

## Introduction to Diffraction

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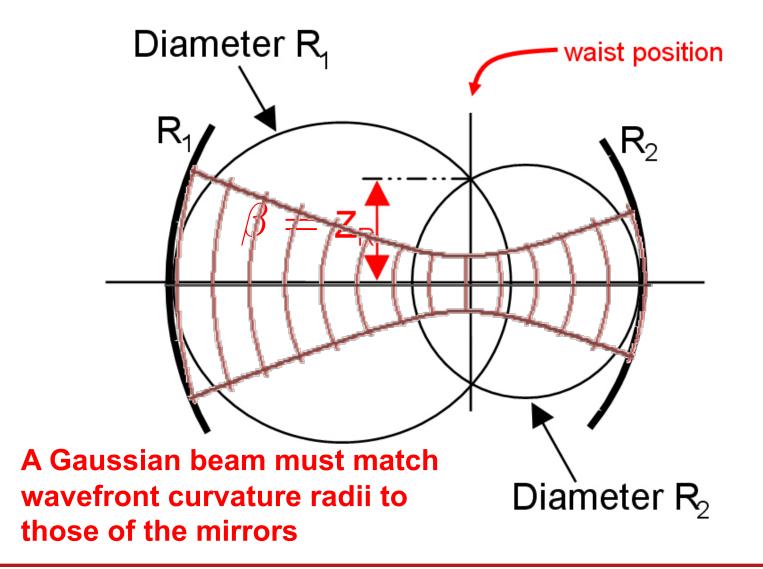
#### **Outline**

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- Gaussian beams in optical resonators (contd.)
- He-Ne laser spectrum (contd.)
- Introduction to diffraction
- Fresnel vs. Fraunhofer diffraction



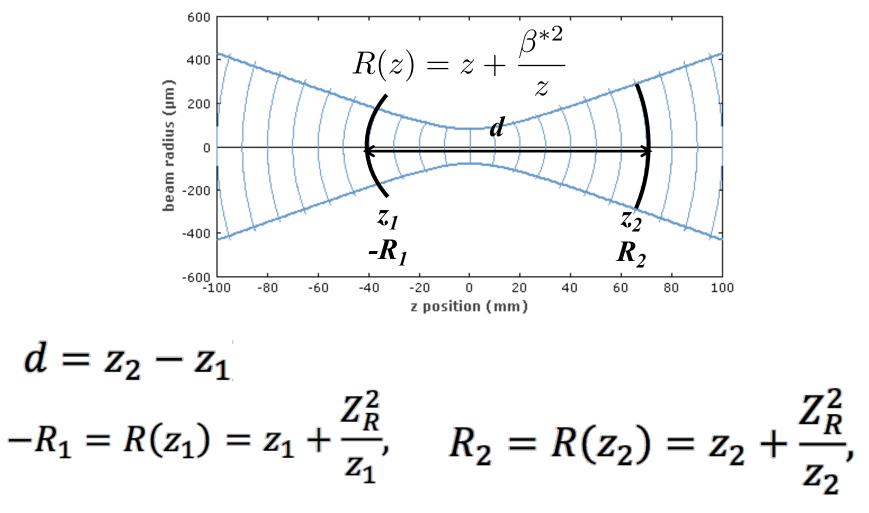
### **Two-mirror resonator**





## Gaussian mode in 2-mirror cavity

#### Recall Gaussian mode curvature:





## Solution: Rayleigh range & waist position in the cavity

$$z_{1} = \frac{d(d - R_{2})}{R_{1} + R_{2} - 2d'}$$

$$z_{2} = \frac{d(R_{1} - d)}{R_{1} + R_{2} - 2d'}$$

$$Z_{R} = \frac{d(R_{1} + R_{2} - d)(R_{1} - d)(R_{2} - d)}{(R_{1} + R_{2} - 2d)^{2}},$$

Recall that the rms waist is given by:  $\sigma_x = \sqrt{Z_R \epsilon_x} =$ 

$$\sqrt{Z_R \frac{\lambda}{4\pi}}$$

## Actual wave equation to solve

$$\nabla^{2}\mathbf{E} - \frac{1}{c^{2}} \frac{\partial^{2}\mathbf{E}}{\partial t^{2}} = 0 \underbrace{\mathbf{E}(t, \mathbf{r})}_{Fourier}$$
Or equivalently Helmholtz equation:
$$k = \omega/c \quad \nabla^{2}\mathbf{E} + k^{2}\mathbf{E} = 0 \underbrace{\mathbf{E}(\omega, \mathbf{r})}_{\mathbf{E}(\omega, \mathbf{r})}$$

This is usually solved using the *paraxial approximation* 

$$r = \sqrt{x^2 + y^2 + z^2} \approx z + \frac{x^2 + y^2}{2z}$$



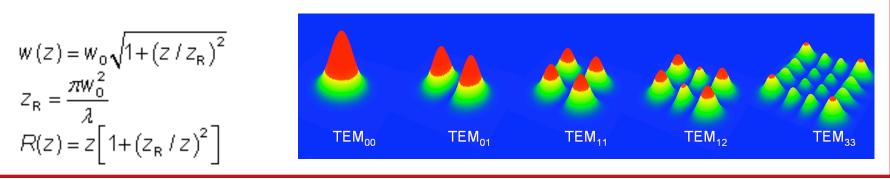
There are *infinite solutions* to the Helmholtz equation in free space.

E.g. Hermite-Gaussian TEM<sub>nm</sub> modes (E-field, not intensity!)

$$E_{nm}(x,y,z) = E_0 \frac{w_0}{w(z)}$$

$$\cdot H_n\left(\sqrt{2} \frac{x}{w(z)}\right) \exp\left(-\frac{x^2}{w(z)^2}\right) \cdot H_m\left(\sqrt{2} \frac{y}{w(z)}\right) \exp\left(-\frac{y^2}{w(z)^2}\right)$$

$$\cdot \exp\left(-i\left[kz - (1+n+m)\arctan\frac{z}{z_R} + \frac{k\left(x^2 + y^2\right)}{2R(z)}\right]\right)$$



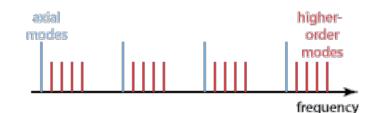


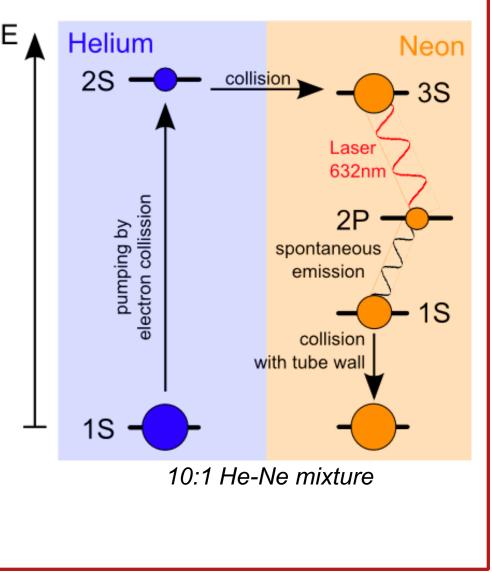
## He-Ne laser energy levels

95% of laser power is in the  $TEM_{00}$  (Gaussian mode)

Width of the resonance (medium gain) ~1.5 GHz, i.e. only 2 axial modes are typically present

Can use Michelson's interferometer to "see" the individual axial modes

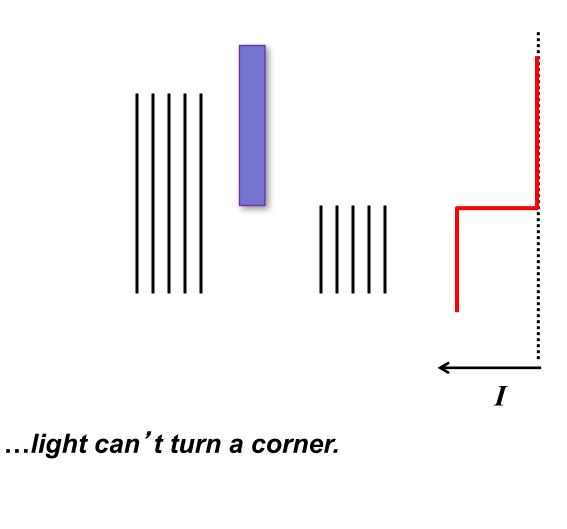








#### Geometrical optics...







#### Physical optics...



Francesco Maria Grimaldi (1618 - 1663)

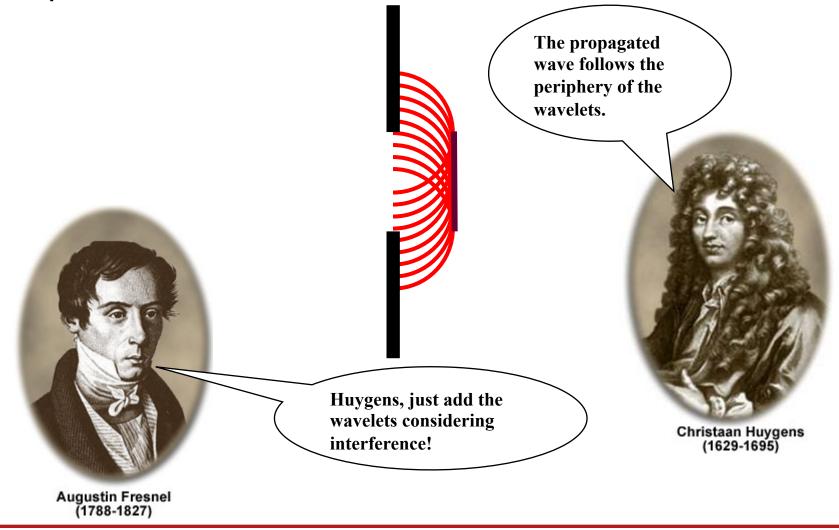
#### ...actually, it can.

Ι



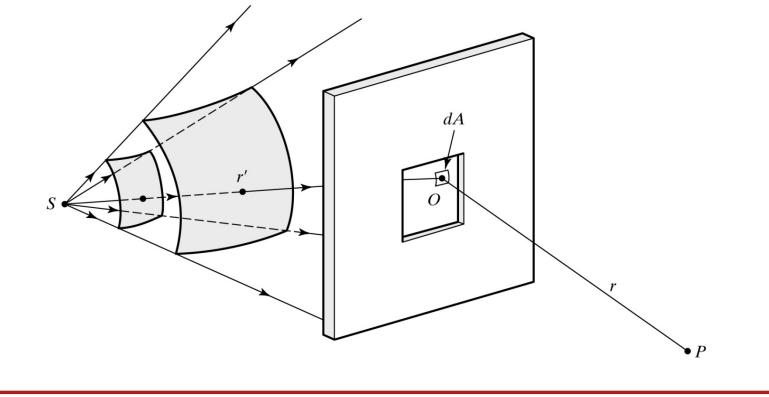
## Hyugens-Fresnel principle

every point on a wavefront may be regarded as a secondary source of spherical wavelets





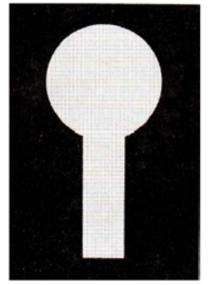
If one perturbs a plane wavefront, the Huygens wavelets will no longer constructively interfere at all points in space. Adding the wavelets by physical optics explains why light can turn corners and create fringes around images of objects.





## Diffraction: a generic wave phenomenon

#### Key-hole Incoherent

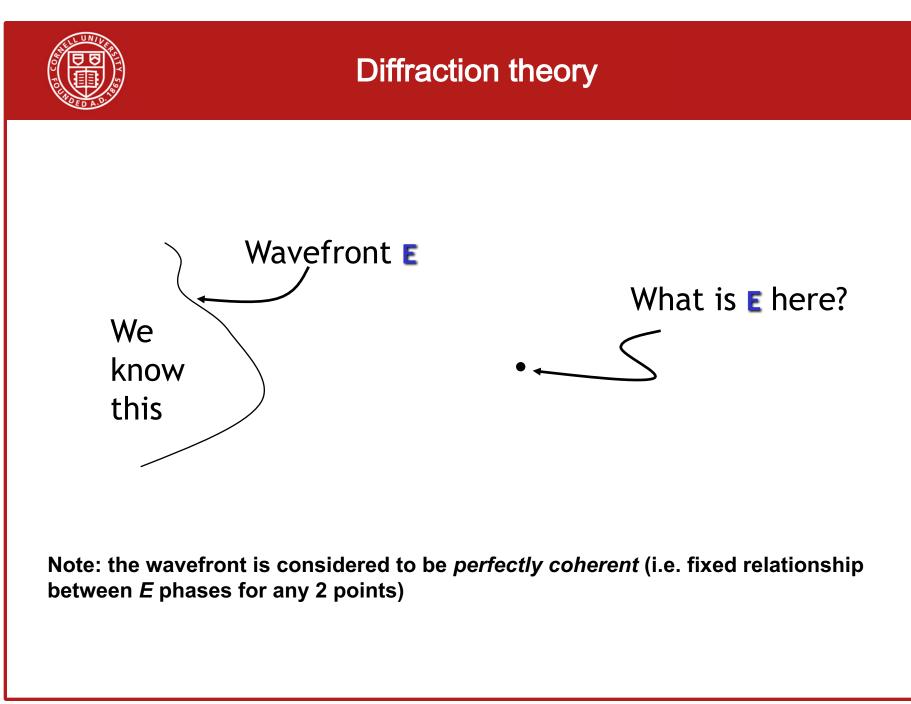


#### Coherent illumination



#### Sea waves

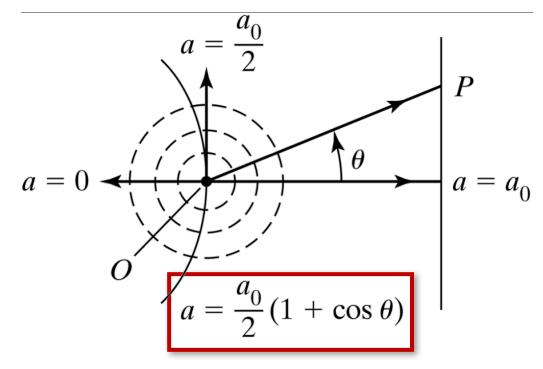






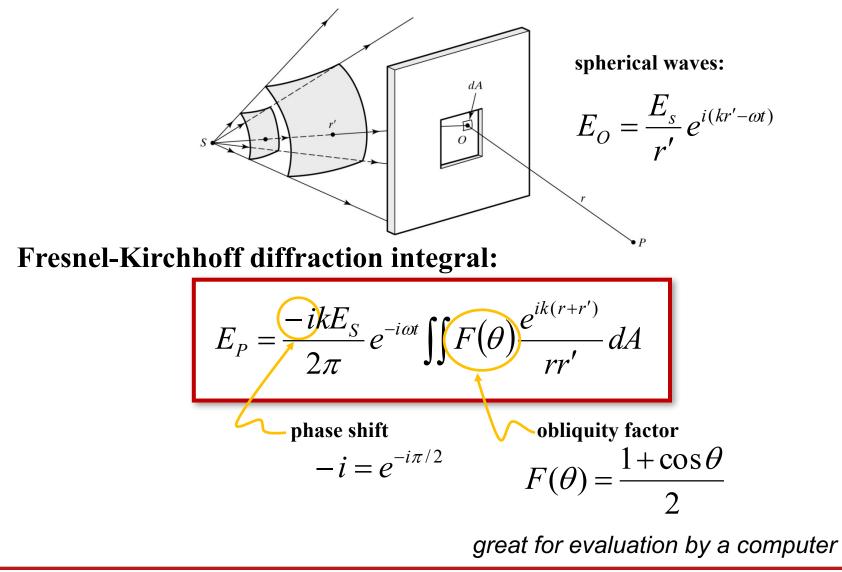
# **Obliquity factor**

- wavelets propagate isotropically—in forward and reverse directions
- to use the Huygens approach, modify amplitude of wavefront as a function of  $\theta$ :





## Calculating the diffracted wave amplitude





### Fresnel vs. Fraunhofer



Augustin Fresnel (1788-1827)



Joseph von Fraunhofer (1787-1826)

#### Contemporaries, but not collaborators (nor competitors).



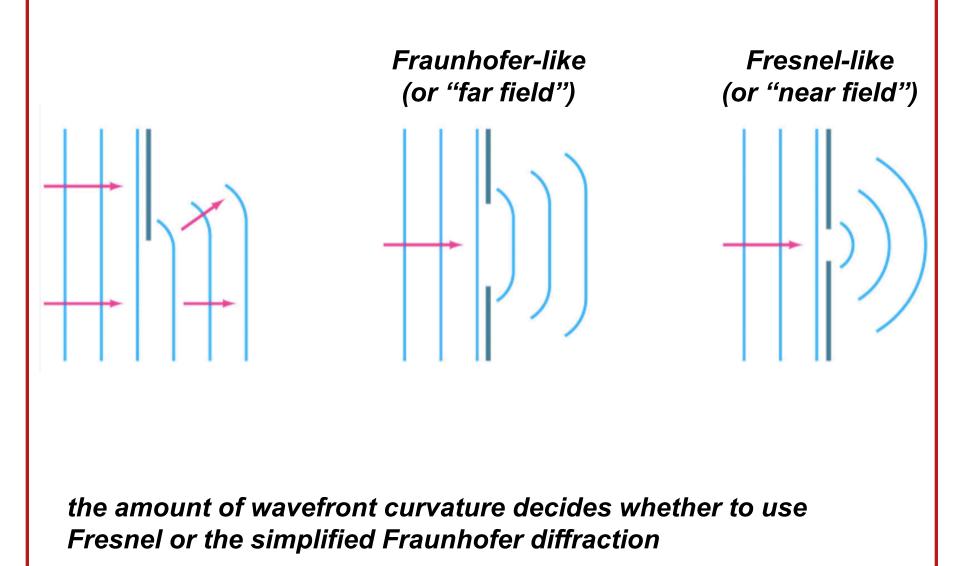
# Fresnel vs. Fraunhofer diffraction

- Fresnel regime is the nearfield regime: the wave fronts are curved, and their mathematical description is more involved.
- Very far from a point source, wavefronts almost plane waves.
- Fraunhofer approximation valid when source, aperture, and detector are all very far apart (or when lenses are used to convert spherical waves into plane waves)



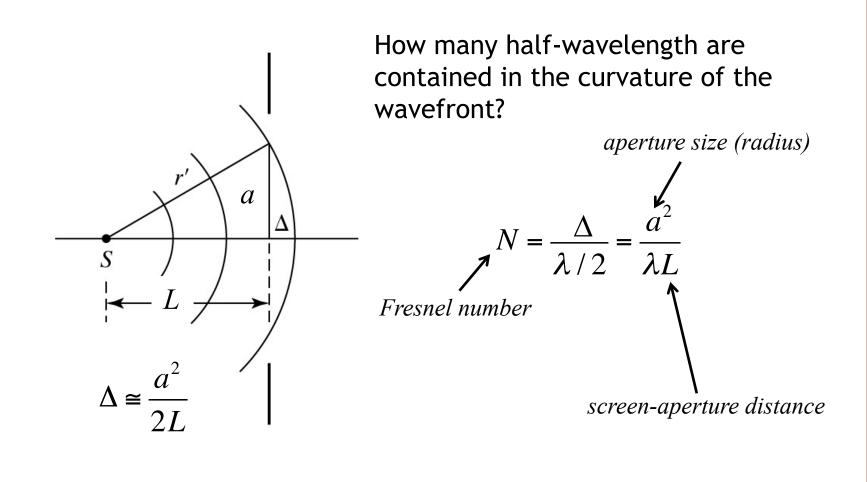


### Aperture size matters



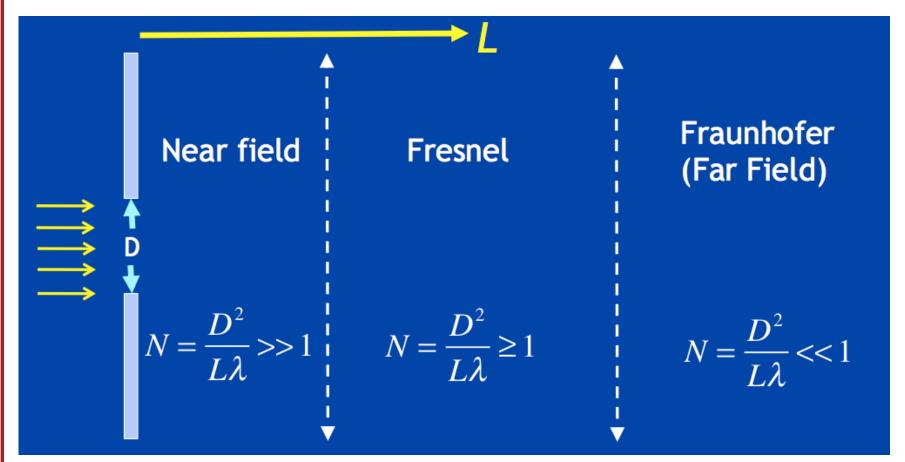


### **Fresnel number definition**





## Regions of validity for diffraction calculations

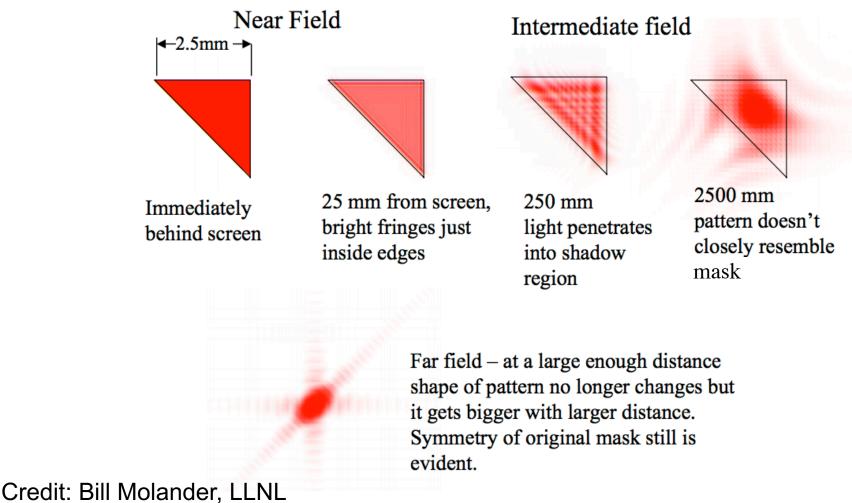


Math simplifies:  $E_p$  is simply a 2D FT of the  $E_s$  at the aperture!



## An example

# Pattern on screen at various distances





## Links/references

http://www.optiqueingenieur.org/en/courses/OPI\_ang\_M01\_C03/co/Grain\_OPI\_ang\_M01\_ C03.html

http://edu.tnw.utwente.nl/inlopt/overhead\_sheets/Herek2010/week7/13.Fr esnel%20diffraction.ppt

http://www.ucolick.org/~max/289/Lectures%20Final%20Version/Lecture%203%20Physical%20Optics/Lecture3%20Physical%20Optics\_v2.ppt