alpha = -\frac{1}{2} \beta', \quad \gamma = \frac{1+\alpha^2}{\beta}$$

$$a = r' = \sqrt{2J \frac{mc}{p}} \left[-\frac{2\alpha + \beta \frac{p'}{p}}{2\sqrt{\beta}} \sin(\psi + \phi_0) + \frac{\beta \psi'}{\sqrt{\beta}} \cos(\psi + \phi_0) \right]$$

$$a' \approx -\frac{1}{p} \left[r(pK + \frac{1}{2} p'') + ap' \right]$$

Quad + cavities

$$a' = -\sqrt{2J \frac{mc}{\beta p}} \begin{pmatrix} \frac{(\beta \psi')^2 + \alpha^2}{\beta} + \alpha' - \alpha \frac{p'}{p} + \beta \frac{p''}{2p} - \beta \frac{3p'^2}{4p^2} \\ 2\alpha \psi' + \beta \frac{p'}{p} \psi' - \beta \psi'' \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix}$$

$$= -\sqrt{2J \frac{mc}{\beta p}} \begin{pmatrix} \beta(K + \frac{1}{2} \frac{p''}{p}) - (\alpha + \beta \frac{p'}{2p}) \frac{p'}{p} \\ \beta \frac{p'}{p} \psi' \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix}$$

$$\Rightarrow \psi' = \frac{A}{\beta}, \text{ choice: } A = 1 \quad \left\{ \begin{pmatrix} r \\ a \end{pmatrix} = \sqrt{2J \frac{mc}{p}} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha + \beta \frac{p'}{2p}}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix} \right.$$

$$\alpha' + \gamma = \beta \left[K + \left(\frac{p'}{2p} \right)^2 \right]$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Beta functions in accelerating cavities

$$\begin{pmatrix} r \\ a \end{pmatrix} = \sqrt{2J_n \frac{mc}{p}} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\tilde{\alpha}}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \cos(\psi + \phi_0) \\ \sin(\psi + \phi_0) \end{pmatrix}, \quad \tilde{\alpha} = \alpha + \beta \frac{p'}{2p}$$

For systems with changing energy one uses the normalized Courant-Snyder invariant $J_n = J \mathbf{b}_r \mathbf{g}_r$

$$\begin{pmatrix} r & a \end{pmatrix} \begin{pmatrix} \frac{1}{\sqrt{\beta}} & \frac{\tilde{\alpha}}{\sqrt{\beta}} \\ 0 & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\tilde{\alpha}}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} r \\ a \end{pmatrix} \frac{p}{2mc} = \begin{pmatrix} r & a \end{pmatrix} \begin{pmatrix} \frac{1+\tilde{\alpha}^2}{\beta} & \tilde{\alpha} \\ \tilde{\alpha} & \beta \end{pmatrix} \begin{pmatrix} r \\ a \end{pmatrix} \frac{p}{2mc} = J_n$$

Reasons:

- (1) J is the phase space amplitude of a particle in (x, a) phase space, which is the area in phase space (over 2π) that its coordinate would circumscribe during many turns in a ring. However, $a = p_x/p_0$ is not conserved when p_0 changes in a cavity. Therefore J is not conserved.
- (2) $J_n = J p_0/mc$ is therefore proportional to the corresponding area in (x, p_x) phase space, and is thus conserved.



Cornell University

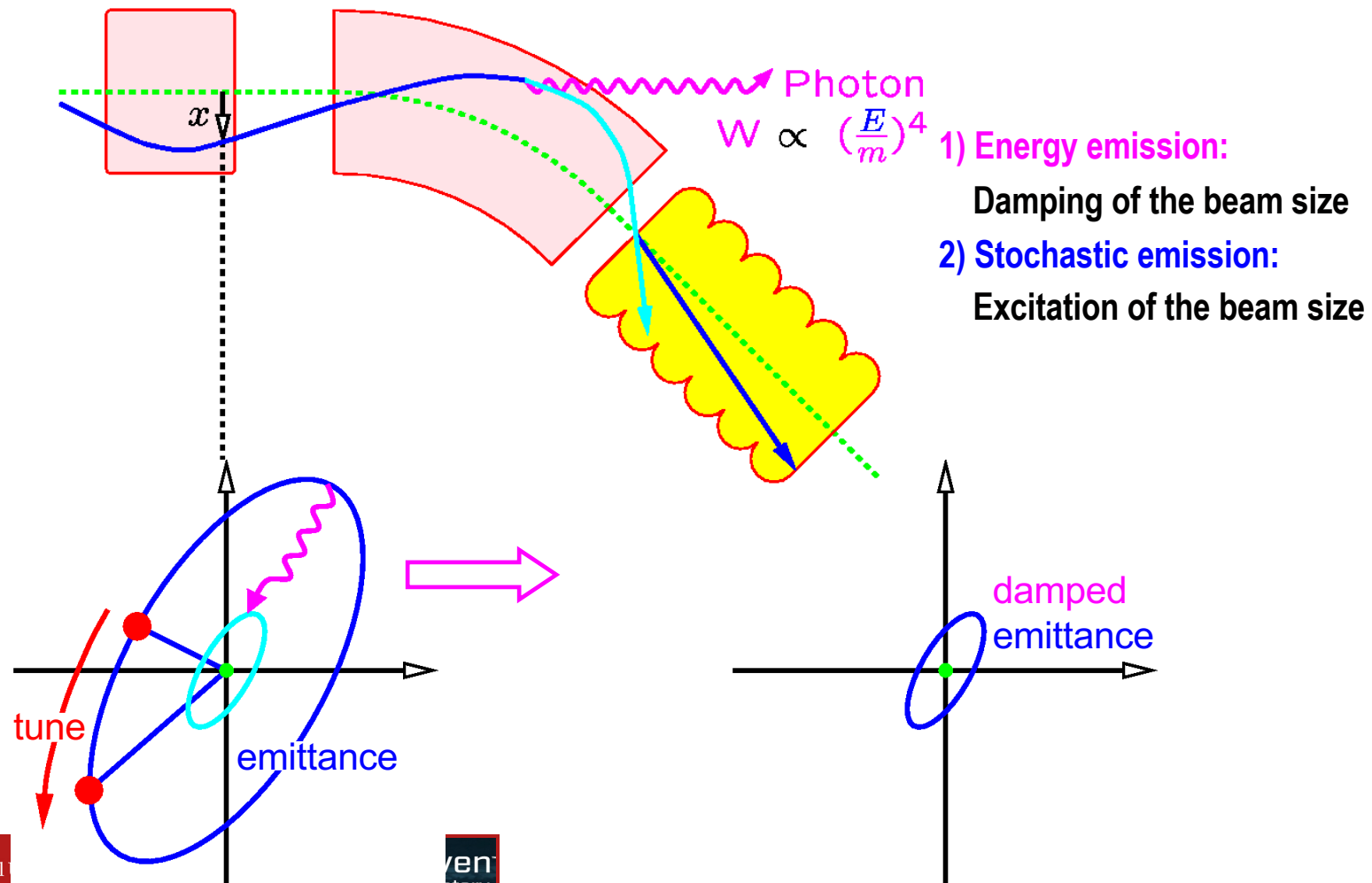


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Radiative damping of the transverse emittance



Cornell

SLAC

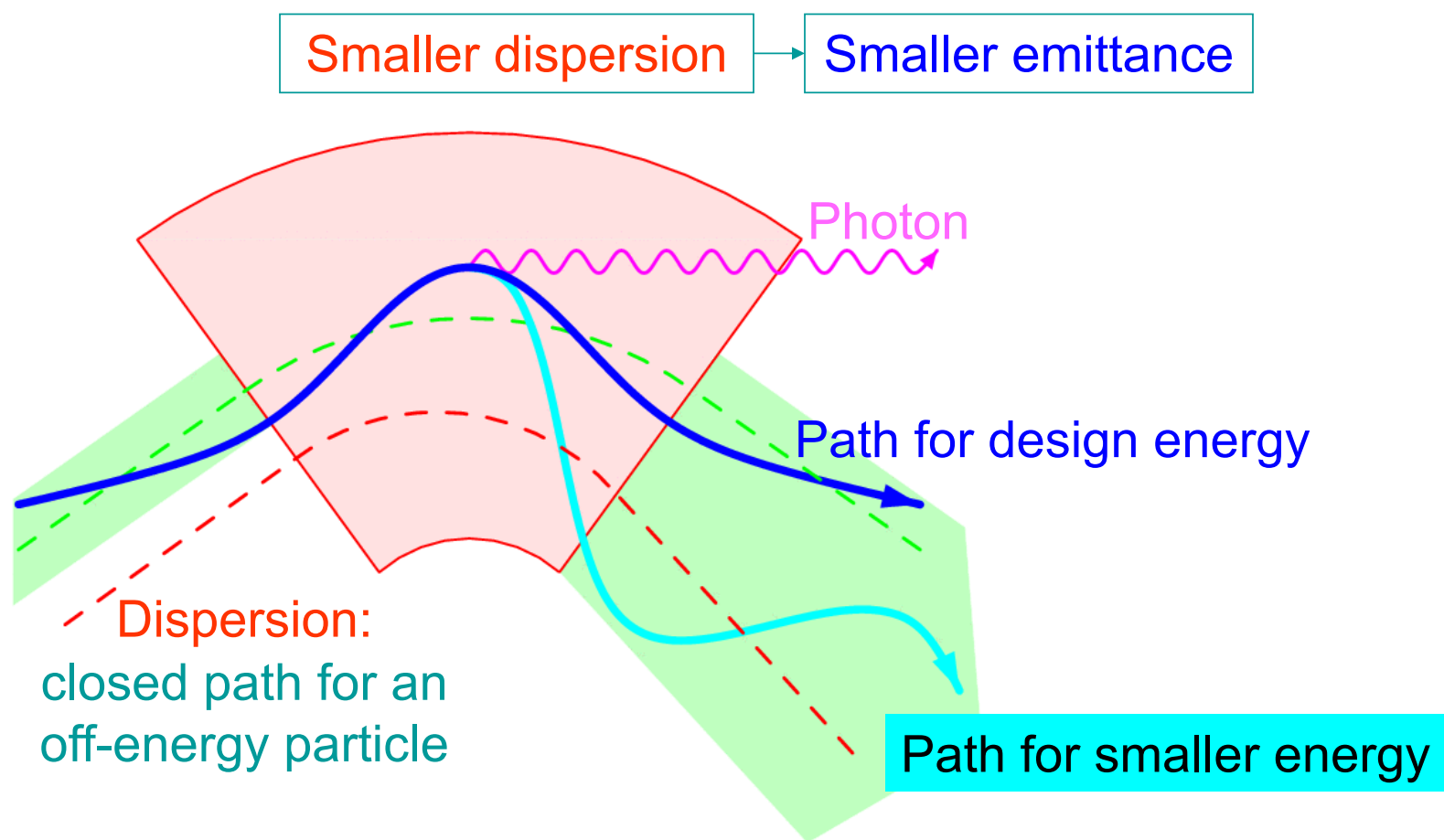
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Radiative excitation of the transverse emittance



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023
Georg.Hoffstaetter@Cornell.edu

Radiative damping of the longitudinal emittance

- 1) **Energy emission:** Particles with larger energy radiate more, leading them closer to the average energy.
- 2) **Stochastic emission:** Random noise in the energy of emitted photons lead to an energy spread.



The equilibrium of the two effects leads to an equilibrium longitudinal emittance.

The damping time is the time it takes to radiate off the energy of the beam, while it is kept at constant energy with RF cavities. It is usually a few 100 revolutions.

During this damping time the beam forgets its history, particle coordinates are reshuffled within the beam.



Cornell University



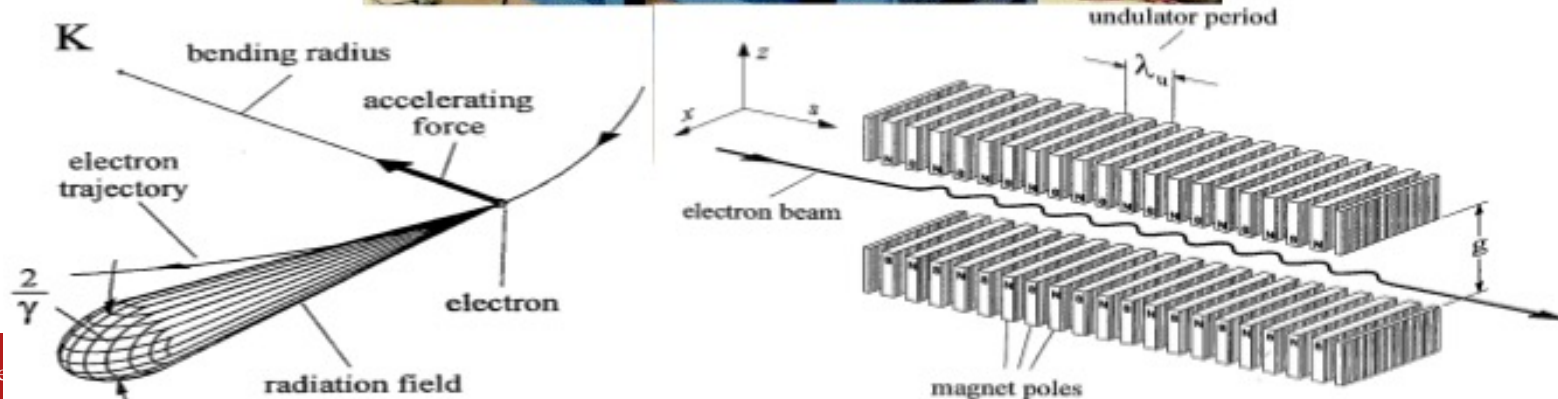
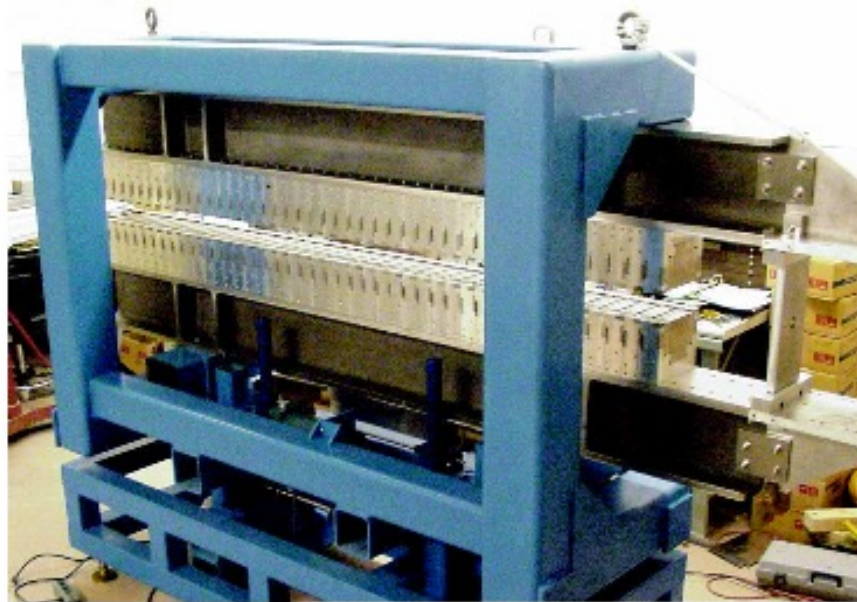
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

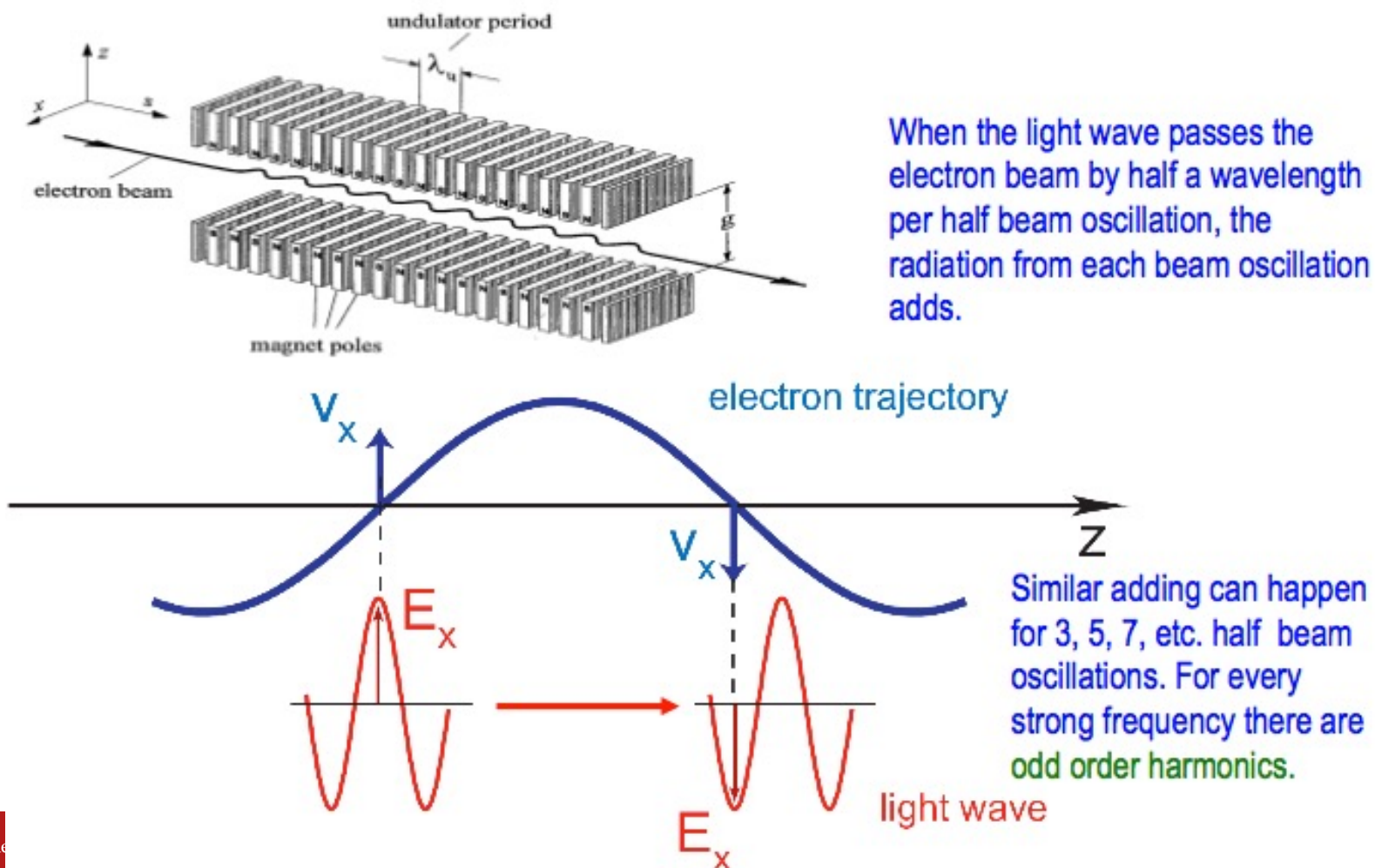
USPAS summer 2023

Radiative production by electrons

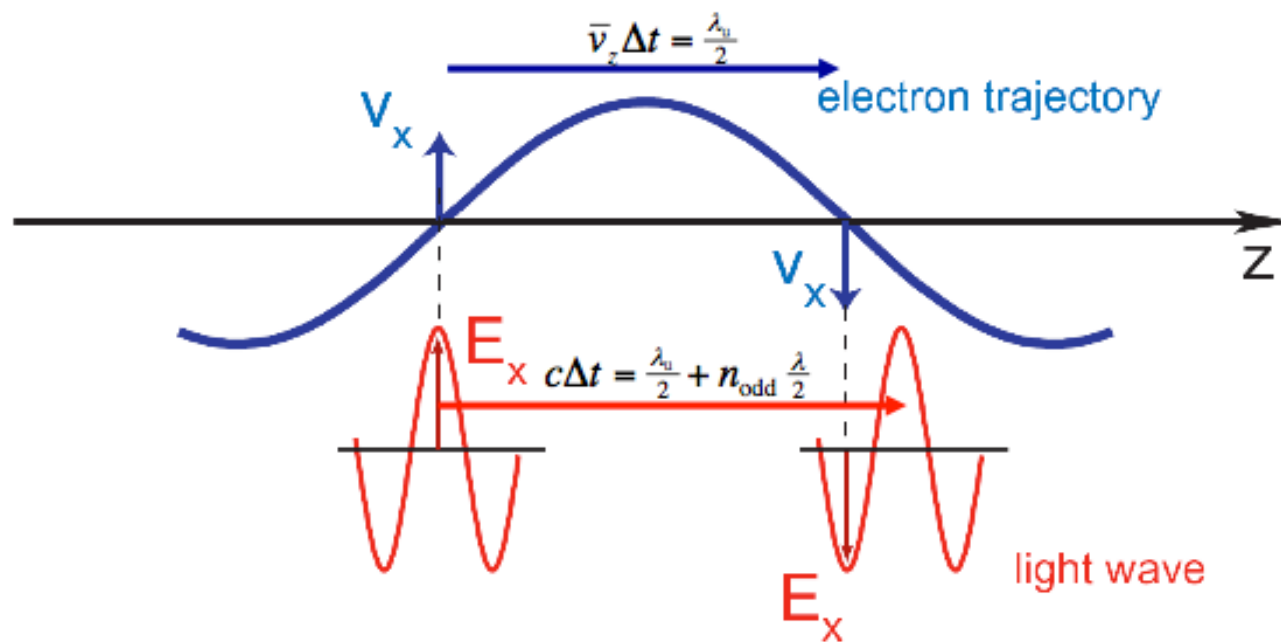


Cornell

Radiative production in undulators



Coherent addition of radiation



$$\frac{c}{\bar{v}_z} \frac{\lambda_u}{2} = \frac{\lambda_u}{2} + n_{\text{odd}} \frac{\lambda}{2} \Rightarrow \boxed{\lambda = \frac{1}{n_{\text{odd}}} \lambda_u \left(\frac{1}{\beta_z} - 1 \right)}$$



Cornell University



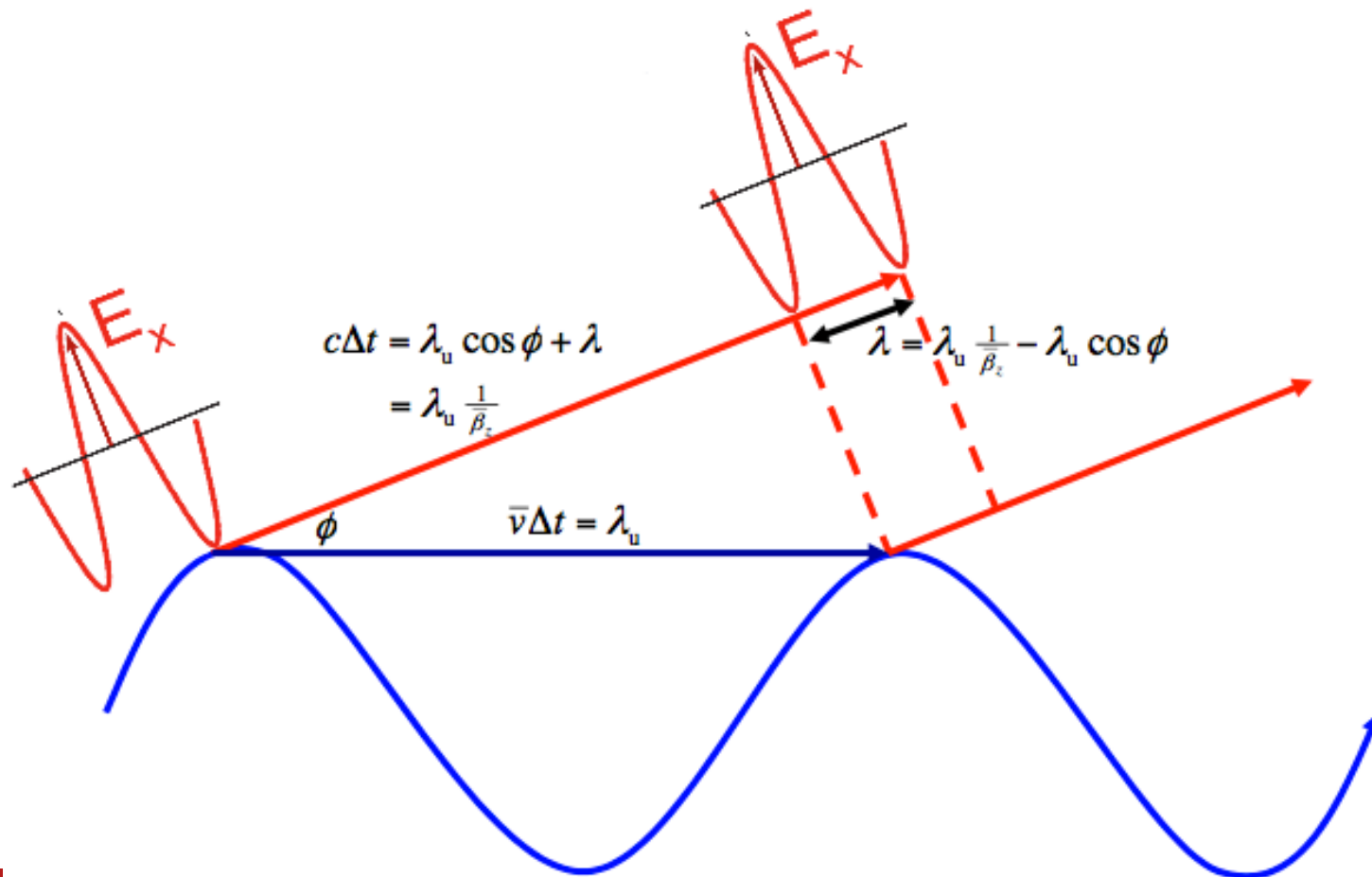
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Coherent addition at angles



Cornell University



Georg.Hoffstaetter@Cornell.edu

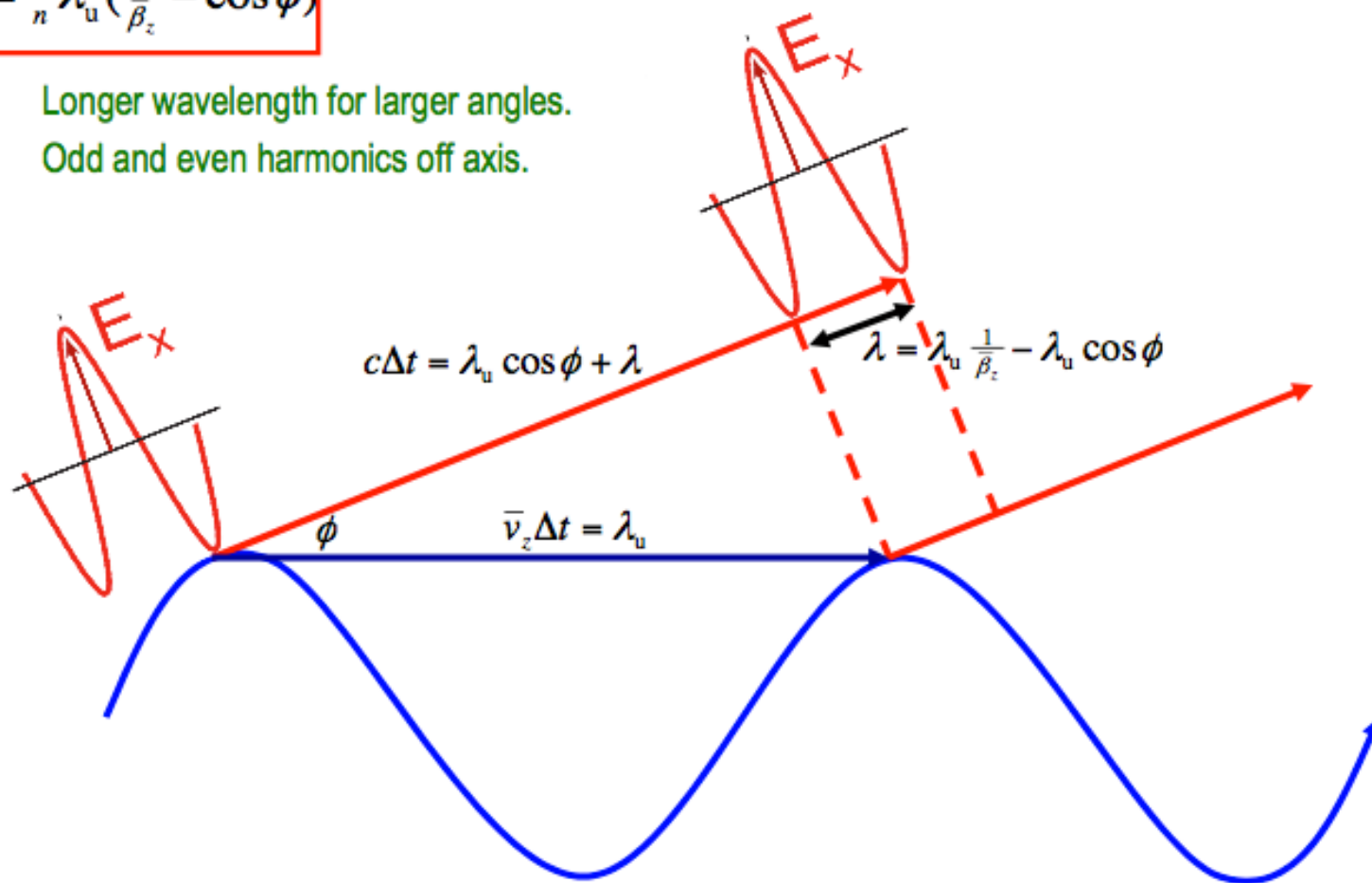
Graduate Accelerator Physics

USPAS summer 2023

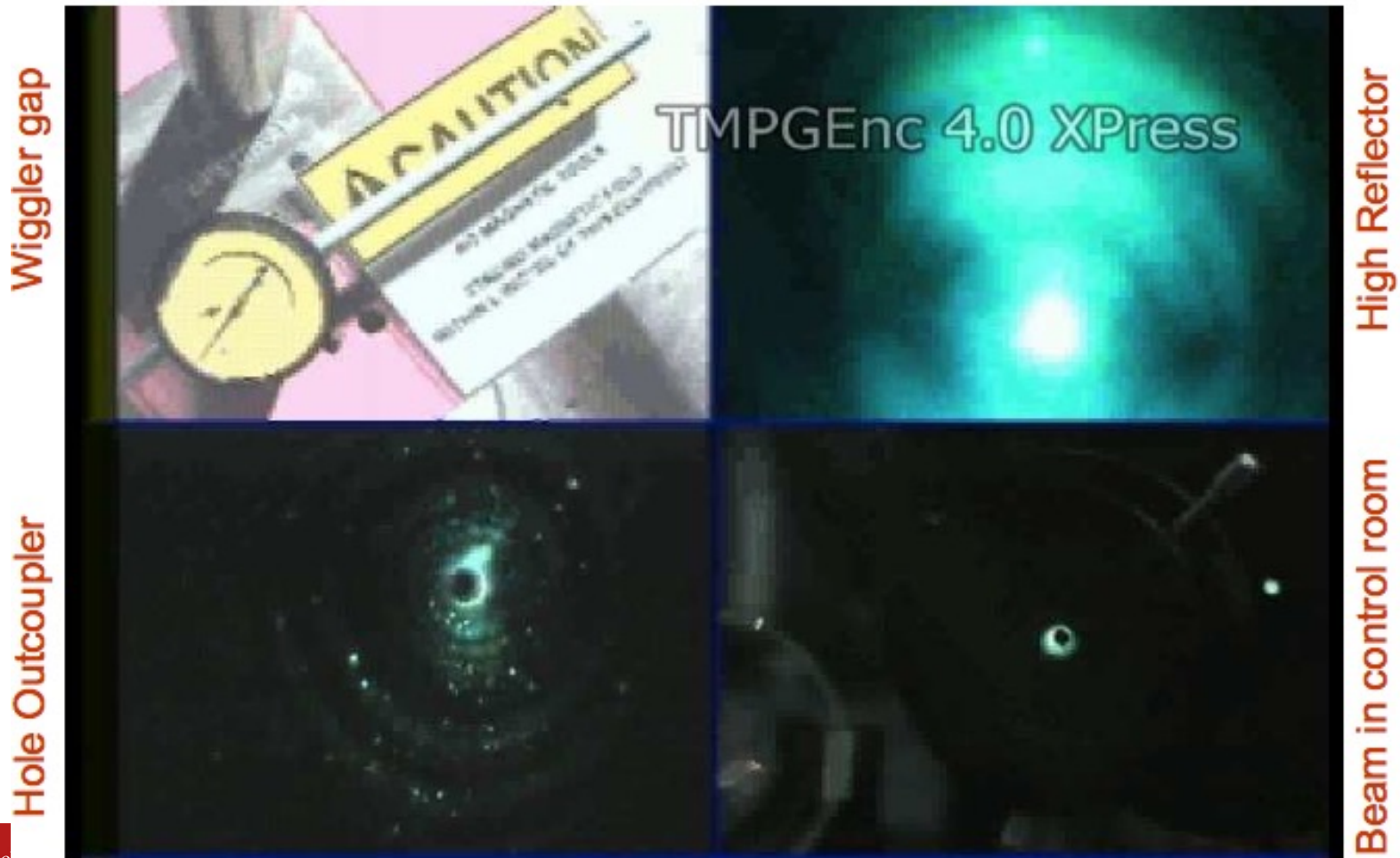
Radiation production at angles

$$\lambda = \frac{1}{n} \lambda_u \left(\frac{1}{\beta_z} - \cos \phi \right)$$

- 1) Longer wavelength for larger angles.
- 2) Odd and even harmonics off axis.



Lasing at the JLAB FEL



Cornell University



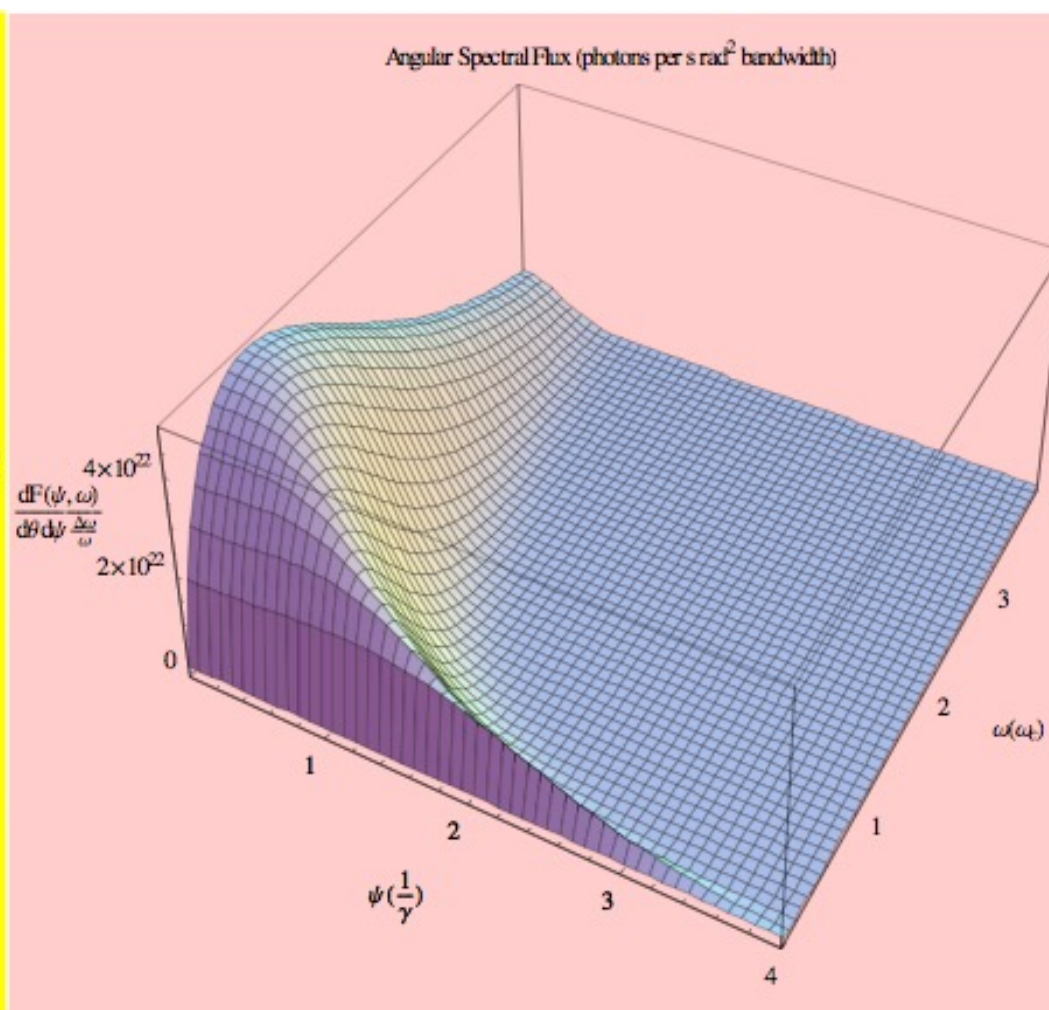
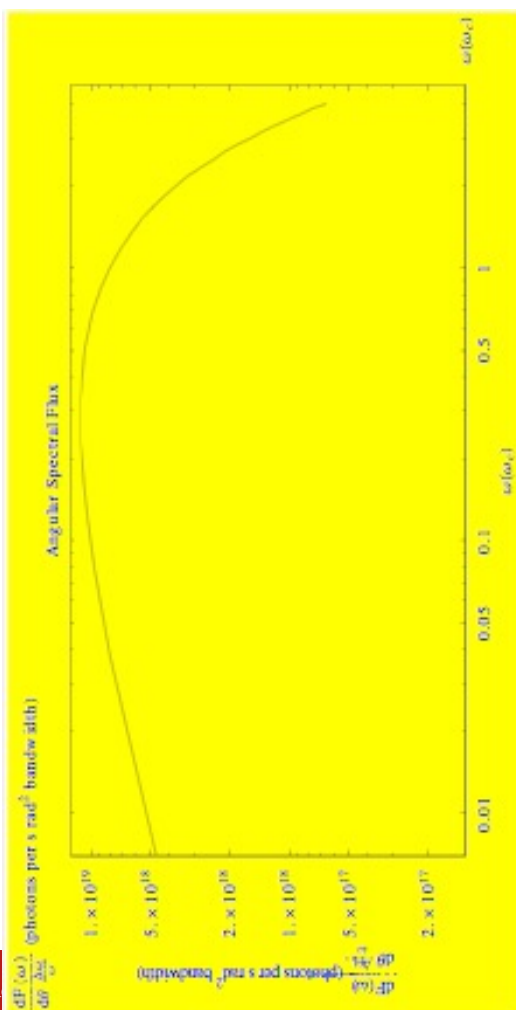
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

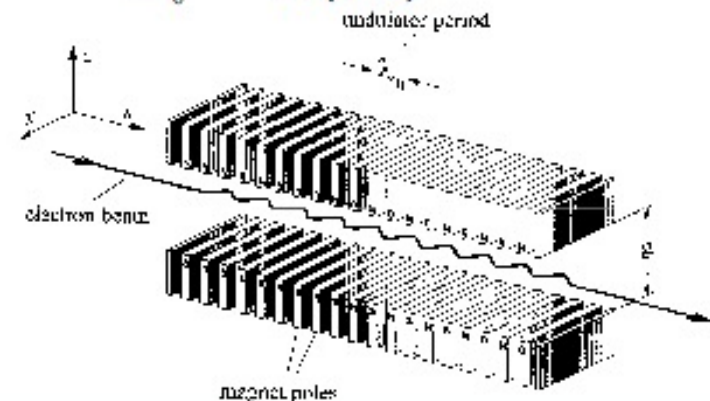
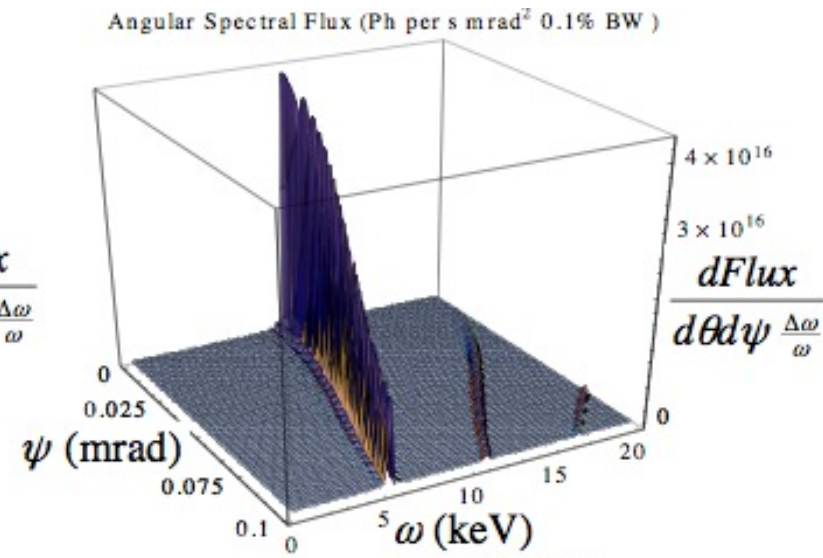
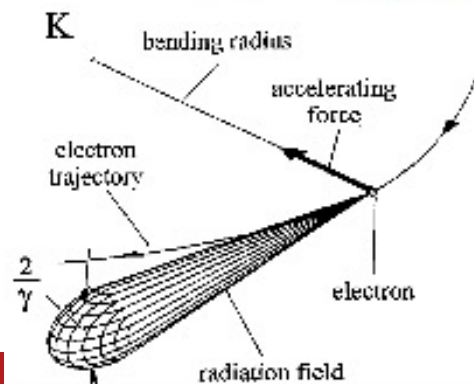
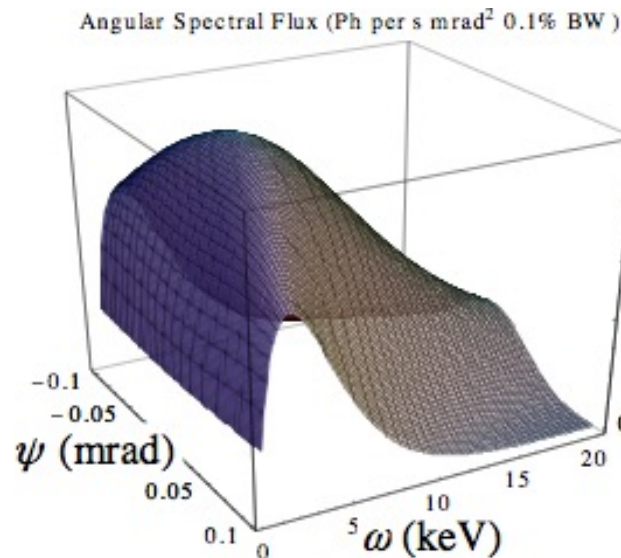
USPAS summer 2023

Radiative from bending magnets



Cornell

Photon flux from bends and undulators



Cornell University



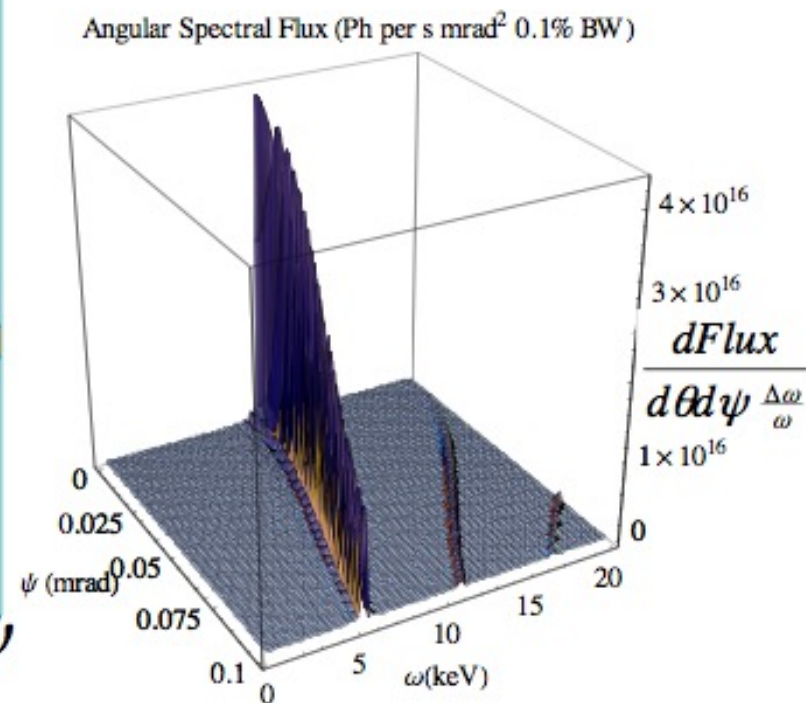
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

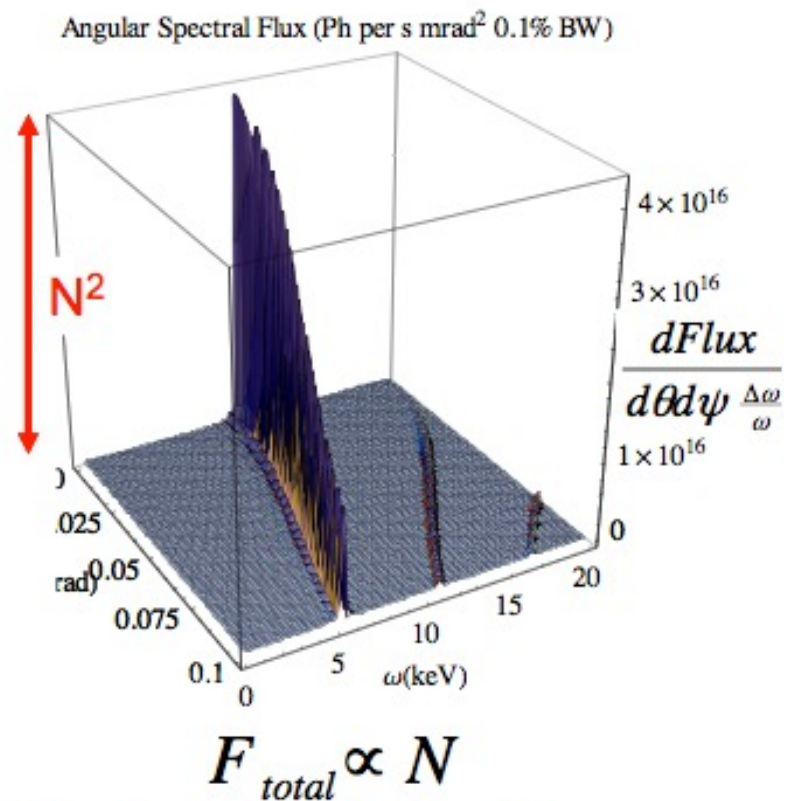
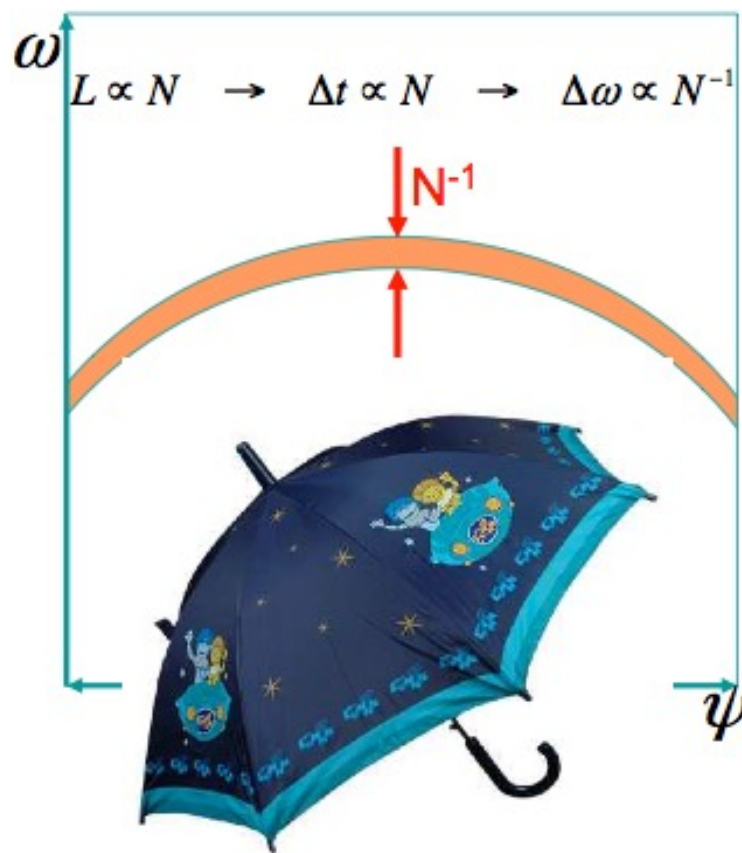
The umbrella of N-pole undulator radiation



$$F_{total} \propto N$$

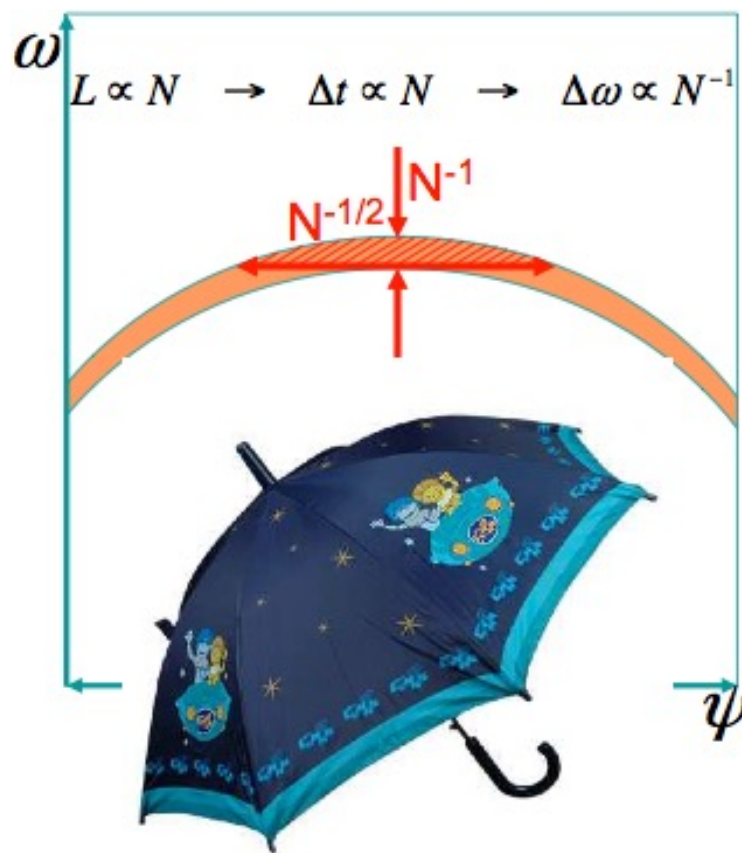
Flux from N poles is N times
the flux from one pole

The umbrella of N-pole undulator radiation

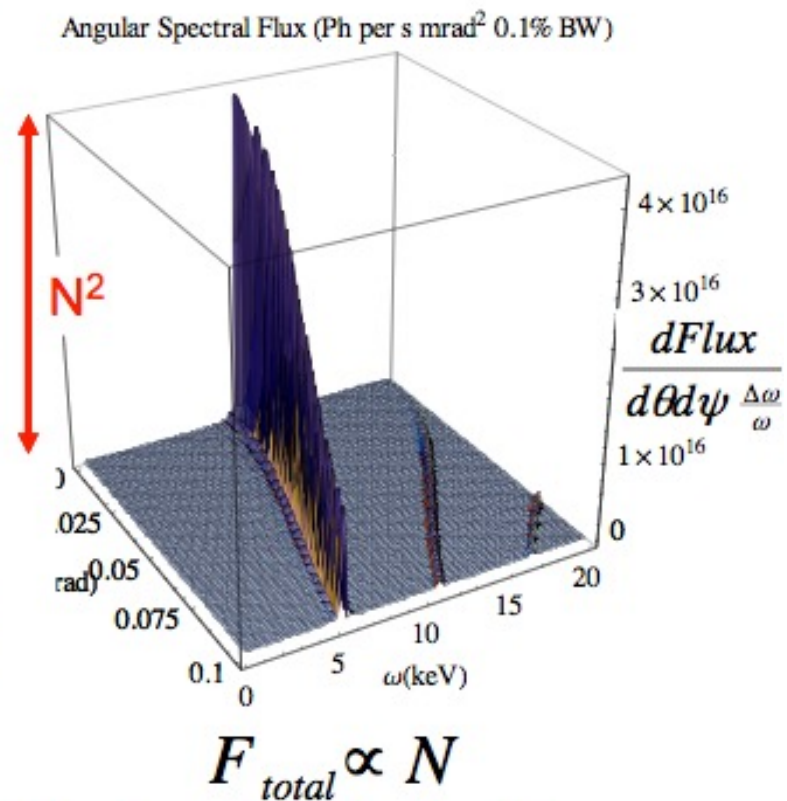


Flux from N poles is N times
the flux from one pole

The umbrella of N-pole undulator radiation

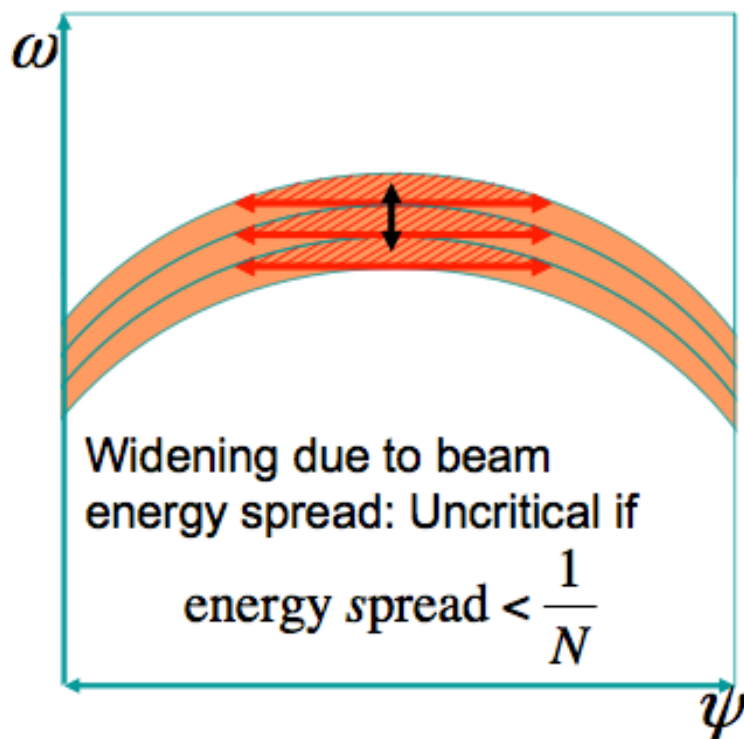


The power in the central cone is independent of N

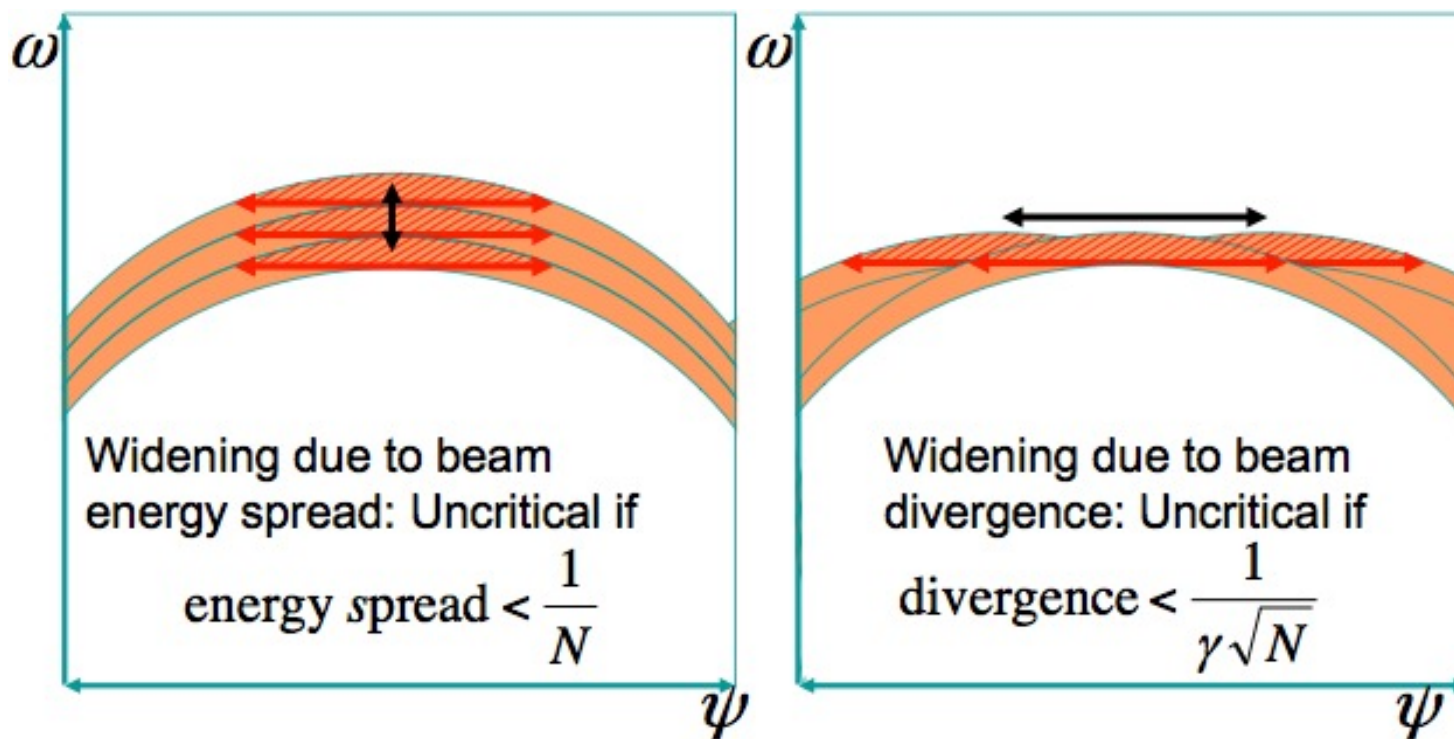


Flux from N poles is N times the flux from one pole

Brightness reduction by beam properties

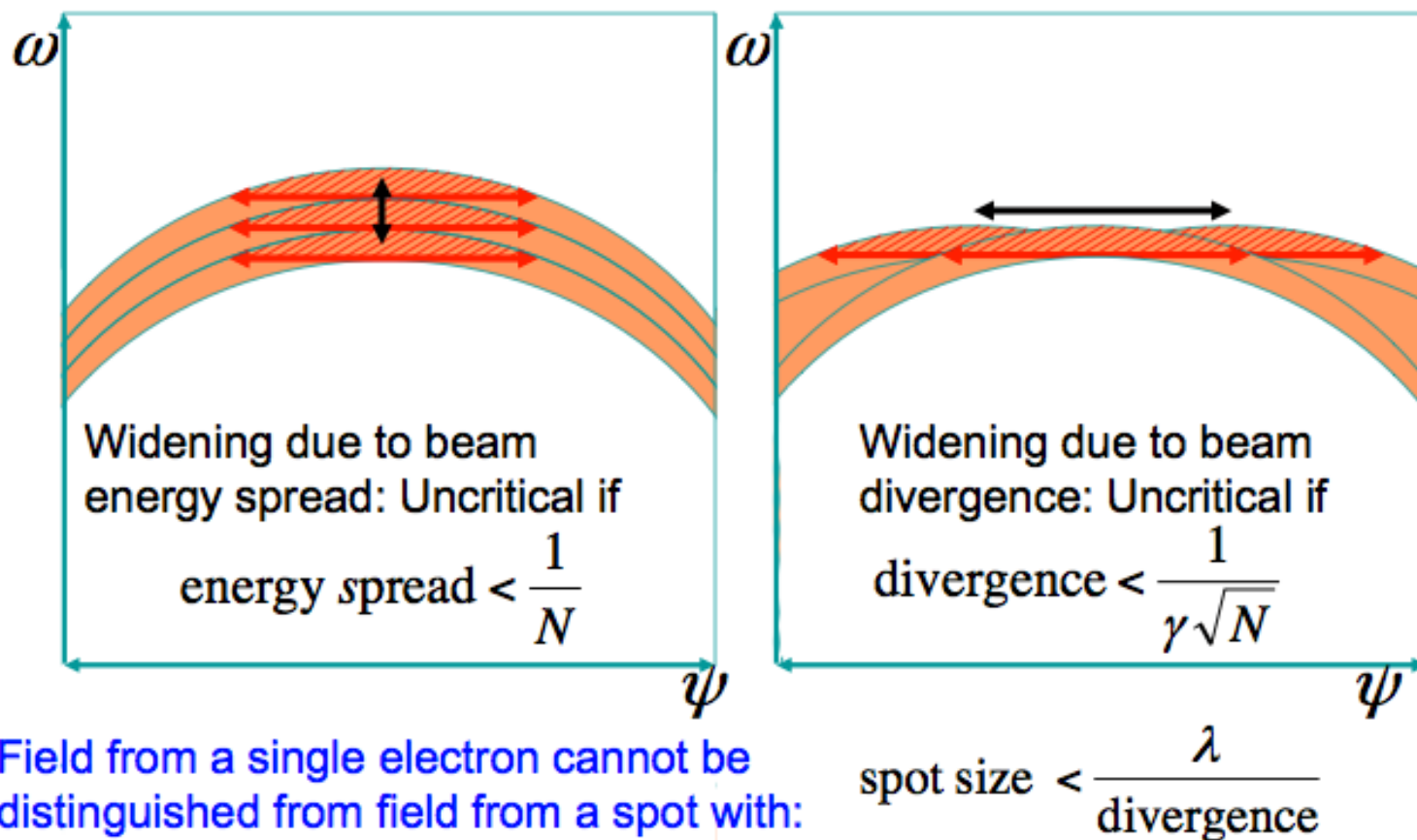


Brightness reduction by beam properties



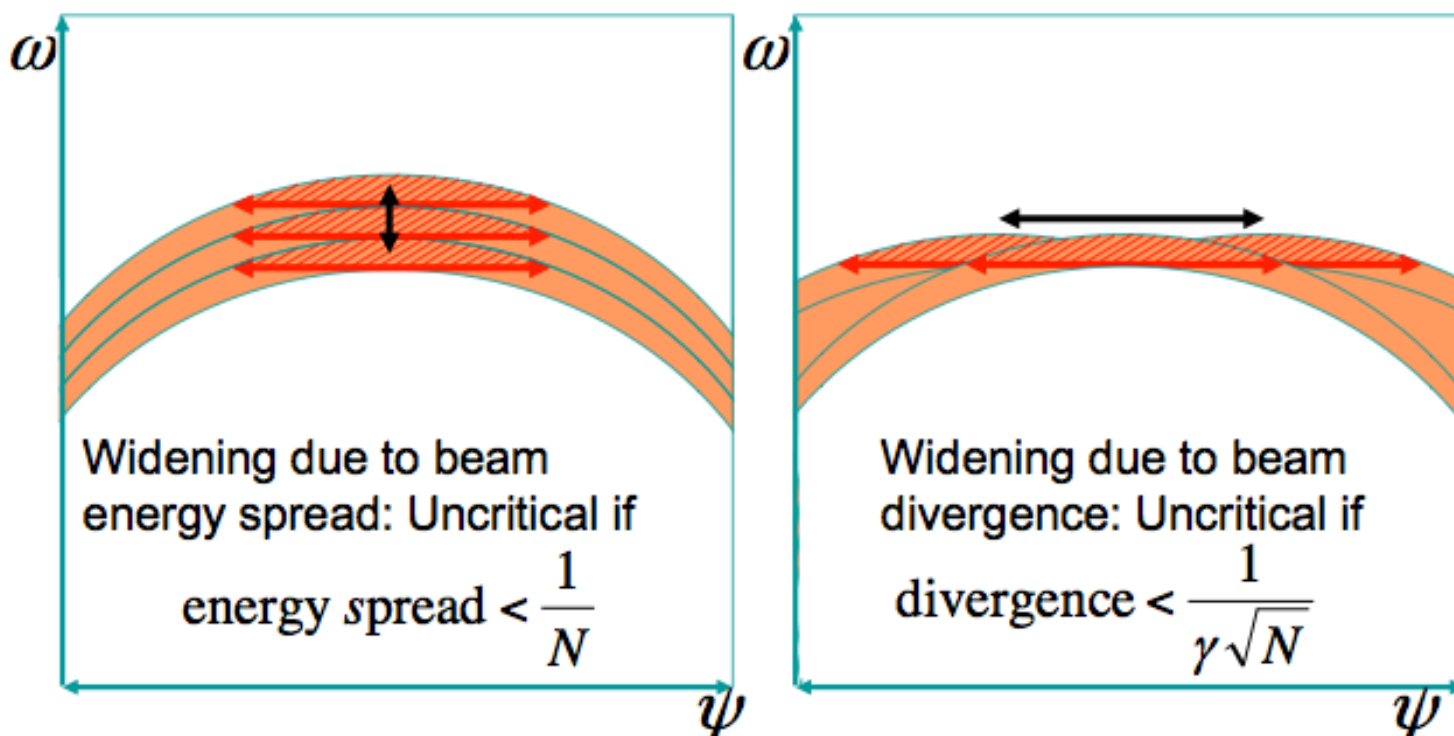
Cornell

Brightness reduction by beam properties



Cornell

Brightness reduction by beam properties



Field from a single electron cannot be distinguished from field from a spot with:

$$\text{spot size} < \frac{\lambda}{\text{divergence}}$$

To take advantage of many undulator poles, the electron beam needs to have little energy spread, little divergence, and small beam size.



Cornell

Dynamical Systems

$$\vec{z}(s) = \vec{M}(s; s_0, \vec{z}_0)$$

dynamical variable \vec{z}
Flow, transport map \vec{M}

By referring to a reference trajectory, transport maps in accelerators become
origin preserving: $\vec{M}(s; s_0, \vec{0}) = \vec{0}$

Flows build a group under concatenation:

$$\vec{M}(s; s_1, \cdot) \circ \vec{M}(s_1; s_0, \vec{z}_0) = \vec{M}(s; s_1, \vec{M}(s_1; s_0, \vec{z}_0)) = \vec{M}(s; s_0, \vec{z}_0)$$

- 1) Identity element: $\vec{M}(s; s_0, \vec{z}) = \vec{z}$
- 2) Inverse element of $\vec{M}(s; s_0, \vec{z}) = \vec{M}^{-1}(s; s_0, \vec{z})$ is $\vec{M}(s_0; s, \vec{z})$

In physics, the flow is often given as a solution of a first order ODE $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$

(Note that an nth order ODE can be rewritten as an n-dimensional first order ODE.)



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Uniqueness

Note that not all ODEs $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$
 have a unique solution $\vec{z}(s)$
 through a given point $\vec{z}(0) = \vec{z}_0$

Picard-Lindelöf:

A unique solution through (\vec{z}_0, s_0) exists for $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$ if $\vec{f}(\vec{z}, s)$ is Lipschitz continuous and bounded,
 i.e. it is continuous, bounded, and there is a number N such that

$$|\vec{f}(\vec{z}_1, s) - \vec{f}(\vec{z}_2, s)| < N |\vec{z}_1 - \vec{z}_2|$$

Example: $H = \frac{1}{2} p^2 + V(q), \quad V(q) = -8\sqrt{|q|}^3 \Rightarrow \dot{q} = p, \quad \dot{p} = 12\sqrt{|q|}$

There are two solutions through the point $(q, p, t) = (0, 0, 0)$

1. $(q(t), p(t)) = (0, 0)$
2. $(q(t), p(t)) = (t^4, 4t^3) \Rightarrow (\dot{q}, \dot{p}) = (4t^3, 12t^2)$

In our following treatments we do require uniqueness of solutions.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Linear Systems

Linear ODEs in N dimensions $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$ have $\vec{f}(\lambda \vec{z}, s) = \lambda \vec{f}(\vec{z}, s)$

$$\frac{d}{ds} \vec{z} = \underline{L}(s) \vec{z}$$

There are N linearly independent solutions. For example $\vec{z}_n(s)$ going through (\vec{z}_0, s_0)

With $z_{0i} = 0$ for $i \neq n$ and $z_{0n} = 1$

$$\vec{z}_0 = (0, \dots, 1, \dots, 0)^T \Rightarrow \vec{z}_n(s)$$

One speaks of N fundamental solutions.

Superposition for linear ODEs:

If z_1 is a solution and z_2 is a solution, then

any linear combination $Az_1 + Bz_2$ is also a solution

$$\frac{d}{ds} \vec{z}_1 = \underline{L}(s) \vec{z}_1 \quad \& \quad \frac{d}{ds} \vec{z}_2 = \underline{L}(s) \vec{z}_2 \quad \Rightarrow \quad \frac{d}{ds} (A\vec{z}_1 + B\vec{z}_2) = \underline{L}(s)(A\vec{z}_1 + B\vec{z}_2)$$

Therefore any solution through (\vec{z}_0, s_0) can be written as $\vec{z}(s) = \sum_{n=1}^N \vec{z}_n(s) z_{0n}$

$$\vec{z}(s) = \vec{M}(s; s_0, \vec{z}_0) = \underline{M}(s, s_0) \vec{z}_0 \quad \underline{M}(s, s_0) = (\vec{z}_1(s), \dots, \vec{z}_N(s))$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Nonlinear Systems

Nonlinear ODEs in N dimensions $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$

Have no fundamental solutions. Each solution has to be determined separately for each initial condition.

Examples: Plasma, Galaxies

$$H(\dots, \vec{r}_j, \dots, \dots, \vec{p}_j, \dots) = \sum_j \frac{\vec{p}_j^2}{2m_j} + \sum_{k \neq j} \frac{q_j q_k}{|\vec{r}_j - \vec{r}_k|}$$

Finding a general solution, flow, or transport map can be very hard.
This has not even been possible for the 3 body problem.

$$\vec{z}(s) = \vec{M}(s; s_0, \vec{z}_0)$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Weakly Nonlinear Systems

Weakly nonlinear ODEs $\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s)$

Have a right hand side that can be

approximated well by a truncated Taylor expansion

$$\vec{f}(\vec{z}, s) \approx \underline{L}(s) \vec{z} + \sum_{j,k} \vec{f}_{jk} z_j z_k + \sum_{j,k,l} \vec{f}_{jkl} z_j z_k z_l + \dots + \sum_{\vec{k}, \text{order } O} \vec{f}_{\vec{k}} \vec{z}^{\vec{k}} + \dots$$

$$\vec{z}^{\vec{k}} = \prod_{n=1}^N z_n^{k_n}, \quad \sum_{\vec{k}, \text{order } O} \dots = \sum_{n=1}^N \sum_{k_n} \dots \quad \text{with} \quad \sum_{n=1}^N k_n = O$$

By solving the Taylor expanded ODE one tries to find a Taylor expansion of the transport map:

$$\vec{M}(s; s_0, \vec{z}_0) \approx \underline{M}(s, s_0) \vec{z}_0 + \dots + \sum_{\vec{k}, \text{order } O} \vec{M}_{\vec{k}} \vec{z}_0^{\vec{k}} + \dots$$

Note:

While this approach is usually chosen, it is not certain that a **transport map of the Taylor expanded ODE** is a Taylor expansion of the **transport map of the original ODE**. One therefore often speaks of “formally” finding the Taylor expansion of the transport map.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Aberrations and Sensitivities

$$\vec{M}(s; s_0, \vec{z}_0) \approx \underline{M}(s, s_0) \vec{z}_0 + \dots + \sum_{\vec{k}, \text{order } O} \vec{M}_{\vec{k}} \vec{z}_0^{\vec{k}} + \dots$$

The Taylor coefficients are called aberrations of order O and are denoted by

$$(z_i, z_1^{k_1} \dots z_6^{k_6}) \equiv M_{\vec{k}, i}, \quad \text{order } O = \sum_{n=1}^6 k_n$$

Parameter dependences lead to **sensitivities**:

$$\frac{d}{ds} \vec{z} = \vec{f}(\vec{z}, s, \varepsilon) \Rightarrow \vec{z}(s) = \vec{M}(s, \varepsilon; s_0, \vec{z}_0)$$

$$\vec{M}(s, \varepsilon; s_0, \vec{z}_0) \approx \underline{M}(s, s_0) \vec{z}_0 + \underline{M}^1(s, s_0) \vec{z}_0 \varepsilon + \dots + \sum_{\vec{k}, n, \text{order } O} \vec{M}_{\vec{k}}^n \vec{z}_0^{\vec{k}} \varepsilon^n$$

$$(z_i, z_1^{k_1} \dots z_6^{k_6} \varepsilon^n) \equiv M_{\vec{k}, i}^n, \quad \text{order } O = n + \sum_{j=1}^6 k_j$$

How can all these Taylor coefficients be computed?



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

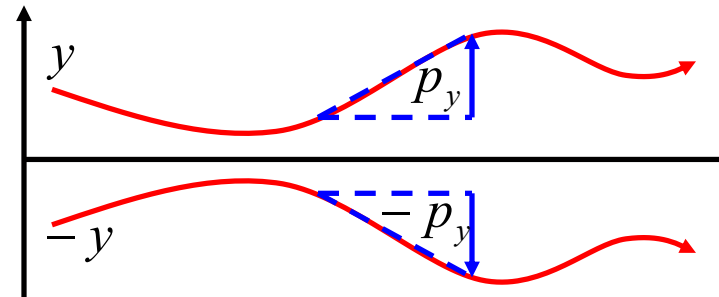
Horizontal Midplane Symmetry

This is the most important symmetry in nearly all accelerators.

$$\vec{r}^{\oplus} = (x, -y, z)$$

$$\vec{p}^{\oplus} = (p_x, -p_y, p_z)$$

$$\frac{d}{dt} \vec{p} = \vec{F}(\vec{r}, \vec{p}) \Rightarrow \frac{d}{dt} \vec{p}^{\oplus} = \vec{F}(\vec{r}^{\oplus}, \vec{p}^{\oplus})$$



$$\vec{z} = (x, a, y, b, \tau, \delta)$$

$$\vec{z}(s) = \vec{M}(s, \vec{z}_0)$$

$$\vec{z}^{\oplus} = (x, a, -y, -b, \tau, \delta)$$

$$\vec{z}^{\oplus}(s) = \vec{M}(s, \vec{z}_0^{\oplus})$$

$$M_i(s, \vec{z}_0^{\oplus}) = M_i(s, \vec{z}_0) \quad \text{for } i \in \{1, 2, 5, 6\}$$

$$M_i(s, \vec{z}_0^{\oplus}) = -M_i(s, \vec{z}_0) \quad \text{for } i \in \{3, 4\}$$

$$(x, x_0^{k_1} \dots \delta_0^{k_6}) = 0 \quad \text{for } k_3 + k_4 \text{ is odd} \quad \text{similarly for } a, \tau \text{ and } \delta$$

$$(y, x_0^{k_1} \dots \delta_0^{k_6}) = 0 \quad \text{for } k_3 + k_4 \text{ is even} \quad \text{similarly for } b$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Double Midplane Symmetry

In addition to midplane symmetry, some elements are symmetric around the vertical plane, e.g. quadrupoles, glass lenses

$$\vec{z} = (x, a, y, b, \tau, \delta)$$

$$\vec{z}(s) = \vec{M}(s, \vec{z}_0)$$

$$\vec{z}^{\oplus} = (x, a, -y, -b, \tau, \delta)$$

$$\vec{z}^{\oplus}(s) = \vec{M}(s, \vec{z}_0^{\oplus})$$

$$\vec{z}^{\otimes} = (-x, -a, y, b, \tau, \delta)$$

$$\vec{z}^{\otimes}(s) = \vec{M}(s, \vec{z}_0^{\otimes})$$

$$M_i(s, \vec{z}_0^{\oplus}) = M_i(s, \vec{z}_0) \quad \text{for } i \in \{1, 2, 5, 6\}$$

$$M_i(s, \vec{z}_0^{\oplus}) = -M_i(s, \vec{z}_0) \quad \text{for } i \in \{3, 4\}$$

$$M_i(s, \vec{z}_0^{\otimes}) = -M_i(s, \vec{z}_0) \quad \text{for } i \in \{1, 2\}$$

$$M_i(s, \vec{z}_0^{\otimes}) = M_i(s, \vec{z}_0) \quad \text{for } i \in \{3, 4, 5, 6\}$$

$$(x, x_0^{k_1} \dots \delta_0^{k_6}) = 0 \quad \text{for } k_1 + k_2 \text{ is even or } k_3 + k_4 \text{ is odd} \quad \text{similarly for } a, \tau \text{ and } \delta$$

$$(y, x_0^{k_1} \dots \delta_0^{k_6}) = 0 \quad \text{for } k_1 + k_2 \text{ is odd or } k_3 + k_4 \text{ is even} \quad \text{similarly for } b$$

Rotational Symmetry

Some optical elements are completely rotationally symmetric in the x-y plane,
e.g. solenoid magnets, many glass lenses

$$w = x + iy, \quad \alpha = a + ib$$

$$\vec{z} = (w, \bar{w}, \alpha, \bar{\alpha}, \tau, \delta)$$

$$\vec{z}^\oplus = (e^{i\varphi} w, e^{-i\varphi} \bar{w}, e^{i\varphi} \alpha, e^{-i\varphi} \bar{\alpha}, \tau, \delta)$$

$$\vec{z}(s) = \vec{M}(s, \vec{z}_0)$$

$$\vec{z}^\oplus(s) = \vec{M}(s, \vec{z}_0^\oplus)$$

$$M_i(s, \vec{z}_0^\oplus) = e^{i\varphi} M_i(s, \vec{z}_0) \quad \text{for } i \in \{1, 3\}$$

$$M_i(s, \vec{z}_0^\oplus) = e^{-i\varphi} M_i(s, \vec{z}_0) \quad \text{for } i \in \{2, 4\}$$

$$M_i(s, \vec{z}_0^\oplus) = M_i(s, \vec{z}_0) \quad \text{for } i \in \{5, 6\}$$

$$(w, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq 1 \quad \text{similarly for } \alpha$$

$$(\bar{w}, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq -1 \quad \text{similarly for } \alpha^*$$

$$(\tau, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq 0 \quad \text{similarly for } \delta$$

$$(w, |w_0|^2 w_0), (\bar{\alpha}, |w_0|^2 \bar{w}_0), (\alpha, w_0^2 \bar{\alpha}_0), (\tau, |w_0|^2) \quad \text{can all be non-zero}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

C_n Symmetry

Some optical elements have C_n symmetric in the x-y plane,
e.g. C₂ for quadrupole, C₃ for sextupoles, etc.

$$w = x + iy, \quad \alpha = a + ib$$

$$\vec{z} = (w, \bar{w}, \alpha, \bar{\alpha}, \tau, \delta)$$

$$\vec{z}^\oplus = (e^{i\frac{2\pi}{n}} w, e^{-i\frac{2\pi}{n}} \bar{w}, e^{i\frac{2\pi}{n}} \alpha, e^{-i\frac{2\pi}{n}} \bar{\alpha}, \tau, \delta)$$

$$\vec{z}(s) = \vec{M}(s, \vec{z}_0)$$

$$\vec{z}^\oplus(s) = \vec{M}(s, \vec{z}_0^\oplus)$$

$$M_i(s, \vec{z}_0^\oplus) = e^{i\frac{2\pi}{n}} M_i(s, \vec{z}_0) \quad \text{for } i \in \{1, 3\}$$

$$M_i(s, \vec{z}_0^\oplus) = e^{-i\frac{2\pi}{n}} M_i(s, \vec{z}_0) \quad \text{for } i \in \{2, 4\}$$

$$M_i(s, \vec{z}_0^\oplus) = M_i(s, \vec{z}_0) \quad \text{for } i \in \{5, 6\}$$

$$(w, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq jn + 1 \quad \text{similarly for } \alpha$$

$$(\bar{w}, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq jn - 1 \quad \text{similarly for } \alpha^*$$

$$(\tau, w^{k_1} \dots \delta^{k_6}) = 0 \quad \text{for } k_1 - k_2 + k_3 - k_4 \neq jn \quad \text{similarly for } \delta$$

$$(w, \bar{w}_0), (\bar{\alpha}, |w_0|^2 w_0), (\alpha, \bar{w}_0^2 \alpha_0), (\tau, |w_0|^2) \quad \text{can all be non-zero for } C_2$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Symplecticity

$[\vec{\partial} \vec{M}^T]^T \underline{J} [\vec{\partial} \vec{M}^T] = \underline{J}$ Symplecticity leads to the requirement that sums over certain products of aberrations must be either 0 or 1.

Separation into linear and nonlinear part of the map:

$$\vec{M}(\vec{z}) = \underline{M}_1(\vec{z} + \vec{N}(\vec{z}))$$

$$\underline{M}(\vec{z}) = [\vec{\partial} \vec{M}^T]^T = [(1 + \vec{\partial} \vec{N}^T) \underline{M}_1^T]^T = \underline{M}_1(1 + \underline{N}(\vec{z}))$$

$$(1 + \underline{N})^T \underline{M}_1^T \underline{J} \underline{M}_1 (1 + \underline{N}) = \underline{J} \Rightarrow \underline{M}_1^T \underline{J} \underline{M}_1 = \underline{J}, \quad \underline{N}^T \underline{J} + \underline{J} \underline{N} = -\underline{N}^T \underline{J} \underline{N}$$

For the leading order n-1 (the first order that appears in \underline{N}): $\underline{N}^T \underline{J} + \underline{J} \underline{N} = 0 + O^n$

\underline{N} is a Hamiltonian matrix up to order n and can thus be written up to order n as: $\vec{N}(\vec{z}) = \underline{J} \vec{\partial} f(\vec{z}) + O^{n+1}$

$$\begin{aligned} w(s_f) &= (x, x_0) \partial_a f + (x, y_0) \partial_b f + i[(y, x_0) \partial_a f + (y, y_0) \partial_b f] \\ &= (w, x_0) \partial_a f + (w, y_0) \partial_b f \\ &= \frac{1}{2}[(w, w_0) + (w, \bar{w}_0)][\partial_\alpha f + \partial_{\bar{\alpha}} f] - \frac{1}{2}[(w, w_0) - (w, \bar{w}_0)][\partial_\alpha f - \partial_{\bar{\alpha}} f] \\ &= (w, w_0) \partial_{\bar{\alpha}} f + (w, \bar{w}_0) \partial_\alpha f = (w, w_0) \partial_{\bar{\alpha}} f \end{aligned}$$



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Special Aberrations

Dispersion (for δ as parameter of 4-dimensional motion) $\vec{z} = \underline{M}(s) \vec{z}_0 + \vec{D}(s) \delta$

Chromatic aberrations $(x, \dots \delta^n), \quad n \neq 0$

Geometric aberrations $(x, x^{k_1} a^{k_2} y^{k_3} b^{k_4} \dots), \quad \sum_{i=1}^4 k_i \neq 0$

Purely Geometric aberrations $(x, \dots \delta^n), \quad n = 0$

Opening aberrations $(x, x^{k_1} \dots y^{k_2} \dots), \quad k_1 + k_2 = 0$

Field aberrations $(x, x^{k_1} \dots y^{k_2} \dots), \quad k_1 + k_2 \neq 0$

Spherical imaging systems: $(w, \alpha) = 0$

Spherical aberration for rotational symmetry $(w, \alpha | \alpha|^2)$

Coma line $(w, w | \alpha|^2)$

Coma circle $(w, \bar{w} \alpha^2)$

Astigmatism $(w, w^2 \bar{\alpha})$

Curvature of Image $(w, |w|^2 \alpha)$

Distortion $(w, w | w|^2)$



Cornell University



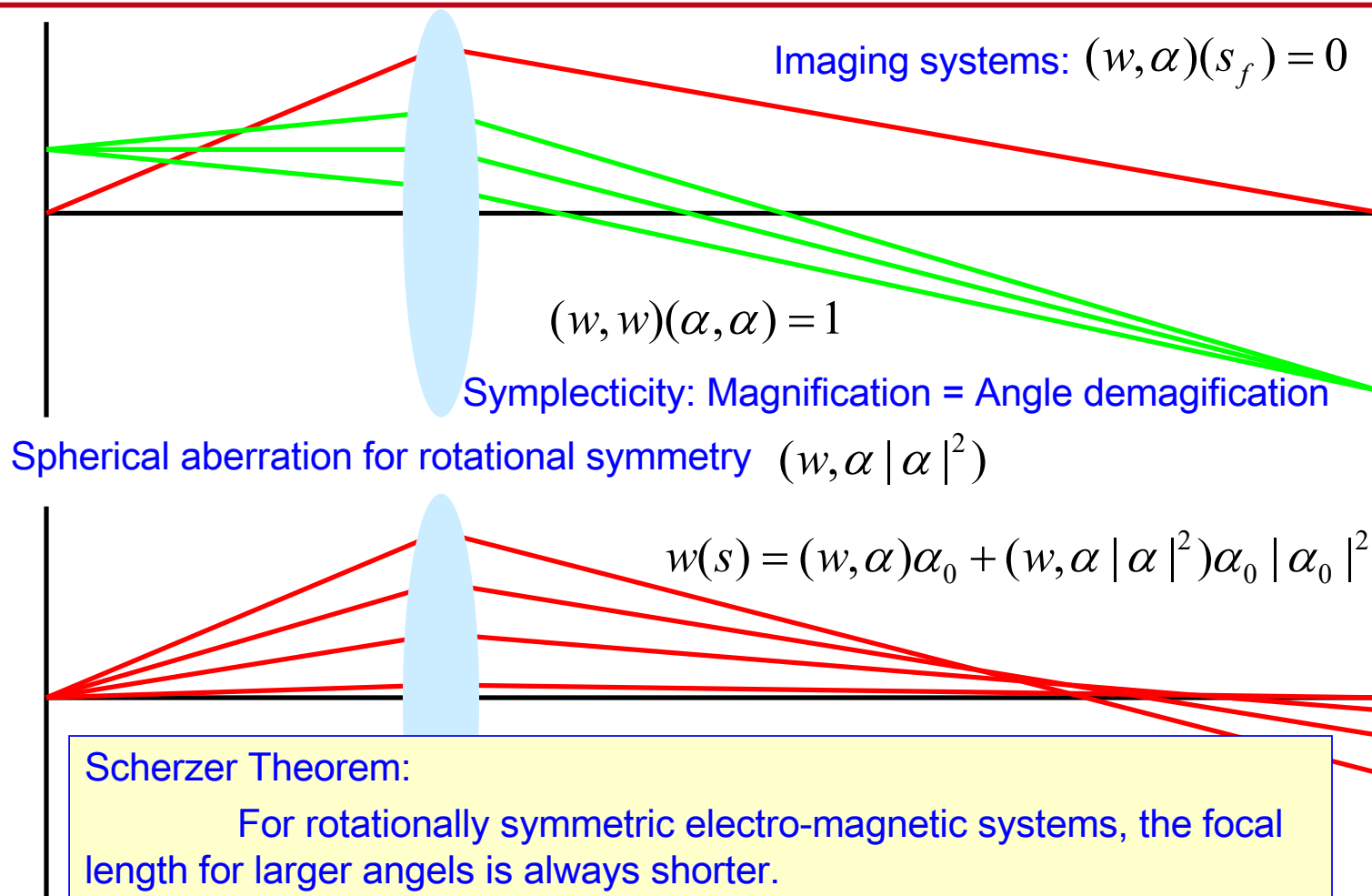
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Aberrations for rotational symmetry



Cornell University



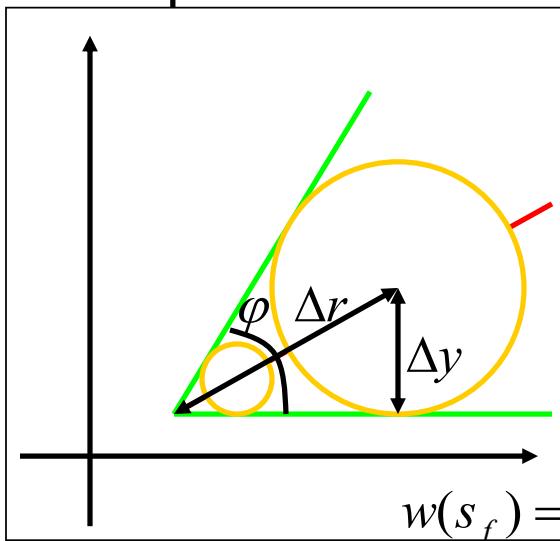
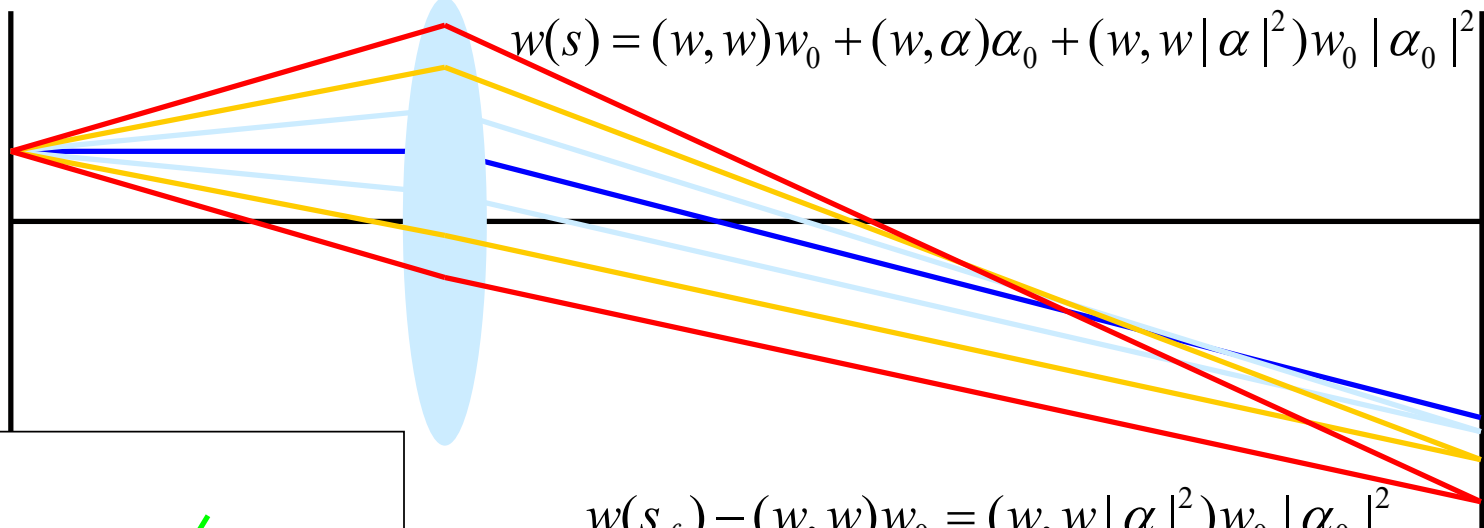
Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Koma Line and Koma Circle



$$w(s_f) - (w, w)w_0 = (w, w|\alpha|^2)w_0|\alpha_0|^2 + (w, \bar{w}\alpha^2)\bar{w}_0\alpha_0^2$$

$$\varphi = 2 \arcsin\left(\frac{\Delta y}{\Delta r}\right) = 2 \arcsin\left(\frac{(w, \bar{w}\alpha^2)}{(w, w|\alpha|^2)}\right)$$

Symplecticity yields:

$$(w, w|\alpha|^2) = 2(w, \bar{w}\alpha^2) \Rightarrow \varphi = 60^\circ$$

Since:

$$w(s_f) = (w, w_0)\partial_{\bar{\alpha}}[\dots + \text{Re}\{Kw\alpha\bar{\alpha}^2\}] = (w, w_0)[\dots + Kw\alpha\bar{\alpha} + \frac{1}{2}\bar{K}\bar{w}\alpha^2]$$



Cornell University

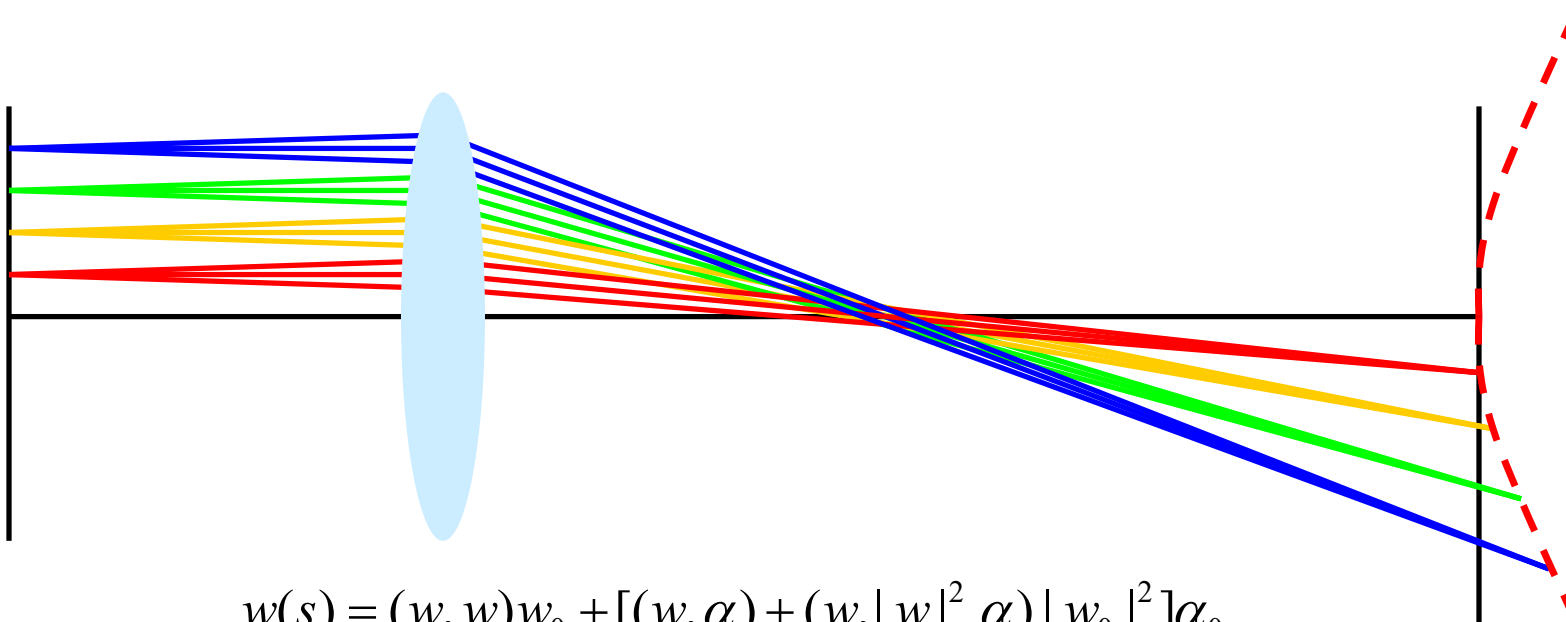


Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Curvature of Image



$$w(s) = (w, w)w_0 + [(w, \alpha) + (w, |w|^2 \alpha) |w_0|^2] \alpha_0$$

The focus occurs at $(w, \alpha)(s_f) + (w, |w|^2 \alpha)(s_f) |w_0|^2 = 0$



Cornell University



Brookhaven
National Laboratory

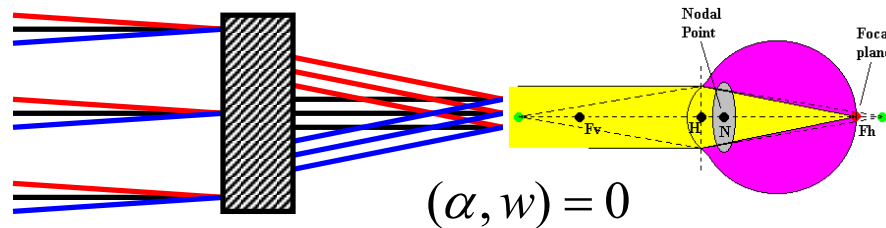
Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

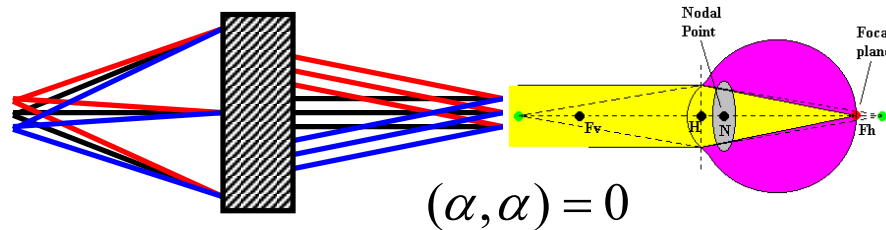
Other special systems

Telescope:
parallel to parallel system



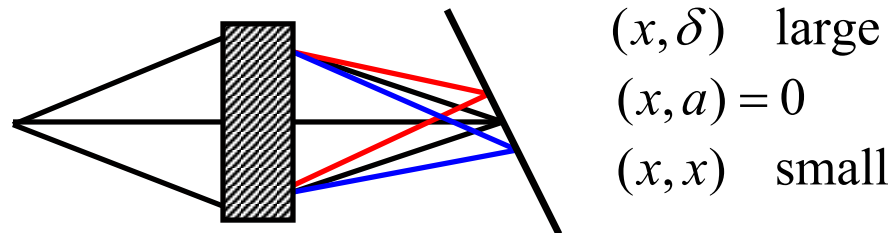
Nonlinearly corrected telescope: $(\alpha, w^n) = 0$

Microscope:
point to parallel system



Nonlinearly corrected microscope: $(\alpha, \alpha^n) = 0$

Spectrograph:
point to parallel system



Nonlinearly corrected spectrograph: $(x, a^n b^m) = 0$

Tilt of focal plane: $(x, a\delta) \neq 0$ the focus is at $(x, a)(s_f) + (x, a\delta)(s_f)\delta = 0$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Variation of constants

$$\vec{z}' = \vec{f}(\vec{z}, s)$$

$$\vec{z}' = \underline{L}(s)\vec{z} + \Delta\vec{f}(\vec{z}, s) \quad \text{Field errors, nonlinear fields, etc can lead to } \Delta\vec{f}(\vec{z}, s)$$

$$\vec{z}'_H = \underline{L}(s)\vec{z}_H \Rightarrow \vec{z}_H(s) = \underline{M}(s)\vec{z}_{H0} \quad \text{with} \quad \underline{M}'(s)\vec{a} = \underline{L}(s)\underline{M}(s)\vec{a}$$

$$\vec{z}(s) = \underline{M}(s)\vec{a}(s) \Rightarrow \vec{z}'(s) = \underline{M}'(s)\vec{a} + \underline{M}(s)\vec{a}'(s) = \underline{L}(s)\vec{z} + \Delta\vec{f}(\vec{z}, s)$$

$$\vec{a}(s) = \vec{z}_0 + \int_0^s \underline{M}^{-1}(\hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s}$$

$$\vec{z}(s) = \underline{M}(s) \left\{ \vec{z}_0 + \int_0^s \underline{M}^{-1}(\hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s} \right\}$$

$$= \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s})\Delta\vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s}$$

Perturbations are propagated
from s to s'



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Iteration of Aberrations

$$\vec{z}(s) = \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}(\vec{z}(\hat{s}), \hat{s}) d\hat{s}$$

$$\vec{z}_1(s) = \vec{z}_H(s)$$

$$\vec{z}_2(s) = \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}(\vec{z}_1(\hat{s}), \hat{s}) d\hat{s}$$

\vdots

$$\vec{z}_n(s) = \vec{z}_H(s) + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}(\vec{z}_{n-1}(\hat{s}), \hat{s}) d\hat{s}$$

Taylor expansions: $\Delta \vec{f}(\vec{z}, s) = \Delta \vec{f}_2(\vec{z}, s) + \Delta \vec{f}_3(\vec{z}, s) + \dots$, $\Delta f_O = \sum_{\vec{k}, \text{order } O} \vec{f}_{\vec{k}} \vec{z}^{\vec{k}}$

$$\vec{z}_1(s) = \underline{M}(s) \vec{z}_0$$

$$\vec{z}_2(s) = \underline{M}(s) \vec{z}_0 + \int_0^s \underline{M}(s, \hat{s}) \Delta \vec{f}_2(\vec{z}_1(\hat{s}), \hat{s}) d\hat{s}$$

$$\vec{z}_3(s) = \underline{M}(s) \vec{z}_0 + \int_0^s \underline{M}(s, \hat{s}) \{ [\Delta \vec{f}_2(\vec{z}_2(\hat{s}), \hat{s})]_3 + \Delta \vec{f}_3(\vec{z}_1(\hat{s}), \hat{s}) \} d\hat{s}$$

Poisson Bracket

The Poisson Bracket is defined as

$$[f(\vec{z}), g(\vec{z})] = \sum_i \partial_{q_i} f \partial_{p_i} g - \partial_{p_i} f \partial_{q_i} g = \vec{\partial}^T f \underline{J} \vec{\partial} g$$

The Poisson Bracket can be viewed as a product on the vector space of phase space functions. It is:

- 1) Linear: $[f, ag] = [af, g] = a[f, g], \quad a \in \mathbb{R}$
- 2) Distributive: $[f, g + h] = [f, g] + [f, h]$

This turns the vector space into an **algebra**.

The multiplication is furthermore:

- 1) Anti-commutative: $[f, g] = -[g, f]$
- 2) Has a Jacobi-identity: $[f, [g, h]] + [g, [h, f]] + [h, [f, g]] = 0$
as can be proven by the product rule: $[f, gh] = g[f, h] + [f, g]h$

This turns the algebra into a **Lie algebra**.

Example: $\vec{a} \times \vec{b}$ turns \mathbb{R}^3 into a Lie algebra.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Map computation by Lie Algebra

The Poisson-Bracket operator of $f, :g:$ is defined as $:g:h = [g, h]$

$$\underline{:H:g} = [H, g] = -[g, H] = -\vec{\partial}^T g \underline{J} \vec{\partial} H = -\vec{\partial}^T g \underline{\frac{d}{ds}} \vec{z} = -\underline{\frac{d}{ds}} g(\vec{z})$$

$$\underline{\frac{d}{ds}} \vec{z} = \vec{f}(\vec{z}, s) \Rightarrow -:H:z_j = \underline{\frac{d}{ds}} z_j = f_j(\vec{z}, s), \quad -:H:f_j = \underline{\frac{d}{ds}} f_j - \frac{\partial}{\partial s} f_j$$

In the main field region where $\underline{\frac{d}{ds}} \vec{z} = \vec{f}(\vec{z}) \Rightarrow -:H:f_j = \underline{\frac{d}{ds}} f_j = \underline{\frac{d^2}{ds^2}} z_j$

If $g(\vec{z}) = \frac{d^n}{ds^n} z_j$ then $-:H:g = \underline{\frac{d}{ds}} g = \underline{\frac{d^{n+1}}{ds^{n+1}}} z_j \Rightarrow (-:H:)^n z_j = \underline{\frac{d^n}{ds^n}} z_j$

Propagator: $e^{-\Delta s:H:} \vec{z} = \sum_{n=0}^{\infty} \frac{(-\Delta s:H:)^n}{n!} \vec{z} = \sum_{n=0}^{\infty} \frac{\Delta s^n}{n!} \underline{\frac{d^n}{ds^n}} \vec{z} = \vec{M}(s + \Delta s; s, \vec{z})$

$$\begin{aligned} \vec{M}_2 \circ \vec{M}_1(\vec{z}_0) &= \vec{M}_2(\Delta s_2, \vec{z}(\Delta s_1)) = \sum_{n=1}^{\infty} \frac{(-\Delta s_1:H_1:)^n}{n!} \vec{M}_2(\Delta s_2, \vec{z}_0) \\ &= e^{-\Delta s_1:H_1(\vec{z}_0):} e^{-\Delta s_2:H_2(\vec{z}_0):} \vec{z}_0 \end{aligned}$$

$$\vec{M}(s; \vec{z}_0) = \vec{M}_n \circ \dots \circ \vec{M}_2 \circ \vec{M}_1(\vec{z}_0) = e^{-\Delta s_1:H_1:} \dots e^{-\Delta s_n:H_n:} \vec{z}_0$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Poisson Bracket Invariance

The Poisson Bracket is invariant under a symplectic transfer map

$$[f(\vec{M}(\vec{z})), g(\vec{M}(\vec{z}))] = \vec{\partial}^T f \Big|_{\vec{M}} \underline{M} \underline{J} \underline{M}^T \vec{\partial} g \Big|_{\vec{M}} = [f(\vec{z}), g(\vec{z})]_{\vec{M}(\vec{z})}$$

For nonlinear expansions, one writes the transport map as a linear matrix
and a nonlinear Lie exponent,

$$\vec{M}_1(\vec{z}) = \underline{M}_1 e^{H_1(\vec{z})} \vec{z} = \underline{M}_1 \sum_{n=0}^{\infty} \frac{H_1^n}{n!} \vec{z}$$

since a linear Lie exponent requires infinitely many terms in the power sum, but the nonlinear exponent terminates when a finite order expansion is sought.

$$\begin{aligned} (\underline{M}_2 e^{H_2(\vec{z})} \vec{z}) \circ (\underline{M}_1 e^{H_1(\vec{z})} \vec{z}) &= \underline{M}_2 e^{H_2(\underline{M}_1 e^{H_1(\vec{z})} \vec{z})} \underline{M}_1 e^{H_1(\vec{z})} \vec{z} \\ &= \underline{M}_2 \underline{M}_1 e^{H_2(\underline{M}_1 e^{H_1(\vec{z})} \vec{z})} e^{H_1(\vec{z})} \vec{z} = \underline{M}_2 \underline{M}_1 e^{H_1(\vec{z})} e^{H_2(\underline{M}_1 \vec{z})} \vec{z} \end{aligned}$$

When these equations are used to compute and manipulate transfer maps, one speaks of the Lie algebraic method.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Computing Taylor Expansions

$$f(x) \approx \sum_{n=1}^{\text{order } O} \frac{x^n}{n!} \partial^n f|_0$$

But taking this approach for complicated functions would be very cumbersome:

1. $f(x) = \frac{1}{1 + \sin x} - 1, \quad f(0) = 0, \quad \partial f|_0 = \left. \frac{-\cos x}{(1 + \sin x)^2} \right|_0 = -1,$

$$\partial^2 f(x) = \left. \frac{\sin x(1 + \sin x) + 2 \cos^2 x}{(1 + \sin x)^3} \right|_0 = 2, \quad \underline{f(x) \approx -x + x^2 + O^3}$$

2. $f(x) = \frac{1}{1 + \sin x} - 1,$

This approach is formalized in the field of
automatic differentiation using a
Differential Algebra.

$$f(x) \approx \frac{1}{1 + x - \frac{1}{6}x^3 + O^4} - 1$$

$$\approx -(x - \frac{1}{6}x^3) + (x - \frac{1}{6}x^3)^2 - (x - \frac{1}{6}x^3)^3 + O^4$$

$$\approx -x + x^2 - \frac{5}{6}x^3 + O^4$$

Computations with TPSA(n)

Computation of a function in \mathbb{R} is done by a finite number of elementary operations (+, -, x) and elementary function evaluations (sin, cos, exp, 1/x, ...).

$$f(x) = \frac{1}{1 + \sin x} - 1$$

1. $x \in \mathbb{R}$
2. sin
3. 1 +
4. 1/
5. -1

If $g_n(x)$ is the truncated power series of order n of $g(x)$ and $h_n(x)$ is that of $h(x)$ we can look for elementary operations (“+”, “-”, “x”) so that

$g_n + h_n$ is the TPS(n) of $g+h$

$g_n - h_n$ is the TPS(n) of $g-h$

$g_n \cdot h_n$ is the TPS(n) of $g \cdot h$

Similarly we can look for elementary functions (“sin”, “cos”, “exp”, “1/x”, ...) so that

“sin”(g_n) is the TPS(n) of $\sin(g)$, “exp”(g_n) is the TPS(n) of $\exp(g)$, etc.

Evaluating all elementary operations and elementary functions in $f(x)$ in terms of “+”, “-”, “x”, “sin”, “cos”, “exp”, “1/x”, ... starting with the TPS(n) of x , leads to the TPS(n) of $f(x)$.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Automatic Differentiation with TPSA(n)

Example: computing the TPS(3) of $f(x) = \frac{1}{1 + \sin x} - 1$

1. TPS(3) of x is x
2. TPS(3) of " \sin " x is $x - \frac{1}{6}x^3$
3. $1 + x - \frac{1}{6}x^3 = 1 + x - \frac{1}{6}x^3$
4. $i(x) = \frac{1}{1+x}$, " i "(x) = $1 - x + x^2 - x^3$, " i "($x - \frac{1}{6}x^3$) = $1 - x + x^2 - \frac{5}{6}x^3$
5. $1 - x + x^2 - \frac{5}{6}x^3 - 1 = -x + x^2 - \frac{5}{6}x^3$

This automatically (i.e. with a computer) leads to derivatives of $f(x)$:

$$f(0) = 0, f'(0) = -1, f''(0) = 2, f'''(0) = -4$$

Truncated power series can be added "+" and multiplied "x" and there is a neutral element of multiplication (i.e. 1). Therefore the vector space of TPS(n) forms an algebra. It is called the **Truncated Power Series Algebra** TPSA(n).



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Differential Algebra $_1D_1$

An addition and multiplication with a scalar leads to a vector space over \mathbb{R}^2 :

$$\{a_0, a_1\}, \{b_0, b_1\} \in \mathbb{R}^2, t \in \mathbb{R}$$

$$\{a_0, a_1\} + \{b_0, b_1\} = \{a_0 + b_0, a_1 + b_1\}$$

$$t\{a_0, a_1\} = \{ta_0, ta_1\}$$

The introduction of a multiplication $\{a_0, a_1\} \{b_0, b_1\} = \{a_0 b_0, a_0 b_1 + a_1 b_0\}$ leads to an algebra if it is:

- 1) Distribut. $\{a_0, a_1\}(\{b_0, b_1\} + \{c_0, c_1\}) = \{a_0, a_1\} \{b_0, b_1\} + \{a_0, a_1\} \{c_0, c_1\}$
- 2) Has a neutral element: $\{a_0, a_1\} \{1, 0\} = \{a_0, a_1\}$

and additionally to a ring if it is

- 3) Commutative: $\{a_0, a_1\} \{b_0, b_1\} = \{b_0, b_1\} \{a_0, a_1\}$
- 4) Associative: $\{a_0, a_1\}(\{b_0, b_1\} \{c_0, c_1\}) = (\{a_0, a_1\} \{b_0, b_1\}) \{c_0, c_1\}$

All these properties are clearly given, since first order power expansion

have this multiplication: $(a_0 + a_1 x)(b_0 + b_1 x) = a_0 b_0 + (a_0 b_1 + a_1 b_0)x + O^2$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Differential Algebra ${}_1D_1$

By the introduced addition and multiplication we created an **algebra**, since the multiplication is commutative and associative we also created a **ring**, but **not a field**. Complex numbers are a field since there is a multiplicative inverse for all numbers except 0.

$$\{a_0, a_1\} \{b_0, b_1\} = \{a_0 b_0, a_0 b_1 + a_1 b_0\} \Rightarrow \{a_0, a_1\} \left\{ \frac{1}{a_0}, -\frac{a_1}{a_0^2} \right\} = \{1, 0\}$$

We further introduce a **differentiation**: $\partial\{a_0, a_1\} = \{a_1, 0\}$

It is a differentiation since it satisfies a **product rule**:

$$\partial(\{a_0, a_1\} \{b_0, b_1\}) = \{a_0 b_1 + a_1 b_0, 0\} = (\partial\{a_0, a_1\}) \{b_0, b_1\} + \{a_0, a_1\} (\partial\{b_0, b_1\})$$

By adding a differentiation we have created a **Differential Algebra (DA)**.

Differentiation of Polynomials: $f(x) = 2 + x^2 \Rightarrow f'(x) = 2x$

$$f(\{2, 1\}) = \{2, 0\} + \{4, 4\} = \{6, 4\} = \{f(2), f'(2)\}$$

Since $\{f, f'\} + \{g, g'\} = \{(f + g), (f + g)'\}$, $\{f, f'\} \{g, g'\} = \{(fg), (fg)'\}$

Every polynomial: $P(\{f, f'\}) = \{P(f), [P(f)]'\}$ and $P(\{x, 1\}) = \{P(x), P'(x)\}$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Elementary Functions in \mathbf{D}_1

$$e(a_0 + a_1 x) = e(a_0) + e'(a_0) a_1 x + O^2$$

leads to

$$e(\{a_0, a_1\}) = \{e(a_0), a_1 e'(a_0)\}$$

$$\sin(\{a_0, a_1\}) = \{\sin a_0, a_1 \cos a_0\}$$

$$\cos(\{a_0, a_1\}) = \{\cos a_0, -a_1 \sin a_0\}$$

Since $\{f, f'\} + \{g, g'\} = \{(f+g), (f+g)'\}$, $\{f, f'\} \{g, g'\} = \{(fg), (fg)'\}$

and $e(\{f, f'\}) = \{e(f), [e(f)]'\}$

Therefore $F(\{f, f'\}) = \{F(f), [F(f)]'\}$ and $F(\{x, 1\}) = \{F(x), F'(x)\}$

So that **automatic differentiation** works not only for Polynomials but for any function that is constructed from a finite number of operations and elementary functions.

Computer programs that have differential algebra elements as data types can evaluate any function or algorithm in this data type and obtain derivatives of the function or derivatives of the algorithm.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The Differential Algebra $_nD_v$

The concept of $_1D_1$ can be extended to truncated power series of order n and to v variables. This leads to the differential algebra $_nD_v$. For each coefficient in the n th order expansion there is one dimension in the vectors of $_nD_v$.

Power expansions for v variables have extremely many expansion coefficients:

A polynomial of order n in v variables has $\dim(_nD_v) = \frac{(n+v)!}{n!v!}$ coefficients since

$$\underbrace{\dim(_nD_v) - \dim(_{n-1}D_v)}_{z_1^{k_1} \dots z_v^{k_v}, \sum_{j=1}^v k_j = n} = \underbrace{\dim(_nD_{v-1})}_{z_1^{k_1} \dots z_{v-1}^{k_{v-1}}, \sum_{j=1}^v k_j \leq n}, \quad \frac{(n+v)!}{n!v!} - \frac{(n-1+v)!}{(n-1)!v!} = \frac{(n+v-1)!}{n!(v-1)!}$$

and iteration of $_nD_v$ starts with the correct conditions: $\dim(_nD_1) = n+1 = \frac{(n+1)!}{n!}$

Example: $\dim(_{10}D_6) = 8008$ $\dim(_0D_v) = 1 = \frac{v!}{v!}$

Computer programs that have differential algebra elements as data types produce the n th order power expansion of v -dimensional functions or algorithms automatically.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Equivalence classes in $_nD_v$

A TPS(n) of a function $f(x)$ defines the equivalence class of all functions that have the same TPS(n).

$$\text{Def: } f =_n g \quad \text{if} \quad \vec{\partial}^{\vec{k}} f(0) = \vec{\partial}^{\vec{k}} g(0) \quad \forall \vec{k} \text{ with order } \leq n$$

$=_n$ is an equivalence relation since it has

- 1) the identity property $f =_n f \quad \forall f$
- 2) the symmetry property $f =_n g \quad \text{if} \quad g =_n f$
- 3) the transitivity property $f =_n h \quad \text{if} \quad f =_n g \text{ and } g =_n h$

Equivalence classes: Def: $[f]_n = \{g \mid g =_n f\}$

Arithmetic of equivalence class:

$$[f]_n + [g]_n \equiv [f + g]_n$$

$$[f]_n [g]_n \equiv [f g]_n$$

Those operations generate a differential algebra.

$$t[f]_n \equiv [t f]_n$$

$$\partial_j [f]_n \equiv [\partial_j f]_{n-1}$$

$$e([f]_n) \equiv [e(f)]_n$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Concatenation of maps

$$f(x), g(x) \quad \text{and} \quad [f]_n, [g]_n \in_n D_1$$

$$[f(g(x))]_0 = [f(g(0))]_0$$

$$[f(g(x))]_1 = [f(g(0)) + g'(0)f'(g(0))x]_1$$

The composition of two TPS(n) can only be computed if the first one is origin preserving, then

$$[f]_n \circ [g]_n \equiv [f(g(x))]_n$$

If two maps that are known to order n and the first one is origin preserving, then the composition of the maps is known to order n.

$$[\vec{M}_1]_n, [\vec{M}_2]_n \in_n D_v^v$$

$$[\vec{M}_2]_n \circ [\vec{M}_1]_n = [\vec{M}_2(\vec{M}_1(\vec{z}))]_n$$

Therefore the reference trajectory is always chosen as origin for the maps accelerator elements.



Cornell University



Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Inversion of Maps

The nth order inverse of an origin preserving function can be computed within the differential algebra (DA):

$$\vec{M}(\vec{z}) = \underline{M}_1 \vec{z} + \vec{N}(\vec{z})$$

$$\vec{M} \circ \vec{M}^{-1}(\vec{z}) = \underline{M}_1 \vec{M}^{-1} + \vec{N} \circ \vec{M}^{-1} = \vec{z}$$

$$\vec{M}^{-1} = \underline{M}_1^{-1}(\vec{z} - \vec{N} \circ \vec{M}^{-1})$$

$$[\vec{M}^{-1}]_n = \underline{M}_1^{-1}[\vec{z} - \vec{N} \circ \vec{M}^{-1}]_n = \underline{M}_1^{-1}(\vec{z} - [\vec{N}]_n \circ [\vec{M}^{-1}]_{n-1})$$

Iterative computation of the inverse:

$$[\vec{M}^{-1}]_1 = [\underline{M}_1^{-1} \vec{z}]_1$$

$$[\vec{M}^{-1}]_2 = \underline{M}_1^{-1}(\vec{z} - [\vec{N}]_2 \circ [\underline{M}_1^{-1} \vec{z}]_1)$$

$$[\vec{M}^{-1}]_3 = \underline{M}_1^{-1}(\vec{z} - [\vec{N}]_3 \circ (\underline{M}_1^{-1}(\vec{z} - [\vec{N}]_2 \circ [\underline{M}_1^{-1} \vec{z}]_1)))$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Generating Functions

The motion of particles can be represented by **Generating Functions**

Each **flow** or **transport map**: $\vec{z}(s) = \vec{M}(s, \vec{z}_0)$

With a **Jacobi Matrix** : $M_{ij} = \partial_{z_{0j}} M_i$ or $\underline{M} = \left(\vec{\partial}_0 \vec{M}^T \right)^T$

That is **Symplectic**: $\underline{M} \underline{J} \underline{M}^T = \underline{J}$

Can be represented by a **Generating Function**:

$$F_1(\vec{q}, \vec{q}_0, s) \quad \text{with} \quad \vec{p} = -\vec{\partial}_q F_1, \quad \vec{p}_0 = \vec{\partial}_{q_0} F_1$$

$$F_2(\vec{p}, \vec{q}_0, s) \quad \text{with} \quad \vec{q} = \vec{\partial}_p F_2, \quad \vec{p}_0 = \vec{\partial}_{q_0} F_2$$

$$F_3(\vec{q}, \vec{p}_0, s) \quad \text{with} \quad \vec{p} = -\vec{\partial}_q F_3, \quad \vec{q}_0 = -\vec{\partial}_{p_0} F_3$$

$$F_4(\vec{p}, \vec{p}_0, s) \quad \text{with} \quad \vec{q} = \vec{\partial}_p F_4, \quad \vec{q}_0 = -\vec{\partial}_{p_0} F_4$$

6-dimensional motion needs only **one function** ! But to obtain the transport map this has to be **inverted**.



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

Computation of Generating Functions

For any map for which the TPS(n) is known, a TPS(n+1) of a generating function that produces this map can be computed. For example, looking for

$$F_1(\vec{q}, \vec{q}_0, s) \quad \text{with} \quad \vec{p} = -\vec{\partial}_q F_1(\vec{q}, \vec{q}_0, s) \quad , \quad \vec{p}_0 = \vec{\partial}_{q_0} F_1(\vec{q}, \vec{q}_0, s)$$

$\vec{z} = \vec{M}(\vec{z}_0)$ is given as TPS(n)

$$\begin{pmatrix} \vec{q} \\ \vec{q}_0 \end{pmatrix} = \begin{pmatrix} \vec{M}_q(\vec{z}_0) \\ \vec{q}_0 \end{pmatrix} = \vec{l}(\vec{z}_0) \quad , \quad \begin{pmatrix} \vec{p}_0 \\ \vec{p} \end{pmatrix} = \begin{pmatrix} \vec{p}_0 \\ \vec{M}_p(\vec{z}_0) \end{pmatrix} = \vec{h}(\vec{z}_0) = \underline{J} [\vec{\partial} F_1(\vec{q}, \vec{q}_0)]_{\vec{l}(\vec{z}_0)}$$

$$\vec{\partial} F_1 = -\underline{J} \vec{h} \circ \vec{l}^{-1} \quad \Rightarrow \quad F_1 = -\underline{J} \int_0^{(\vec{q}, \vec{q}_0)} \vec{h} \circ \vec{l}^{-1}(\vec{Q}) d\vec{Q}$$

$$[\vec{M}]_n \quad \Rightarrow \quad [\vec{l}]_n, [\vec{h}]_n \quad \Rightarrow \quad [\vec{l}^{-1}]_n, [\vec{h}]_n \circ [\vec{l}^{-1}]_n$$

$$[F_1]_{n+1} = -\underline{J} \int_0^{(\vec{q}, \vec{q}_0)} [\vec{h}]_n \circ [\vec{l}^{-1}]_n d\vec{Q}$$

Particle coordinates (q0,p0) are propagated by such generating functions when zeros of the following equations are found numerically:

$$\vec{p} + \vec{\partial}_q F_1(\vec{q}, \vec{q}_0, s) = \vec{0} \quad \text{and} \quad \vec{p}_0 - \vec{\partial}_{q_0} F_1(\vec{q}, \vec{q}_0, s) = \vec{0}$$



Cornell University



Brookhaven
National Laboratory

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023

The next Generation of Accelerator builders



Cornell University

Georg.Hoffstaetter@Cornell.edu

Graduate Accelerator Physics

USPAS summer 2023