

Atomic spectra & multi-electron states

review: spin \Rightarrow S-electron levels split into 2!

to explain Stern-Gerlach magn. dipole moment of electron:

$$M_z = \frac{-e}{2m} g S_z + \text{orbital contribution}$$

classical $g=1$
actual for spin $g=2$

Pauli eqn: need two wavefns to fully describe e^- : $\begin{pmatrix} \Psi_{\uparrow}(\vec{r}) \\ \Psi_{\downarrow}(\vec{r}) \end{pmatrix}$ "spinor"

$$|\Psi\rangle = |\Psi_{\uparrow}\rangle | \uparrow \rangle + |\Psi_{\downarrow}\rangle | \downarrow \rangle$$

$$\hat{S} = \frac{\hbar}{2} \vec{\sigma}$$
 with Pauli matrices for $\frac{1}{2}$ spin

$$S = \frac{1}{2}, m_s = \pm \frac{1}{2} \quad (\text{only 2 states needed for arb. spin orientation})$$

joint eigenbasis for \hat{S}_z, \hat{S}^2 :

$$\hat{S}^2 |S, m_s\rangle = \hbar^2 S(S+1) |S, m_s\rangle \quad \text{with } |\frac{1}{2}, \frac{1}{2}\rangle \equiv |\uparrow\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

$$\hat{S}_z |S, m_s\rangle = \hbar m_s |S, m_s\rangle \quad |\frac{1}{2}, -\frac{1}{2}\rangle \equiv |\downarrow\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

① Detour: perturbation theory basics

$$\hat{H} = \hat{H}_0 + \hat{K}_p \quad \text{with } \hat{K}_p \text{"small"}$$

\hat{H}_0 basis $|\Psi_n\rangle$ still "good enough" for \hat{H}

$$E_n \rightarrow E_{0,n} + \Delta E_n$$

$$\text{with } \Delta E_n \approx \langle \Psi_n | \hat{K}_p | \Psi_n \rangle$$

② spectral structure

line splitting without ext. fields

gross structure $\sim E_n$ (meV)

fine structure $\sim 10^{-4} \text{ eV}$

"two magnet bars"

hyperfine structure $\sim 10^{-6} \text{ eV}$

the Lamb shift $= 4.3 \times 10^{-6} \text{ eV}$

splitting of S, p levels of H

term in the Hamiltonian

- $V(r)$ nucl. potential

- relativistic correction

- spin-orbit coupling

- nucleus' spin-orbit

- QED

mass renormalization $+1017 \text{ MHz}$
 vacuum polarization -27 MHz
 anomalous magn. moment $+68 \text{ MHz}$

total $+1058 \text{ MHz}$

③ Spin-orbit coupling

a) field due to nuclei : $\vec{E} = \frac{1}{c} \nabla V(r) = \frac{1}{c} \frac{dV(r)}{dr} \frac{\vec{r}}{r}$
 in e^- moving frame, this in part becomes B-field

$$\vec{B} = -\frac{\vec{s} \times \vec{E}}{c^2} = -\frac{\vec{s} \times \vec{r}}{ec^2 r} \frac{dV(r)}{dr} = \frac{\vec{L}}{emc^2} \frac{1}{r} \frac{dV(r)}{dr}$$

$$V_{SO} = -\vec{p}_s \cdot \vec{B} \text{ where } \vec{p}_s = \frac{-e}{am} q \vec{S}, \Rightarrow \text{replace } \vec{S} \cdot \vec{L} \text{ in Q.M.}$$

classically $V_{SO} = \frac{1/2}{m^2 c^2} \frac{1}{r} \frac{dV(r)}{dr} \vec{S} \cdot \vec{L}$ correct treatment gives addl. $1/2$

b) from e^- 's point of view: (e^- is not in inertial frame)
 nucleus is spinning around with \vec{L}
 B-field $\propto \vec{L}$; energy $V_{SO} \propto \vec{S} \cdot \vec{L}$ Internal Zeeman

c) classically: magn. dipoles interact, \Rightarrow
 \vec{L} and \vec{S} are no longer conserved
 (though L^2 and S^2 are)

Q.M.: cannot use M_L and M_S to label quant. states

\Rightarrow solution: use total angular momentum $\vec{J} = \vec{L} + \vec{S}$
 (must still be conserved)

④ Total angular momentum

eigenvalues of \hat{J}^2 : $\hbar^2 j(j+1)$

$$\hat{J}_z: \hbar m_j \quad m_j = \underbrace{-j, \dots, 0, \dots, j}_{2j+1 \text{ values}}$$

\Rightarrow use n, l, j, m_j to label quant. state

Q: what values can j take? $l-s \leq j \leq l+s$ (if $l \geq s$)

e.g. $j=l \pm \frac{1}{2}$ for H-like

⑤ going from $|n, l, m_L, m_S\rangle$ to $|n, l, j, m_j\rangle$ is "basis change"

check that we did not lose states

fixed l, S : M_L, M_S can take $2(2l+1)$ values

n $^{2s+1} L_j$
 spectr. notation

$$j, m_j \text{ can take } [2 \underbrace{(l-\frac{1}{2})+1}_{j_1}] + [2 \underbrace{(l+\frac{1}{2})+1}_{j_2}] = 2(2l+1)$$

⑥ other effects

relativity: $KE \neq \frac{p^2}{2m}$, instead $KE = \sqrt{(pc)^2 + (mc^2)^2} - mc^2$
 $\approx \frac{p^2}{2m} - \frac{p^4}{8m^3c^2} + \dots$

can show that generally $E_n \rightarrow E_{n'}$

hyperfine structure

nuclei have magnetic dipole:

dipole-dipole interaction $\propto \vec{\mu}_p \cdot \vec{\mu}_e$

solved using same techniques as spin-orbital coupling

$$\Delta E_{\text{hyper}} \sim \alpha^2 \frac{m_e}{m_p} \text{ in atomic units}$$

H-atom (most dominant element in the Universe) nuclear spin ↑ electron spin ↑

$$5.9 \times 10^{-6} \text{ eV}$$

or 1420 MHz

21 cm wavelength

very important in radio-astronomy "21 cm line"